Differences in gas venting from ultramafic-hosted warm springs: the example of Oman and Voltri Ophiolites

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Abstract:

Serpentinisation of mantle rocks, leading to natural venting of hydrogen and methane, has been reported to occur at the global scale, wherever fluids percolate in ultramafic formations. Here we compare gas composition from two on-land, low-temperature, and hyper-alkaline springs hosted on ultramafic rocks in the ophiolite massifs of the Sultanate of Oman and the Ligurian Alps (Voltri Group, Genoa region, Northern Italy). These two settings exhibit similar chemical and mineralogical features but show diverse styles of gas venting. Commonly to all hyper-alkaline springs, gases are characterised by relatively high N₂, very low O₂ and CO₂ concentrations, and a strong enrichment in H₂ and CH₄. The comparison between Oman and Liguria highlights a high variability of the H₂/CH₄ ratios whereby the gas phase of the Oman Ophiolite is enriched in H₂ whereas being CH₄ enriched in the Voltri Ophiolite. These results combined with literature data define three groups that may reflect different stages of serpentinisation producing fluids from hydrogen-dominated to methane-dominated. The origin of these distinct groups might lie in the difference of the mineralogical composition of the rocks within which the fluids circulate, on the degree of alteration of the rocks and finally on the geological/metamorphic history of the ophiolite.

Keywords: Methane ; hydrogen ; gas composition ; hyper-alkaline springs ; Oman ; Voltri Massif

1. Introduction

Natural venting of hydrogen (H2) and associated methane (CH4) is a common feature of fluid rock interaction, and in particular of serpentinisation of ultramafic rocks. In the last ten years, this phenomenon has gained significant interest in the scientific and industrial communities, from prebiotic chemistry and astrobiology to energy and mineral resource exploration. Serpentinisation reactions – i.e. the alteration of mantle peridotites by fluids of various origin (e.g. meteoric, seawater) – may naturally produce high amounts of H2, which leads subsequently to the abiotic formation of CH4 and possibly higher hydrocarbon (HC) chains via Fischer-Tropsch-Type (FTT) reactions. This FTT process occurs when faults allow fluid penetration within the peridotites and have been identified in different settings: along fracture zones, passive margins, at the wall of rift valleys, at mid-ocean ridges (e.g., Cannat et al., 2010; Charlou et al., 2010), and also in deep igneous and metasedimentary rocks (Sherwood Lollar et al., 2006).

51 It is now clear that wherever ultramafic rocks are percolated by fluids, alkaline springs 52 are present. In the Deep Sea, a few hydrothermal vents lying on ultramafic rocks are known today, including either high-temperature vent fields (>350°C) such as Rainbow, Logatchev 53 54 and Ashadze (Charlou et al., 2002; Fouquet et al., 2008; Lein et al., 2000), or low-temperature 55 ones (~100°C) such as Lost City vent field (Kelley et al., 2005; Proskurowski et al., 2008). In 56 recent years, several ones were identified and/or suspected on ultra-slow spreading ridges, e.g. 57 Knipovich, Gakkel and Cayman ridges (Bach et al., 2002; Connelly et al., 2012). On-land, 58 percolation of meteoric water in obducted ophiolite massif produces low-temperature hyper-59 alkaline springs, which are considered as "analogues" of the Lost City vent field off the Mid-60 Atlantic Ridge. They are reported in Oman (Chavagnac et al., 2013a, 2013b; Neal and Stanger, 61 1983; Sano et al., 1993), Philippines (Abrajano et al., 1988), Turkey (Etiope et al., 2011; 62 Hosgormez, 2007), Italy (Chavagnac et al., 2013a, 2013b; Cipolli et al., 2004, Schwarzenbach 63 et al., 2013), California (Blank et al., 2009), Cyprus (Neal and Shand, 2002), Greece (Etiope 64 et al., 2013a), Portugal (Etiope et al., 2013b; Marques et al., 2008), British Columbia (Power 65 et al., 2007), New Zealand (Lyon et al., 1990) and New Caledonia (Boulart et al., 2012; 66 Launay and Fontes, 1985), amongst others. The main differences between submarine vents 67 and on-land warm springs are the nature of the percolating water, i.e. seawater and meteoric 68 fluids, respectively, as well as venting temperature, which is lower than 100°C in the 69 continental settings. This reveals that the development of hydrogen and methane venting in 70 ultramafic-hosted environment may encompass a broad range of conditions (pressure, 71 temperature, water/rock ratio, etc...) leading to various degrees of advancement in the 72 chemical reactions for the production of natural H₂ and CH₄.

Serpentine results from the interaction of water with Fe-Mg rich minerals, i.e. olivine and pyroxene. During the serpentinisation reactions, Fe^{2+} contained in olivine is oxidised to Fe³⁺ (which may lead to the formation of secondary magnetite) while water reduction produces hydrogen and aqueous hydroxyl responsible for high pH values. Hydrocarbon gases may be generated abiotically via FTT reactions (i.e. Sabatier reaction) through the reduction of carbon dioxide by H_2 to form CH₄ and longer-chain HC (Holm and Charlou, 2001; Ingmanson and Dowler, 1977; Konn et al., 2009a, 2009b; Shock, 1990). Serpentinisation is thus accompanied by the production of methane and light alkanes (Ague, 2000). Note that low-temperature hyper-alkaline waters have lost all their dissolved inorganic carbon (DIC) during the course of serpentinisation reactions (Chavagnac et al., 2013a; Sader et al., 2007).

83 Large quantities of CH₄ and lesser amounts of light alkanes can be produced by FTT 84 reactions when they are catalysed by magnetite, chromite or awaruite at temperatures around 85 300°C (Berndt et al., 1996; McCollom and Seewald, 2001, 2007), but also at lower 86 temperatures (de Boer et al., 2007; Etiope et al., 2013a; Hosgormez, 2007, Neubeck et al., 87 2011). Nevertheless, the sole abiotic origin of HC in ultramafic environments remains 88 questionable in the sense that biologically produced methane can be present at the same 89 locations and it is difficult to distinguish between the two processes (Bradley and Summons, 90 2010; Etiope et al., 2011; Konn et al., 2009a, 2009b; Sherwood Lollar et al., 2006).

91 In this contribution, we report new data on the gas chemical composition from two on-92 land, low-temperature, and hyper-alkaline warm springs hosted on ultramafic rocks of the 93 ophiolite massifs of the Sultanate of Oman and the Ligurian Alps (Voltri group, Genoa region, 94 Northern Italy). The present dataset complements the ones of Neal and Stanger (1983) and 95 Sano et al. (1993), who first reported the composition of the gases venting in the hyper-96 alkaline springs of Oman. The composition of gases bubbling in the Ligurian springs is given by Boschetti et al. (2013), Cipolli et al. (2004) and Schwarzenbach (2011). Alteration of these 97 98 two ophiolites by meteoric waters generates hyper-alkaline springs with a pH > 9 and similar 99 chemical and mineralogical features (Chavagnac et al., 2013a, 2013b) but with great 100 differences in their gas chemical composition.

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102 **2.** Geological Settings

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104 2.1.Oman Ophiolite

The Semail Nappe in the Northern Mountain of the Sultanate of Oman (Figure 1) is one of the 105 largest (~ 30.000 Km²) and best exposed (thanks to semi-arid climate) fragment of the oceanic 106 107 Tethys lithosphere on land (Glennie et al., 1973). Its geological evolution includes magmatic 108 accretion along an oceanic spreading centre in a short time laps from ~97 to ~95 Ma (Goodenough et al., 2010; Rioux et al., 2013; Tilton et al., 1981; Tippit et al., 1981) 109 110 immediately followed by intra-oceanic thrusting in a near-ridge environment (Boudier et al., 111 1985; Lanphere, 1981; Montigny et al., 1988). Its obduction onto the Arabian margin 112 occurred about 70 m.y. ago (Glennie et al., 1973). Petrological and geochemical signatures of 113 igneous rocks of this ophiolite are clearly MORB-related in several districts (Benoit et al., 1996; Ceuleneer et al., 1996; Python and Ceuleneer, 2003; Python et al., 2008). Elsewhere 114 115 parent melts have a more ambiguous signature that can be attributed to nascent subduction 116 (Pearce et al., 1981) or to incorporation of very-high temperature hydrothermal water in a 117 spreading setting (inducing hydrated melting and modifying liquid lines of descent). A thick 118 pile of neritic carbonates deposited on the top of the ophiolite after subsidence of the Oman 119 margin (Glennie et al., 1973). The present exposure of the ophiolite is related to recent (Mio-120 Pliocene) uplift of the Oman mountains, likely a prelude to the future collision with Eurasia 121 (Fournier et al., 2006). Alkaline springs are essentially located in two particular settings: i) at 122 the basal contact between autochthonous sediments and mantle peridotites, and ii) at the 123 contact between the mantle peridotites and the cumulates from the crustal section (the 124 paleo"Moho"). Rare springs are located within the mantle or within the crustal section, away 125 from these major geological discontinuities. When present, these springs and gas vents are 126 located along faults and shear discontinuity zones in the serpentinised ultramafic section, but 127 also in less altered peridotites and gabbroic sections. Release of gases can be directly seen 128 from bubbling within the wadi bed (local Arabic name for small rivers). Gas discharge is also 129 evidenced by pockmarks forming in the precipitates accumulating in the riverbed (Chavagnac 130 et al., 2013a). Neal and Stanger (1983) estimated the gas flows from 10 mL/sec up to 10 L/sec 131 at Nizwa and B'lad (station 29) for the highest flows (Figure 1).

Although the present climate in Oman is quite dry (117 mm/year average annual rainfall; Kwarteng et al., 2009), the region was characterised rather recently (a period centred on 8.200 years BP) by a much more humid climate with intense rain events related to the end of the last glacial period (Burns et al., 2001). It is likely that a large proportion of the underground aquifers of Oman were refilled during this period, enabling the perennial flow of hyper-alkaline waters (Clark and Fontes, 1990; Dewandel et al., 2005).

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139 2.2. Voltri Group, Liguria, Italy

140 In Liguria, the studied area is located in the Voltri Massif, which is part of the Penninic 141 ophiolites (Internal Ligurides) and forms the transition between the Alps and the Appenine 142 (Figure 2). It is the largest ophiolite massif in the Alps-Appenine system (Brouwer et al., 2002) 143 composed of three main units: i) a calcschist unit consisting of blueschist to eclogite-facies 144 calcareous metasediments, metavolcanics and slices of serpentinites (Voltri-Rossiglione Unit), 145 ii) a serpentinite unit with highly-altered peridotites including eclogitic metagabbros and 146 metabasalts (Beigua Unit) and, iii) a lherzolite unit (Erro-Tobbio Unit) (Brouwer et al., 2002; 147 Bruni et al., 2002; Capponi et al., 1994; Cipolli et al., 2004; Vignaroli et al., 2010).

148 The Voltri group is derived from the Western Tethys ocean that separated European 149 and Adriatic plates in the Middle Jurassic (Bill et al., 2001). It is considered as on-land 150 analogues of oceanic lithosphere originated at slow-spreading ridges (Barrett and Spooner, 151 1977). Indeed, the chemistry of the Beigua metagabbros indicates a tholeitic origin typical of
152 mid-ocean ridges magmatism (Brouwer et al., 2002). The Piedmont-Ligurian Ocean was later
153 (Early Cretaceous) subject to subduction and collision leading to the emplacement of oceanic
154 units onto the European continental crust (Brouwer et al., 2002).

Most of the hyper-alkaline springs found by Cipolli et al. (2004) are located within the Beigua Unit composed of metagabbros and metabasalts that underwent multistage metamorphic evolution in blueschist to eclogite facies. Only a couple of these springs are hosted on altered lherzolites of the Erro-Tobbio Unit (Figure 2). These springs emerge along faults and fractures through serpentinites and related rocks.

160 Springs are usually located in the riverbeds, just above the water level and often 161 harnessed with tubing. Direct gas bubbling was observed only at 3 locations (GOR34, GOR35,

162 L43; Figure 2) but the discharge was very irregular, sometimes absent for several minutes.

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164 **3.** Sampling and analyses

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Sampling was performed in June 2010 in Liguria and in January 2011 in Oman. Sampling sites were chosen from a database kindly provided by the Ministry of Water Resources of Oman, while in Liguria, sampling sites were chosen from the database published in Cipolli et al. (2004). Details on locations, hosting lithology/mineralogy, as well as basic chemical composition of the spring waters (pH, temperature, major cations and anions) are summarised in Table 1 and is extensively discussed in Chavagnac et al. (2013b).

172 21 gas samples from 9 sites were collected in Oman directly from the springs (Figure
173 1). In Liguria, a total of 12 samples were collected from the 3 sites where active bubbling was
174 observed (L43, GOR34, GOR35, Figure 2). All gas samples were collected into crimp-sealed
175 gas glass bottles by water displacement. Dissolved gas samples were also collected at all sites

176 visited in 2010 (7 sites, 15 samples in Liguria) and 2011 (12 sites, 23 samples in Oman) into 177 poisoned (HgCl₂), crimp-sealed glass bottles. Dissolved gases were then analysed using the 178 headspace method (Kolb and Ettre, 1997) followed by gas chromatography in the same way 179 as the gas samples, using a SRI 8610C gas chromatograph, fitted with a Flame Ionization 180 Detector/Methanizer (FID-M) for the detection of methane (CH_4), small alkanes (C_1 - C_5), CO 181 and CO₂, and a Helium Ionization Detector (HID) for the detection of hydrogen (H₂), oxygen 182 (O_2) , and nitrogen (N_2) . Methane gas is eluted together with CO and CO₂ on a 3' Molecular Sieve Packed Column, using Hydrogen as the carrier gas, while H₂, O₂, and N₂ are eluted on 6' 183 184 Silica Gel Packed Column with Helium as the carrier gas.

In both cases, gas compositions are given as non-dried gases, taken into account thewater vapour phase.

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- 188 **4. Results and discussion**
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190 4.1. Water chemistry and gas compositions

191 The full chemical composition of the hyper-alkaline fluids collected in Oman and Liguria and 192 the composition of the associated mineralogical assemblages formed therein are presented in 193 Chavagnac et al. (2013a, 2013b). For clarity we here summarise the composition, temperature 194 and pH, together with the precipitate mineralogy and geological context (Table 1). 195 Composition of the gas phase and the concentration of dissolved gases in the alkaline waters 196 are given, respectively, in Tables 2 and 3. All the springs investigated in this study discharge 197 hyper-alkaline waters, whose pH values range from 9.5 to 11.7 in Liguria, and from 10.1 to 198 11.9 in Oman. These waters all belong to the Ca-OH hydro-chemical group characterised by 199 very low magnesium and DIC contents (Table 1) (Chavagnac et al., 2013b). The Ca-OH type 200 - also described as mature waters as they evolved from rainwater to neutral Mg-HCO₃ waters and to high-pH Ca-OH type waters (Bruni et al., 2002) – is typical of waters circulating in a deep and closed ultramafic rocks system affected by present-day, low-temperature serpentinisation (Bruni et al., 2002; Chavagnac et al., 2013b; Etiope et al., 2013a; Neal and Shand, 2002, Schwarzenbach et al., 2013). According to these results, the gases, analysed in this study, are venting from waters coming from a deep aquifer, mainly hosted by serpentinised peridotites and related rocks (Bruni et al., 2002).

Temperatures vary from 21.4 to 23.7°C in Ligurian springs while they range from 20.9 to 29.9°C in Oman. These temperatures are very similar (although slightly higher in Oman springs) while the local climatic conditions are obviously very different between Oman and Northern Italy. Chavagnac et al. (2013b) noted that there was no significant influence of the seasonal variation on the spring water temperature both in Oman and Liguria. As indicated by Neal and Stanger (1984), these temperatures indicate that the water cannot penetrate deeper than the base of the ultramafic formation in Oman, i.e. 5 km.

214 All the gas samples at both locations are strongly enriched in N_2 and have very low O_2 215 partial pressures (Tables 2 and 3). This leads to a gas phase N_2/O_2 ratio varying from 130 to 216 440 in Oman and 417 to 848 in Liguria (atmospheric ratio is ~1.9). These values are similar to 217 those found by Etiope et al. (2013a) in the Othrys ophiolite springs (Archani, Greece) as well 218 as other world-wide hyper-alkaline springs (Abrajano et al., 1988; Cipolli et al., 2004; Neal 219 and Stanger, 1983) Dissolved N₂ and O₂ concentrations vary between 656 to 2135 μ M and 5.7 220 to 23.4 µM, respectively in Oman, and from 600 to 3785 µM and from 3.2 to 31.6 µM, 221 respectively in Liguria. Waters are therefore supersaturated in N_2 (the equilibrium 222 concentration at T=25°C and Salinity = 0.5 is ~400 μ M; Hamme and Emerson, 2004) while 223 the dissolved O₂ concentration is below the concentration of a water body at equilibrium with 224 the atmosphere (402 μ M at T=25°C and Salinity = 0.5; Green, 1958). The low O_{2(aq)} content, 225 also observed by Etiope et al. (2013a) in Greece and Boschetti et al. (2013) in Italy, is a

226 characteristic of hyper-alkaline waters and is due to the reducing conditions linked to 227 hydrogen production during serpentinisation process. The actual presence of O_2 in our 228 samples may be due to (air?) contamination during sampling. Similar explanation could be 229 suggested for the high N₂ concentrations; however, contamination alone cannot account the 230 high N₂ content of the gas phase. As reported by Etiope et al. (2013a), variability in dissolved 231 N₂ and O₂ concentrations may rather be related to the meteoric water input into the upwelling 232 spring water, especially in Liguria whereby average annual rainfall is ten times higher than in 233 Oman.

234 Malatesta et al. (2011) argue that the Voltri Massif exhumed within a serpentinite 235 "channel" whereby plutonic and sedimentary rocks metamorphosed at different pressure and 236 temperature conditions are mixed during their pathway to the surface, wrapped by chlorite-237 amphibole-talc schists (e.g. Hoogerduijn Strating, 1991). Within these metamorphic rocks, 238 the low-strain serpentinised peridotites contain higher nitrogen concentration (4-6 ppm) than 239 their high-pressure counterparts (1-3 ppm), but still much lower than metasedimentary rocks 240 (100-1700 ppm) (Busigny et al., 2003; Philippot et al., 2007). In the latter, N₂ occurs as 241 ammonium (NH₄) and is substituting potassium in K-bearing minerals (Busigny et al., 2003). 242 The higher N_2/O_2 and K/Cl ratios of Liguria samples compared to the Oman ones (as shown 243 in Figure 3) suggest that serpentinisation-derived waters were in contact with the 244 metasedimentary formation along their pathway, leaching N_2 from the rocks. Note that the 245 nitrogen in the metasedimentary formation was not released during subduction (Philippot et 246 al., 2007).

Another typical feature of low-temperature hyper-alkaline springs is the low concentration in CO_2 , both in the gas phase (this work) and in the waters (Chavagnac et al., 2013b). Our measurements (Tables 2 and 3) are therefore coherent with most of the findings in other ophiolites in the world (Abrajano et al., 1988; Etiope et al., 2011, 2013a; Neal and Stanger, 1983). In Oman, CO_2 was detected neither in the gas phase nor in the aqueous phase, while in Liguria, CO_2 is significantly present at concentrations up to 22 μ M in the aqueous phase and up to 130 ppm in the vapour phase (actual atmospheric $CO_2 \sim 350$ ppm). The extremely low concentration in CO_2 (below detection limit) in high-pH waters of Oman shows that there is no supply of atmospheric CO_2 to these waters (as suggested by Pfeifer, 1977).

Oman and Voltri hyper-alkaline springs are characterised by a very low DIC content (Chavagnac et al., 2013b, Schwarzenbach et al., 2013; and therefore a very low equilibrium pCO₂) enabling the absorption of atmospheric CO₂ by the waters. This is evidenced by the formation of a thin layer of carbonate precipitates (calcite and/or aragonite) in Oman springs (Table 1, Chavagnac et al., 2013a). At the same time, there may be consumption of CO₂ in waters to form CH₄ and other HC during the FTT reactions (see §4.2).

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$4.2.H_2$ and hydrocarbons (C1-C5)

Enrichment of the gas phase in H_2 and CH_4 (Tables 2 and 3) is typical of hyper-alkaline springs. However, a striking difference between Oman and Liguria is in the amount of $H_{2(g)}$ compared to $CH_{4(g)}$. Comparison between Oman and Liguria shows indeed very different styles of H_2 and CH_4 venting while the main geological setting (ultramafic rocks) and the environmental conditions (T, pH) appear similar.

A comparison of H_2/CH_4 ratios for other ultrabasic formations, as depicted in Figure 4, points to three groups: i) a H_2 -dominated group including Oman, Rainbow, Lost City, Logatchev and Socorro Island (Neal and Stanger, 1983; Charlou et al., 2002; Proskurowski et al., 2008; Schmidt et al., 2008; Taran et al., 2010); ii) a CH₄-dominated including the Tekirova Ophiolite (Turkey) (Etiope et al., 2011), the Othrys Ophiolite (Greece) (Etiope et al., 2013a) together with our samples from the Ligurian Alps; iii) a third group, which includes 276 the Zambales ophiolite (Abrajano et al., 1988) and the Bay of Prony (New Caledonia) 277 (Boulart et al., 2012), is intermediate with a ratio close to 1-2. Interestingly, the H₂-278 dominated group is mostly constituted by vents developing along mid-ocean ridges and 279 characterised by higher temperatures, although Lost City is a low-temperature system 280 compared to Rainbow or Logatchev (Charlou et al., 2002; Proskurowski et al., 2008). 281 Nevertheless, fluid temperatures of the seafloor vents are significantly higher than those of 282 on-land hyperalkaline springs, which are overall below 40°C. The Oman ophiolite is the sole 283 belonging to the H_2 -dominated group. This may be related to a simpler geological history than 284 the other ophiolites, as it was not included in the Alpine collision belt contrary to the CH₄-285 dominated ophiolites. Indeed, the latter have undergone multi-stage metamorphism during the 286 collision event and their obduction onto the continent.

287 We propose that the data depicted in Figure 4 may be related to the influence of 288 geodynamic and geologic context on serpentinisation reactions. The first argument comes 289 from the locations of the hyper-alkaline springs from which gas bubbling is clearly observed. 290 Indeed, in Oman, gas samples were collected from hyper-alkaline springs emerging from the 291 major structural discontinuity at the transition between crustal section and mantle rock (paleo-292 Moho), and none from hyper-alkaline springs located nearby the basal thrust plane, e.g. the 293 metamorphic structural discontinuity (Chavagnac et al., 2013b). In contrast, gas bubbling in 294 Liguria was observed only at three hyper-alkaline springs emerging from various 295 metamorphosed substratum (highly serpentinised metabasalt/metagabbros and serpentinised 296 lherzolites) wrapped by chlorite-amphibole-talc schists (Hoogerduijn Strating, 1991). These 297 schists may have been used as a conduit by meteoric waters to infiltrate and interact with the 298 substratum, explaining the high N₂ gas content of Liguria hyper-alkaline springs.

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300 - Hydrogen

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301 In Oman, H₂ is the second dominant gas after N₂ in the gas phase, ranging from 9.4% (site 27) 302 to 12.1% (site 28) of the total gas volume. The lowest dissolved H₂ concentrations were found 303 in the wadi at site 20 (26.7 to 65 μ mol/l) while the highest values were found at site 28 (377 304 µmol/l). This agrees with the concentrations previously measured (Neal and Stanger, 1983; 305 Sano et al., 1993). Note that H₂ concentrations given by Neal and Stanger (1983) – up to 99% 306 vol – are much higher than ours, probably as a consequence of the different methodologies 307 used for gas measurements and of the fact that Neal and Stanger (1983) values might not have 308 taken into account the water vapour in the total gas volume.

In the Voltri Group, $H_{2(g)}$ concentrations range from 0.011% (L43) to 1.2% (GOR34) of the total gas volume, i.e. 10 to 100 times lower than in Oman. Dissolved H₂ concentrations are relatively low (up to to 5 µmol/l on average) except for two sites (C11 and LER2) where the concentrations are higher (up to 562.6 µmol/l at LER2). A strong variability between samples from the same site can be observed although conditions of sampling were identical from one sample to the other, which may depict a sporadic venting of H₂. This remains to be confirmed by further measurements.

316 The low H₂ content in the Voltri samples, also reported by Boschetti et al. (2013) for 317 the Voltri Group as well as by Etiope et al. (2013a) in another setting in Greece, seems to be 318 uncommon for hyper-alkaline environments. During the reactions of serpentinisation at 319 moderate temperatures ($<300^{\circ}$ C), olivine and pyroxene react with water to form serpentine, 320 brucite, magnetite and various Fe/Ni alloys depending on redox conditions of reactions. Other 321 phases such as talc, amphibole (e.g. tremolite), chlorite and sulfides can complete the 322 parageneses. In fact, as noted by Mevel (2003), this is a succession of simple reactions starting with the silicate dissolution leading to the production of $SiO_{2(aq)}$, which then reacts 323 with Mg^{2+} and OH^{-} to form the serpentine crystal and the secondary minerals. In the same 324 time, water reacts with metals in silicates, i.e. oxidizing Fe^{2+} contained in olivine into Fe^{3+} 325

326 which can be incorporated both in serpentine and magnetite (O'Hanley and Dyar, 1993; 327 Evans, 2008; Andreani et al., 2013), and reducing H₂O to H_{2(g)}. Hence, the assemblage 328 serpentine-brucite determines the silica activity of the system, which subsequently controls the oxidation of Fe^{2+} to Fe^{3+} and therefore $H_{2(g)}$ production (Evans, 2008; Katayama et al., 329 330 2010). Seyfried et al. (2007) also showed in laboratory experiments that a high silica activity 331 could limit the production of H₂. Low H₂ content is then supported by SiO₂ concentrations 332 measured in the Ligurian fluids, which are significantly higher than in Oman (Table 1, 333 Chavagnac et al., 2013b). Our data in Oman and Liguria may therefore suggest a first low-334 silica activity stage producing high amounts of H₂ (Oman) followed by a second stage of 335 higher silica activity derived from the altered mafic rocks, limiting the production of H₂ 336 (Liguria). Episode of a low silica activity fluid metasomatism (rodingitization) is a common feature of ophiolite complexes, taking place primarily during ocean-floor metamorphism and 337 338 evidenced in terms of lithologies by the occurrence of epidote-rich to diopside-rich rocks in or 339 adjacent to serpentinites (e.g. Schandl et al., 1989). Diopsidites and rodingites were identified 340 both in Oman (Python et al., 2007) and in Liguria (e.g. Dal Piaz et al., 1980). Ferrando et al. (2010) suggest that ophiolite complexes (i.e. the Alpine orogeny), which are subducted then 341 342 exhumed, may undergo additional episode(s) of serpentinisation and rodingitization during 343 the prograde to retrograde metamorphic path. This would be evidenced by the occurrence of 344 H₂-rich fluid inclusions (and also a Ca metasomatism), but these are rarely preserved and 345 detected in metamorphic environments. However, Peretti et al. (1992) report the occurrence of 346 fluid inclusions in the Malenco peridotites composed of H_2 (0.3–3 mol.% of the total gas) within brines (5.1 wt.% CaCl₂ + 6.4 wt.% NaCl) and Ferrando et al. (2010) described fluid 347 348 inclusions in Bellecombe meta-ophiolites made up 1.0 and 0.4 mol% of H_2 and CH_4 . respectively, within brines (6 wt.% CaCl₂ + 6 wt.% NaCl). While both studies illustrate a 349 350 second episode of Ca-rich and Si-undersaturated reducing fluid percolation, it is unclear 351 whether this event is related to the prograde or late stage metamorphic history of the Alpine 352 orogeny. Nevertheless, these authors clearly identified that H_2 -rich serpentinisation-related 353 fluid was produced before the final exhumation of the ophiolite complex.

354 In addition, temperature and water-rock ratio are two other parameters that have a 355 strong control on Fe speciation during serpentinisation and therefore on H_2 production, as 356 discussed by Klein et al. (2009). In their model calculations, they indicate that at low 357 temperatures (<150-200°C) and low water-rock ratios (<0.1 to 5), hydrogen production is related to the formation of Fe³⁺-serpentine. Andreani et al. (2013) reported similar results on 358 the role of Fe^{3+} -serpentine on H₂ production based on a μ -XANES study of Fe-speciation in 359 serpentine minerals. Besides, they stated that Fe³⁺ goes predominantly into serpentine, rather 360 361 than into magnetite, for serpentinisation degree up to 75%.

362 As emphasised in Hellevang et al. (2011), the presence of nitrogen, sulfur and carbon 363 compounds in various redox states may also affect the H₂ generation by limiting the degree of reduction of the aqueous solutions. As for instance, Hellevang et al. (2011) simulated the 364 365 effect of carbon species on the potential for H₂ generation from an aqueous solution of olivine 366 (Fo90). In natural aquatic systems, the dissolved CO_2 reacts with H_2O to form H_2CO_3 , $HCO_3^$ and CO32-. However, due to the serpentinisation reactions, conditions are increasingly 367 368 reducing, which thermodynamically favour reduction of inorganic carbon to CH₄. Hellevang 369 et al. (2011) summarised the equilibrium as the following:

$$370 \quad 2H_2O + CO_{2(aq)} \leftrightarrow CH_{4(aq)} + 2O_{2(aq)} \tag{1}$$

According to their experiments, at low carbon content, the aqueous solution is rapidly reduced and allows the formation of significant amounts of H_2 . On the opposite, at higher carbon content, more olivine is required to form significant H_2 quantities. In the same experiments, at constant C content, the CO₂ partial pressure decreases as CO₂ is converted to CH₄, allowing O₂ activity to drop, which favours H_2 formation. In our case, the presence of CO₂ both in the 376 gas and aqueous phase from Liguria indicates that the CO_2 is not completely converted to 377 CH_4 ; hence the aqueous solution may not be significantly reduced to allow H_2 formation at 378 significant levels.

In Oman, the absence of CO_2 in both gas and aqueous phases indicates that all the CO_2 is either converted to CH_4 or consumed by carbonate formation. This is confirmed by the relationship between O_2 and H_2 (Figure 5; Hellevang et al., 2011), which shows that H_2 generation might be also controlled by the introduction of O_2 in the system.

383

384 - Methane and hydrocarbons

385 In Oman, CH_{4(g)}, the third most abundant gas, ranges from 1.5% at sites 11 and 13 up to 8% at 386 site 31. In the dissolved phase, concentrations in CH_4 vary from 7.4 μ mol/l at site 26 to 209.2 387 µmol/l at site 30. In Liguria, CH₄ is the second dominant gas after N₂, sometimes the first. 388 The lowest contents were observed at GOR35 (10.5% of the total gas volume) and the highest 389 at L43 (37%). L43 was also characterised by the highest dissolved CH₄ concentration (649.4 390 µmol/l, the water-atmosphere equilibrium concentration for this station is 2.3 nmol/l), while 391 the lowest was measured at LER2. Intermediate concentrations (~300 µmol/l) were found at 392 C11 and BR1.

393 It is commonly accepted that CH_4 and possibly longer hydrocarbons are forming in 394 hyper-alkaline fluids abiogenically, through the FTT reactions as the following (Abrajano et 395 al., 1988; Charlou et al., 2002; Hosgormez, 2007; Proskurowski et al., 2008):

$$396 \quad 4H_2 + CO_2 \leftrightarrow CH_4 + 2H_2O \tag{2}$$

397 Carbon isotopic composition data previously collected in Oman and Liguria undoubtedly 398 indicate the abiogenic origin of CH_4 ($\delta^{13}C_1 \sim -9\%$; Boschetti et al., 2013; Neal and Stanger, 399 1983). However, one cannot exclude another synthesis pathway, which consists in the 400 reduction of formic acid, formaldehyde, methanol and later CH_4 (Seewald et al., 2006).

401 FTT reactions have been studied for more than a century, to synthesize CH₄ and 402 alkanes from CO at high temperature. However, the efficiency of FTT synthesis from CO₂ (i.e. 403 methanation) is lower under aqueous hydrothermal conditions (Foustoukos and Seyfried, 404 2004). Provided the presence of catalysts (i.e. magnetite, awaruite and chromite) and at high 405 temperature (>300°C), FTT reactions from CO₂ can nevertheless produce high amounts of 406 CH₄ and light alkanes (Berndt et al., 1996; Horita and Berndt, 1999; McCollom and Seewald, 407 2001). More recently, Neubeck et al. (2011) studied the generation of H₂ and CH₄, in the 30-408 70°C range, and demonstrated that the presence of H₂, CO, CO₂ and the necessary catalysts 409 (i.e. magnetite, chromite among others) leads to the generation of CH₄ in a few months, even 410 at low temperature. In fact, in the absence of catalysts, methanation is thermodynamically 411 favoured at temperatures below 100°C although it evolves slowly over time (Etiope et al., 412 2013a). The relationship between dissolved SiO₂ and CH₄ in Liguria (Figure 6) shows that the 413 production of CH₄ is controlled by dissolved SiO₂ and therefore by the dissolution of olivine. 414 This is in line with Neubeck et al. (2011) who showed that the dissolved CH₄ concentration is 415 a proxy for olivine dissolution rates.

416 Besides the role of the catalysts, Oze and Sharma (2005) proposed a possible 417 formation of CH₄ without H₂ mediation during serpentinisation in abundant CO₂ conditions, 418 which is the case in Ligurian springs. However, it is difficult today to establish the role of 419 each parameter, especially since H_2 may be consumed as it is produced, maintaining its 420 concentration at a low level. We may suggest that the high concentration in CH₄ in both the 421 gas and the aqueous phase in Liguria is the result of the FTT reactions, occurring at a later 422 stage of the serpentinisation process while in Oman, the lower amount of magnetite as well as 423 the low content in CH₄ tend to support that the serpentinisation process is at an earlier stage. 424 Nevertheless, we cannot exclude the fact that the availability of carbon in Oman might be a 425 limiting factor for FTT reactions, which is not the case in Liguria (Schwarzenbach et al., 426 2013).

Traces of light alkanes were found in the gas phase both in Liguria and in Oman. Samples from Liguria showed concentrations of C_2 ranging from 20 ppmV (GOR35) to 441 ppmV (L43), but no HC with longer chains were detected. In Oman, C_2 was present in all samples, from 8.2 ppmV in the wadi (site 20) to 108.9 ppmV at site 27. Higher alkanes (up to C_4) were found only at site 11 with variable concentrations. Only at sites 11 and 30, dissolved C_2 was detected (up to 240 nmol/l) in the spring waters. No dissolved alkanes (C_2 to C_4) were detected in the dissolved phase in the Ligurian samples.

434 Although finding light alkanes as in our Oman and Ligurian samples is common for 435 ultramafic fluids, their origin remains subject to debate (Proskurowski et al., 2008; Konn et al., 436 2009a). In the Tekirova ophiolite (Turkey), Etiope et al. (2011) and Hosgormez (2007) 437 showed from the isotopic compositions that the alkanes were not all produced by FTT-438 reactions, but more probably originated from a strong thermogenic component. However, 439 isotopic studies cannot prove the abiogenic origin of longer-chained HC ($>C_5$) since a variety 440 of mostly unknown fractionation steps may occur along the synthesis pathways, with a 441 dominance of biogenic source and/or processes hiding their characteristic signature (Konn et 442 al., 2009a).

Although the carbon isotope composition is not yet available for our gas samples, a first estimate of the light alkane distribution can be made with the Schulz-Flory method (Figure 7; Flory, 1936). Indeed, the product distribution of hydrocarbons formed during the FTT-reactions generally follows the Schulz-Flory distribution, which is expressed as the following equation:

448
$$\frac{W}{n} = (1 - \alpha)^{n-1} \alpha^{n-1}$$
 (3)

449 where W is the molar fraction of HC molecules, n the number of C atoms and α the chain 450 growth probability factor. As a result, the distribution of HC is controlled by the chain growth 451 probability factor, which is typical of abiotic synthesis.

452 Our data in Oman do not fit this model, confirming that part of the light alkanes are 453 probably generated through degradation of organic matter or thermogenesis in the same way 454 as in Tekirova ophiolite (Etiope et al., 2011). This is especially true for higher hydrocarbons 455 (above C_3). We may suggest that this is also the case of the Voltri ophiolite although longerchain HC were not detected with our methodology. However, the amount of C₂ seems to be 456 associated with the amount of CH₄, taking into account that the gas venting from the 457 458 metaperidotites of the Erro-Tobbio coincides with those of the Oman crustal section/peridotite, 459 which are undistinguishable from one another (Figure 8). Adding the data from Boschetti et al. (2013) to Figure 8 confirms this trend with an R^2 above 0.85. This tends to indicate that C_2 is 460 derived from abiogenic polymerisation, which has been experimentally demonstrated earlier 461 462 in Des Marais et al. (1981) by 'spark-discharge' gas phase experiments. However, as noted by 463 Proskurowski (2010), uncatalysed polymerisation remains to be proven in aqueous conditions.

464

465

5. Concluding remarks

466 Serpentinisation is a phenomenon that occurs wherever mantle rocks are altered by aqueous 467 fluids. The alteration of the Oman and Liguria ophiolites by meteoric waters indeed generates 468 hyper-alkaline fluids with pH up to 11-12 at temperatures of 20-40°C at the springs. This is 469 accompanied by the production of a gas phase enriched in hydrogen and subsequently 470 methane. Comparison of the gas composition from Oman and Liguria, and more generally 471 from various places over the world, shows significant differences in terms of H₂/CH₄ ratios. 472 Three groups have been identified likely related to different stages in the serpentinisation 473 reactions, from the hydrogen-dominated to the methane-dominated fluids. The distinction 474 between these groups might lie in the difference of rock mineralogical composition in which 475 the fluids circulate, on the degree of rock alteration and finally on the geological history of the 476 ophiolite. To fully understand the present-day serpentinisation reactions leading to the
477 formation of hydrogen, it is necessary to know the depth and the nature (rock composition) of
478 the reaction zone, which remains an undocumented black box to-date.

479 The comparison of the Oman and Liguria cases with other hyper-alkaline springs 480 worldwide illustrates even more the diversity of gas venting within ultramafic environments 481 and the complexity of the chemical reactions, which may be controlled by the geodynamic 482 and the geological context. It is clear that ophiolites have evolved differently since their obduction, especially in the context of the Alpine collision belt. The resulting metamorphic 483 484 history has therefore an impact on the gas composition of the fluids. Hence, is it relevant to 485 describe the current on-land hyper-alkaline springs as analogues of the MAR Lost City 486 hydrothermal system?

487

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788 Table and Figure Captions

789

Table 1: Chemical composition, geological settings and precipitates of the Oman and Liguria
hyperalkaline springs. Sampling indicates the type of sample collected (G: gas/DG: dissolved
gas).

Table 2: Chemical compositions of the gases collected in Oman and Liguria hyperalkalinesprings. n.d.: not detected.

Table 3: Concentration of the dissolved gases in the Oman and Liguria hyperalkaline waters.

796 Concentrations are given in µmol/l. n.d.: not detected.

797 Figure 1: Simplified geological map of the Oman ophiolite showing the locations of the

investigated springs. For a better comprehension of the surface lithology, the reader may refer

- to Table 1 where details are given.
- 800 Figure 2: Simplified geological map of the Voltri Massif (Liguria, Italy) showing the

801 locations of the investigated springs. Stars indicate the two stations where active gas bubbling802 was observed.

Figure 3: The variation of the K/Cl ratio of the alkaline waters versus the N_2/O_2 ratio of the gases for the Liguria and Oman springs.

Figure 4: Mean H₂/CH₄ ratios for our samples (Oman 2011 and Liguria 2010) compared with 805 806 various ultramafic environments showing three distinct gas composition groups. Data used 807 here are from Abrajano et al, 1988 (Zambales); Charlou et al., 2002 (Rainbow); 808 Proskurowsky et al., 2008 (Lost City); Taran et al., 2010 (Socorro Island); Etiope et al, 2011 809 (Tekirova); Boulart et al., 2012 (Bay of Prony); Etiope et al, 2013 (Othrys); Schmidt et al., 810 2008 (Logatchev); Etiope et al., 2013b (Portugal); Boschetti et al., 2013 (Liguria 2013). Error 811 bars indicate the standard deviation of the gas concentration distribution. Note that Rainbow 812 and Logatchev are considered as high-temperature vent sites (T>350°C) while Lost City is 813 ~90-100°C. The on-land venting sites are considered as low-temperature (T<40°C). In red, 814 sites developing in a mid-ocean ridge context and in blue in a subduction context.

815 Figure 5: The oxygen (O₂) versus hydrogen partial pressure of the Liguria and Oman gases

816 **Figure 6:** The dissolved SiO_2 concentration versus the $CH_4(g)$ partial pressure for the Oman 817 and Liguria springs.

Figure 7: Logarithmic distribution of mole fraction as a function of carbon number for Sites 11, 13, 20 and 27 in Oman showing that the Oman hydrocarbon composition does not fit the Schulz-Flory distribution. Dashed line represent a distribution typical of abiotic synthesis (Sherwood Lollar et al., 2008). Mole fraction is calculated as $x_i = n_i/n_{tot}$ with n_i as the amount of *i* constituent and n_{tot} the total amount of all constituents.

823 **Figure 8**: The CH₄(g) partial pressure versus that of ethane for the Oman and Liguria gases.

33

Figure 1



Figure 2



Figure 3



Figure 4



Figure 5



Figure 6



Figure 7



Figure 8



Table 1

| Sites | Latitude N | Longitude E | Т (°С) | pН | SiO ₂ (mmol/l) | Mg (mmol/l) | Ca (mmol/l) | K (mmol/l) | Na (mmol/l) | Cl (mmol/l) | Sampling | Geology | Precipitates |
|---------|---------------|----------------|-----------|------|------------------------------|----------------|----------------|---------------|----------------|----------------|----------|---|-----------------------------------|
| Oman | | | | | | | | | | | | | |
| Site 11 | 22.9880 | 57.2890 | 28.3 | 11.6 | 0.05 | 0.47 | 0.81 | 0.27 | 6.46 | 5.01 | G/DG | Serpentinite | |
| Site 13 | 23.4033 | 56.8577 | 28.0 | 11.7 | 0.00 | 0.00 | 1.87 | 0.16 | 6.84 | 4.80 | G/DG | peridotite/gabbro, MOHO | Aragonite, Brucite |
| Site 20 | 23.6232 | 57.1092 | 29.4 | 11.2 | 0.04 | 0.82 | 0.15 | 0.18 | 4.82 | 3.83 | G/DG | Bedded gabbro above peridotite (100 m thick) | Aragonite, Brucite |
| Site 26 | 23.6187 | 57.1080 | 31.3 | 11.3 | 0.02 | 0.00 | 0.66 | 0.22 | 5.83 | 3.95 | DG | Peridotite (100 m above the Moho) | Aragonite, Brucite |
| Site 27 | 24.7008 | 56.2692 | 31.2 | 11.1 | 0.12 | 1.40 | 0.51 | 0.08 | 3.02 | 2.12 | G/DG | Wherlite/gabbro, within the crustal section | Aragonite |
| Site 28 | 24.5242 | 56.2845 | 30.7 | 10.1 | 0.08 | 1.13 | 0.61 | 0.09 | 3.17 | 3.17 | G/DG | Bedded gabbro and ultramafic cumulates, within the crustal section | Aragonite, Mg-calcite, Brucite |
| Site 29 | 23.9572 | 56.4377 | 23.8 | 11.6 | 0.00 | 0.00 | 1.03 | 0.15 | 6.64 | 3.83 | DG | Serpentinite with gabbro 100 m above | Aragonite (brucite) |
| Site 30 | 22.9037 | 58.4225 | 27.5 | 11.7 | 0.02 | b.d.l. | 1.96 | 0.14 | 6.65 | 2.81 | DG | Peridotite | Aragonite, Brucite |
| Site 31 | 22.8907 | 58.3902 | 26.4 | 11.6 | 0.16 | 1.83 | 0.66 | 0.07 | 3.49 | 2.87 | G/DG | Peridotite near MOHO | Aragonite (brucite) |
| Liguria | | | | | | | | | | | | | |
| ERR20 | 44.508793 | 8.499098 | 23.7 | 11.1 | 51.31 | 0.00 | 0.33 | 0.05 | 0.74 | 0.50 | DG | Serpentinite + metagabbros | Calcite |
| C11 | 44.469 | 8.695346 | 14.0 | 11.5 | 3.62 | 0.23 | 0.07 | 0.04 | 0.57 | 0.60 | DG | Serpentinite + metagabbros | Calcite |
| GOR34 | 44.596363 | 8.783641 | 23.7 | 11.6 | 10.17 | 0.01 | 0.85 | 0.11 | 0.85 | 0.42 | G | serpentinised lherzolites | Calcite+Aragonite |
| GOR35 | 44.59655 | 8.786923 | 21.4 | 11.2 | 45.10 | 0.12 | 0.46 | 0.05 | 0.32 | 0.22 | G/DG | serpentinised lherzolites | Calcite |
| L43 | 44.458 | 8.768944 | 22.6 | 11.7 | 33.77 | n.a. | 0.76 | 0.13 | 1.26 | 0.45 | G/DG | Serpentinite + metagabbros | Calcite |
| L43Ru | 44.458 | 8.768944 | 20.0 | 11.6 | n.a. | n.a. | n.a. | n.a. | n.a. | n.a. | G | Serpentinite + metagabbros | Calcite |
| BR1 | 44.445473 | 8.778927 | 17.3 | 11.5 | 25.88 | n.a. | 0.75 | 0.08 | 1.05 | 0.50 | DG | Serpentinite + metagabbros | Calcite |
| LER21 | 44.424297 | 8.6582717 | 17.5 | 11.3 | 4.97 | n.a. | 0.72 | 0.02 | 0.43 | 0.55 | DG | Serpentinite + metagabbros | Calcite |
| LER2 | 44.40568 | 8.618569 | 17.3 | 9.5 | 17.84 | 0.01 | 0.56 | 0.10 | 0.45 | 0.56 | DG | Serpentinite + metagabbros | Calcite |

Table 2

| Sites | CH ₄ (%vol) | H ₂ (%vol) | N ₂ (%) | O ₂ (ppmV) | CO (ppmV) | CO ₂ (ppmV) | C ₂ H ₆ (ppmV) | C ₃ H ₈ (ppmV) | C ₄ H ₁₀ (ppmV) | C ₅ H ₁₂ (ppmV) |
|---------------------|---------------------------|--------------------------|-----------------------|-----------------------|--------------|------------------------|--------------------------------------|---|--|--|
| Oman 2011 | | | | | | | | | | |
| Site 11 | 1.5 | 10.6 | 12.4 | 507.6 | n.d. | n.d. | 69.6 | 15.1 | 4.4 | n.d. |
| Site 11 | 1.6 | 10.7 | 12.7 | 543.8 | n.d. | n.d. | 70.3 | 14.7 | 5.2 | n.d. |
| Site 11 | 2.5 | 11.1 | 12.7 | 568.6 | n.d. | n.d. | 70.7 | 14.9 | 5.8 | n.d. |
| Site 13 | 1.5 | 10.5 | 11.2 | 474.6 | n.d. | n.d. | 9.9 | 8.2 | n.d. | n.d. |
| Site 13 | 1.5 | 11.3 | 12.2 | 570.7 | n.d. | n.d. | 9.5 | 7.8 | 5.7 | 8.2 |
| Site 20 | 5.6 | 11.2 | 13.3 | 559.5 | n.d. | n.d. | 10.1 | n.d. | n.d. | n.d. |
| Site 20 | 5.5 | 11.3 | 12.7 | 552.3 | n.d. | n.d. | 11.2 | n.d. | n.d. | n.d. |
| Site 20 | 6.3 | 10.9 | 13.5 | 563.0 | n.d. | n.d. | 13.9 | n.d. | n.d. | n.d. |
| Site 20 | 4.3 | 11.0 | 12.8 | 418.5 | n.d. | n.d. | 8.2 | n.d. | n.d. | n.d. |
| Site 27 | 5.2 | 9.6 | 25.0 | 568.5 | n.d. | n.d. | 84.8 | n.d. | n.d. | n.d. |
| Site 27 | 5.6 | 10.0 | 27.0 | 788.2 | n.d. | n.d. | 108.9 | n.d. | n.d. | n.d. |
| Site 27 | 5.0 | 9.4 | 24.0 | 578.6 | n.d. | n.d. | 92.3 | n.d. | n.d. | n.d. |
| Site 28B | 3.3 | 10.4 | 18.0 | 970.9 | n.d. | n.d. | 83.5 | n.d. | n.d. | n.d. |
| Site 28B | 3.6 | 10.4 | 15.5 | 497.6 | n.d. | n.d. | 72.7 | n.d. | n.d. | n.d. |
| Site 28C | 2.6 | 9.9 | 21.7 | 1639.1 | n.d. | n.d. | 91.1 | n.d. | n.d. | n.d. |
| Site 28C | 3.8 | 10.7 | 15.7 | 616.4 | n.d. | n.d. | 52.6 | n.d. | n.d. | n.d. |
| Site 28D | 4.2 | 11.0 | 15.5 | 480.9 | n.d. | n.d. | 97.1 | n.d. | n.d. | n.d. |
| Site 28D | 4.9 | 11.5 | 17.6 | 610.8 | n.d. | n.d. | 93.4 | n.d. | n.d. | n.d. |
| Site 28D | 4.5 | 12.1 | 17.8 | 673.7 | n.d. | n.d. | 94.1 | n.d. | n.d. | n.d. |
| Site 31 | 7.9 | 10.7 | 15.7 | 491.0 | n.d. | n.d. | 59.7 | n.d. | n.d. | n.d. |
| Site 31 | 8.0 | 10.7 | 15.9 | 532.1 | n.d. | n.d. | 61.9 | n.d. | n.d. | n.d. |
| <u>Liguria 2010</u> | | | | | | | | | | |
| GOR34 | 35.7 | 0.2 | 35.9 | 447.1 | n.d. | n.d. | 233.2 | n.d. | n.d. | n.d. |
| GOR34 | 30.4 | 0.0 | 31.5 | 373.7 | n.d. | 106.7 | 212.8 | n.d. | n.d. | n.d. |
| GOR35 | 10.5 | 0.0 | 33.8 | 459.7 | n.d. | 107.5 | 23.5 | n.d. | n.d. | n.d. |
| GOR35 | 10.7 | 1.2 | 33.8 | 401.5 | n.d. | 106.7 | 24.6 | n.d. | n.d. | n.d. |
| L43Ru | 34.4 | 1.0 | 32.4 | 559.3 | 39.2 | 124.9 | 359.5 | n.d. | n.d. | n.d. |
| L43Ru | 34.4 | 0.0 | 32.0 | 480.3 | n.d. | 129.7 | 353.2 | n.d. | n.d. | n.d. |
| L43Ru | 33.6 | 0.0 | 31.0 | 386.0 | n.d. | 123.5 | 339.7 | n.d. | n.d. | n.d. |
| L43 | 37.0 | 0.0 | 30.7 | 381.0 | n.d. | n.d. | 441.3 | n.d. | n.d. | n.d. |
| L43 | 34.7 | 0.0 | 30.7 | 736.1 | n.d. | 106.2 | 390.1 | n.d. | n.d. | n.d. |
| L43 | 35.6 | 0.1 | 31.3 | 621.4 | n.d. | 106.5 | 418.3 | n.d. | n.d. | n.d. |
| L43 | 36.7 | 0.1 | 30.3 | 373.0 | n.d. | n.d. | 423.4 | n.d. | n.d. | n.d. |
| L43 | 35.8 | 0.0 | 30.2 | 356.5 | n.d. | n.d. | 409.9 | n.d. | n.d. | n.d. |

| Sites | Samples | H_2 | CH ₄ | N_2 | 02 | СО | CO ₂ | C ₂ H ₆ |
|-------------|----------|-------|-----------------|--------|------|------|-----------------|-------------------------------|
| Oman | | | | | | | | |
| <u>2011</u> | | | | | | | | |
| Site 11 | GD1 | 214.6 | 36.3 | 1281.4 | 13.4 | n.d. | n.d. | 0.24 |
| Site 11 | GD2 | 215.7 | 38.4 | 762.6 | 7.4 | n.d. | n.d. | 0.28 |
| Site 13 | GD1 | 333.2 | 25.4 | 848.8 | 7.9 | n.d. | n.d. | n.d. |
| Site 13 | GD2 | 340.8 | 28.0 | 725.1 | 8.7 | n.d. | n.d. | n.d. |
| Site 20 | GD2 GV | 222.3 | 96.5 | 755.2 | 8.8 | n.d. | n.d. | n.d. |
| Site 20 | GD2 PV | 341.7 | 147.6 | 725.7 | 7.2 | n.d. | n.d. | n.d. |
| Site 20 | GD2 Wadi | 26.7 | 15.1 | 995.5 | 12.3 | n.d. | n.d. | n.d. |
| Site 20 | GD3 Wadi | 65.1 | 15.5 | 848.5 | 10.7 | n.d. | n.d. | n.d. |
| Site 26 | GD2 | 168.6 | 7.4 | 846.3 | 6.7 | n.d. | n.d. | n.d. |
| Site 27 | GD1 | 132.1 | 84.7 | 823.6 | 7.4 | n.d. | n.d. | n.d. |
| Site 27 | GD5 | 184.9 | 15.4 | 892.2 | 11.0 | n.d. | n.d. | n.d. |
| Site 27b | GD1 | 179.3 | 71.9 | 821.6 | 7.4 | n.d. | n.d. | n.d. |
| Site 27b | GD2 | 59.4 | 39.5 | 2135.2 | 23.4 | n.d. | n.d. | n.d. |
| Site 28C | GD2 | 274.2 | 64.6 | 1661.4 | 15.1 | n.d. | n.d. | n.d. |
| Site 28D | GD2 | 115.7 | 25.3 | 872.7 | 10.4 | n.d. | n.d. | n.d. |
| Site 28G | GD1 | 377.5 | 70.2 | 656.6 | 5.7 | n.d. | n.d. | n.d. |
| Site 29 | GD1 | 219.5 | 38.0 | 827.7 | 8.2 | n.d. | n.d. | n.d. |
| Site 29 | GD2 | 168.0 | 31.1 | 825.8 | 8.4 | n.d. | n.d. | n.d. |
| Site 30 | GD2 | 338.9 | 209.2 | 794.7 | 8.7 | n.d. | n.d. | 0.15 |
| Site 30 | GD1 | 227.2 | 206.0 | 839.0 | 9.2 | n.d. | n.d. | 0.14 |
| Site 31 | GD1 | 48.0 | 64.4 | 1382.3 | 14.9 | n.d. | n.d. | n.d. |
| Site 31 | GD1 | 225.0 | 68.2 | 976.2 | 12.3 | n.d. | n.d. | n.d. |
| Site 31 | GD1 | 98.5 | 21.1 | 850.3 | 10.5 | n.d. | n.d. | n.d. |
| Liguria 20 | 10 | | | | | | | |
| ERR20 | 2GD | 3.8 | 19.0 | 711.2 | 5.2 | 0.03 | 0.99 | n.d. |
| ERR20 | 1GD | 4.3 | 5.8 | 691.6 | 4.2 | 0.02 | 0.99 | n.d. |
| C11 | 1GD | 1.2 | 384.5 | 1066.5 | 9.7 | n.d. | 1.23 | n.d. |
| C11 | 3GD | 0.5 | 349.7 | 1184.6 | 11.7 | n.d. | 1.23 | n.d. |
| C11 | 2GD | 14.6 | 456.6 | 806.5 | 5.0 | n.d. | 1.18 | n.d. |
| GOR35 | 1GD | 5.8 | 88.1 | 1186.7 | 11.0 | n.d. | 1.09 | n.d. |
| L43 | 1GD | 2.1 | 649.4 | 600.6 | 3.3 | n.d. | 1.00 | n.d. |
| BR1 | 1GD | 3.0 | 279.9 | 1131.8 | 10.8 | n.d. | 1.18 | n.d. |
| BR1 | 2GD | 0.3 | 278.2 | 1113.4 | 10.5 | n.d. | 1.14 | n.d. |
| BR1 | 3GD | 5.8 | 285.8 | 951.4 | 8.3 | n.d. | 1.10 | n.d. |
| LER21 | 1GD | 0.0 | 57.0 | 1134.8 | 10.8 | n.d. | 1.21 | n.d. |
| LER21 | 2GD | 2.5 | 50.8 | 1155.6 | 11.0 | n.d. | 1.25 | n.d. |
| LER21 | 3GD | 4.1 | 54.2 | 1091.4 | 10.1 | n.d. | 1.11 | n.d. |
| LER2 | 1GD | 561.6 | 0.8 | 3784.9 | 31.6 | n.d. | 21.52 | n.d. |
| LER2 | 3GD | 66.7 | 0.0 | 2041.2 | 18.0 | n.d. | 22.19 | n.d. |

Table 3