

The Cryogenian arc formation and successive high-K calc–alkaline plutons of Socotra Island (Yemen)

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Abstract:

The Socotra Island belongs to the southern rifted margin of the Gulf of Aden and occupied in Neoproterozoic times a key position to constrain the age and the nature of the largely hidden Neoproterozoic rocks of the Arabian plate. Our integrated field, petrographic, geochemical and geochronological study in the Neoproterozoic rocks recognises three main successive events: (a) high-temperature ductile deformation and metamorphism forming probably in a compressive or transpressive regime; (b) mafic to intermediate intrusions as vertical sheets, kilometre-scale gabbro laccoliths, mafic dike swarm and lavas which present mainly a depleted arc signature with some evidences of evolution from an enriched-arc signature; (c) felsic intrusions mainly composed of highly potassic calc–alkaline and pinkish granites dated between 840 and 780 Ma. Relationships between the various petrographic types and U–Pb data suggest that these events occurred during a relatively short time span (80 Ma at max). Earlier high-temperature–low-pressure metamorphism stage as well as geochemical signature of mafic rocks show that development of Cryogenian formations of Socotra were controlled successively by an Andean-arc and a back-arc setting. These features cannot be easily reconciled with those of the Arabian–Nubian shield to the west of Socotra and of the Mozambique Belt to the south. We propose that the Socotra basement was developed at an active margin close to the India block in Cryogenian times.

Keywords : Neoproterozoic ; East African–Antartic Orogen ; Arabian–Nubian shield ; Socotra Island ; Andean-type arc ; Back-arc basin

1. Introduction

The East-African-Antartic-Orogen (EAAO, Fig. 1a), one of the hugest orogen of the Earth history, extends over 8000 kilometres from Egypt to the north to Antarctica to the south (Stern, 1994; Meert, 2003; Jacobs and Thomas, 2004, 2010). The EAAO corresponds to a N52 S Neoproterozoic collision zone between proto-East Gondwana represented by the current Arabia, Somalia, Pakistan, India and Madagascar and the proto-West Gondwana represented by the East-Saharan, Congo and Tanzania cratons (Meert, 2003; Collins and Pisarevsky, 2005; Collins, 2006). This orogen presents a great complexity due to the fact that East and West Gondwana did not exist as Neoproterozoic supercontinents in their own right, but correspond to an accretion of various terranes which gathered during the Neoproterozoic time (see Collins and Pisarevsky, 2005, and references therein). The northern part of the EAAO corresponds to the Arabian-Nubian shield (ANS, Fig. 1b) dominated by juvenile mid-Neoproterozoic (Cryogenian) island-arc terranes associated to microcontinental blocks such as the Afif-Abbas composite terrane and the Al Mahfid gneiss terrane (e.g. Stern, 1994; Whitehouse et al., 1998 and 2001; Johnson and Woldehaimot, 2003, Meert, 2003; Collins and Pisarevsky, 2005), both bounded by sutures zones with ophiolites. ANS is characterized by mild accretion at low to medium metamorphic grade (Stern, 1994; Shackleton, 1996) which occurred between 700 and 600 Ma (Nehlig et al., 2002, references therein), and by a final stage of post-orogenic extension marked by late calc-alkaline and alkaline igneous activity and by formation of some core complexes (Blasband et al., 2000; Greiling et al., 1994; Avigad and Gvirtzman, 2009) which began around 600 Ma (Jonhson and Woldehaimot, 2003). By contrast the southern part of the EAAO shows high-grade rocks

70 of Ediacaran-early Cambrian age and is devoid of accreted juvenile Neoproterozoic arc
71 formations (Jacobs and Thomas, 2004, Bingen et al., 2009). This part of the orogen seems to
72 represent a continent-continent collision zone which occurred during the final accretion of
73 various Gondwana blocks (Muhlongo and Lenoir, 1994; Jacobs et al., 1998; Kröner et al.,
74 2001, Jacobs and Thomas, 2004).

75 There is a lack of magnetic data for the period between 1 Ga to 820 Ma to identify
76 clearly the eastern limit of the EAOO and of the composites terranes of the ANS (Li et al.,
77 2008). For instance, new models of late Mesoproterozoic reconstruction proposed that the ANS
78 microcontinental blocks were either placed adjacent to northeast India (e.g. Li and Powell,
79 2001), adjacent and outboard of the Congo-São Francisco cratons (Collins and Pisarevsky,
80 2005), or between India and Sahara (Li et al., 2008). The Socotra Island (Yemen, Fig. 1b)
81 have occupied, prior to the Oligo-Miocene Gulf of Aden opening (e.g. d'Acremont et al.,
82 2006; Leroy et al., 2010a), a key position between the proto-East Gondwana, the high-grade
83 rocks of the EAOO and the ANS (Fig. 1a). It corresponds, together with the basement of
84 Mirbat in Oman, to the only km-scale outcrop of Neoproterozoic rocks in the eastern part of
85 the Arabic plate (Fig. 1b). In this way, the study of the Socotra Neoproterozoic rocks should
86 allow the specification of the boundaries of the various domains of the EAOO and their
87 geological histories. We investigate the Neoproterozoic basement of Socotra and propose,
88 based on an integrated field, petrographic, geochemical and geochronological study, that it
89 corresponds to a well-preserved Cryogenian arc that may have formed above an active margin
90 of the Indian proto East-Gondwana block.

91

92

93 **Geological setting**

94

95 The Socotra Island belongs to the southern rifted margin of the Gulf of Aden which is
96 an active oceanic basin (Leroy et al., 2004). The Gulf of Aden forms by the split away of
97 Somalia plate to the South and Arabian plate to the North resulting from an oblique rifting
98 characterized by a N20°E-striking extension and a N70°E-oriented rift (Bellahsen et al., 2006;
99 Autin et al., 2010). Rifting history started around 35 Ma and oceanic accretion is recorded
100 since 17.6 Ma along the whole Gulf of Aden, *sensu stricto* (e.g. Leroy et al., 2010a and b;
101 Leroy et al., this volume).

102 The Gulf of Aden rifting is expressed in two different ways in the Socotra Island: a
103 western domain made up of two well-expressed Oligo-Miocene tilted blocks (Razin et al.,

104 2010; Fig. 2) and an eastern domain composed of a single and huge tilted block formed by the
105 Haggier 1500 m-high mountains (Fig. 2). These two domains are separated by a NE-SW
106 transfer fault zone dipping toward the northwest. The hanging wall of the transfer zone
107 corresponds to the tilted blocks bounded by N110°E normal faults and possess a Quaternary
108 relief lower than the footwall which is the main basement outcrop of the Island (Fig. 2).

109 Three main basement outcrops can be observed in the island (Fig. 2), which correspond
110 to three basement highs located at the head of tilted blocks. The Qalansya and the Sherubrub
111 areas are located in the western part of the island and both represent respectively *ca.* 65 km²
112 and 50 km². The most voluminous basement outcrop, the Mont-Haggier basement high
113 representing *ca.* 580 km² is located at the eastern part of the island. The Neoproterozoic
114 basement of Socotra displays a great variety of metamorphic, plutonic and volcanic rocks
115 previously studied by Beydoun and Bichan (1970). They have evidenced that the oldest part
116 of this basement is made of amphibolite facies meta-sediments and meta-igneous rocks which
117 have been intruded by synkinematic granites and late-kinematic gabbros. Post-kinematic
118 igneous activity gave rise to a sequence of volcanic rocks, hornblende/biotite and peralkaline
119 granites, gabbros and minor intrusions, which make up the bulk of the Haggier mountains. At
120 the footwall of the major transfer fault of Socotra, Beydoun and Bichan (1970) have evidenced
121 an association of bedded mudstones, sandstone and tuffs belonging to the Hadibo series.
122 Beydoun and Bichan (1970) proposed that: (i) metamorphic events and synkinematic plutonic
123 rocks are Precambrian in age, (ii) the Hadibo series is late pre-Cambrian, and (iii) the
124 peralkaline granites are Early Paleozoic.

125 The reconstruction of the Gulf of Aden shows that the Socotra Island has occupied, prior
126 to the Oligo-Miocene rifting, a position close to the Precambrian Mirbat and Al Halaaniyat
127 Islands outcrops in Oman (Figs. 1b, e.g. d'Acremont et al., 2006; Leroy et al., 2010a). Gass et
128 al. (1990), based on few geochemical and Rb/Sr whole-rock ages ranging from 850 ± 27 Ma
129 to 706 ± 40 Ma, proposed that Precambrian rocks of Oman lie within the Pan-African domain
130 *s.l.* and are not part of an older basement such as that identified in eastern Saudi Arabia (Afif
131 and Al-Mahfid Terranes, Fig. 2b). Mercolli et al. (2006) defined four units in the
132 Neoproterozoic basement of Mirbat. The Juffa group corresponds to alternation of
133 paragneisses, amphibolites and few meta-ultramafic lenses. U-Pb ages on zircon and Pb-Pb
134 ages on garnet clustering around 815 Ma are interpreted as the age of the metamorphism in
135 amphibolite facies (Mercolli et al., 2006). The Sadh Group corresponds to two types of
136 orthogneisses dated by U-Pb *in situ* method on magmatic looking zircons at 816 ± 12 Ma and
137 799 ± 5 Ma (Mercolli et al., 2006). The Tonalite Group including three kilometre-scale calc-

138 alkaline plutons was dated by U-Pb *in situ* method around 780-790 Ma (Mercolli et al., 2006).
139 Finally, the Granite Group comprises: (i) different types of dikes and small bodies of granite
140 dated by step leaching Pb/Pb on garnet between 770 and 750 Ma, (ii) two small granitic
141 plutons without precise ages, (iii) basaltic to rhyolitic Shaat Dike Swarm without precise age.

142

143

144 **Petrographic and microstructural study of the Socotra basement**

145

146 The Neoproterozoic basement of Socotra displays a great variety of metamorphic,
147 plutonic and volcanic rocks. A preliminary map of the various lithological formations was
148 performed by Beydoun et al. (1970). We have completed and modified this map for the three
149 principals inliers of Neoproterozoic rocks by field study and satellite images (Fig. 2 and 3).
150 This section describes the different lithological formations of the Socotra basement.

151

152 *The metamorphic basement*

153

154 The metamorphic basement corresponds to an alternating of quartzites, paragneisses,
155 micaschists, orthogneisses and ortho-amphibolites. Paragneisses and red or white quartzites
156 crop out on large kilometre scale area. Orthogneisses and ortho-amphibolites are observed as
157 metre to hectometre-scale lenses (Fig. 4a) or boudins in paragneisses and quartzites. Finally
158 micaschists are only locally observed.

159 Micaschists consist of quartz, biotite, muscovite and locally andalusite. Paragneisses
160 display paragenesis of quartz, alkali feldspar, plagioclase, biotite, muscovite, sillimanite and
161 locally relictual andalusite (Fig. 4b). These two-micas micaschists and paragneisses seem to
162 be derived from the same metapelitic protolith affected by different grades of metamorphism.
163 Paragneisses show frequently evidences of partial melting and correspond to metatexites (Fig.
164 4b), with 30 % max. of leucosomes. In two areas migmatites are strongly evolved and
165 correspond to diatexites (e.g. Menhert, 1968), which contain an important amount of
166 cordierite. Orthogneisses are composed mainly by porphyric alkali feldspar, quartz and biotite
167 (= augengneisses). Amphibolites correspond to former sills of mafic rocks (Fig. 4a), probably
168 gabbros as shown by observations of relictual orthopyroxene and plagioclase laths (Fig. 4c),
169 affected by a HT-LP metamorphism with neo-formation in the amphibolite-facies of
170 Hb+Ep+Ab and coeval deformation.

171 Rocks of the metamorphic basement show a well-defined foliation associated to multi-
172 scale folding. Foliations and contact between the different metamorphic formations are
173 generally steeply dipping and oriented on average at *ca.* N80°E. Syn-kinematic neo-formed
174 sillimanite which replaces andalousite observed as relic in paragneisses (Fig. 4b), and
175 spectacular anatexis phenomena with leucocratic/melanocratic banding parallel to the
176 foliation show that an important HT-LP metamorphism under amphibolite facies conditions
177 affected the oldest rocks of Socotra during the main deformation event.

178

179 *The volcanic series*

180

181 To the south of the Mont-Haggier area the metamorphic unit is overlain by an
182 association of pyroclastic and effusive rocks belonging to the Southern Haggier volcanic
183 series. Contact between these two units is subhorizontal and devoid of evidence of
184 deformation. The effusive rocks correspond to an alternating of rhyolitic, andesitic, dacitic and
185 basaltic lavas. Pyroclastic rocks, together with thin lava flows, are mainly located to the west
186 of the volcanic formation (Beydoun and Bichan, 1970) and show an alternation of breccias
187 (Fig. 4d), agglomerates, and tuffs (Fig. 4e). Although volcanic rocks are devoid of evidence
188 of deformation and of large metamorphism overprint, we observed locally at the base of the
189 volcanic complex some evidences of HT metamorphic overprint marked by replacement in
190 basaltic lava of pyroxene minerals by hornblende.

191 At the footwall of the transfer fault, close to Hadibo, we have observed very peculiar
192 yellowish basalt (sample SP15A, localization in figure 2) with a doleritic texture. This basalt
193 (Fig. 4f) is intruded by the Haggier pink granite (see below). Relationships between this basalt
194 and the Southern-Haggier volcanic series are difficult to determine.

195

196 *The mafic to intermediate plutonic sheets*

197

198 To this group belong several hectometre to metre thick sheets of tonalites, granodiorites
199 and diorites that have intruded the metamorphic basement in Qalansya, Haggier and
200 Sherubrub areas, and that were successively deformed under amphibolite facies conditions.

201 We studied in detail this group in the western part of the Qalansya basement high, where
202 we observed on a hundred metres long banks along the road, several sheets of more or less
203 differentiated plutonic rocks on a short distance (Fig. 5A). Petrographically fine-grained
204 tonalites and medium to coarse-grained granodiorites are composed of 30 to 40 % of
205 plagioclase, 30 to 40 % of quartz, 10 to 25 % of alkali felspar and 10 % of mafic minerals

206 (green hornblende and biotite). We have observed also magnetite, zircon, apatite and sphene.
207 Diorites are fine to medium-grained and composed of 40 to 60 % of plagioclases, 30 to 40 %
208 of green hornblende. We have observed also quartz and accessory minerals corresponding
209 mainly to magnetite, allanite, zircon, apatite and sphene. Locally these rocks are affected by
210 intense metamorphic recrystallization under amphibolite facies conditions as shown by the
211 growth of secondary hornblende, albite and epidote (Figs. 5 B-a and B-b).

212 Emplacement of diorites, tonalites and granodiorites occurred probably by successive
213 pulses as shown by the sharp contacts between the different sheets of these rocks (Fig. 5 B-c
214 and B-d). We observed also sometimes mingling between diorites and granodiorites
215 suggesting that emplacement of successive pulses occurred rapidly (Fig. 5 B-e). These
216 diorites and granitoids display a discrete steeply dipping foliation oriented N80°E and a
217 subvertical stretching lineation marked by disposition of neo-formed hornblende grains along
218 their long axis (Fig. 5 B-b and B-f). Dikes of microgranite, aplite and basalt have later
219 intruded the sheets of mafic to intermediate plutonic rocks.

220

221 *The layered gabbros*

222

223 Two km-scale bodies of layered gabbros occur in Socotra. The Haggier layered gabbros
224 intruded the metamorphic basement, covers *ca.* 40 km² of low rusty-looking hills to the E-SE
225 of Hadibo. These gabbros are devoid of any solid-state deformation and present a magmatic
226 layering (Fig. 6a) with a rhythmic alternation of pyroxene and plagioclase-rich layers. These
227 layers vary from a few millimetres, where individual bands can be traced over a distance of
228 several metres, to several centimetres in thickness. Petrographically, medium to coarse-
229 grained gabbros show a cumulate texture and consist of 60 to 80 % of zoned labradorite and
230 40 to 20 % of mafic minerals (Fig. 6b). Mafic minerals consist of clinopyroxene,
231 orthopyroxene and more rarely olivine replaced by hornblende and locally by chlorite and
232 tremolite-actinolite that attests that a middle grade metamorphic event, followed by a low-
233 grade event, have affected these gabbros. Accessory minerals correspond mainly to magnetite.

234 The Sherubrub gabbros cover a surface of *ca.* 6 km² in the western part of the area and
235 show the same petrographic characteristics as the Haggier layered gabbros. These gabbros are
236 locally associated with intermediate rocks as tonalites and granodiorites and present magmatic
237 layering (Fig. 6c and d).

238

239 *Granites*

240

241 Two kilometre-scale bodies of granitic rocks crop out in Socotra: the Sherubrub pluton
242 and the Mont-Haggier pluton.

243

244 The Sherubrub pluton and its satellites cover half the Sherubrub area on *ca.* 25 km². It is
245 made up mainly of pink medium-grained hypersolvus granite which contains 40 % of quartz,
246 30 % of microcline, 20 % of perthitic orthoclase, biotite and green hornblende (Fig. 7a and b).
247 Accessory minerals correspond mainly to magnetite, zircon, apatite and euhedral sphene.
248 Myrmekite is generally abundant at the contact between plagioclase and K-feldspar. Many
249 metre to hectometre-scale enclaves and raft of country rocks have been observed in this
250 granite (Fig. 7c). We also found in the Sherubrub pluton some bodies of grey and medium-
251 grained granite (Fig. 3). The pluton is associated in its periphery with a spectacular swarm of
252 pegmatites which are intrusive in paragneisses and gabbros (Fig. 7d). The Sherubrub pluton
253 shows sharp and subvertical contacts with its country rocks. Microstructural study indicates
254 that the Sherubrub granites present a progressive evolution from the core of the pluton where
255 sub-solidus deformation occurred (Fig. 7a) to the contact where high-temperature solid state
256 deformation occurred (protomylonites) and where we have measured sub-vertical E-W
257 foliation planes (Fig. 7b)

257

258 The Mont-Haggier pluton consists of two major facies covering *ca.* 250 km². The
259 Northern-Haggier biotite granite covers *ca.* 60 km² around Hadibo. It corresponds to a light
260 grey and medium-grained granite which contains 40 % of quartz, 20 to 30 % of zoned
261 plagioclase, 20 to 30 % of myrmekitic alkali feldspar and 10 % of biotite and green hornblende
262 (Fig. 7e). Accessory minerals correspond to apatite, zircon, magnetite and other opaque
263 minerals. Granite is associated with many mafic enclaves. Locally granites are more
264 leucocratic and present rare muscovite. Microstructural study indicates that the Northern-
265 Haggier biotite granite has suffered a very weak solid-state deformation marked by internal
266 deformation and reorientation of quartz grains (Fig. 7e). The Haggier pink granite covers *ca.*
267 200 km² in the centre of the Haggier basement high and can be studied on a natural vertical
268 cross-section of 1,500 metres. This granite is pinkish to reddish (Fig. 7f and 8a), and generally
269 coarse-grained. Quartz (~ 30%) and perthitic alkali feldspar (~ 60%) with subordinate albite
270 grains are the main components. We observed also rare arfvedsonite (Fig. 8b), zircon,
271 monazite and magnetite. The Haggier pink granite is generally devoid of internal deformation,
even though few thin-sections display quartz grains deformed under high-temperature

272 conditions (chessboard texture, Fig. 8c). The eastern contact of the Haggier pink granite with
273 the layered gabbros, which correspond to the bottom of the pluton, is diffuse, characterized by
274 spectacular magmatic “breccias” (Fig. 8a), and dip moderately at map scale. Magmatic
275 “breccias” are characterized by strongly angular and numerous enclaves of gabbro in granites
276 showing that the gabbros were crystallized at the time of granite intrusion. The southern and
277 western limits of the pluton show sharp and subvertical contacts with the Southern Haggier
278 volcanic series that may be followed for several kilometres. Study of the various contacts
279 suggest that the pluton were developed mainly in the Southern Haggier volcanic series at the
280 top of the Haggier layered gabbros. Finally, we observed from top to bottom of the batholith,
281 an evolution of the texture of the granites, coarse-grained in the major part of the outcrop, and
282 fine-grained in the western part of the Mont-Haggier area close to the contact with the
283 Cenozoic and Mesozoic series.

284

285 *The Haggier dike swarm*

286

287 A spectacular amount of dikes has been observed in the Mont-Haggier area. We have
288 not performed an exhaustive mapping of these dikes. Only the largest ones have been
289 represented on the petrographic map (Fig. 2). We distinguish three generations: (i) dikes of
290 microgranites, (ii) basaltic and andesitic dikes, (iii) dacitic and rhyolitic dikes.

291 Dikes of pink to red and fine grained granite are clearly associated with the pink
292 Haggier granite intrusion and are ubiquitous to the west and to the north of this intrusion. Metre
293 to hectometre-scale microgranitic dikes present a microgranular porphyric texture with the
294 same mineral association as the Haggier pink granite. While measurements of dikes of pinkish
295 microgranites show some disparities (Fig. 9a), a major family can be easily identified,
296 oriented on average around N33°E and subvertical.

297 Numerous volcanic dikes are intrusive in all the formations of the Mont-Haggier area.
298 The pink Haggier granite is devoid of mafic dikes and is only cut by numerous rhyolitic dikes.
299 Mafic dikes are homogeneously oriented (mean at N59°E, Fig. 9b) and correspond
300 dominantly to basaltic dikes, rich in green hornblende and plagioclase, with a doleritic texture
301 (Fig. 8d and e). Rhyolitic dikes (Fig. 8f) are homogeneously oriented (mean at N41°E, Fig.
302 9c) and display a microlithic-porphyric texture with orthoclase megacrysts.

303

304

305 **Major and trace elements**

306

307 Several mafic and felsic rocks from Mont-Haggier, Sherubrub and Qalansya basement
308 were sampled for whole-rock analysis (Table 1). Mafic rocks were picked up at Mont Haggier
309 and Qalansya areas whereas analysed felsic rocks were selected from Mont Haggier and
310 Sherubrub samples.

311

312 *Mafic rocks*

313

314 According to geochemical data, mafic rocks from Socotra Island suffered different
315 degrees of alteration. The Haggier layered gabbros are the least altered rocks with Loss on
316 Ignition (LOI) below 1.6 wt.%. Most Basaltic dikes belonging to the Haggier dike swarm
317 (SP10 and SP18), basaltic lava (S20) and tuff (SP16B) from the Southern Haggier volcanic
318 series are significantly altered with LOI up to 3.8 wt.% (Table 1). Diorites (CS14A and
319 CS18A) from the mafic to intermediate plutonic sheets of Qalansya are almost fresh rocks
320 (LOI = 2.01 wt.% and 0.98 wt.%). The yellowish basalt cropping out in the Northern part of
321 the Mont-Haggier area is highly altered (SP15A: LOI = 5.99 wt.%).

322

323 Major elements composition of volcanic and subvolcanic rocks from Mont-Haggier are
324 in good agreement with those of basalts, basaltic andesites and andesites (Table 1), even
325 though MgO and CaO concentrations are significantly low most likely due to strong
326 differentiation process. According to the SiO₂ vs K₂O diagram (Fig. 10), three volcanic and
327 subvolcanic rocks (SP10, SP16bB and SP18A) are calc-alkaline in composition whereas the
328 yellowish basalt SP15A belong to the high-K calc-alkaline to shoshonitic field. However, due
329 to high LOI values, such discrimination should be used with caution. For example, sample
330 S20 falls in the arc tholeiites series field most likely due to Large-Ion Lithophile Element
331 (LILE) leaching during alteration process. Diorites of Qalansya are calc-alkaline (CS18A) or
332 high-K calc-alkaline (CS14A) in composition. Mafic to intermediate rocks from Mont-
333 Haggier show a wide range of Al₂O₃ (13.52-26.32 wt.%), Fe₂O₃ (2.71-13.99 wt.%), CaO
334 (5.20-14.78 wt.%), MgO (2.43-7.55 wt.%), TiO₂ (0.18-2.56 wt.%) and Mg# (17-56) reflecting
335 difference between gabbro cumulates and non-cumulative rocks. Mafic to intermediate rocks
336 from Qalansya display almost homogeneous geochemical composition except for SiO₂ (49.38
337 vs. 54.55 wt.%) and Fe₂O₃ (8.48 vs. 11.15 wt.%). According to trace elements data, two
338 different signatures can be distinguished from non-cumulative rocks (Fig. 11a and Fig. 11c;
Table 1). The first signature recorded in the basalt SP15A and in the diorites of Qalansya, is

339 characterized by low Nb and Ta negative anomalies ($La_N/Nb_N = 3-4$) indicative of
340 suprasubduction setting and by low LILE and LREE enrichment (e.g. $Ce/Yb < 15$) underlined
341 by a relatively flat multi-element patterns in Figure 9a. The second signature, which is
342 recorded in the volcanic and subvolcanic rocks of Mont-Haggier area shows a large negative
343 Nb and Ta anomaly ($La_N/Nb_N > 7$) and shows significant enrichment in LILE and LREE (e.g.
344 $Ce/Yb \gg 15$) characteristic of volcanic arc basalts. In Figure 11d, the first group belongs to
345 the arc calc-alkaline basalts field and the second group plots in the transitional arc basalts
346 field and is close to arc tholeiite basalts (i.e. immature arc basalt) and back-arc basalt field.
347 Gabbro cumulates (Fig. 11a and Fig. 11b) are geochemically related to the second group as
348 underlined by chondrite-normalized rare earth elements diagram (Fig. 11b) which emphasizes
349 an evolution by fractional crystallization process. As a consequence, major and trace elements
350 composition of mafic to intermediate rocks highlights existence of two distinct geochemical
351 compositions in Socotra basement (Figure 11a, 11c and 11d; Table 1): (i) a signature
352 diagnostic of a depleted-arc or back-arc basin setting, and (ii) a signature characteristic of
353 enriched-arc setting.

354

355 *Felsic rocks*

356

357 Felsic rocks from the Haggier dike swarm, and the Haggier and Sherubrub plutons
358 display a moderate range of chemical composition (Table 1) with high SiO_2 content (71.67-
359 78.50 wt.%), low to moderate contents of Al_2O_3 (11.22-14.34 wt.%) and Fe_2O_3 (0.91-3.47
360 wt.%), low levels of MgO and TiO_2 (< 0.53 wt.% and < 0.37 wt.% respectively), and
361 moderate to high alkalis concentrations ($Na_2O : 2.95-5.33$ wt.%; $K_2O : 3.02-5.81$ wt.%). In
362 Figure 12a, felsic rocks are distributed between peraluminous, metaluminous and peralkaline
363 fields although most analyses are characterized by a metaluminous composition. Trace
364 elements data underlines wide variations in MREE, HREE, Zr and Hf, whereas LILE (except
365 Ba) and LREE composition exhibit a relative narrow range of composition (Fig. 12b)
366 resulting in relatively low LREE/HREE ratios (e.g. $La_N/Yb_N < 6.7$). Nb and Ta may be highly
367 variable but always display negative anomalies on silicate Earth-normalized multi-elements
368 diagram (Fig. 12b). Based on the plot of $10000 * Ga/Al$ vs. $Zr + Nb + Y + Ce$, felsic rocks
369 from both Mont Haggier and Sherubrub areas fall in the field of A-type granite except for two
370 samples which belong to the field of I-M-S-type granites (Fig. 12c). However, in similar plots
371 (i.e. $10000 * Ga/Al$ vs. Nb; $10000 * Ga/Al$ vs. Y; $10000 * Ga/Al$ vs. Ce; not shown) analyses
372 show a complex and transitional signature as they fall between fractionated I-type granite and

373 A-type granite fields. On a discrimination diagram (Fig. 12d), felsic rocks plot both within the
374 volcanic arc/syn-collision and oceanic ridge granites fields, except for one sample located in
375 the within plate granite field. Thus, geochemistry of felsic rocks from Socotra is not entirely
376 consistent with anorogenic (ORG and WPG) and orogenic (VAG) signature and their
377 emplacement in Cryogenian in a back-arc setting cannot be ruled out.

378

379

380 **U/Pb data**

381

382 Three samples were collected for zircon U-Pb dating. HAEU-1 and HAEU-2
383 (localization in fig.2) were collected in the Mont-Haggier area. HAEU-1 is representative of
384 the Haggier pink granite. HAEU-2 corresponds to an enclave of gabbro in the Haggier pink
385 granite, showing the same petrographic characteristics of the Haggier layered gabbros. Other
386 sample (SHEU-1, localization in fig.3) collected in the Sherubrub basement high, is
387 representative of the pinkish granite of the Sherubrub pluton.

388 Zircon grains from each sample were selected by hand-picking from 50-300 μm
389 fractions and mounted in epoxy resin together with fragments of the 91,500 standard zircon,
390 dated at 1062.4 ± 0.4 Ma [Ontario, Canada; Wiedenbeck et al., 1995]. Mounts were then
391 polished and gold coated. U-Pb isotopic compositions were determined in situ using the
392 CRPG-CNRS Cameca IMS-1270 ion microprobe (Nancy, France). Further information on
393 instrumental conditions and data reduction procedures is given in Deloule et al. (2002).
394 Concordia diagrams and discordia lines were constructed using the Isoplot program (Ludwig,
395 2003). In the forthcoming paragraphs, we use $^{206}\text{Pb}/^{238}\text{U}$ ages for concordant data, rather than
396 $^{207}\text{Pb}/^{235}\text{U}$ ages, which are more sensitive to common lead contribution (the ^{207}Pb ion signal is
397 about ten times lower than the ^{206}Pb ion signal). Discordant data were used on the concordia
398 diagrams to calculate possible discordia lines.

399 Eleven anhedral to squat zircon grains were analysed for sample HAEU-1 and for each
400 crystal core and rim were analysed (Table 2). Analyses show very small common lead
401 contribution but significant Pb losses. One zircon grain provides a concordant $^{206}\text{Pb}/^{238}\text{U}$ age
402 at 858 ± 11 Ma and is interpreted as an inherited core. The remaining ten analyses are
403 subconcordant or discordant. Regression curve calculated from these zircon grains gives a
404 discordia line (Fig. 13a) with an upper intercept at 816 ± 12 Ma (MSWD = 0.49).

405 Eleven zircon grains were analysed for sample HAEU-2 and for each crystal core and
406 rim were analysed (Table 2). Based on morphology, two types of zircons can be identified: (i)

407 euhedral and acicular crystals and (ii) anhedral to subhedral and squat crystals. Data show
408 very small amount of common lead ($Pbc < 1\%$) but some exhibit slight Pb or U losses. One
409 zircon grain gives concordant $^{206}\text{Pb}/^{238}\text{U}$ age at 856 ± 12 Ma and most likely corresponds to
410 an inherited core. Seven of the remaining eleven analyses are scattered along discordia line
411 (Fig. 13b) defining an upper intercept at 825 ± 18 Ma (MSWD = 0.49). Mean $^{206}\text{Pb}/^{238}\text{U}$ age
412 of 821 ± 16 Ma (MSWD = 0.53) was calculated from four subconcordant analyses. Since
413 gabbro enclaves are intruded by some granitic veins, we assume the presence of two
414 population of zircon: (i) one associated with the gabbro enclaves and (ii) and the other
415 crystallized from the granitic intrusions and veins. However, taking into account the low
416 zirconium concentrations of gabbros and the very low abundance of zircon observed in thin
417 sections, we assume that most zircon crystals separated from HAEU-2 belong to granitic
418 intrusions and veins. Even if some zircon grains from gabbros enclaves have been analysed,
419 homogenous distribution of data observed in Fig.13b indicates a relative contemporaneity
420 (within $^{206}\text{Pb}/^{238}\text{U}$ age errors) of mafic and felsic magmatism. Whatever, calculated U-Pb
421 age is identical within errors to age obtained for the sample HAEU-1 and is interpreted as the
422 age of emplacement of the Haggier pink granite.

423 In the Sherubrub granite (sample SHEU-1), zircon grains are euhedral, prismatic and
424 acicular. One zircon analysis displays slight loss of uranium linked to a reopening of the U-Pb
425 system. The remaining seven analyses are subconcordant to discordant. Regression of these
426 zircon analyses provides a discordia line (Fig. 13c) with an upper intercepts at 812 ± 27 Ma
427 (MSWD = 1.8). This age is in good accordance with ages acquired from the only concordant
428 analysis on a Tera-Wasserburg diagram (not shown) i.e. 809 ± 13 Ma ($^{206}\text{Pb}/^{238}\text{U}$ age) and
429 815 ± 11 Ma ($^{207}\text{Pb}/^{235}\text{U}$ age).

430

431

432 **Discussion**

433

434 *The high-temperature metamorphic stage of the Socotra basement*

435

436 The oldest rocks of Socotra correspond mainly to an association of paragneisses and
437 quartzites forming the metamorphic basement. Protolith of these rocks probably correspond to
438 an alternation of metapelite, sandstone and argillaceous sandstone which could correspond to
439 delta or continental shelf deposits. These rocks were metamorphosed and deformed under
440 amphibolite facies conditions reaching partial melting, and were intruded coevally by gabbros

441 and granites which now form interfoliated lenses or boudins of ortho-amphibolites and
442 orthogneisses. These metamorphic rocks show many lithological similarities with the Juffa
443 group in Oman (Mercolli et al., 2006) which corresponds to an alternation of paragneisses,
444 micaschists and ortho-amphibolites, and the Banded Gneiss Complex which corresponds to an
445 association of augengneisses and biotite-hornblende gneisses. Steep foliation planes and
446 contacts between the various units of the metamorphic basement of Socotra suggest that
447 metamorphism and deformation occurred in a compressive or in a transpressive regime. Our
448 U-Pb data in granites show scarcity of inherited cores represented by two zircons with U-Pb
449 ages at *ca.* 860 Ma. These measurements suggest that the orthogneisses, migmatites and
450 orthoamphibolites of Socotra are younger than 860 Ma. This is consistent with both datation
451 of emplacement of the protolith of the orthogneisses of the Banded Gneiss Complex and
452 datation of the metamorphic event in the Juffa Group and the Banded Gneiss Complex around
453 815 Ma obtained by Mercolli et al. (2006). However we have no elements to precise the age
454 of the protoliths of the paragneisses and quartzites that could be of Mesoproterozoic age as
455 the Juffa Group in Mirbat (Mercolli et al., 2006).

456

457 *Mafic intrusions on the Cryogenian active margin of Socotra*

458

459 Metamorphic rocks of the Socotra basement have been intruded by voluminous mafic
460 and felsic magmatic bodies. Field relationships show that felsic intrusions occur generally
461 later than the mafic to intermediate one. Mafic to intermediate bodies of Socotra show some
462 evidences of metamorphic overprint in amphibolite facies conditions. However these rocks
463 have not suffered the strong solid-state deformation recorded by the metamorphic basement
464 and are devoid of evidence of partial melting. Emplacement of mafic to intermediate
465 intrusions is thus late-kinematic and post-date metamorphic peak. The mafic to intermediate
466 plutonic sheets have been affected by an incipient solid-state deformation and show discrete
467 steeply dipping foliation planes parallel to those of the metamorphic basement. These sheets
468 correspond probably to the first mafic intrusions.

469 Mafic rocks were characterized by two distinct signatures. A signature diagnostic of a
470 depleted-arc or back arc basin ($Ce/Yb > 15$; $La_N/Nb_N > 7$) is recorded in the mafic rocks
471 belonging to the Mont-Haggier area as the Haggier layered gabbros, the basaltic lavas of the
472 Southern Haggier volcanic series and the mafic dikes of the Haggier dike swarm. The Haggier
473 mafic system allows thus the observation of a well-preserved upper part of a juvenile arc or a
474 back-arc basin (Fig. 14), with a magmatic chamber characterized by crystallization of

475 cumulates, volcanic effusions in surface and formation of a dike swarm which allowed the
476 magmas to be transferred from the magmatic chamber to the surface. The second signature of
477 mafic to intermediate rocks is characteristic of enriched-arc setting ($Ce/Yb > 15$; $La/Nb >$
478 6.5) and is evidenced in the yellowish basalt SP15A and in the diorites of the mafic to
479 intermediate plutonic sheets of Qalansya. The low contents of HREE compared to the others
480 REE of the sample SP15A (e.g. $La/Yb > 80$) could indicate the presence of garnet in the
481 source of this basalt and thus a genesis at high depth.

482 Geochemical compositions of the mafic rocks of Socotra indicate clearly a
483 suprasubduction context during the middle Neoproterozoic times (Cryogenian). The first
484 mafic rocks were emplaced in a context of an enriched-arc setting forming probably in an
485 Andean-type arc. By contrast the last emplaced mafic rocks of Socotra are characteristic of a
486 depleted-arc context which could correspond to the formation of either (i) an intra-oceanic
487 island-arc, (ii) or an island-arc above a thinned continental crust, or (iii) a back-arc basin. By
488 considering that the basement of Socotra is affected by an important HT-LP metamorphism
489 prior and coevally with the formation of the depleted-arc, we suggest that emplacement of
490 mafic rocks presenting a depleted-arc signature occurred during the opening of a back-arc
491 basin. This hypothesis provides also an explanation to understanding the contrasted structural
492 features between the gabbros cumulates characterized by moderately dipping magmatic
493 layering that can be formed during local extensional regime, and steeply dipping calc-alkaline
494 sheets that probably intruded the metamorphic basement during compressive regime.
495 Homogeneous orientation of basaltic dikes around $N60^{\circ}E$ indicates that the back-arc basin (or
496 juvenile arc) of Socotra is probably aligned along the same direction during Cryogenian times
497 (Fig. 14).

498 The sheets of mafic to intermediate rocks of Socotra correspond probably to an
499 equivalent of the Mahall Complex of Mirbat (Mercolli et al., 2006), which intruded the Juffa
500 group around 800 Ma and corresponds to original dioritic and tonalitic plutons that have
501 been deformed and recrystallised under amphibolite facies conditions. The juvenile arc
502 system of Socotra including cumulates, basaltic dikes and tholeiitic lavas could correspond to
503 an equivalent of the Tonalite Group in Mirbat which have been dated between 790 and 780
504 Ma and have never suffered solid-state deformation (Mercolli et al., 2006). The Tonalite
505 Group comprises two small bodies of hornblende bearing layered olivine gabbroic intrusions
506 which are petrographically similar to the Haggier layered gabbro body, and two kilometre-
507 scale tonalite to granodiorite layered plutons, the Hadbin and Fusht Complex. Al-Kathiri
508 (1998) suggested that the layered gabbros of Mirbat represent the parental magmas from

509 which the dioritic-tonalitic rocks of the Fusht and Hadbin complex evolved by hornblende
510 dominated fractionation. Although in Mirbat volcanic rocks that could correspond to the
511 Southern Haggier volcanic Series were never found, volcanic clasts are very abundant in the
512 sediments of the Ediacaran clastic Mirbat formation (Mercolli et al., 2006). This suggest that
513 the volcanic series representing the upper crustal level during Cryogenian times were largely
514 eroded during Ediacaran in the Mirbat basement as in the western part of the Socotra Island.

515

516 *The Cryogenian High-K calc-alkaline intrusions of Socotra*

517

518 The Sherubrub pluton intrusive in the Sherubrub gabbro and in the metamorphic
519 basement shows systematic vertical and sharp contacts with it country rocks. While
520 emplacement of this pluton is clearly post-metamorphic peak, strong high to middle
521 temperature deformation at the contact with it country-rocks associated with a sub-vertical
522 solid-state foliation suggests a syn-collisional emplacement in compressive regime. The
523 Haggier pink granite was clearly intrusive at the top of the Haggier layered gabbros (Fig. 14)
524 when these gabbros were crystallized as shown by the presence of magmatic breccia. These
525 granites have suffered a very weak deformation and emplaced at the top of the upper crustal
526 domain as shown by their textural evolution from the top to the base of the pluton. Finally the
527 Northern Haggier biotite granite crops out in a low-relief zone and we were not able to study
528 relationship between this granite and its country-rocks. This granite is intruded by some
529 basaltic dikes (Fig. 14) presenting the same geochemical affinity of the Haggier layered
530 gabbros. We can thus suggest that emplacement of the Northern Haggier biotite granite is
531 coeval with the final stage of emplacement of mafic rocks. Whereas U-Pb datation on the
532 Haggier pink granite and the Sherubrub granites present important uncertainties, ages are very
533 close and suggest in accordance with geochemical data that felsic plutons of Socotra were
534 emplaced during the same magmatic activity which occurred between 840 and 780 Ma.
535 Granites and felsic dikes present geochemical characteristics which underline their
536 transitional character although with a certain affinity with A-type granites. Also the
537 geochemical diagrams do not associate them with a distinct geodynamic setting, however
538 field relationships suggest that their emplacement occurred by the end of the Cryogenian
539 active margin formation. Orientations of microgranitic and rhyolitic dikes associated with the
540 Haggier pink granite are very close to the orientation of the basaltic dikes, and are thus
541 consistent with this hypothesis (Fig. 14). Moreover structural data in the Sherubrub pluton
542 suggest that their emplacement occurred during compressive regime. In this way, emplacement

543 of these granites could correspond to the final stage of the orogenic system evolution during
544 the closure of a back-arc system in an Andean type margin.

545 The Mirbat block presents a later calc-alkaline acidic magmatism event characterized by
546 some hectometric-scale bodies of granite and by the developement of a pegmatitic dike
547 swarm (Mercolli et al., 2006). Mirbat is thus less or not affected by the voluminous high-
548 potassic acidic magmatic event that we have observed in Socotra. Subsequently to these
549 Cryogenian events, the various units of Mirbat recorded fast exhumation (Mercolli et al.,
550 2006). The mirbat basement rocks are overlain by clastic sediments of the Mirbat formation
551 which is probably Ediacaran in age (Mercolli et al., 2006). The Hadibo series (Beydoun and
552 Bichan, 1970) are composed mainly by sandstones. These clastic series are restricted to the
553 footwall of the main transfer fault of Socotra (Fig. 2) and overlain the Cryogenian basement.
554 These series recorded probably as the Mirbat formation, Ediacaran denudation of the
555 Cryogenian basement of Oman.

556

557 *The Socotra and Mirbat basements in the case of the East-African-Antartic-Orogen*

558

559 Our study of the Neoproterozoic Socotra basement underlines formation of a juvenile
560 arc system and/or back-arc basin with some evidences of evolution from an Andean-type arc.
561 Formation of this system is late compared to a stage of HT-LP metamorphism that was
562 recorded in the older metasedimentary rocks of Socotra. Although we have insufficient
563 geochronological data to constrain the ages of the earliest events recorded in the Socotra
564 basement, we can claim on the basis of our data on Socotra and by analogy with the Mirbat
565 block in Oman: (i) that the basement of Socotra is affected by a high temperature
566 metamorphic stage whose minimum age is 815 Ma; (ii) that the metamorphic basement is first
567 intruded by mafic to intermediate rocks related to an Andean type arc and secondly by mafic
568 rocks related to a juvenile arc/or back-arc basin; and (iii) that the arc system is intruded by
569 large acidic plutons with a transitional affinity, these plutons being characterized by a
570 minimum age at 780 Ma.

571 The comparison of these characteristics with those of the various domains of the EAAO
572 provides more information. Between 700 and 600 Ma, the ANS suffered a low to medium
573 grade metamorphism linked to terrane accretion. This metamorphic stage is younger than the
574 juvenile arc formation which occurred mainly between 800 and 700 Ma (e.g. Nehlig et al.,
575 2002; Johnson and Woldehaimot, 2003). The Mozambic belt in the central and southern part
576 of the EAAO is characterized by coeval deformation and high-temperature metamorphism,

577 marked by formation of granulites, between *ca.* 600 and 540 Ma (Jacobs and Thomas, 2004).
578 The terranes of the continental Yemen show the transition between the ANS and the high-
579 grade rocks of the EAAO. The Al-Mafid and Abas gneiss terranes (Fig. 1b) recorded a major
580 *ca.* 760 Ma metamorphism and deformational event linked to terranes accretion and marked
581 by Pb loss in zircons in Archean gneisses and by zircon crystallization ages of granitic
582 gneisses (Whitehouse et al., 1998). Moreover $^{40}\text{Ar}/^{39}\text{Ar}$ ages obtained from intrusive rocks at
583 the boundary between Al-Mafhid and Abas gneiss terranes provide a lower limit of 615 Ma
584 upon terrane assembly (Whitehouse et al., 1998). The basements of Socotra and Mirbat
585 present a high-grade metamorphic stage earliest than that recorded in the pan-african EAAO.
586 The magmatic rocks of the ANS present generally a typical volcanic-arc signature while the
587 magmatism of Socotra shows a complex evolution in a short time span with very peculiar
588 geochemical signatures, that suggest an evolution from an Andean-type arc to a juvenile
589 arc/back-arc basin with later emplacement of large high-potassic plutons. Finally, structural
590 studies indicate that the major foliation planes are homogeneously oriented between N70°E
591 and N80°E in Socotra and Mirbat (Mercolli et al., 2006), whereas the main directions
592 underlined in the ANS are NS to NW-SE (Johnson and Woldehaimot, 2003, references
593 therein).

594 These comparisons suggest that the basement of Mirbat and Socotra were formed in a
595 different geodynamic setting from those of the ANS and the high-grade Mozambic belt of the
596 EAAO. Though there is a lack of paleomagnetic data to have a detailed paleogeographic
597 reconstruction of this period, we consider following Li et al. (2008) that the microcontinental
598 blocks of Afif-Abbas and Al-Mahfid were placed between India and Sahara in Cryogenian
599 times. Many evidences highlight that a Cryogenian active Andean-type margin was developed
600 on the western part of the Greater India block (Li et al., 2008). This active margin has been
601 identified by using integrated geochronological, geochemical and paleomagnetic analysis in
602 Seychelles Islands (Torsvik et al., 2001a; Tücker et al., 2001; Ashwal et al., 2002), in
603 northwestern India with the Malani Igneous Suite (Torsvik et al., 2001b) and in northeastern
604 Madagascar (Tucker et al., 1999a and b). In the Seychelles Islands this active margin is
605 marked by emplacement of two groups of granites dated around 750 Ma: the Mahé group
606 which corresponds to greyish granites and the Praslin group which corresponds to redish to
607 pinkish granites (Ashwal et al., 2002). These two granites present many petrographic and
608 geochemical similarities with both the greyish and pinkish granites of Socotra, which
609 indicates that the Neoproterozoic rocks of Socotra could also form on the Cryogenian

610 Andean-type margin of the Indian block. This hypothesis would need additional structural,
611 isotopic and more geochronological data to be tested.

612

613

614 **Conclusion**

615

616 The basement of the Socotra Island presents an early high-temperature metamorphism
617 event associated with a strong deformation and then a voluminous mafic and felsic
618 emplacement event. U-Pb data indicate that these successive events occurred during a relative
619 short time span, between 860 and 780 Ma at maximum. Mafic magmatism shows the
620 evolution from an Andean type arc to a juvenile arc/or back-arc basin. Felsic magmatism is
621 characterized by emplacement of voluminous highly potassic calc-alkaline granites forming
622 large plutons and which occurred probably at the final stage of arc history. These granites are
623 not anorogenic as it has been previously published by analogy to the peralkaline granites of
624 the ANS which were emplaced at *ca.* 600 Ma.

625 These features cannot be easily reconciled with those of the Arabian-Nubian Shield to
626 the west of Socotra and with the Mozambique Belt to the south. We propose that the Socotra
627 basement was developed on an active margin located near the Indian block in Cryogenian
628 times.

629

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634

635

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637

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778

779 **Figure caption**

780

781 Fig. 1: a) Reconstitution of the late Neoproterozoic EAAO (modified from Jacobs and
782 Thomas, 2004). Abbreviations: EF = European fragments, M = Madagascar, S = hypothesis
783 for the paleoposition of the Socotra Island at the end of the Neoproterozoic. b) Geological
784 map of the Arabian-Nubian Schield (modified from Whitehouse et al., 1998, Nelhig et al.,
785 2002; Jonhson and Woldehaimot, 2003).

786

787 Fig. 2: Geological and topographic map of Socotra. Topographic contours have been realized
788 by using the SRTM data (2001). Oligo-Miocene structures are from Razin et al. (2010).

789

790 Fig. 3: Detailed geological map of the Sherubrub area.

791

792 Fig. 4: Field photographs, microphotographs (scale bar = 2 mm); a) Ortho-amphibolitic
793 boudins in white quartzites (Sherubrub area); b) Bt+Sil+Afs±And (see Whitney and Evans,

794 2010 for mineral abbreviations) paragenesis in migmatitic paragneiss (metatexite) at the
795 contact between leucosome and melanosome (plane polarized light); c) Gabbro metamorphosed
796 in the amphibolite-facies (planar polarized light, Mont Haggier area); d) Pyroclastic basaltic
797 breccia; e) Basaltic tuff with breccia levels (planar polarized light); f) Intrusive contact of the
798 Haggier pink granite in the doleritic basalt SP15A.

799

800 Fig. 5: A: Geological transect representing the studied area which correspond to a single bank
801 and location of the photographs B. B = Field photographs or microphotographs (scale bar =
802 2mm) or field sketch: a) Hornblende-bearing medium-grained diorites with stretching
803 lineation; b) Schematic illustration of microphotograph a; c) Alternating of sheets of tonalites
804 and diorites; d) Straight contact between tonalites and granodiorites; e) Mingling between
805 dioritic and granodioritic magmas; f) Vertical stretching lineation in diorites.

806

807 Fig. 6: Field photographs or microphotographs (scale bar = 2mm): a) Haggier layered gabbro;
808 b) Cpx+Pl+Ol layered gabbro (planar polarized light, Mont Haggier area); c) Intrusion of
809 granodiorite sills in the Sherubrub gabbros; d) Schematic illustration of photograph c.

810

811 Fig. 7: Field photographs or microphotographs (scale bar = 2mm): a) Sherubrub granite with
812 weak sub-solidus deformation (as = automorphic sphene, crossed polars); b) Protomylonitic
813 microstructure in Sherubrub granite (crossed polars); c) Amphibolitic rafts in pinkish
814 Sherubrub granite; d) Pegmatite swarm in gabbros at the periphery of the Sherubrub granite;
815 e) Quartz, zoned plagioclase and biotite in the Northern Haggier biotite granite (crossed
816 polars); f) Texture of the Haggier pink granite (planar polarized light).

817

818 Fig. 8: Field photographs or microphotographs (scale bar = 2mm): a) Magmatic breccia with
819 element of gabbro in Haggier pink granite matrix; b) Arfvedsonite (in blue) in the Haggier
820 pink granite (planar polarized light); c) Chessboard texture in quartz of the Haggier pink
821 granite (crossed polars); d) Basaltic dike intrusive in the Northern-Haggier Bt granite; e)
822 Doleritic texture in basaltic dikes formed mainly by hornblende and plagioclase; f) Rhyolitic
823 dikes intrusive in the Haggier layered gabbros.

824

825 Fig. 9: Stereograms of dikes orientation in the Mont-Haggier area (Schmidt lower
826 hemisphere), a) Dikes of pinkish microgranite (36 measurements), b) Rhyolitic dikes (33), g)
827 Basaltic dikes (17).

828

829 Fig. 10: Total alkali silica diagram (LeBas et al., 1986) for magmatic rocks from Socotra
830 Island. For the Mont-Haggier samples squares represent the mafic lavas and dikes, circles
831 represent the granites and the triangle correspond to the basalt SP15A.

832

833 Fig. 11: A and B. Patterns of mafic cumulates and related volcanic and subvolcanic rocks
834 from Mont-Haggier area normalised to Silicate Earth (McDonough and Sun, 1995) and
835 Chondrites (Sun and McDonough, 1989), respectively. C. Patterns of diorites from Qalansya
836 area and of the basalt SP15A from the Mont-Haggier area normalised to Silicate Earth
837 (McDonough and Sun, 1995). D. Discrimination diagram (Cabanis and Lecolle, 1989) for
838 subvolcanic to volcanic mafic rocks of Mont-Haggier area and the diorite of Qalantsya.

839

840 Fig. 12: A. ANK vs ACNK (Maniar and Piccoli, 1989) discrimination diagram for granitic
841 bodies of Sherubrub area (grey circles) and Mont-Haggier area (dark circles). B. Patterns of
842 granitic rocks from Socotra Island normalised to Silicate Earth (McDonough and Sun, 1995).
843 C. Zr+Nb+Y+Ce vs $1000 \times \text{Ga}/\text{Al}$ diagram (Whalen et al., 1987) of granitic rocks from Socotra
844 island. D. Discrimination diagram (Pearce et al., 1984) of granitic rocks from Socotra island.

845

846 Fig. 13: U-Pb concordia diagrams for (a) Haggier pink granite (sample HAEU-1), (b) gabbro
847 enclaves in Haggier pink granite (sample HAEU-2) and (c) Sherubrub granite (sample SHEU-
848 1).

849

850 Fig. 14: Interpretative block-diagram of the Mont-Haggier area.

851

852 Table 1: Major and trace elements compositions of selected magmatic rocks from Socotra
853 island. < D.L. below detection limit. $\text{Mg}\# = 100 \times \text{MgO}/(\text{MgO} + \text{FeO}_{\text{Total}})$, on a molar basis.

854

855 Table 2: Analytic results of U-Pb dating using ion microprobe (CAMECA IMS 1270) on
856 zircons from samples HAEU-1, HAEU-2 and SHEU-1.

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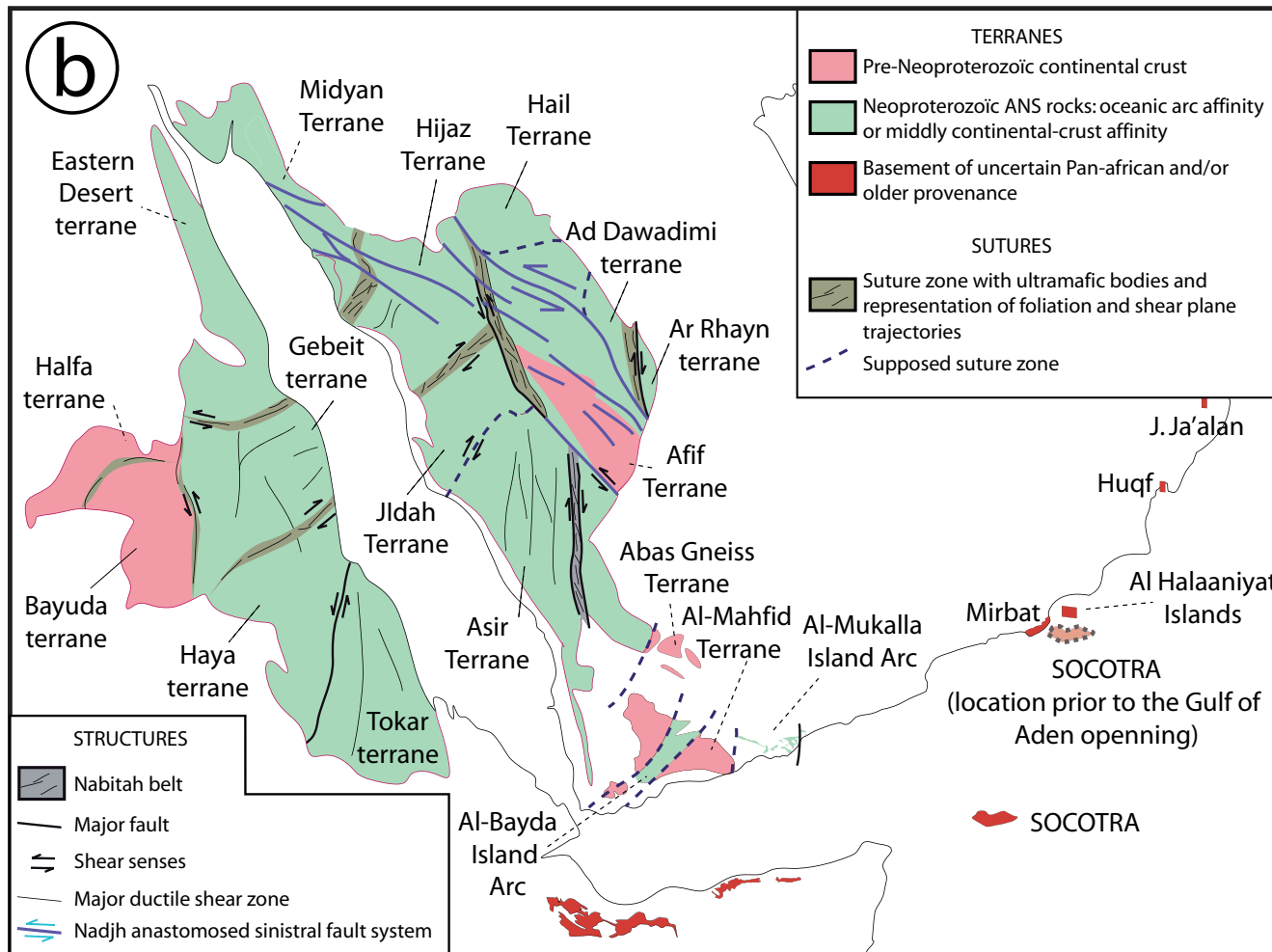
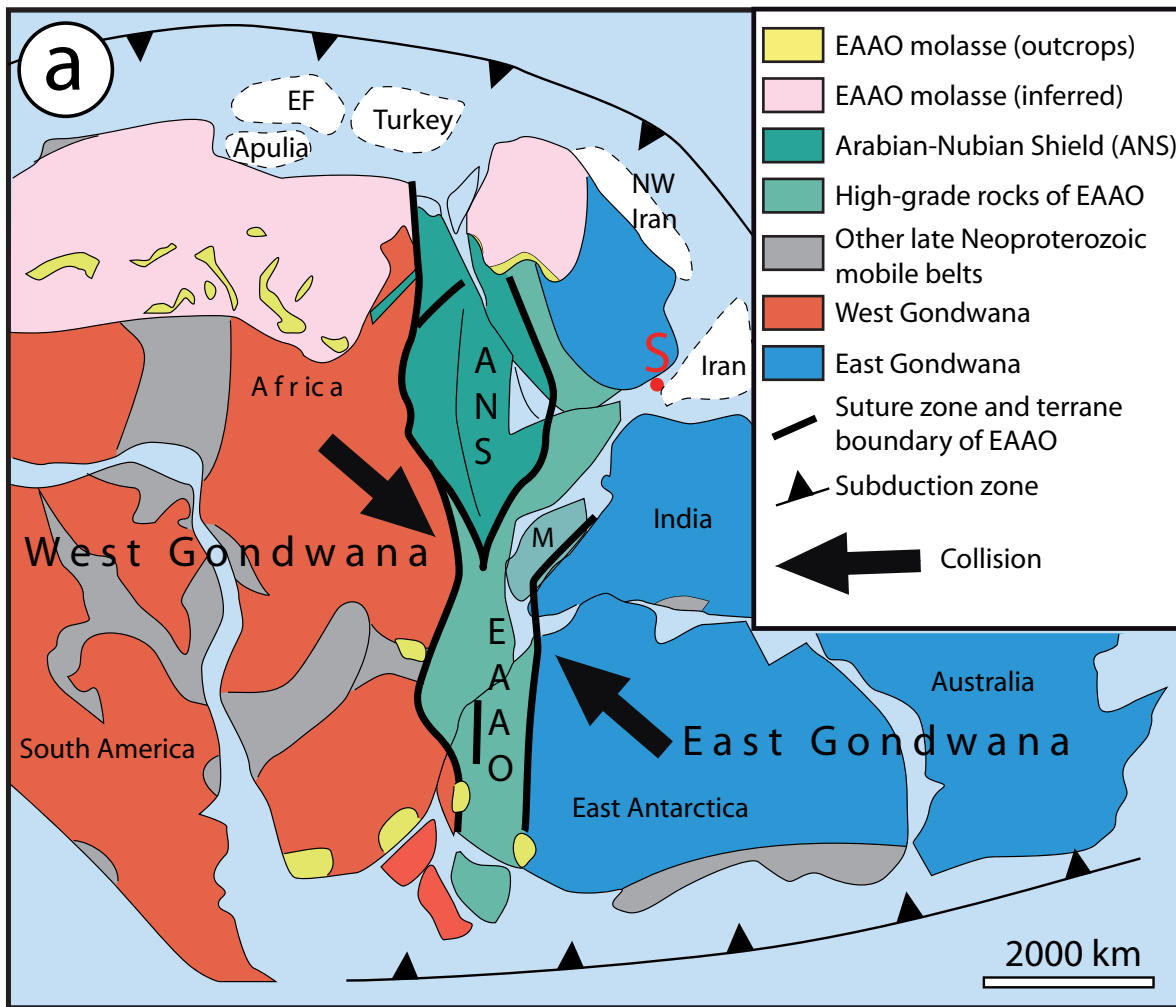
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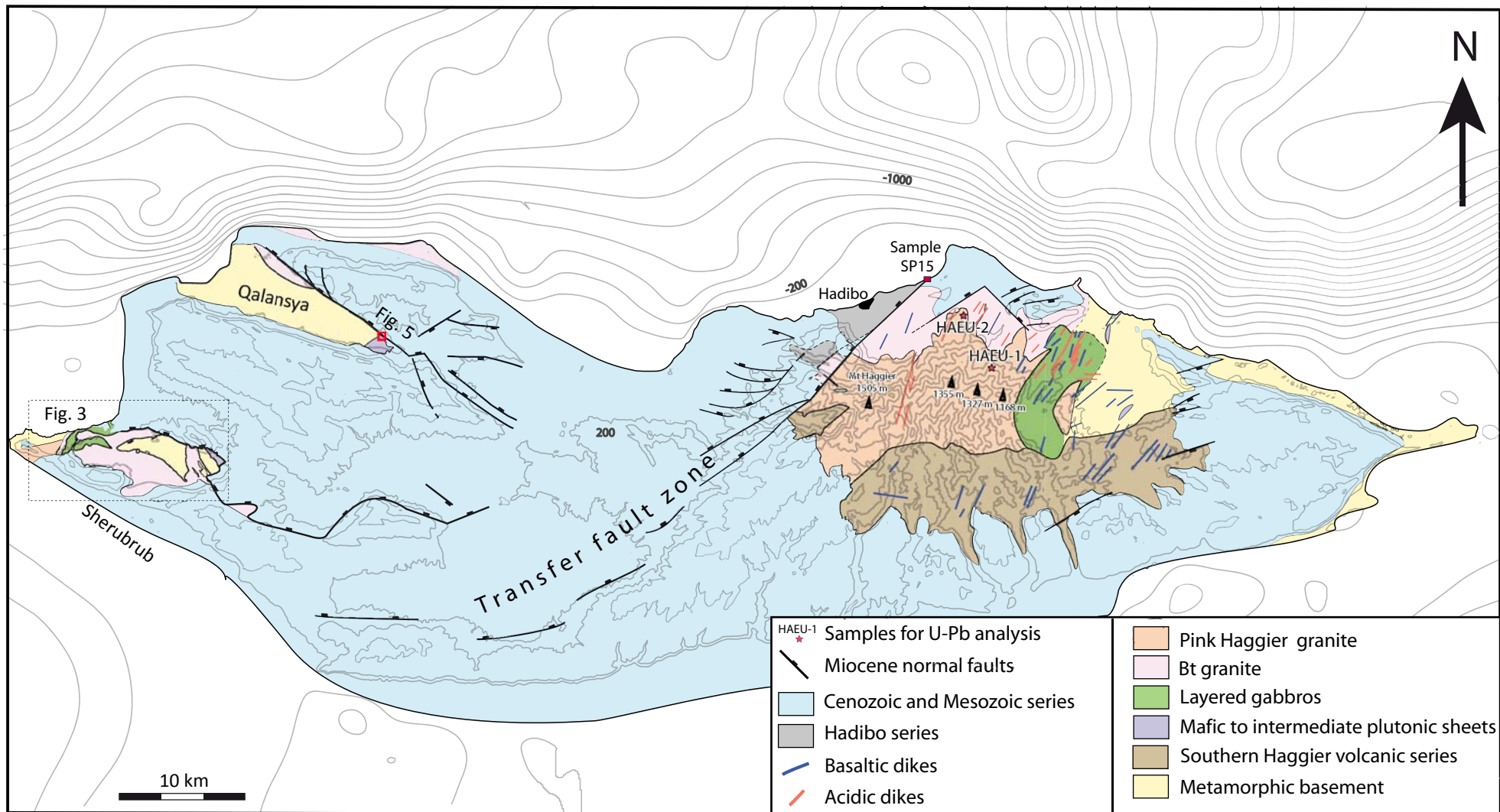
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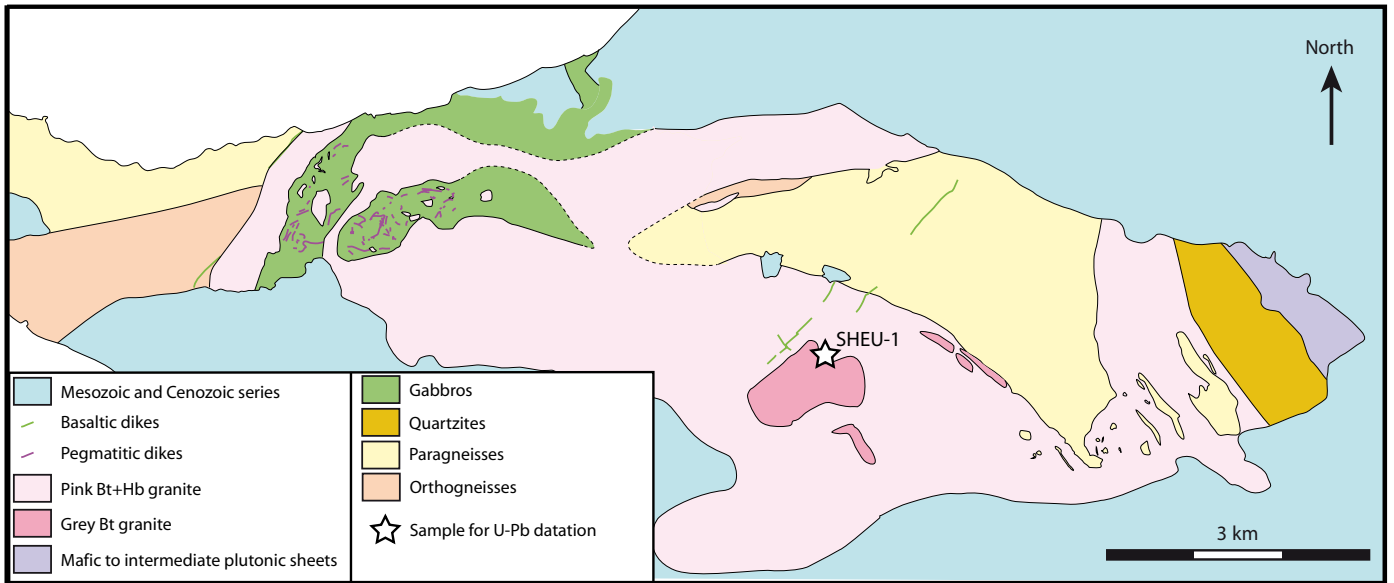
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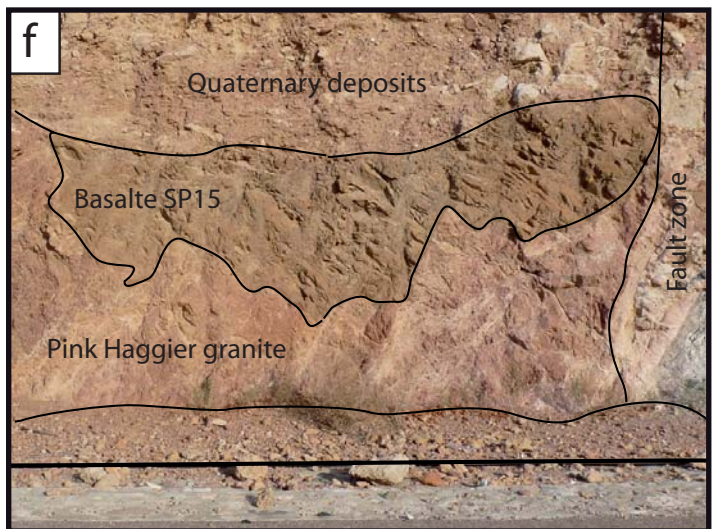
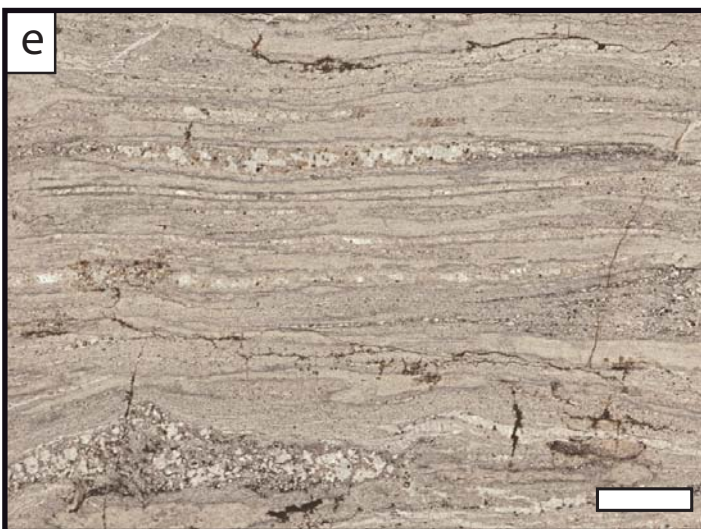
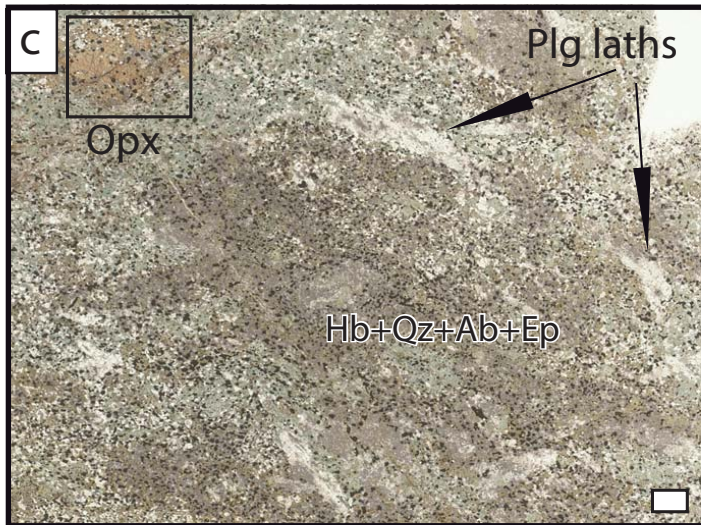
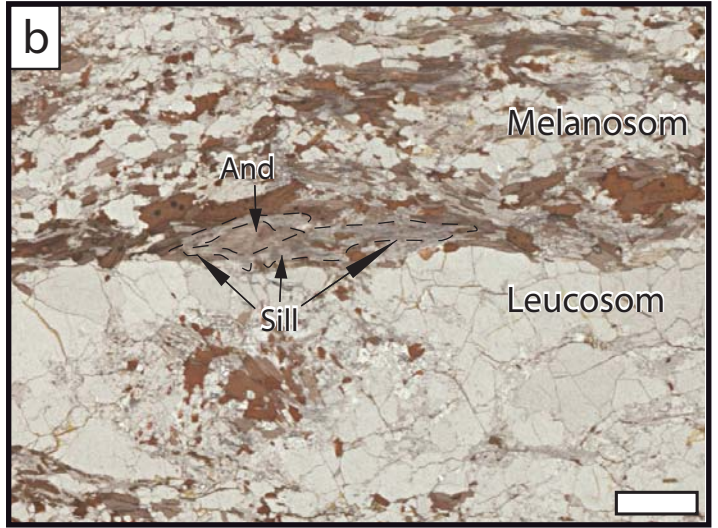
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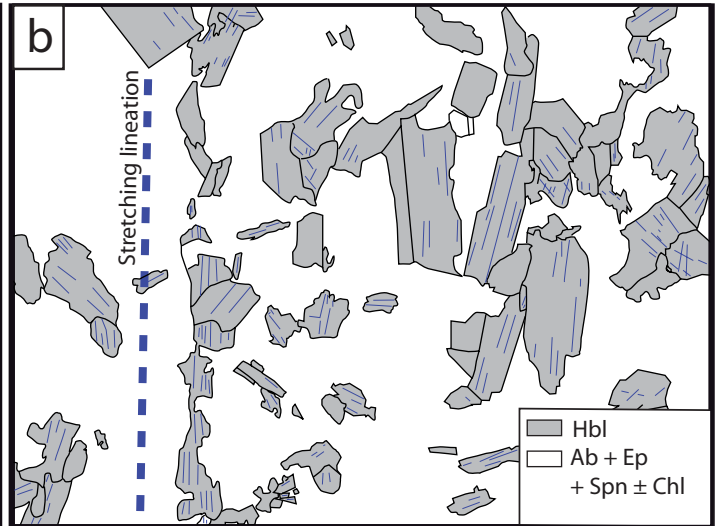
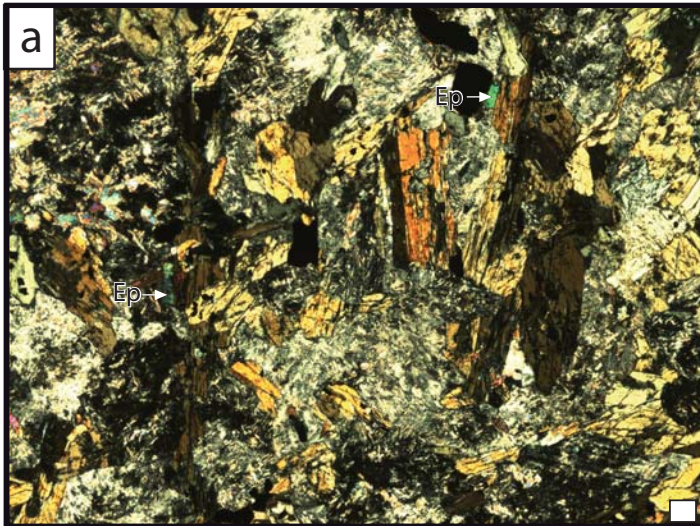
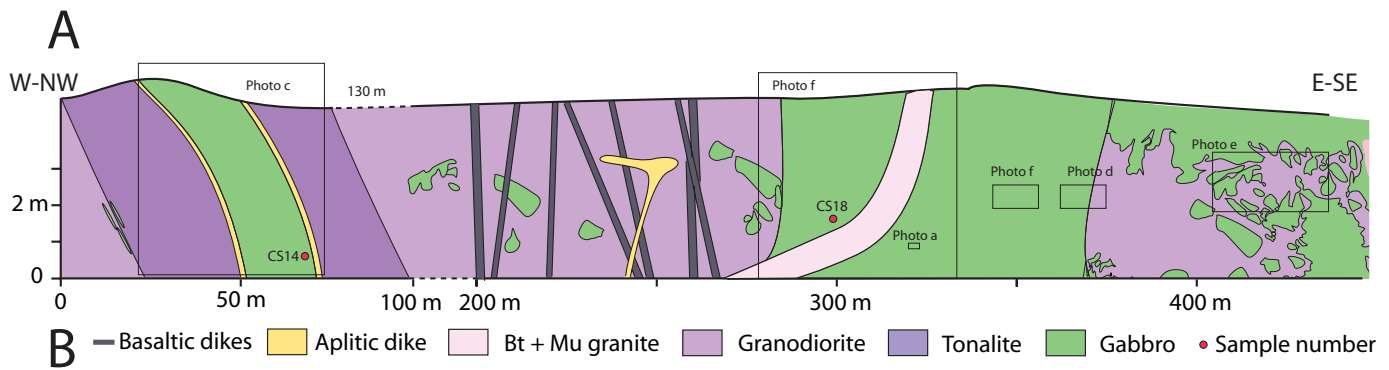
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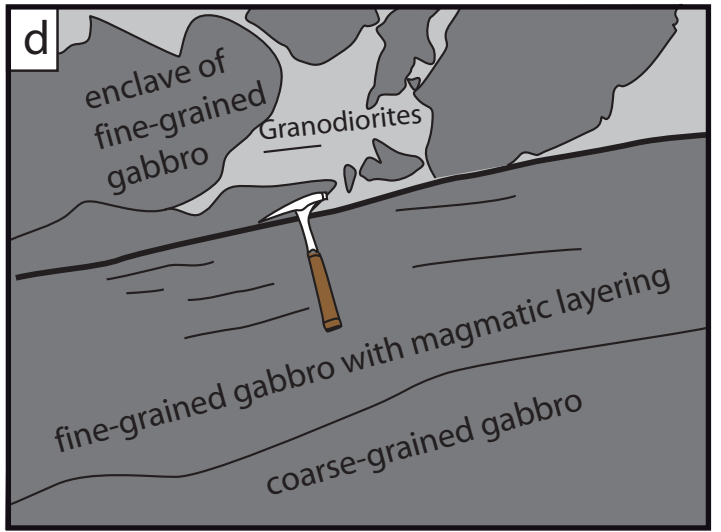
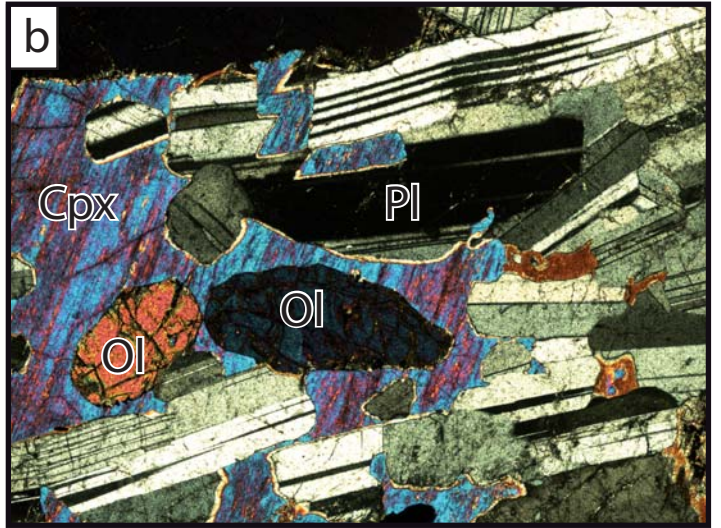


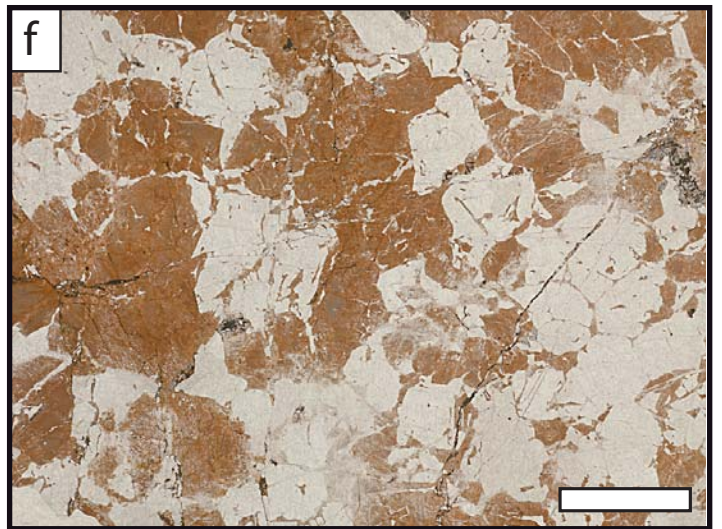
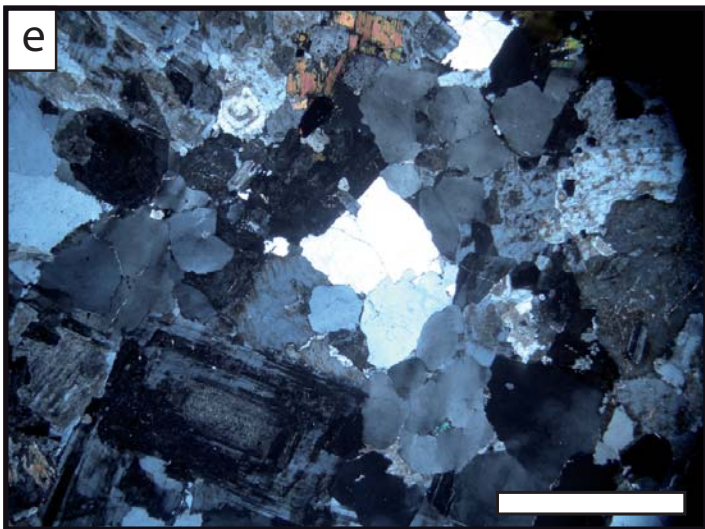
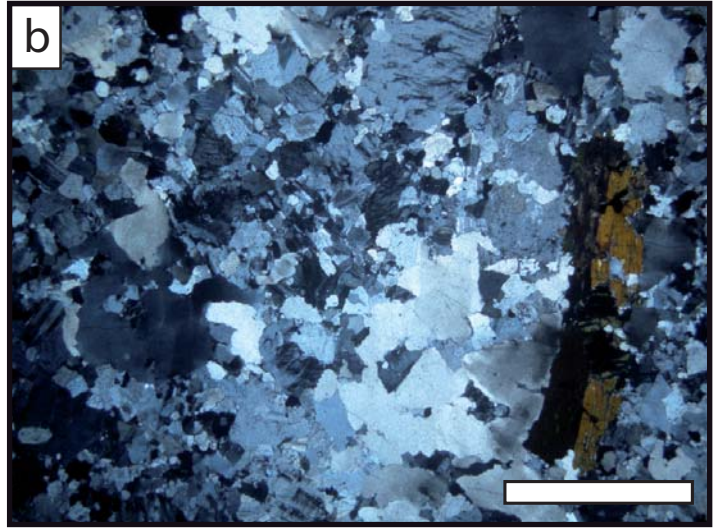
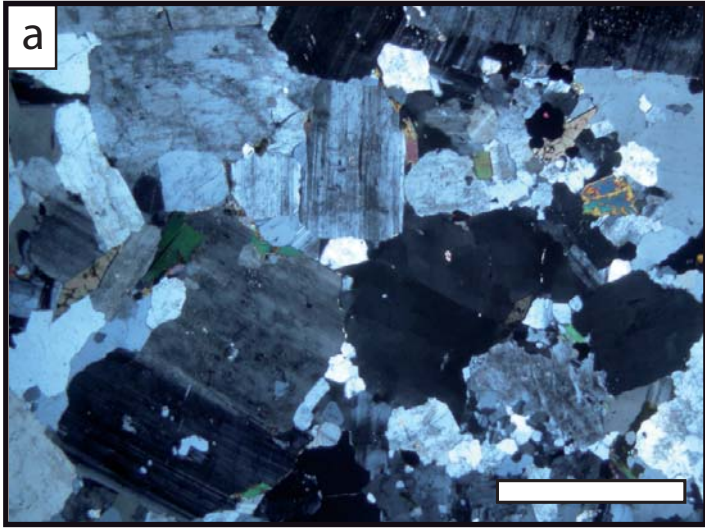


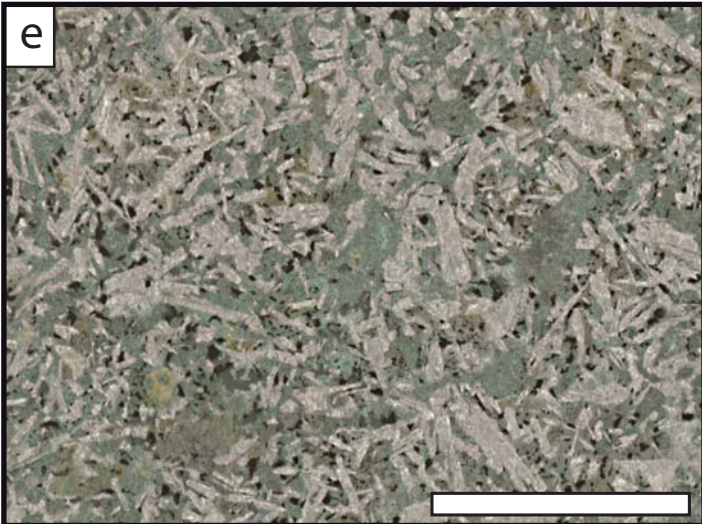
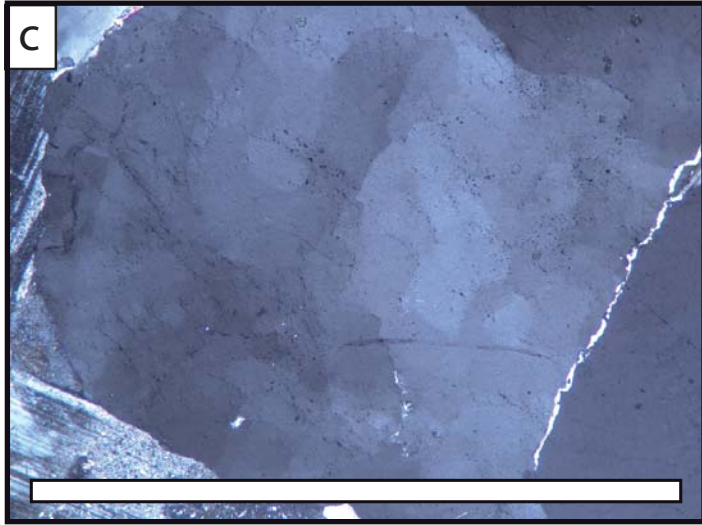
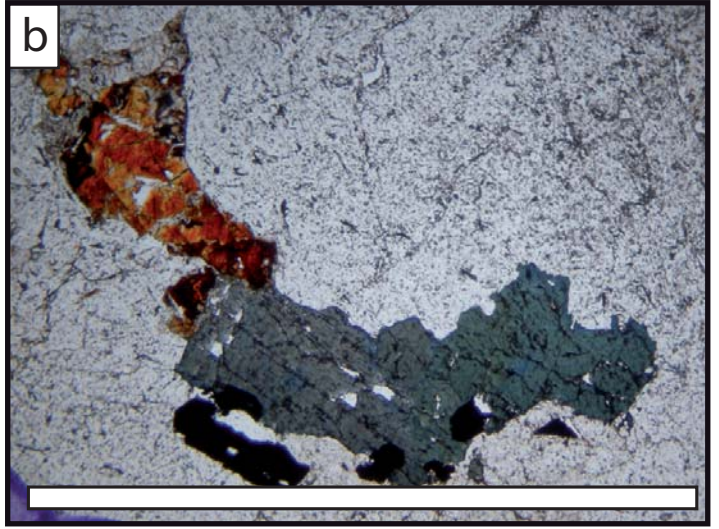


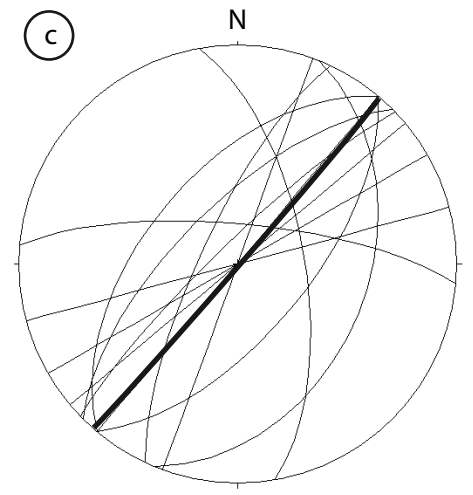
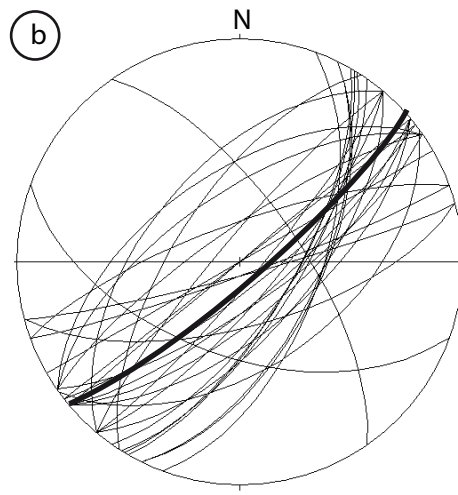
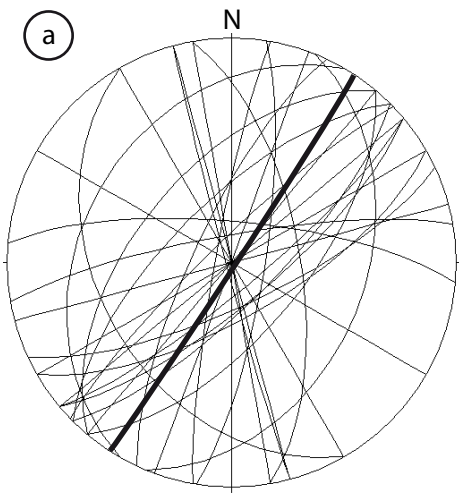


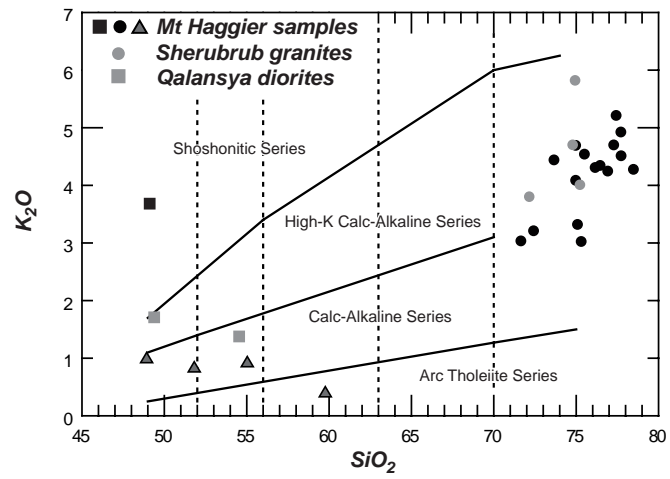


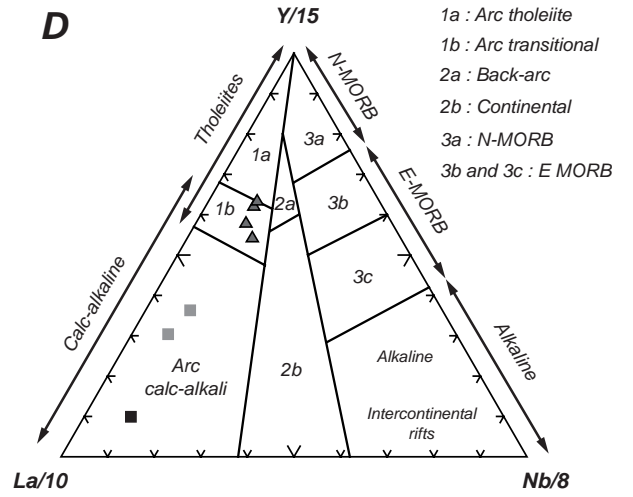
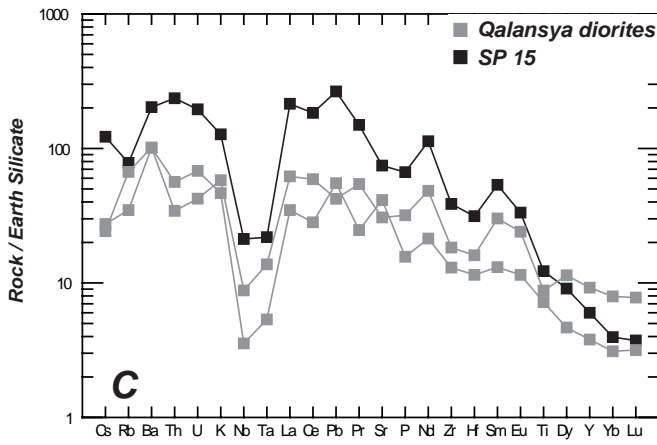
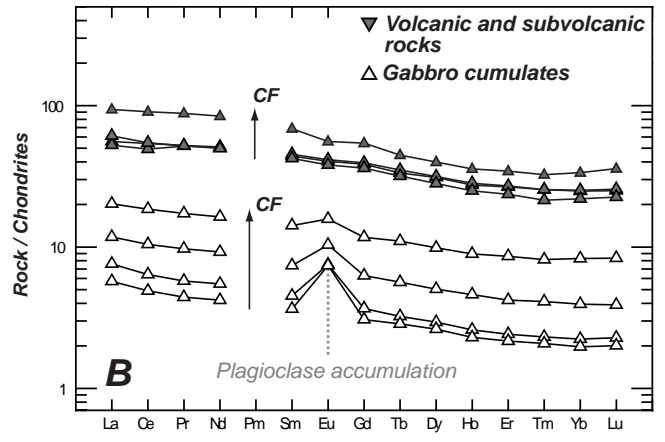
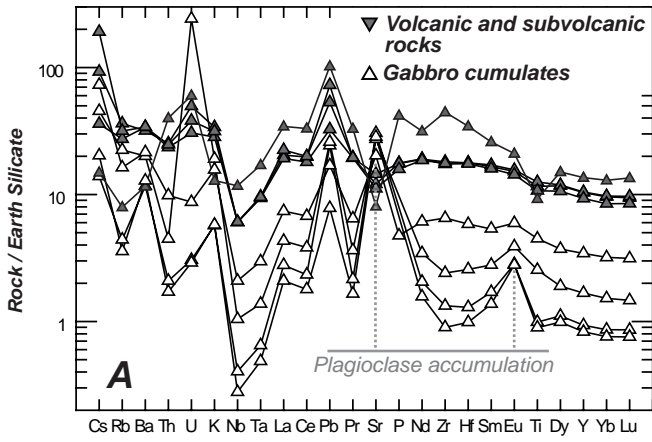


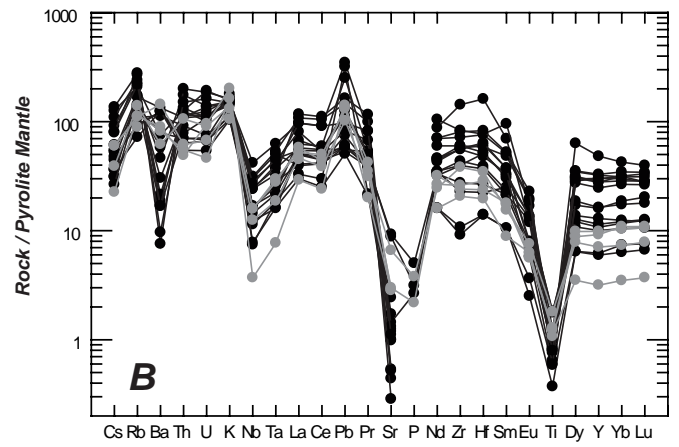
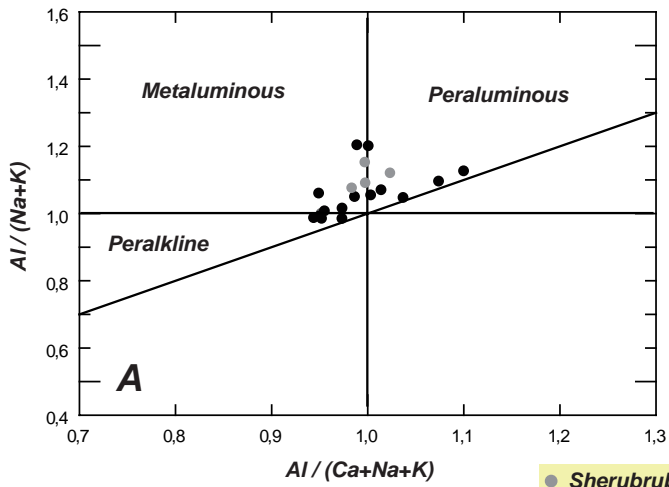




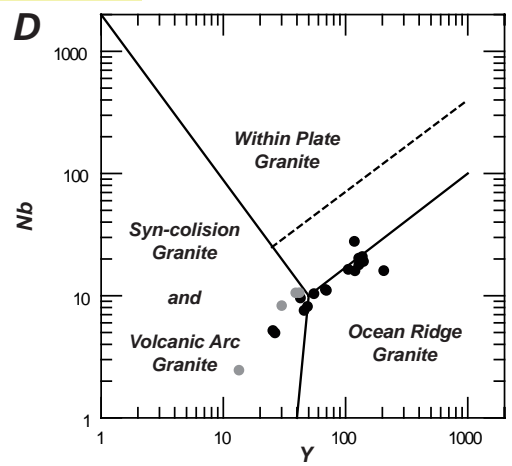
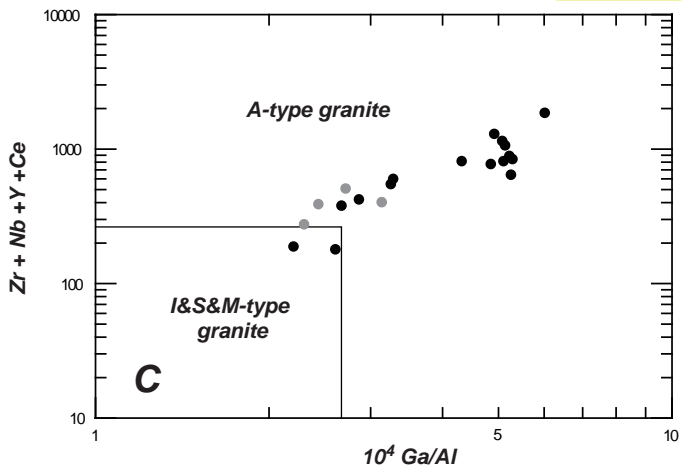


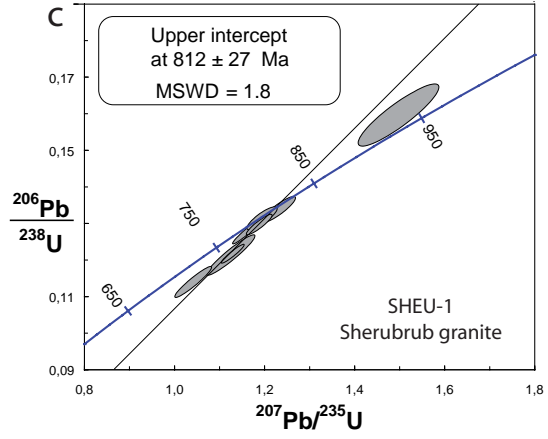
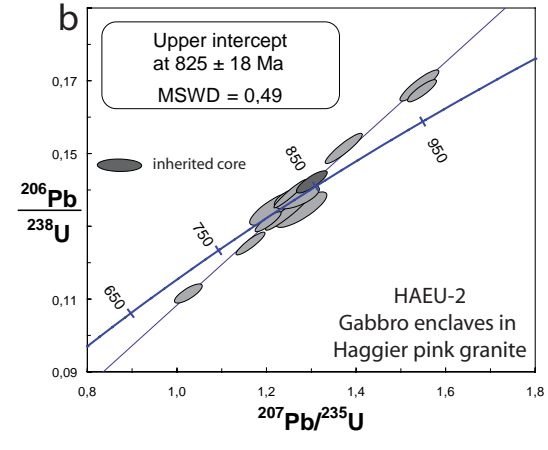
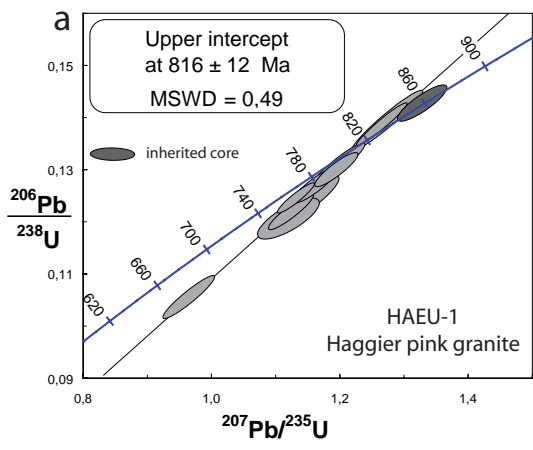





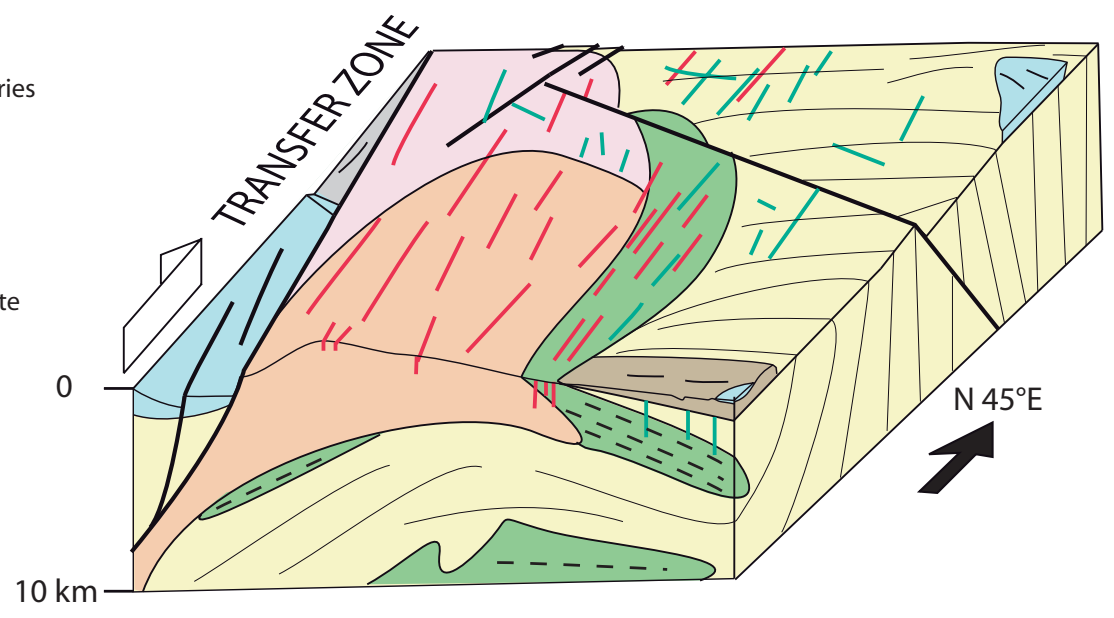


● Sherubrub granites
● Mt Haggier granites and felsic dikes





-  Tertiary fault zones
-  Cenozoic and Mesozoic series
-  Hadibo series
-  Acidic dikes
-  Haggier pink granite
-  Northern-Haggier Bt granite
-  Basaltic dikes
-  Southern volcanic series
-  Haggier layered gabbros with magmatic layering
-  Metamorphic basement with foliation trajectories



Echantillon	Domain	Contents (ppm)				Measured $^{204}\text{Pb}/^{206}\text{Pb}$	% Pbc	Corrected ratios						Ages (Ma)				
		Pb	U	Th	Th/U			$^{207}\text{Pb}/^{206}\text{Pb}$	$\pm \sigma$	$^{206}\text{Pb}/^{238}\text{U}$	$\pm \sigma$	$^{207}\text{Pb}/^{235}\text{U}$	$\pm \sigma$	Err. Corr	$^{206}\text{Pb}/^{238}\text{U}$	$\pm \sigma$	$^{207}\text{Pb}/^{235}\text{U}$	$\pm \sigma$
HAEU1-a	Core	5	44	34	0.77	2.46E-04	0.23%	0.0667	0.0204	0.1379	0.0022	1.2680	0.0329	0.62	833	13	831	15
HAEU1-a	Rim	6	52	41	0.79	1.04E-04	0.00%	0.0690	0.0171	0.1344	0.0032	1.2779	0.0378	0.82	813	18	836	17
HAEU1-b	Core	47	394	416	1.06	1.70E-05	0.00%	0.0655	0.0058	0.1388	0.0025	1.2534	0.0240	0.95	838	14	825	11
HAEU1-b	Rim	52	452	432	0.95	3.28E-04	0.54%	0.0653	0.0156	0.1346	0.0030	1.2121	0.0326	0.82	814	17	806	15
HAEU1-c	Core	18	159	145	0.92	6.67E-05	0.04%	0.0666	0.0079	0.1311	0.0018	1.2038	0.0193	0.87	794	10	802	9
HAEU1-d	Core	18	123	109	0.89	5.50E-05	0.03%	0.0663	0.0083	0.1686	0.0028	1.5412	0.0288	0.90	1005	16	947	11
HAEU1-d	Rim	41	315	387	1.23	4.29E-05	0.04%	0.0658	0.0081	0.1513	0.0028	1.3719	0.0274	0.91	908	15	877	12
HAEU1-e	Core	25	177	233	1.32	3.05E-05	0.00%	0.0671	0.0068	0.1672	0.0020	1.5474	0.0215	0.87	997	11	949	9
HAEU1-e	Rim	44	407	673	1.65	1.97E-04	0.31%	0.0675	0.0081	0.1251	0.0021	1.1645	0.0215	0.90	760	12	784	10
HAEU1-f	Rim	17	175	199	1.14	1.39E-04	0.15%	0.0669	0.0123	0.1113	0.0017	1.0257	0.0204	0.79	680	10	717	10
HAEU1-f	Core	27	222	274	1.23	1.29E-04	0.17%	0.0664	0.0097	0.1421	0.0020	1.3012	0.0226	0.83	856	12	846	10
HAEU2-a	Core	17	146	85	0.58	1.53E-05	0.00%	0.0662	0.0054	0.1387	0.0027	1.2654	0.0257	0.96	837	15	830	11
HAEU2-a	Rim	12	116	58	0.50	4.51E-05	0.03%	0.0659	0.0060	0.1245	0.0019	1.1317	0.0187	0.93	756	11	769	9
HAEU2-b	Core	14	133	75	0.56	1.03E-04	0.13%	0.0671	0.0138	0.1242	0.0030	1.1496	0.0323	0.87	755	17	777	15
HAEU2-b	Rim	15	170	76	0.45	1.43E-04	0.20%	0.0664	0.0097	0.1054	0.0027	0.9645	0.0266	0.94	646	16	686	14
HAEU2-c	Core	11	103	48	0.47	2.22E-04	0.32%	0.0674	0.0184	0.1205	0.0026	1.1202	0.0318	0.76	734	15	763	15
HAEU2-d	Core	10	80	40	0.50	9.80E-05	0.09%	0.0675	0.0099	0.1427	0.0023	1.3284	0.0252	0.85	860	13	858	11
HAEU2-d	Rim	13	112	58	0.51	6.50E-05	0.05%	0.0663	0.0079	0.1394	0.0038	1.2744	0.0363	0.96	841	22	834	16
HAEU2-e	Core	13	115	61	0.53	1.10E-04	0.13%	0.0664	0.0083	0.1309	0.0027	1.1981	0.0268	0.93	793	15	800	12
HAEU2-e	Rim	13	120	61	0.51	9.22E-05	0.09%	0.0667	0.0083	0.1299	0.0022	1.1948	0.0222	0.90	787	12	798	10
HAEU2-f	Core	16	145	88	0.61	1.63E-04	0.24%	0.0664	0.0092	0.1268	0.0022	1.1604	0.0227	0.88	769	12	782	11
HAEU2-f	Rim	18	169	101	0.59	2.66E-04	0.42%	0.0669	0.0083	0.1218	0.0023	1.1236	0.0231	0.91	741	13	765	11
SHEU1-a	Rim	24	217	252	1.16	8.90E-06	0.00%	0.0660	0.0051	0.1280	0.0025	1.1644	0.0236	0.97	776	14	784	11
SHEU1-b	Rim	8	68	16	0.23	1.41E-04	0.15%	0.0660	0.0098	0.1312	0.0021	1.1940	0.0222	0.85	795	12	798	10
SHEU1-d	Core	9	88	67	0.77	8.26E-06	0.00%	0.0674	0.0048	0.1215	0.0017	1.1293	0.0170	0.95	739	10	767	8
SHEU1-d	Rim	15	134	110	0.82	6.13E-06	0.00%	0.0666	0.0067	0.1295	0.0019	1.1885	0.0189	0.91	785	11	795	9
SHEU1-e	Core	11	108	118	1.09	5.02E-04	0.80%	0.0641	0.0298	0.1233	0.0028	1.0909	0.0406	0.60	750	16	749	20
SHEU1-e	Rim	15	141	137	0.97	5.68E-04	0.92%	0.0650	0.0573	0.1277	0.0022	1.1448	0.0685	0.29	775	13	775	32
SHEU1-f	Core	12	101	89	0.87	1.06E-05	0.00%	0.0668	0.0093	0.1336	0.0024	1.2314	0.0245	0.88	809	13	815	11
SHEU1-f	Rim	13	97	83	0.86	5.07E-06	0.00%	0.0681	0.0176	0.1596	0.0056	1.4971	0.0591	0.90	954	31	929	24