

# FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM

B. BIJU-DUVAL <sup>(1)</sup>, J. DERCOURT <sup>(2)</sup> and X. LE PICHON <sup>(3)</sup>

## ABSTRACT

We propose a series of paleogeographic maps of the western alpine system, from Triassic to Present, based on the geologic structure and evolution of the Mediterranean, its adjacent folded belts and the associated basins, and on the evolution of the North Atlantic ocean according to the magnetic anomaly informations. The kinematic framework is relatively precise for the Cenozoic but the accuracy is quite poorer for the Mesozoic. Between Triassic and Upper Cretaceous, the reconstructed continental margins can be compared to present passive or active margins. The geologic evolution can be described in terms of plate tectonics. After Upper Cretaceous, the oceanic areas are quite limited and the geologic evolution is dominated by the continental collision processes. For each of eight different key geologic periods (including Present), Africa and Europe are positioned according to the kinematics deduced from the opening of the Atlantic ocean. Three major intermediate plates are identified : Iberia, Apulia and Anatolia. The dimensions of the initial cratons are obtained through a palinspastic restoration taking into account the basement nappes and the effect of superposed tectonics. It is shown that, prior to Lower Jurassic, an ocean (the Tethys) separated Eastern Europe from Africa-Arabia. It is being consumed during Mesozoic because of

the left lateral motion of Africa with respect to Europe, which induces the separation of Apulia and Anatolia from Africa and their accretion to Europe. Distension prevails in the wake of the motion of Apulia-Anatolia, especially between Tunisia and Arabia, which is related to the formation of a new oceanic surface, called the Mesogea. The collision of Apulia occurs at different times depending on the morphology of the European southern margin. It is initiated in Uppermost Jurassic, with a resulting fracturation of Apulia in three blocks ; and ends during Uppermost Cretaceous to the east and west of Rhodope in the median portion. The consumption of the Tethys is accompanied by the creation of marginal basins : the Carpathian flyschs trough and the Black Sea. The motion of Moesia toward the southwest, related to the opening of the Black Sea, increases the Carpathian arcuation. After Upper Cretaceous, the change in the motion of Africa will lead to the following facts. New strike-slip faults appear in the Apulian plate. The Mesogea is being consumed as well as its different narrow troughs which extended toward the Tethys and the Atlantic. This consumption is accompanied by the formation of marginal basins. In the Western Mediterranean, where the consumption is complete, the marginal seas are well developed. In the Eastern Mediterranean, where the consumption is still only partial, they are barely initiated.

(1) Institut Français du Pétrole, 1-4 avenue de Bois-Préau, 92506 Rueil-Malmaison, France

(2) Université des Sciences et Techniques de Lille, B.P. 36, 59650 Villeneuve d'Ascq, France

(3) Centre National d'Exploitation des Océans, 39 avenue d'Iéna, 75116 Paris ; Centre Océanologique de Bretagne, B.P. 307, 29273 Brest, France

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## 1. INTRODUCTION

This paper is an intermediate stage in a systematic attempt at reconstructing the evolution from the early Mesozoic Tethys ocean to the present Mediterranean seas, within the framework of plate tectonics. In an earlier publication (Biju-Duval et al, 1976a), we outlined the way in which we approach this reconstruction, which incorporates, in a synthetic fashion, the very large amount of data now available on the continental margins as well as on the folded belts surrounding the Mediterranean area (Biju-Duval et al. 1977). The main result of this effort is a series of paleogeographic maps at key geologic periods which were prepared with the collaboration of V. Apostolescu, for the distribution of the main geologic facies, and of J.C. Sibuet, for the plate kinematic reconstructions (Biju-Duval et al., 1976 b). These maps are presented in such a way that they can be easily used, for addition or modification, by other scientists.

The purpose of this paper is to give to the reader the necessary elements to have a clear knowledge of the basic methodology followed by us and to outline the main conclusions, the major problems and the directions along which we intend to proceed in future stages.

We have limited, rather arbitrarily, the present reconstruction to the western part of the alpine system which surrounds the Mediterranean seas, Black sea included. It will be obvious to the reader that a full understanding of the eastern Mediterranean area would require a broader framework including the Caspian sea area and extending to the Oman sea to the south-east.

Starting from the present structure of the Mediterranean region, (Ryan et al, 1971, Biju-Duval, 1974, Biju-Duval et al, 1974), we tried to investigate how the geologic evidence for past distension, for the distribution of epicontinental and pelagic facies, and for the superposition of tectonic deformations in the folded belts fit within the pattern of relative motion between Africa and Eurasia since early Mesozoic time, as given by plate kinematics.

In the last few years the structure of the Mediterranean seas has been intensively studied (Auzende et al., 1972 ; Rehault et al., 1974 ; Finetti and Morelli, 1973, 1974 ;

Biju-Duval et al., 1977 ; etc ...). Plate I\*\* shows that it is characterized by narrow zones of thin crust inserted within the alpine folded belt, along the fairly broad Africa-Eurasia plate boundary which is outlined by a diffuse but intense seismicity (e.g. Mc Kenzie, 1972). The limited size of these oceanic basins has been the source of considerable controversy. Actually, only a small part of the surface of the Mediterranean basins are located more than 100 km away from the continental shelves. The continental margins, which mark the transition from continental to oceanic crust have an average width of 100 km or more (e.g. Worzel, 1965) and their deep structure is characterized by a rapid evolution in space and time. It cannot be easily related to either oceanic or continental crust. As a result, the nature of the crust, in the Mediterranean basins, where opposite continental margins can be considered to be contiguous, is necessarily complex and difficult to characterize clearly as either oceanic or continental.

However, we consider that the oceanic nature of the crust underlying the abyssal plains of the western Mediterranean and the southern Tyrrhenian seas is well established. The very large accumulations of sediments over the deepest parts of the eastern Mediterranean and the Black seas indicate that the igneous crust is very thin (Biju-Duval et al, 1976 a). In addition, the existence of deep seismic zones under the Tyrrhenian and Hellenic arcs (e.g. McKenzie, 1972) indicates that active subduction of crust has proceeded there for the last few million years. We consequently also assume an oceanic nature for these portions of crust. The age of formation of the basins can be estimated on the basis of a study of the continental margins which surround them. Although they have often been affected by tectonics, it is possible to recognize that some of them are quite old (Mesozoic) whereas others are much more recent. We estimate on this basis that the eastern Mediterranean basins were formed in Jurassic time, the Black sea in Cretaceous-Eocene, the Western Mediterranean in Oligo-Miocene (Le Pichon et al., 1971 ; Hsü, Montadert et al., 1977), the Tyrrhenian since Upper Miocene (e.g. Morelli, 1975). The present Pannonian basin (see Channell and Horwath, 1976), although not of an oceanic nature, was also formed by distension in Neogene (Boccaletti et al., 1976, Biju-Duval et al, 1976 a).

\*\* Plates I to VIII are inserted in a paper pocket placed at the end of the volume.

## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

The notion of resorbable oceanic lithosphere is fundamental to any reconstruction of folded belts. It requires however to be able to clearly identify traces of former continental margins, whether active or passive. Unfortunately, our ability in this domain is still very limited as shown by the numerous differences between the interpretations made by different authors.

In addition, it becomes ever clearer that considerable shortening may also occur through folding and thrusting of continental crust. For example, in the Alps (Trumpy, 1975) and in the Hellenides (Dercourt et al. 1976), the basement nappes obviously originated from a much wider continental platform than the present folded belt. The numerous studies made within the perimediterranean basins have shown that, to restore the original basement platform, it is necessary to unfold the present tectonic structures, due to a superposition of tectonics, over distances which often exceed 500 km. Thus, the reconstruction of the original cratonic plates requires such a palinspastic restoration.

Finally, as vividly illustrated by McKenzie (1972), the present pattern of the Africa-Eurasia collision is dominated by the fragmentation of the continental crust through a wide belt of deformation. This state of fragmentation through intracontinental collision processes was probably initiated in Upper Cretaceous. It is the present complex pattern of fragmentation which probably led to the identification of numerous (20) plates which evolved independently since early Mesozoic in the Dewey et al. (1973) reconstruction. This reconstruction of the western alpine system is probably the most widely used. But we prefer to consider that the present complexity is mostly the result of the late Mesozoic and Cenozoic intracontinental collision ; the main independent plates should be identified only on the basis of an association of unambiguous evidences for the existence of former plate margins.

As was done in the pioneering work of Dewey et al (1973), we have placed our reconstruction within the framework of plate kinematics. The solution adopted for the Africa-Eurasia relative motion was essentially derived from the work of Pitman and Talwani (1972) from anomaly 31 (68 M.y.) on, whereas the earlier kinematics is obtained from the work of Le Pichon et al (1977). It

will not be discussed in detail here.

Similarly, the bases for our choice on the present structure of the Mediterranean basins have been discussed by Biju-Duval et al (1977) and will not be presented again here.

Finally, due to space limitations, we will not comment in detail the geological data which are used in the paleogeographic maps of (PL. II to VIII). We will only give brief general comments and will postpone a more elaborate discussion to a later paper.

Many of the earlier attempts at reconstructing the alpine evolution are partial attempts which only consider portions of the Mediterranean system (Laubscher, 1969, 1971 ; Dercourt, 1970 ; Alvarez 1972, 1974 ; Boccaletti et al. 1974, etc ...). Others ignore a considerable part of the available information in one or several of the three major domains : plate kinematics, present geophysical structure, geological structure (1) (Smith, 1971 ; Bosellini and Hsü, 1973 ; Channell and Horwath, 1976, etc ...). Our attempt, on the contrary, is built on a broad base in each of these three domains.

But there is one major set of data which we have not used for the reconstruction. These are the so-called "absolute" position data which relate the position of a plate to the axis of rotation of the earth. We feel that there is still too much uncertainty to use these data to position the intermediate plates. We did, however, position the paleolatitudes on our paleogeographic maps modified from the paleomagnetic work of Van Der Voo and French (1974).

A last word should be said about the nomenclature adopted. We call Tethys the ocean which existed between Europe and Africa during early Mesozoic. We call Mesogea the fairly broad ocean which was created during Mesozoic, and we reserve the name of Mediterranean to the present basins, parts of which are relics of the Mesogea.

(1) Numerous reconstructions have been recently proposed : Herz and Savu, 1974 ; Dimetrijevic and Dimetrijevic, 1973 ; Géczy, 1973 ; Hawkesworth et al, 1972 ; Bernoulli and Laubscher, 1972 ; Boccaletti et al. 1974 a, 1974 b ; Dimetrijevic, 1974 ; Hadzi et al. 1974 ; Radulescu and Sandulescu, 1973 ; Dal Piaz, 1974 ; Dal Piaz et al, 1972 ; Hsü, 1971 ; Boccaletti et al., 1971, 1972 ; Elter and Pertusati, 1973 ; Mattauer and Proust, 1976 ; etc ...

## II. METHODOLOGY

1. The framework of our attempt is the plate kinematic evolution of the Africa/Europe plate system obtained from an identification of the Vine and Matthews (1963) isochrons in the Atlantic ocean between Africa and North America and between Europe and North America. The first systematic attempt at identifying the magnetic anomalies on a large scale over these two areas is the one by Pitman and Talwani (1972) which was the basis of the Dewey et al (1973) kinematic solution. Unfortunately, mostly because of fairly slow spreading rates, the anomalies are difficult to identify and the Pitman and Talwani (1972) identifications cannot be considered as final. Several modifications or additions have been proposed since, which include, Williams (1975), Kristoffersen and Talwani (1977) and Kristoffersen (1977) for the Europe-America system, Hayes and Rabinowitz (1975), Barret and Keen (1976) for the Africa-America system and others will undoubtedly come in the near future.

In addition, the kinematic solution is obtained by fitting together corresponding anomalies from both sides of the ridge crest. However, in general, we do not have a good definition of the isochrons over a sufficient length to obtain a very precise fit. As a result, the Africa-Europe relative motions which are obtained by composition of the Africa-America and America/Europe motions cannot be expected to be known with a good precision. For example, Francheteau (1973), using the anomalies identified by Pitman and Talwani (1972) as well as informations on fracture zones trends, has shown that a somewhat different kinematic solution could be obtained for the Africa-America motion. Kristoffersen and Talwani (1977) and Kristoffersen (1977) have obtained a different kinematic solution for the America/Europe motion mostly because of additional anomaly identifications to the north of 50°N.

The complexity of the kinematic problem is increased by the existence of two relatively mobile zones on each side of the Açores-Gibraltar line. Iberia has had an independent motion from Europe in Late Mesozoic and part of Early Cenozoic and the anomalies between Iberia and America cannot be used to obtain the America/Europe motion without allowing for this independent displacement. Similarly there is some indication that Morocco has been affected by limited

displacement with respect to Africa during the opening (e.g. Le Pichon et al., 1977) and the anomalies between Morocco and America should only be used with caution to obtain the Africa-America motion.

Finally, there is some significant uncertainty on the absolute ages of these Vine and Matthews anomalies. The most recent Cenozoic polarity time scale by Tarling and Mitchell (1976), which we have adopted, differs considerably from the Heirtzler et al (1968) time scale or even the one proposed by Larson and Pitman in 1972. For example, according to Larson and Pitman (1972), anomaly 24 is 60 M.y. old (Middle Paleocene) whereas it is 49 M.y. old (Lower Eocene) for Tarling and Mitchell (1976). For the Mesozoic, we have adopted the time scale proposed by Thierstein (1977). The maximum difference between the time scales may reach 11 M.y. for a given anomaly, as shown in the example above.

However, the greatest difficulty in the reconstruction of the Africa/Europe kinematics is related to the early stages for which we do not have Vine and Matthews isochrons but rather have to rely on a reassembly of the continents prior to the break-up. Yet this reassembly is difficult to obtain on purely objective ground. It becomes abundantly clear that no given isobath can be considered to reflect accurately the edge of the original continental crust break-up (Le Pichon et al, 1977). As the break-up between Europe and America most probably did not occur before 95 - 90 My (Cenomanian), all reconstructions prior to that time are plagued by the fairly large uncertainty in the Europe-America reconstruction. This is illustrated in figure 1 which shows the large differences produced by the different Europe-America solutions prior to 90 M.y. At one extreme, there is the solution of Bullard et al. (1965) which fit at the 1000 m isobath. It is not acceptable as it is now known that large portions of the sea-floor, down to a water depth of 4000 m in places, are continental. For example, between Europe and North America, this is true for the Rockall plateau in its entirety, the Orphan Knoll area, the area just south of the Porcupine sea-bight, the Galicia bank etc .. (e.g. Le Pichon et al., 1977). At the other extreme there is the most recent solution by Kristoffersen (1977). Actually, the initial fits obtained from a study of the

FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

ocean floor features and continental margins have regularly tended to get looser from Bullard et al (1965) to Dewey et al (1973) to Le Pichon et al (1977) and Kristoffersen (1977). We feel that the actual position is probably within 200 km of the Le Pichon et al (1977) solution for Europe-North America and that the resulting position for Africa - Europe may be eventually in error by as much as 300 km to the west and 500 km to the east.

The precision, however, gets considerably better posterior to 90 M.y., as shown in figure 1. It is of the order of 200 km to the west and somewhat more to the east. It may be fairly poor for the Campanian-Maestrichtian as Kristoffersen (1977) has suggested that Pitman and Talwani (1972) generally identified anomaly 31 (which is Maestrichtian, 68 M.y.) where it should be 33 (which is Campanian, 76 M.y.).

In this paper, we do not propose any new magnetic anomaly identifications or kinematic solutions. Rather, we have chosen among the solutions proposed what seems to us to be the most probable, recognizing that it may be illusory at the present time to expect a precision in the position of Africa with respect to Europe better than 300 - 500 km prior to 90 M.y. ago and 200 - 300 km later than that. Although these possible errors are large, they only represent about 20 % of the motion of Africa with respect to Europe since the corresponding time.

Table I indicates the solution adopted and figure 2 illustrates the resulting kinematics. The solution chosen is essentially based on the work of Pitman and Talwani (1972) from anomaly 31-33 on. However, between Africa and North America, we have used the slightly different finite rotation angles proposed by Francheteau (1973). Between Europe and North America, we have also used the parameters proposed by Kristoffersen (1977) for anomaly 24 (49 M.y., Ypresian) and anomaly 31-33. Prior to anomaly 31-33, we adopt the solution proposed by Le Pichon et al (1977) for the initial reconstruction and the Oxfordian reconstruction. The Barremian position (anomaly M2) is not very precise and the Cenomanian position which is an interpolation, even less precise. The latter one has not been used in this paper.

2. Within this kinematic framework, we

have chosen seven key geologic periods for which there was a fair control of the relative position of Africa and we have established a paleogeographic map for each of these periods (figure 3 and plates II to VIII). Starting from the present time, these are :

- Tortonian, 9 M.y., anomaly 5 (plate VIII) : just before the Messinian event
- Stampian, 35 M.y., anomaly 13 (plate VII) : just before the opening of the western basin.
- Lutetian, 44 M.y., anomaly 21 (plate VI), which corresponds to an important tectonic event in the whole Mediterranean area (Pyrenean phase)
- Campanian - Maestrichtian, 76-68 M.y., anomaly 33-31 (plate V), which represents the main period of evolution of the future eastern Mediterranean area
- Tithonian-Berriasian, 140 M.y., interpolated between anomaly M 25 and M 2 (plate IV) because of the very important sedimentary and tectogenetic processes (first obduction of ophiolites)
- Dogger, 165 M.y., interpolated between anomaly M 25 and the initial reconstruction (Plate III) during which the main isopic zones are defined as a result of the liassic distension
- Upper Triassic, about 200 M.y., initial fit (plate II) corresponding to the first distensional phases between the Tethys and the future Atlantic ocean.

In addition, plate I gives the present situation. We have not made the Cenomanian reconstruction (112 M.y.), although it is a key geologic period, as we felt that the relative position of Africa was poorly known.

However, it is obvious that, for a given geologic period, the major events that should be illustrated are not synchronous over the whole area. It should be stressed, further, that the dating of geologic facies or events may still be quite imprecise. For example, plate IV shows data concerning the Campanian as well as the Maestrichtian era. Actually, each map shows the general geological pattern in the vicinity of the chosen period and it may include events which occurred immediately before or after the date chosen.

## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

Thus, we often find the same range of possible errors in the chronology of geologic data as in the kinematic data, which confirms that it is probably illusory at the present time to look for much better accuracy in such a reconstruction. It is quite clear that the precision of our paleogeographic reconstructions gets worse with increasing age. The following features have been shown on each map :

The repartition of continental and oceanic areas with a transitional zone which corresponds to the continental margins. This continental margin area distribution has been determined in a very schematic fashion. We must keep in mind, as previously said, that their deep structure is characterized by a rapid evolution in space and time, especially in the case of the active margins. So the size and the shape of such margins have been drawn with some modifications from one map to the other. Moreover, we have not attempted to illustrate systematically the possible existence of marginal basins which have since disappeared. Although we do not exclude this possibility, we have only shown on the plates those for which the evidence seems to us to be unambiguous.

The repartition of major facies in each zone based on both our field experience and a detailed bibliographic study which cannot be exposed in full in this paper for lack of space. Taking in account the previous remarks about the precision of dating and the fact that many rapid changes of facies are known to occur in space and time, we have only drawn in a schematic fashion the main facies, many fine variations being impossible to show at the scale of the maps.

To show the lateral variations we had to adopt a certain number of hypotheses on the initial positions of the different paleogeographic and (or) structural units within the different belts. For example the Rif, Kabyle and Calabrian basement units have been considered to be parts of the southern margin of Iberia to the north of the "Mauretania-Massylian" trough. Our choice is illustrated in figure 4 which should be used as a reference map for plates II to VIII. To establish these reconstructions one needs to reposition at each stage the paleogeographic zones ; this requires analysing each part of the Alpine belt in which superimposed tectonics are known. Because of lack of space it is not possible to mention here the data we have used on the different branches of

the Alpine belt, data which are necessary to understand, and to explain the paleogeographic or tectonic hypotheses chosen. This will be done in a future paper. Nevertheless figure 4 localizes the key paleogeographic zones which existed after the liassic distension. It shows also the main bibliographic references in a condensed form (author, year) even if we have not adopted the conclusions of the mentioned authors. The figuration of zones within stable areas is not difficult but to represent the former margins now included in "internal zones" is not easy : they correspond to the most tectonized, polyphased, granitized and metamorphosed areas. The study of present margins show that they are formed by horsts and grabens which size and shape are difficult to define in these former margins. It is also known that the type of deposits varies enormously over short distances on present continental margins (closeness of clastics sources, relative bathymetry-carbonate compensation depth, currents, ..).

Moreover we have used a study of the major lineaments revealed by the Landsat I satellite to define major fractures related to continental collision (Biju-Duval et al, 1976).

A possible or probable kinematic pattern, although in many cases it would be possible to reconstruct fairly different kinematic patterns. The main rationale was one of minimum complexity. It is clear that onland geological data have been used in priority. For example the relation between facies, tectonic development, overthrusting of ophiolites, calcoalcaline volcanism in north Anatolia during Late Cretaceous leads us to suggest the existence of a subduction zone. In a few cases clear geological data do not exist but the kinematic evolution obliges us to infer the existence of a plate boundary. The minimum complexity criterion has led us to identify only three main plates between Africa and Europe, which are Iberia, Apulia and Anatolia. The size and shape of these intermediate plates is obtained from a rough palinspatic restoration for Apulia and Iberia. For Anatolia the geological data are often so imprecise that we simply chose a size roughly compatible with the inferred kinematic evolution ; it should in particular result in continental collision by Late Cretaceous.

The major volcanic manifestations and the main phases of tectonic emplacement of ophiolites. Because the knowledge of the geochronology, petrography and geochemistry is still

## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

quite variable from one area to another a synthetic view of the volcanic activity all around the Mediterranean area is not yet well established (see Bellon and Letouzey, this volume). Nevertheless we have mentioned the main occurrences of widespread volcanism (including resedimented volcanic deposits). We have assumed in this paper that the obduction of ophiolites was related to the last stage of disparition of an oceanic area. This is not the only possibility (see for example Dewey, 1976) and we actually chose a different one in our preliminary publication (Biju Duval et al, 1976 a).

Finally, we note that these maps are prepared in such a way that they can be easily adapted or modified by others in order to obtain a better fit to other sets of data. The solution chosen is only one among several possible at the present time and we are still seriously considering an alternative solution in which the continental collision between Apulia and Europe only occurs in Late Cretaceous. All the maps are at the same scale in an oblique Mercator projection (pole at 52° N, 150° E) with Europe in its present position. The scale of the map is given by the latitude marks on the side. The paleolatitudes are figured modified from Van Der Voo and French (1974).

### III. PROPOSED MODEL OF EVOLUTION

Figures 2 and 3 illustrate schematically the displacements of Africa with respect to Europe (arbitrarily kept fixed in its present position) since the beginning of the opening of the Atlantic ocean. Three main phases can be recognized.

a) From Mesozoic to Late Cretaceous, Africa is moving left laterally at a fairly high rate (2 to 4 cm/yr).

b) In Upper Cretaceous, the motion is slowed, most probably because of the first major intracontinental collision between Arabia, Anatolia and Eurasia. A relatively modest right lateral motion of Africa follows between Upper Cretaceous and Ypresian (49 M.y.), being related to the initiation of spreading between Europe and North America.

c) Finally, from Ypresian on, there is a general state of intracontinental collision which is characterized by a trigonometric rotation of Africa around a eulerian pole situated in the vicinity of Morocco. Thus the motion is very limited to the west but

increases rapidly to the east where it is essentially north-south. Note that the pole of rotation tends to migrate with time to the west in the Atlantic ocean, which results in increasing compression in the whole western Mediterranean area.

The corresponding paleogeographic evolution is shown in figure 3 and plates II to VIII. During the first phase (prior to Upper Cretaceous), the geologic structure is mostly the result of the plate tectonic evolution of the three main intermediate plates, Iberia, Apulia and Anatolia. From Early Cretaceous on, the geologic structure corresponds to the fragmentation of the intermediate plates, with progressive disparition of the remaining oceanic areas. It is a process of continental collision where the concept of individual plates is difficult to use any more.

In this paper, the motion of Iberia has been obtained from kinematic criteria in the Bay of Biscay, as proposed by Le Pichon et al (1970, 1971), Le Pichon and Sibuet (1971) and Choukroune et al (1973). We do not have such criteria to obtain the motion of Apulia. We chose a solution which is compatible with the Africa-Europe motion and results in collision with Europe in Late Jurassic at the time of emplacement of ophiolites in the Hellenides (Dercourt et al, 1976). We still consider as a viable alternative a solution where collision only occurs in Upper Cretaceous and ophiolites obduction is not related to this continental collision. Note that the formation of the Bay of Biscay and of the Mesogea is not of a "marginal sea" type but rather is due to the displacements of intermediate plates induced by the motion of the main plates. On the other hand, the Black Sea and the Western Mediterranean probably represent fossil marginal basins.

The paleogeographic map will be commented starting from Early Mesozoic. However, their reconstruction was done starting from the present situation (plate I\*) and moving back and forth in time.

#### 1. Upper Triassic (plate II)\*

The opening fit, which is pre-middle-Jurassic, is most probably valid for the whole Triassic time. We have chosen to draw the paleogeography of the Upper Triassic (200 M.y.). We previously discussed the

\* Plates I to VIII are inserted in a paper pocket placed at the end of the volume.

## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

fairly large inaccuracy of this reconstruction, especially in the Iberia-Morocco area.

a) Plate II shows that a broad Tethys ocean exists to the northeast although. Argyriadis (1975) has argued that there is no geologic evidence for oceanic facies in this area. Nevertheless, the distribution of fauna may suggest the existence of such an oceanic hiatus (Enay, 1972 ; Hirsch, 1976). It could be argued further that the distribution of the epicontinental facies is in fair agreement with the existence of an eastward Tethys. For example, north of Apulia, platform carbonate facies merged into deep water deposits (Bernoulli and Jenkyns, 1974). Sediments associated with pillow-lavas have been described but their present tectonic setting is complicated ; so it is still difficult to define exactly the relation between oceanic and marginal facies (Terry, 1971 ; Juteau, 1970). Evaporite germano - andalousian facies extend of Europe and over North Africa ; Nubia west facies are found to the south east ; and carbonate platform deposits are found all along the margins of the Tethys and along the areas of future rifting (Bernoulli and Jenkyns, 1974).

b) Upper Triassic is the period when the first distension occurred (Fallot, 1945, 1948 ; Jenkyns, 1970 ; Ginzburg et al, 1975 etc ...) along the future continental margins of the Mesogea and the future Atlantic ocean. However, the distension on the eastern margins of Apulia, which already existed at that time, is not properly explained by this model, unless one relates it to the formation of marginal seas (Blanchet, this volume). It would accordingly be interesting to be able to distinguish between the different types of volcano-sedimentary deposits in the now tectonized Dinarides - Hellenides folded belts (Circic, 1966 ; see Cadet et al, this volume)

c) In plate II, the gap between Apulia (still connected to Africa) and the African margin is too large. It could be reduced in part with a different more compact fit of the now dissociated parts of Iberia (including Corsica-Sardinia, Calabria, East end West Kabylas, Rif etc ...) such as the one proposed by Bayer et al. (1973) and Le Pichon et al (1977) or with a different position of Northern Africa. In this paper, we have adopted the more conservative reconstruction of Corsica and Sardinia proposed by Biju-Duval et al (1977). The

gap would be greatly increased, in any case, if the position for Iberia is the one proposed by Bullard et al (1965). We feel that this gap may not be significant in view of the large inaccuracies involved in such a reconstruction.

d) Another problem is the interpretation of the Isparta line of Turkey (Brunn et al., 1971) as corresponding to an ancient Transform Fault separating Apulia from Arabia - Anatolia, in spite of very little field evidence for it.

e) A better knowledge of the Cimmerian folding phase in Dobroudgea, Crimea, Caucasus would probably lead to a modification of the scheme proposed in the north-eastern area where an active margin already existed. Continuous deformations and thrusting occurring NW of the Pontides until Triassic time and including so-called "green rocks" (Fourquin, 1975) could indicate the existence of a subduction zone south of Europe.

### 2. Dogger (pl. III)

The distension gets generalized in Liassic time during which the left-lateral rotation of Africa, related to the opening of the Central Atlantic, has begun

a) We have chosen a position close to the position corresponding to M 25 (150 M.y., Oxfordian) in the absence of other magnetic anomaly data.

b) The left-lateral motion of Africa results in distension in Southern Iberia ("dorsale") (Fonbote et Quintero, 1960 ; Busnardo et Chenevoy, 1962 ; Paquet, 1974), in the Alps (Briançonnais) (Bourbon et al, this volume), in East Sicily and along the Levantine margin (Gvirstman, this volume) which can be considered as passive continental margins. These correspond to the formation of zones of Transform Faulting where the oceanic space is small (Alps) or even absent (Pyrénées, South Iberia) or to the formation of the future Mesogea (Monod, 1975), in the wake of the motion of Apulia away from Africa.

c) The left - lateral motion of Africa also results in the beginning of consumption of the Tethys with corresponding formation of marginal seas (Eastern Black Sea - Caspian, Dobroudgea sphenochasm). Important volcanism and granitisations known in the Pontides and Caucasus (Bergougnan, 1975 ;



## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

Khain, 1963) could be related to a such phenomenon.

d) As a result of this evolution, the sedimentary facies get greatly differentiated around Apulia and in areas of subsidence near the margins of Iberia, Africa and Europe ("Citrabetic and Betic troughs" in Iberia, Azema et al. 1974, Atlas trough, Busson 1972, etc ..). Broad epicontinental seas extend on the cratons.

e) As discussed previously, the positions of Apulia and Anatolia are fairly arbitrary. The size of the Mesogea depends of course on the positions adopted. The connection between Apulia and Anatolia is also quite speculative. The absence of known oceanic facies between the two plates leads us to consider that they were separated by a Transform continental Fault with left- lateral motion at that time.

### 3. Tithonian - Berriasian (pl. IV)

The position of Africa was obtained by interpolation between the position corresponding to anomaly M 25 (Oxfordian) and anomaly M2 (Barremian).

a) We consider that this is the time of a first limited closure of the Tethys between Apulia and the Moesian area because this is the main time of tectonic emplacement of ophiolites in the Hellenides (Dercourt, 1970 ; Aubouin et al, 1970 ; Juteau et al, 1973 ; Terry, 1974) which we interpret here as due to incipient continental collision (Maxwell, 1970 ; Hynes et al, 1972). We previously noted that we consider as possible another solution where collision only occurs in Upper Cretaceous time in this area (Biju-Duval et al, 1976 a). In our present scheme, this collision will initiate the fragmentation of Apulia.

b) The oceanic realms, to the east (North Anatolia<sub>n</sub> trough) and west (Alpine Margins) (Bezzi et Picardo, 1971 ; Decandia and Elter, 1972 ; Dietrich and Scandone, 1972 ; Contesogno et al., 1975) of the continental collision (Rhodopian), are fairly restricted. Two marginal basins have been created, the eastern Black Sea and the Carpathian flysch trough ("Siret Ocean" of Herz and Savu, 1974 ; Boccaletti et al. 1974 b) which are affected by important volcanic activity.

c) This very important event is followed by the wide-spread deposition of flyschs during Lower Cretaceous. We note the occurrence of such clastic turbiditic deposits not only along the Hellenides - Dinarides tectonized margins (Blanchet et al. 1969 ; Charvet, 1973 ; Terry et Mercier, 1971 ; Celet et al. 1976) but also in other areas (Durand-Delga et Lambert, 1955) extending to the Atlantic ocean. The flysch deposition may be, at least partly, related to a major change in the paleogeography of the continental areas (Purbeckian facies in England, France, Spain, etc ...) accompanied by a general regression (see Histoire structurale du golfe de Gascogne, Technip edit.). This is also the time during which the Bay of Biscay and the North African trough (Bouillin et al., 1970 ; Bouillin et Kornprobst, 1974 ; Obert, 1974) were initiated.

d) The size of the Mesogea depends on the position chosen for Apulia. It is possible that an active margin started at this time along the southern border of Anatolia, inducing the formation of a marginal sea (the future Troodos - Kizil Dag).

### 4. Campanian - Maestrichtian (pl. V)

The Barremian (anomaly M 2) and Campanian - Maestrichtian (anomaly 31 - 33) positions being fairly inaccurate, we have not used the interpolated Cenomanian position, in spite of the important paleogeographic changes occurring at that time. The Cenomanian saw an early tectonic event in the Alps, the initiation of new flyschs and a broad epicontinental transgression.

However, the Uppermost Cretaceous is a major period in the evolution leading from the Tethys to the Mediterranean, as it is the time of the final consumption of the Tethys. The left-lateral motion of Africa is stopped due to continental collision between Europe-Anatolia and Africa - Arabia. This continental collision, which is the first true collision between the Africa - Arabia and Europe plates, will continue from then on.

a) The consumption of the Tethys is marked by a broad phase of obduction of ophiolites and mélanges (accretionary prisms) south of the Tethys (Pontides, Bergougnan, 1975 ; Apuseni , Lupu, 1974) and north of Africa - Arabia (from Cyprus to Oman; Gass, 1968 ; Lapierre, 1976, Ricou, 1971). An important calco-alcaline volca-

## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

nism from the Balkans to the Lesser Caucasus (Adamia, 1975 ; Boccaletti et al., 1974) accompanies the final oceanic subduction in this area ; the completion of the opening of the Black Sea is probably related to this event (see Letouzey et al., this volume). The Carpathian Sea continues to evolve as a marginal basin.

b) The narrow oceanic trough which ran from the Tethys to the Atlantic north of Apulia has completely disappeared along the Alps (See Tollman, 1969 ; Haccard et al., 1972). To the south, on the contrary, the Mesogea, which is broadly opened toward the Atlantic, has reached its maximum size. Flysch continues to be deposited on both margins of the North African trough (from Lower Cretaceous to Eocene) (see Paquet, 1974). The tectogenesis of the Pyrénées, related to the final stage of opening of the Bay of Biscay, is now engaged (Choukroune et al, 1973 ; Mattauer et Henry, 1974).

c) This general collision to the east results in the fragmentation of Apulia and Anatolia.

The motion of Africa will now change and, until Ypresian (49 M.y.), will correspond to right lateral strike slip, because of the faster opening between Europe and North America than between Africa and North America. After this limited right-lateral motion, Africa will start rotating trigonometrically around a pole near Morocco, resulting in pure compression to the east (fig. 3).

### 5. Lutetian (plate IV)

a) To the west, this is the time of consumption of the narrow western extension of the Mesogea between Africa and Iberia, attested by overthrusts which include fragments of ophiolites (Durand-Delga, 1971 ; Andrieux et Mattauer, 1963 ; Bouillin et Kornprobst, 1974). Paquet (1974) has suggested that transverse wrench faults cut the southern margin of Iberia at this time. The small maximum size of this oceanic area and the slow rate of consumption might explain the absence of calco-alcaline volcanism.

b) This is also the time of the main phase of tectogenesis in the Pyrénées

(Mattauer et Henry, 1974) and the Alps (see Debelmas et Lemoine, 1970 ; Caby 1973). The fragmentation of Apulia may help to explain the arcuate development of the Alps, as proposed by Laubscher (1969, 1971).

c) To the east, the oceanic area gets smaller, mostly by Transform faulting along the northern border of the Mesogea. Anatolia is now fully accreted to Arabia but a new active margin is initiated south of the Caucasus (see Khain, this volume).

d) The collision in the Rhodopian - Moesian area generates underthrusting associated with volcanism (Boccaletti et al., 1974) and new continental fragmentation with movement toward the west, to the west, and toward the east, to the east.

### 6. Stampian (plate VII)

a) The process of continental collision with progressive consumption of the oceanic areas continues to the east. In particular, the old marginal Carpathian basin (initiated in Jurassic) begins to get reduced in size (Contescu, 1974).

b) The present Western Mediterranean still does not exist (Wezel, 1974). However, to the north, the rifting will take place during Upper Oligocene - Aquitanian (Le Pichon et al., 1971 ; Hsu, Montadert et al. 1977). To the south, we assume the formation from Gibraltar to Calabria of a marginal basin located north of the continental fragments of Iberia (Kabylies, Peloritian and Calabria) along the North African consuming plate boundary. The Numidian flysches are deposited both in internal and external positions as proposed by Caire (1971), Wezel (1970), Hoyez (1974).

It is interesting to note that the fragmentation of Iberia corresponds to the migration of the Africa-Europe pole of rotation toward the west.

c) To the east, the Mesogea is being consumed from the west to the east. This will lead to the future development of the Tyrrhenian arc and the opening of the Western Mediterranean followed by the opening of the Tyrrhenian sea. It is the beginning of the Apennines orogenesis (Roman, 1970 ; Boccaletti et al., 1976 ; Haccard et al. 1972).

# FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

## 7. Tortonian (plate VIII)

The paleogeography gets very similar to the present one (Biju-Duval et al., 1977).

a) The Mesogea is being consumed along two arcs, the Calabrian and Aegean arcs and is progressively being reduced to the size of the present Eastern Mediterranean.

b) The opening of the Red Sea corresponds to a faster northward movement of Arabia. As a result, the fragmentation of Anatolia is accentuated. The North Anatolian fault is initiated and allows a migration of Anatolia to the west (e.g. McKenzie, 1972).

c) The Carpathian arc continues to evolve with the formation of the Pannonian basin, as a continental marginal basin (Stegena et al, 1975 ; Contescu, 1974).

d) The subduction ends to the west in North Africa (Magne et Raymond, 1974) and along Italy. It is the age of the main nappes emplacement. The western basin has essentially its present configuration although the Tyrrhenian is smaller than today. Its basins are deep, subsidence is continuous, and Messinian salt deposition will occur in the deepest area, the margins being then affected by erosion (Montadert et al. 1977).

## IV. CONCLUSIONS

The evolution from the Tethys to the present Mediterranean occurred in two main stages

### 1. From Triassic to Uppermost Cretaceous

Africa is moving left laterally with respect to Europe. The paleogeography is essentially controlled by the plate tectonic evolution of three intermediate plates, Iberia, Apulia and Anatolia, with respect to the two major plates, Africa and Eurasia. We adopted the hypothesis of a marginal basin for the Carpathian flysch basin and for the Black Sea. We did not however try to reconstruct the evolution of the Hellenides in terms of marginal basins, although we do not exclude this possibility which is compatible with our scheme.

In Early Mesozoic, the paleogeography is characterized by a wide Tethys ocean

between Europe and Africa. Its exact surface is difficult to define precisely. Previous studies had minimized the extent of the basement overthrusts in the Alps, the Dinarides and the Hellenides etc... Recent studies tend to increase them considerably (see Aegean Congress, Paris ; Tollmann, 1969 ; Derycke et al., 1974 ; Ferrière, 1974). The analysis of the different superposed metamorphic units of alpine age leads to an increasing surface estimate for the now tectonized initial cratons. Is it possible to go further than we have done and to completely eliminate the oceanic surface extending from the Carpathians to the Alps ? In any case, the existence of the Eastern Tethys between the Apulian margin, now incorporated into the Dinarides and Hellenides, the Anatolian margin, now incorporated into the Pontides, and the Rhodope and European margins, is difficult to doubt. The presence of the Eastern Tethys may be inferred from the volcanic activity in Pontides and Caucasus which suggests its subduction under Europe since Triassic. The geometry of the Tethys is partly controlled by the position of North Africa with respect to the old African shield. We did not take into account a possible deformation due to intra-African fractures, in particular the so-called "South Atlas" fault system (Dubourdieu, 1962).

The kinematics of Apulia cannot be directly deduced from the kinematics of the Africa-Europe plates system. Does Apulia collide with Europe as early as Uppermost Jurassic or does the corresponding Jurassic ophiolites obduction occur on the margin of an oceanic basin ? We have chosen, in this paper, the first solution whereas we developed the second one in an earlier paper (Biju-Duval et al., 1976 a). The large amount of clastic turbiditic deposits in the alpine belts during Upper Jurassic - Berriasian seems to us to be easier to interpret if high reliefs have been produced by continental collision. The alternative is still open.

From Triassic to Jurassic, a period of distension, which manifests itself in particular in the formation of continental margins affects the whole area. This is the time during which the Mesogea was formed in the wake of Apulia. After the Upper Jurassic-Lower Cretaceous collision of Apulia with Europe, intracontinental fractures cut Apulia in three sections.

## FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

During Cretaceous, the portions of Apulia situated to the east and to the west of the Upper Jurassic collision zone also collide with Europe. Consequently, the initial Tethys has been completely consumed in Maestrichtian time whereas the Mesogea has reached its maximum size.

### 2. From Uppermost Cretaceous to Present

Africa ceases to move left laterally and starts a right-lateral motion of limited extent until Ypresian. Then it rotates trigonometrically about a pole of rotation close to Morocco, thus resulting in north-south compression to the east. This is a time of intracontinental collision during which the concept of plates is difficult to use any more. The collision of Apulia and Anatolia with Europe results in an intracontinental fragmentation accompanied by large strike-slip faulting which allows the formation of important overthrusts in the Alps.

The Mesogea starts being consumed ; this consumption is accompanied by the creation of heterochronous marginal basins (the Balearic and Tyrrhenian basins). The remnants of the Eastern Mesogea correspond to the present Eastern Mediterranean where the Aegean marginal basin is barely initiated. The consumption of the Carpathian flysch marginal basin is closely related to the Pannonian basin distension which is marked by very active volcanicity.

We have not taken into account the paleomagnetic data for the reconstruction of the positions of the intermediate plates, as we feel that they still are inadequate for this purpose. Due to a lesser quality of geological data, the reconstructions of Anatolia are less precise than those of the other parts of the Mediterranean area.

We are aware that we did not have enough constraints to reconstruct unambiguously the position of Apulia and Anatolia. As we pointed out earlier, we previously chose another solution in which Apulia collides later with Eurasia with a much smaller rotation during Jurassic. If this is so, we would have only to change the Dogger and Upper Jurassic paleogeographic maps. It is clear that the answer to this problem lies primarily in a better knowledge of the evolution of the paleomagnetic poles for the intermediate plates.

However, although this model is far

from being final, it suggests some interesting new geologic problems.

a) Great differences exist between the Mesozoic and Cenozoic sedimentary history and tectonic development. During the Mesozoic, in spite of great uncertainties in the plates positions, the size of the oceans and the exact locations of subduction zones, a comparison can be attempted with present oceans and the tectonic evolution of their associated marginal basins. But, in Cenozoic, after the continental collision occurring in Late Cretaceous between Africa and Europe, the tectonic development changes completely because of reduced oceanic areas and intracontinental fragmentation with large strike-slip faulting which makes comparison with present oceanic margins quite dubious.

b) The former Tethys ocean has completely disappeared. The present Mediterranean basins are either remnants of the Mesozoic Mesogea created by rifting in the wake of the motion of Apulia - or marginal type basins (whether oceanic or intracontinental) associated to a subduction zone.

c) The recent history of the Mediterranean basins (which includes subsidence, the Messinian evaporitic episode, the continuous Africa-Europe collision) masks differences.

d) The importance given to intracontinental strike-slip faulting suggests a new examination of different portions of the Alpine belt.

e) The presence of a Mesogea which is being consumed to the north and west suggests a revision of the schemes of evolution which have been proposed for the Hellenides, and of the relationship between Calabria and the Apennines.

f) The major roles given to the Dobroudgea and Isparta geologic features should be confirmed by more detailed studies.

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FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

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FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

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TABLE I

## PARAMETERS USED FOR RECONSTRUCTION

Anomaly	Age in M.y.	Geologic Age	Rotation applied to Africa			Movement AF/NA	Movement NA/EU
			Latitude	Longitude	Angle		
	Initial	-	52.	- 0.4	- 49.6	Initial fit (1)	Initial fit (1)
M 25	148	Oxfordian	50.	- 2.1	- 41.4	Quiet zone fit(1)	Initial fit (1)
M 2	112	Barremien	45.7	2.6	- 31.5	(2)	Initial fit (1)
-	90	Cenomanian-Turonian	30.3	- 3.6	- 18.4	Interpolation	Initial fit (1)
31-33	76-68	Campanian Maestrichtian	20.1	- 11.3	- 16.0	(3)	(4)
24	49	Ypresian	31.9	- 16.6	- 14.8	(3)	(4)
21	44	Lutetian	33.7	- 18.8	- 12.9	(3)	(5)
13	35	Stampian	29.3	- 26.6	- 7.8	(3)	(5)
5	9	Tortonian	27.1	- 27.9	- 2.4	(3)	(5)

(1) Le Pichon et al (1977)

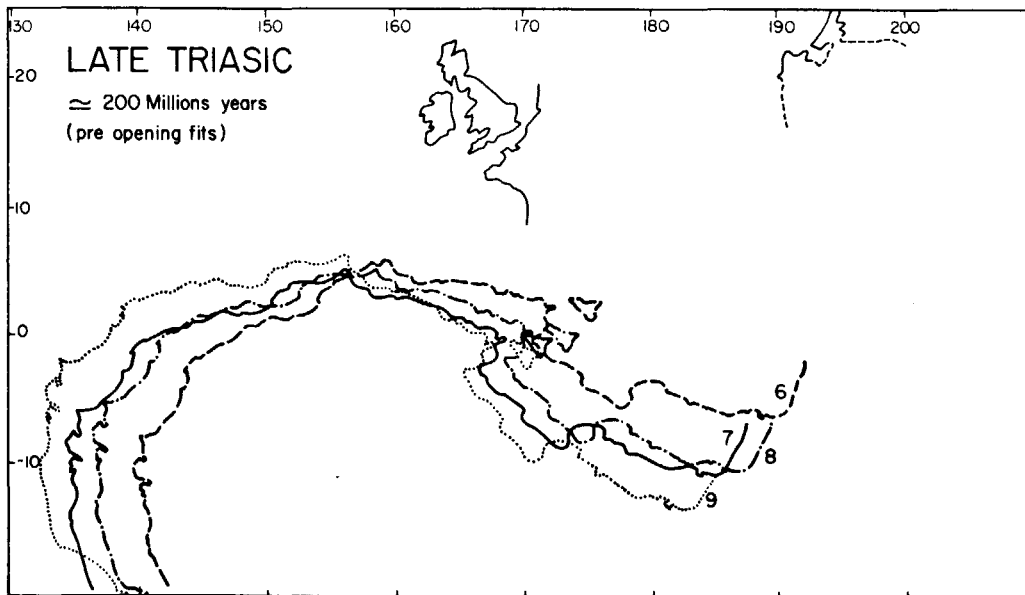
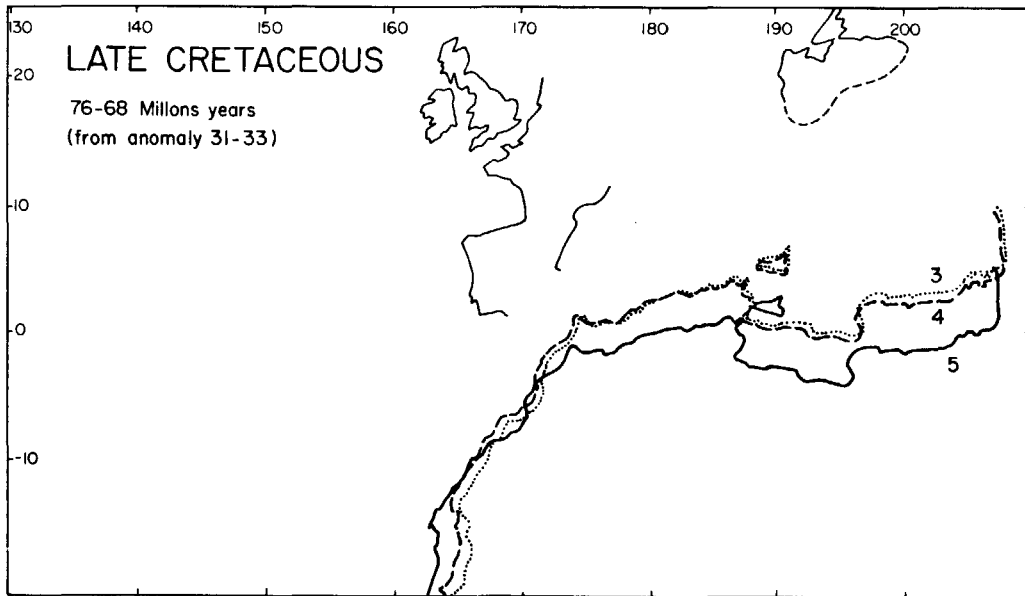
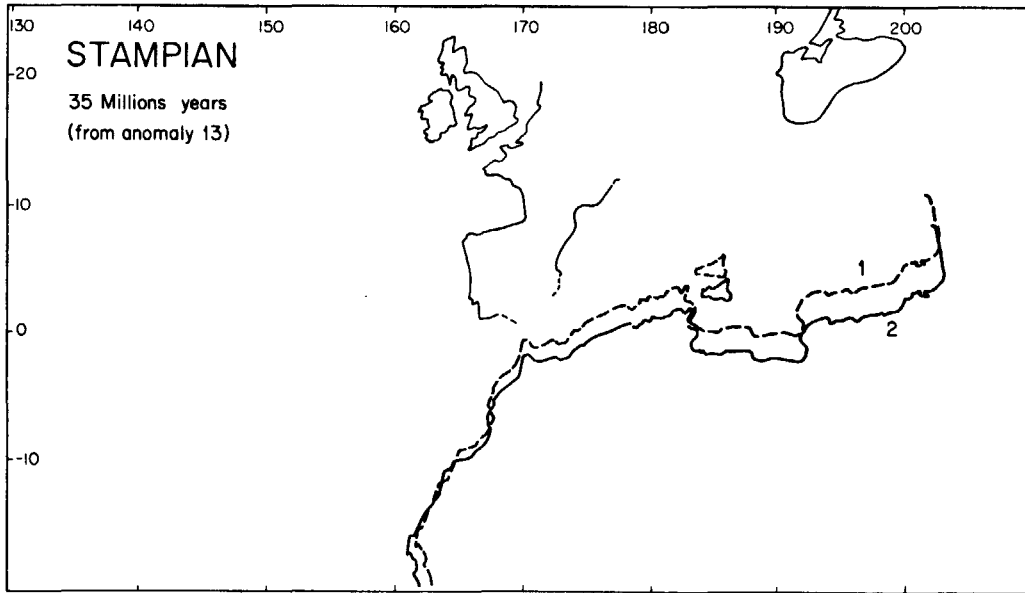
(2) Hayes and Rabinowitz (1975), Vogt et al (1971), Larson and Pitman (1972)

(3) Francheteau (1973)

(4) Kristoffersen (1977)

(5) Pitman and Talwani (1972)

Ages according to Tarling and Mitchell (1976) for anomaly 31 on, according to Thierstein (1977) prior to that. Positive latitude for north, positive longitude for east, convention for sign is negative when seen clockwise from pole of rotation.



FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.

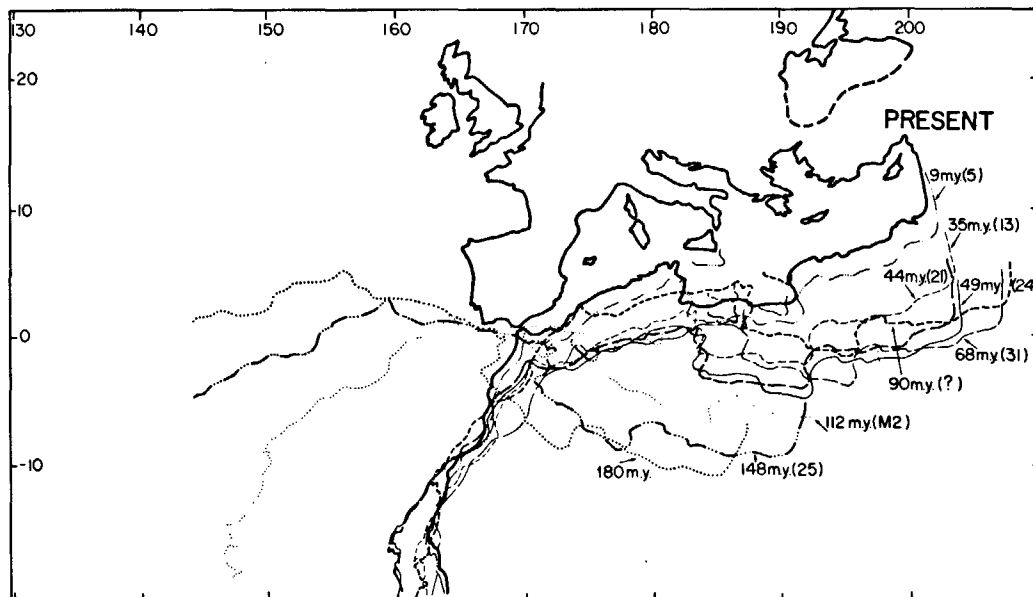
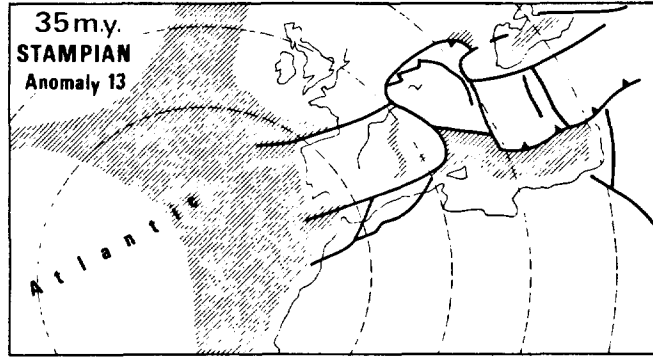
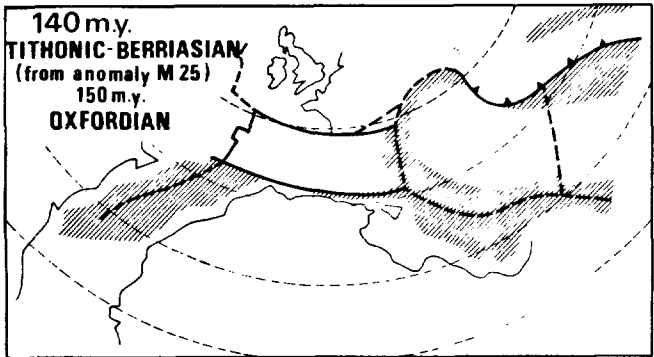
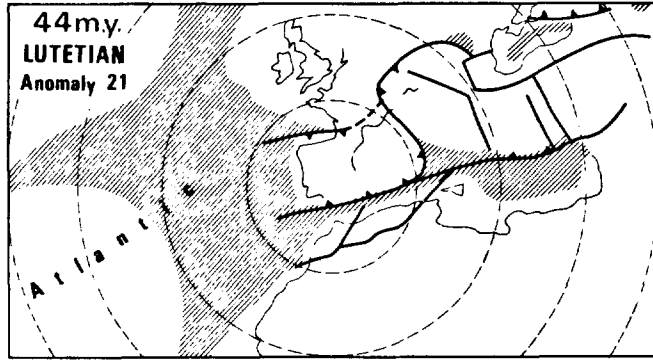
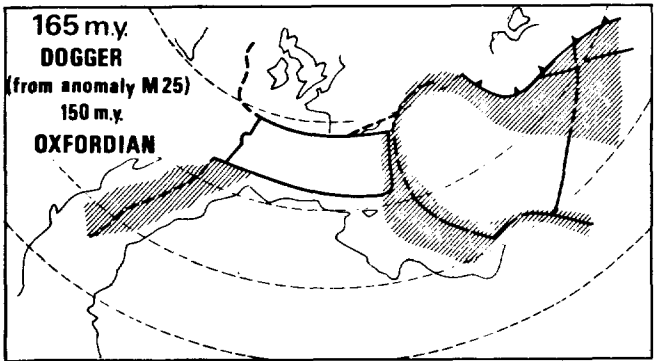
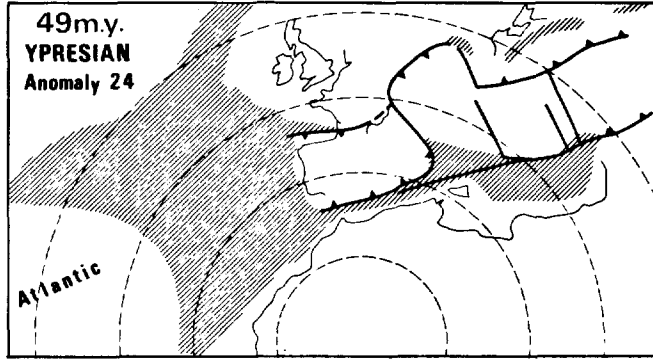
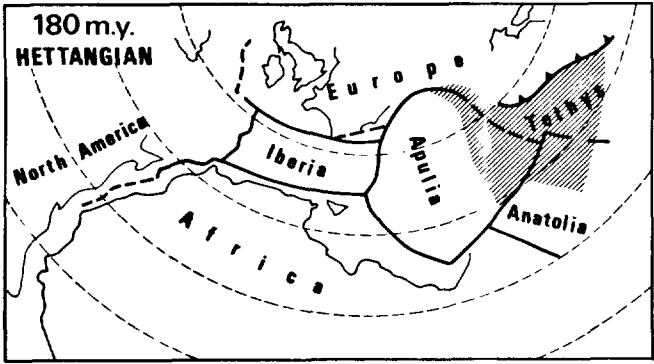


Figure 2 : Plate Kinematics of the Africa-Europe plate system. Parameters are given in table I.

← Figure 1 - Comparison of different kinematic models for the Africa displacements with respect to Europe in its present position. Oblique Mercator projection with pole at 52° N and 150° E as in all the maps of this paper. Numbers refer to references of table I : 1) - ref. 5 ; 2) - ref. 3 for Africa-America and ref. 5 for Europe-America ; 3) - ref. 3 for Africa-America and ref. 5 for Europe-America ; 4) - ref. 5 for both ; 5) - ref. 5 for Africa-America and ref. 4 for Europe-America ; 6) - Bullard et al. 1965 ; 7) - Dewey et al., 1973 ; 8) - ref. 1 for both ; 9) - ref. 1 for Africa-America and ref. 4 for Europe-America.

FROM THE TETHYS OCEAN TO THE MEDITERRANEAN SEAS : A PLATE TECTONIC MODEL OF THE EVOLUTION OF THE WESTERN ALPINE SYSTEM.



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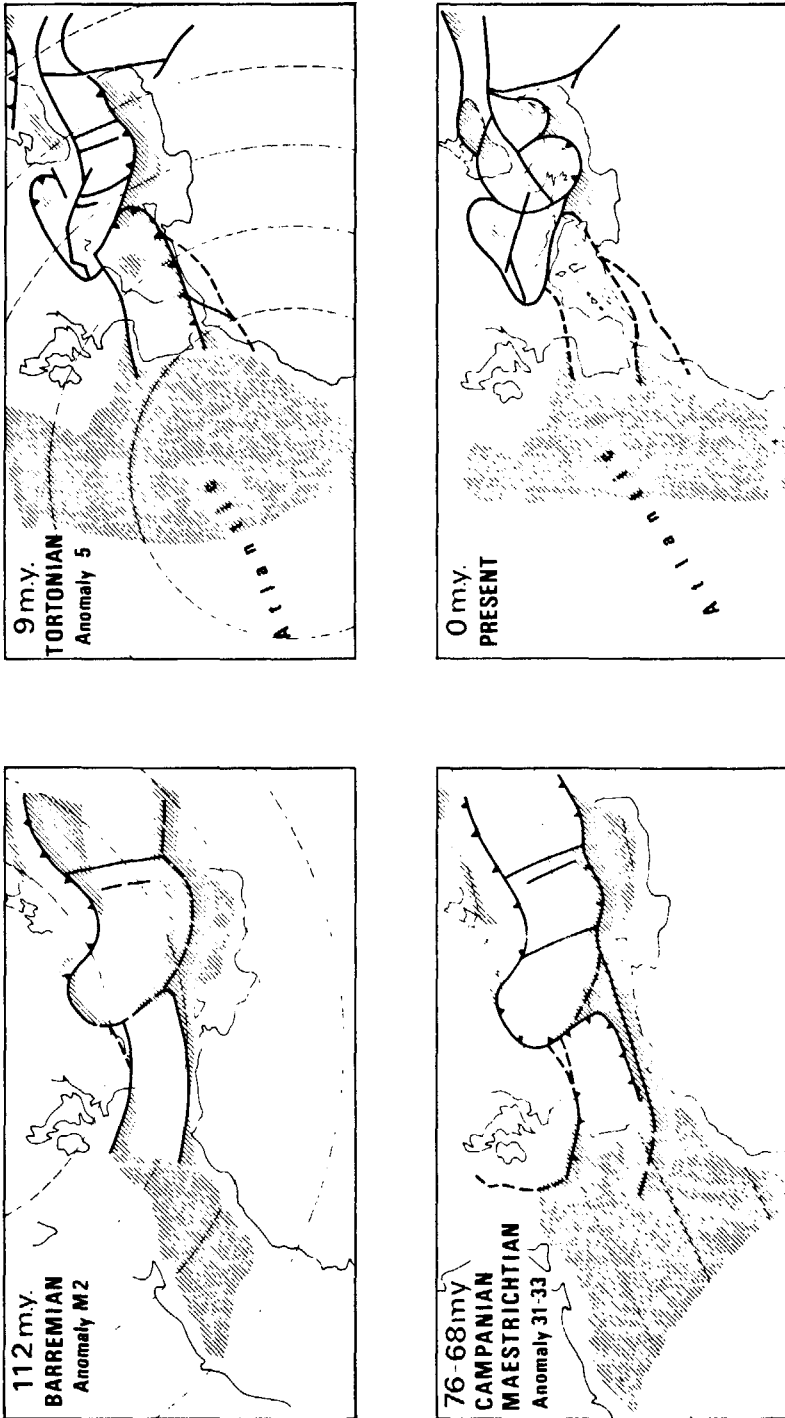


Figure 3 : Proposed evolution from the Tethys to the Mediterranean basins. Parameters are given in table I for positions of Africa. For positions of intermediate plates, see text. Small dotted circles are about pole of rotation of Africa with respect to Europe during the immediately following period.

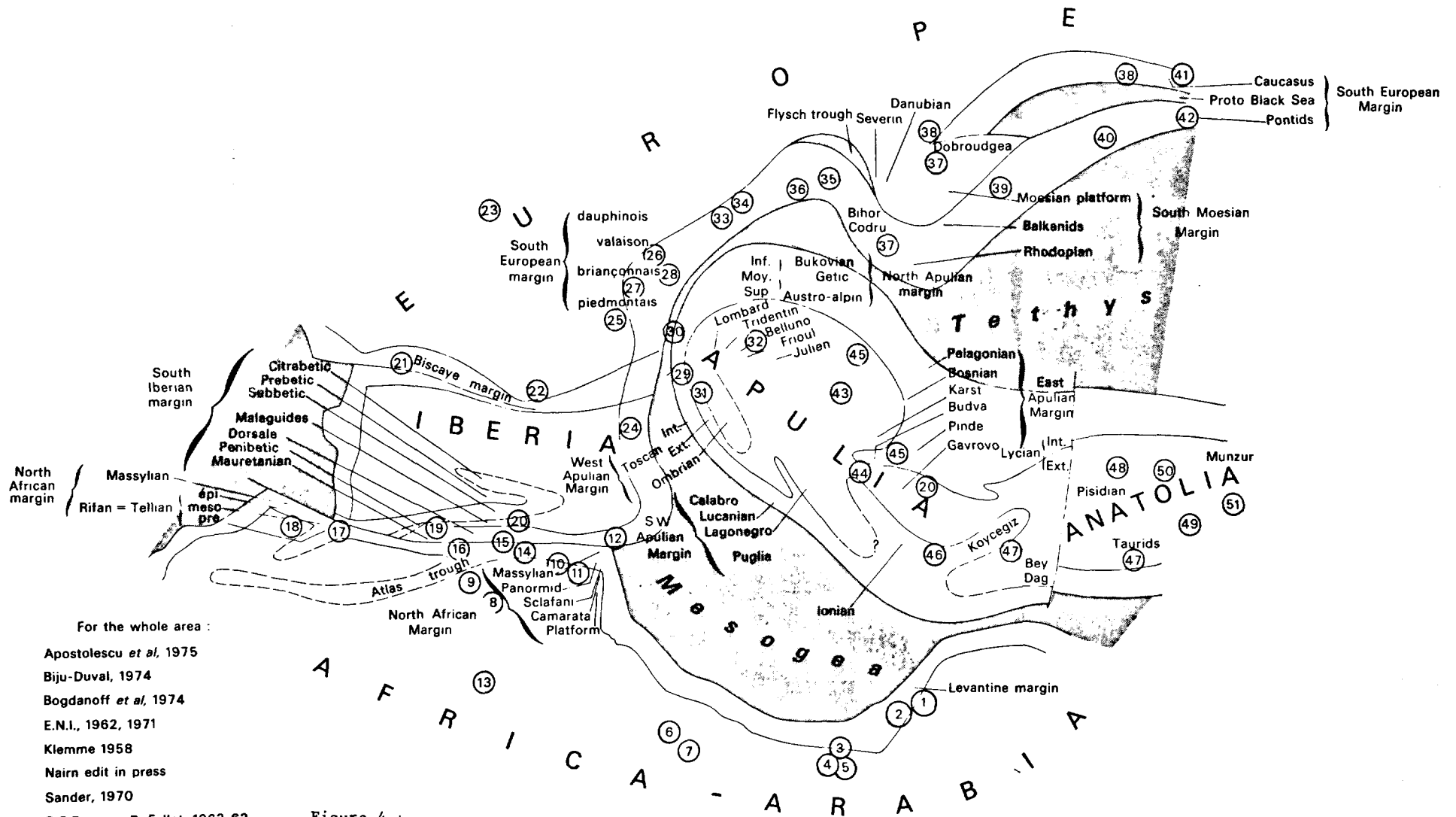


Figure 4

Main paleogeographic zones and general geological bibliography used in this paper illustrated on the Dogger reconstruction (see plate III). Numbers refer to following references (see detailed bibliography) : 1 - Ginzburg et al., 1975 ; 2 - Bein, 1976 ; 3 - Salem, 1976 ; 4 - Saïd, 1962 ; 5 - Youssef, 1968 ; 6 - Carlyle Gray edit, 1971 ; 7 - Conant and Goudarzi, 1967 ; 8 - Livre Jub. M. Solignac, 1973 ; 9 - PESL, Martin edit., 1967 ; 10 - Caire, 1971 ; 11 - PESL, Alvarez and Gahrbandt edit., 1970 ; 12 - Bousquet, 1973 ; 13 - Busson, 1972 ; 14 - Raoult, 1975 ; 15 - Bouillin et al., 1970 ; 16 - Durand Delga, 1969 ; 17 - S.G.F., 1973 ; 18 - Andrieux, 1971 ; 19 - Paquet, 1974 ; 20 - S.G.N., Mem. centenaire, 1970 ; 21 - Hist. Golfe Gascogne, Technip edit., 1971 ; 22 - BRGM, 1974 ; 23 - Ziegler, 1975 ; 24 - S.G.F., 1976 b ; 25 - Debelmas et Lemoine, 1970 ; 26 - Trumpy, 1975 ; 27 - Caby, 1973 ; 28 - Trumpy et Haccard, 1969 ; 29 - Haccard et al., 1972 ; 30 - S.G.F., 1975 ; 31 - Abbate et al., 1970 ; 32 - Aubouin, 1963 ; 33 - Tollmann, 1969 ; 34 - Geyssant et Tollmann, 1966 ; 35 - Sandulescu, 1975 ; 36 - Boccaletti et al., 1976 ; 37 - Mahel edit., 1975 ; 38 - Dzotsenide, 1968 ; 39 - Fourquin, 1975 ; 40 - Bergougnan, 1975 ; 41 - Khain, 1963 ; 42 - Adamia, 1975 ; 43 - Aubouin, 1960 ; 44 - Aubouin et Ndojaj, 1964 ; 45 - S.G.F., 1970 ; 46 - S.G.F., 1976 a ; 47 - PESL, Campbell edit., 1971 ; 48 - Brinkmann, 1972 ; 49 - Ozgül, 1976 ; 50 - Okay and Dileköz edit., 1974 ; 51 - Kestin and Demisman, 1971.