
A Geophysical Model Function for Wind Speed Retrieval from C-Band HH-polarized Synthetic Aperture Radar

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Abstract :

Synthetic aperture radar (SAR) imagery is routinely acquired at HH-polarization in high latitude areas for measuring surface wind over the ocean. However, in the contrary of VV-polarization, there is no HH-polarization geophysical model function (GMF) exists to directly retrieve wind speed from SAR images. In general, HH-polarized Normalized Radar Cross Section (NRCS) is thus converted into VV-polarization and then conventional CMOD functions are used with auxiliary wind direction information for wind speed retrieval. In this letter, we propose a new GMF for SAR ocean surface wind speed retrieval, called CMODH, which relates the C-band NRCS acquired at HH-polarization over the ocean, to the 10-m height wind speed, incident angle and relative wind direction. We first use more than 220,000 ENVISAT ASAR radar backscatter measurements collocated with ASCAT winds to derive the CMODH coefficients. Subsequently, 1459 RADARSAT-2 (RS-2) and 428 Sentinel-1A/B (S1-1A/B) HH-polarized SAR acquisitions under different wind speeds are matched to in situ buoy observations to validate CMODH. The statistical comparisons between SAR-observed and simulated NRCS show a bias of -0.07 dB and a root mean square error of 1.62 dB for RS-2, and -0.01 dB and 2.48 dB for S1-1A/B. These results suggest that the proposed CMODH has the potential to directly retrieve ocean surface wind speeds using C-band SAR images acquired at HH-polarization, with no need for NRCS transformation by using various empirical and theoretical polarization ratio models.

Keywords : HH-polarization, ocean surface wind speed, synthetic aperture radar (SAR)

SYNTHETIC aperture radar (SAR) is a very useful microwave sensor for remote sensing of ocean surface features with ability for high-resolution and almost all-weather observations including night and day. Spaceborne SARs are on onboard polar-orbiting satellites and thus have the capability to observe high-latitude areas with shorter revisit time than middle-and low-latitude regions. In particular, they are routinely used to monitor seasonally or annual variations in sea-ice in the Arctic, Antarctic, and other high-latitude areas such as Chukchi Sea, the Labrador Sea, or the Beaufort Sea. When only one co-polarized channel is available, HH-polarization is preferred for sea-ice application because the contrast with ocean is more pronounced than at VV-polarization [1]. As a consequence, the acquisition strategy at high latitudes gives priority to HH-polarization (and cross-polarization when available). In the contrary, VV is widely at mid-latitude for ocean applications, including wind field measurements.

Moreover, the high-resolution ocean surface wind is not only a key complementary geophysical parameter to investigate sea-ice motion and melt but also the coupling between ocean and the marginal ice zone (MIZ) in high-latitude regions. Long time-series satellite-tracked sea-ice motion observations reveal that changes in wind speed are linked to the observed increase in ice drift speeds in the Center Arctic, on climate timescales [2]. Antarctic sea-ice motions observed at large scale with satellites also suggest that significant trends in Antarctic ice drift can be statistically related to local winds in most geographic areas [3]. For example, in coastal East Antarctica, persistent katabatic winds from the ice sheet's interior cause widespread snow erosion and enhance the melt of blue ice and firn snow [4].

SAR ocean surface wind speed retrieval is usually based on geophysical model functions (GMFs) [5]–[7], which are derived from VV-polarized scatterometer observations. Conventional scatterometer-derived GMFs have been demonstrated to be applicable for moderate wind speed retrieval with an accuracy of 1.5 m/s from ERS, ENVISAT ASAR, Sentinel-1A, or RADARSAT-2 VV-polarized SAR images [8]–[12]. Mouche and Chapron [13] pioneered the idea of relying directly and massively on SAR data to describe the relationship between NRCS and ocean surface wind in both co-polarizations and possibly derive GMF from SAR data.

72 Recently, a new GMF has been proposed for C-band SAR
 73 coastal wind speed mapping [14], by using SAR measurements
 74 acquired in VV-polarization collocated with buoy observa-
 75 tions. However, no similar well-documented wind retrieval
 76 model exists to derive wind speed from the HH-polarized
 77 SAR imagery. Therefore, various empirical and theoretical
 78 polarization ratio (PR) models have been developed to trans-
 79 form normalized radar cross section (NRCS) from HH into
 80 VV-polarization [15]–[20], along with CMOD GMFs to
 81 retrieve wind speed. A systematic comparison of PR mod-
 82 els for ocean surface wind retrieval using ENVISAT ASAR
 83 images shows a small wind speed underestimation or overesti-
 84 mation at very small or large incident angles, respectively [21].

85 Although the CMOD GMF-PR approach constitutes a
 86 practical hybrid model for wind speed retrieval from SAR
 87 images acquired at HH-polarization, the NRCS transfor-
 88 mation certainly induces error in wind speed estimation.
 89 In order to directly retrieve wind speeds using HH-polarized
 90 SAR imagery, we develop a new GMF relating NRCS at
 91 HH-polarization to radar incident angle, wind speed at 10 m
 92 height, and relative wind direction. Finally, to complement
 93 the VV-polarization, the next European scatterometer mis-
 94 sion onboard MetOp-SG will also operate at HH-polarization
 95 (and VH-polarization) [22]. The proposed GMF should, thus,
 96 directly benefit to this forthcoming satellite mission. This letter
 97 is organized as follows. Data sets are described in Section II.
 98 GMF development and validation are presented in Section III.
 99 Summary and conclusion are given in Section IV.

100 II. DATA SET

101 In this letter, more than 2700 ENVISAT/ASAR images
 102 acquired in wide swath mode (WSM) in 2009 are collocated
 103 with MetOp/ASCAT and ECMWF winds to develop the
 104 HH-polarized GMF. The incident angles of WSM are between
 105 16° and 42° in the near and far ranges, respectively. SAR
 106 acquisitions in WSM cover about 400 km in the range direc-
 107 tion and can be more than 1000 km in the azimuth direction.
 108 We processed all ASAR WSM products to derive the NRCS,
 109 incident angles, and azimuth angles at a spatial resolution
 110 cell of 12.5 km. These measurements are then matched to
 111 ASCAT wind speeds and directions. This results in more than
 112 220 000 collocations. The ENVISAT ASAR and ASCAT orbits
 113 and field of view enable temporal intervals of less than 1 h,
 114 for collocation. This collocated data set has already been used
 115 to analyze wind sensitivity for different incident angles and
 116 azimuth angles [13].

117 In order to validate the proposed GMF, we collect 1352
 118 RS-2 SAR images acquired in quad-polarization (HH + HV +
 119 VH + VV) and 107 dual-polarization (HH + HV) imag-
 120 ing modes between October 2008 and April 2013, and
 121 428 Sentinel-1A/B (S1-1A/B) images acquired in extra wide
 122 swath (EW) mode between October 2014 and April 2018.
 123 A summary of the parameters for the RS-2 quad-and dual-
 124 polarization modes, including incident angles, spatial resolu-
 125 tions, swaths, and noise-equivalent sigma-zero (NESZ) values,
 126 is given in Table I. The EW mode of S1-1A/B data has the
 127 capability to observe the ocean surface with dual-polarization
 128 (HH + HV) channels. Table II contains the major parameters

TABLE I
MAJOR PARAMETERS OF THE RS-2 QUAD-AND
DUAL-POLARIZATION IMAGING MODES

Parameter	quad-polarization mode	dual-polarization ScanSAR Narrow mode
Polarization	HH+HV+VH+VV	VV+VH
Incident angles	20° – 49°	20° – 46°
Azimuth resolution	8 m	60 m
Range resolution	5.4 m	79.9–37.7 m
Swath	25 km	300 km
NESZ	-36.5 ± 3 dB	28.5 ± 2.5 dB

TABLE II
MAJOR PARAMETERS OF THE S1-1A/B IW AND EW IMAGING MODES

Parameter	Extra wide swath (EW) arrow mode
Polarization	HH+HV
Incident angles	19° – 47°
Azimuth resolution	40 m
Range resolution	20 m
Swath	400 km
NESZ	-22 dB

129 for EW mode. All RS-2 and S1-1A/B SAR images are col-
 130 located with 102 *in situ* National Data Buoy Center (NDBC)
 131 buoys in the Gulf of Alaska, off the East and West coasts
 132 of the USA, and the European Seas. The time interval for
 133 the collocation is smaller than 30 min. We process the NRCS
 134 at HH-polarization at a spatial resolution of 10 km to match
 135 the resolution used by Mouche and Chapron [13], where
 136 medium (instead of high) resolution was used to be in line
 137 with collocated wind field from ASCAT. This approach results
 138 in 1527 and 461 collocated data pairs for RS-2 and S1-1A/B,
 139 respectively, which are used to compare simulated and
 140 SAR-measured NRCS.

141 III. MODEL DEVELOPMENT AND VALIDATION

142 Similar to NRCS at VV-polarization, HH-polarized NRCS
 143 also depend on radar incident angle, wind speed, and direction
 144 (see [13, Fig. 3(b) and (d)]). Therefore, we describe NRCS
 145 at HH-polarization by a nonlinear mapping function of the
 146 incident angle, wind speed, and relative wind direction (the
 147 angle between the true wind direction and the radar observa-
 148 tion direction), which is given as

$$149 \sigma_0^{\text{hh}}(v, \phi, \theta) = (B_0(v, \theta)[1 + B_1(v, \theta) \cos(\phi) \\ 150 + B_2(v, \theta) \cos(2\phi)])^p \quad (1)$$

151 where the isotropic B_0 , the upwind/downwind B_1 , and the
 152 upwind/crosswind B_2 amplitude terms are the functions of
 153 wind speed v and incident angle θ . The superscript p is a
 154 constant with a value of 1.6. The superscript p and the transfer
 155 functions used to define B_0 , B_1 , and B_2 are adopted from
 156 CMOD5 [6] for use in this study. The proposed form of the
 157 GMF is almost the same as that of CMOD5 excepted for the
 158 p exponent, on B_0 .

159 For each incident angle and wind speed bin, using the collo-
 160 cated ENVISAT ASAR and ASCAT data set, we first analyze
 161 the NRCS at HH-polarization as a function of azimuth angle

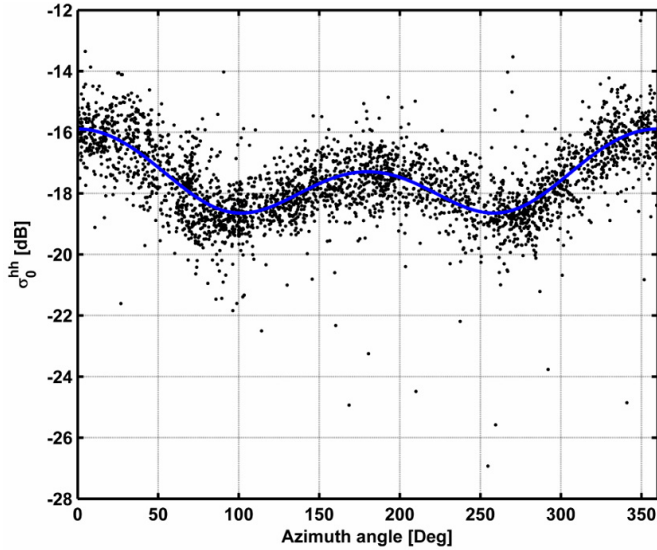


Fig. 1. HH-polarized NRCS as a function of azimuth angle, for a 7 m/s wind speed and 35° incidence angle. Black dots are SAR-observed NRCS at HH-polarization. Blue line represents regression results.

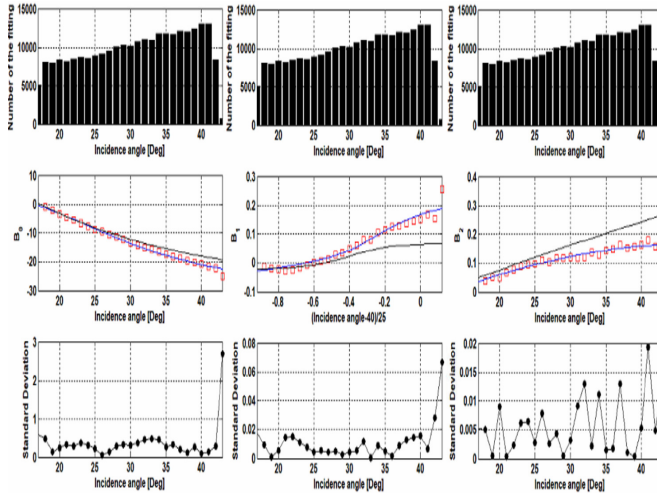


Fig. 2. Number of the data for (Top) B_i coefficients derivation, (Middle) B_i coefficients, and (Bottom) standard deviation of the regression as a function of incident angle for 7 m/s wind speed. B_i coefficients derived from (1), the regression analysis approach, and the CMOD5 are represent as red squares, and blue and black solid lines, respectively.

to determine the corresponding B_i ($i = 0, 1, 2$) coefficients by using a regression analysis approach. Fig. 1 presents the ASAR-observed NRCS (black dots) versus azimuth angles for 7 m/s wind speed and 35° incident angle. Fig. 1 clearly shows the azimuth modulation of NRCS at HH-polarization. Estimated NRCS values using (1) are indicated in the blue solid line in Fig. 1, which are in good agreement with the radar observations, suggesting that (1) is appropriate to describe the proposed GMF.

Subsequently, to further assess the rationality of the functional form, B_i coefficients are also analyzed as a function of wind speed and incident angle. The top three panels in Fig. 2 show the number of data used for B_i coefficients derivation. For 7 m/s wind speed and different incident angles, B_i coefficients derived from (1), the regression analysis

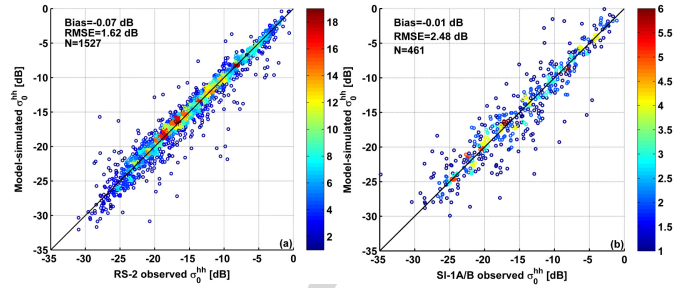


Fig. 3. Comparisons of observed NRCS in HH-polarization (x-axis) from (a) RS-2 and (b) S1-A/B with CMODH model simulations (y-axis). The color bar indicates the scatter density.

approach, and the CMOD5 function are represented as red squares, and blue and black solid lines, in the middle three panels of Fig. 2.

Overall, the trend of B_i coefficients obtained by our analysis and those provided by CMOD5 GMF are in very good agreement excepted at the largest incident angle of 43°. The B_0 values from the regression analysis and from CMOD5 estimation are almost identical at low incident angles (<23°). As incident angles increase, the discrepancies between B_i coefficients obtained for HH from our regression analysis and from CMOD5 gradually increase. This is in line with a known different sensitivity to ocean sea-surface backscattering between VV- and HH-polarizations. Indeed, at C-band, polarization effect is strongly incident angle-dependent. In particular, our analysis confirms that B_1 values at VV are smaller than at HH, while B_0 and B_2 , are larger at VV.

The standard deviations of the regression analysis for each incident angle bin are illustrated in the bottom three panels in Fig. 2. The standard deviation for the large incident angle bin of 43° is much larger than those for other incident angle bins, which is possibly caused by the very few data numbers in this bin. Consequently, we do not use NRCS for this large incident angle bin to develop our HH-polarized GMF. In this study, B_i coefficients are determined via a nonlinear weighted regression approach by minimizing the standard deviation between ENVISAT ASAR observations and regression analysis for all wind speeds and incident angle bins. The number of the data in each wind speed and incident angle bin is considered as a weighted factor for the regression analysis. The final model formulation and 28 coefficients for HH-polarization are given in the Appendix. Since our collocated data sets both involve NRCS at HH- and VV-polarization, we also use (1) to derive the coefficients of GMF for VV-polarization, which are shown in the Appendix. Therefore, one can estimate NRCS at HH-and VV-polarization using (1) and coefficients with given incident angle, wind speed, and direction.

To assess the accuracy of the proposed GMF, we make statistical validations using collocated RS-2 and S1-1A/B observations and buoy measurements. Fig. 3(a) shows the simulated and SAR-measured NRCS at HH-polarization, with a bias of -0.07 dB and a RMSE of 1.62 dB. For comparison between model simulations and S1-1A/B observations, the bias and RMSE are -0.01 and 2.48 dB, which are shown in Fig. 3(b). For NRCS comparisons, the RS-2 data have smaller

222 RMSE than the S1-1A/B data. The reason for this result may
 223 be associated with two factors: 1) RS-2 has lower NESZ than
 224 the S1-1A/B and 2) the collocations between S1-1A/B and
 225 buoy data are less accurate than that of RS-2. For S1-1A/B
 226 collocation, some buoys are outside the SAR image which
 227 leads to the large spatial interval of the collocation. However,
 228 for RS-2 collocation, each buoy locates in the SAR image and
 229 thus there is almost no spatial match error.

230 IV. CONCLUSION

231 Accurate wind field information in high latitudes and polar
 232 regions plays an important role in sea-ice motion and melt
 233 simulations and predictions. The high sensitivity of SAR
 234 backscatter at HH-polarization to sea-ice makes it more
 235 preferable for sea-ice monitoring, rather than VV-polarization.
 236 However, unlike CMOD GMFs, no HH-polarized GMF can be
 237 directly used to retrieve ocean surface wind speed. As a result,
 238 various empirical and theoretical PR models are used to trans-
 239 form NRCS from HH into VV-polarization, which induces
 240 inevitable errors in the retrieved wind speeds. In order to
 241 solve this problem, we develop a new HH-polarized GMF for
 242 C-band SAR, taking advantage of a large number of C-band
 243 SAR data and collocated scatterometer wind measurements.
 244 The proposed GMF has potential to improve SAR wind speed
 245 retrieval accuracy and can be used to directly retrieve wind
 246 speed in mixing areas of ice and water, like the MIZ.

247 The new GMF relates NRCS at HH-polarization to radar
 248 incident angle, wind speed at 10 m height, and relative wind
 249 direction. We use a nonlinear weighted regression method
 250 and the collocated data set (ENVISAT ASAR backscatter
 251 and ASCAT winds) to derive all the coefficients in the
 252 GMF. Furthermore, we use an independent matching of
 253 data (RS-2 and S1-1A/B backscatter and buoy measurements)
 254 to validate the proposed GMF. The statistical comparisons
 255 between radar observations and GMF simulations show a bias
 256 of -0.07 dB and a root-mean-square error of 1.62 dB for RS-2,
 257 and -0.01 dB and 2.48 dB for S1-1A/B, respectively. Based on
 258 this new GMF, we do not need to make a NRCS transformation
 259 with a PR model before wind speed retrieval. The scatter-
 260 omer (SCA) onboard EUMETSAT Polar System-Second
 261 Generation (EPS-SG) mission shall provide observations with
 262 VV-, HH-, or VH-polarization. The addition of HH- or
 263 VH-polarization is considered as an option for extending
 264 the upper limit of the dynamic measurement range. There-
 265 fore, the proposed GMF can be a benefit for EPS-SG SCA
 266 wind retrieval. Moreover, this also opens new perspectives
 267 for the forthcoming launch of the RADARSAT Constellation
 268 Mission (RCM) with new polarization configurations.

269 APPENDIX

270 CMODH MODEL FORMULATION AND COEFFICIENTS

271 The form of the CMODH model is

$$272 \sigma_0^{\text{hh}}(v, \phi, \theta) = (B_0(v, \theta)[1 + B_1(v, \theta) \cos(\phi) \\ 273 + B_2(v, \theta) \cos(2\phi)])^p \quad (\text{A1})$$

274 where p is constant with a value of 1.6 , B_0 , B_1 , and
 275 B_2 are functions of wind speeds v and the incidence

TABLE III
CMODH COEFFICIENTS

Function	Coefficients	HH	VV
B_0	c_1	-0.72722756511	-0.13393789593
	c_2	-1.1901195406	-0.74081314533
	c_3	0.33968637656	0.34811480603
	c_4	0.086759069544	0.019382338942
	c_5	0.003090124916	-0.008066293463
	c_6	0.011761378188	0.006426074015
	c_7	0.129158495658	0.096343783534
	c_8	0.083506931034	0.042280179737
	c_9	4.092557781322	5.007750349297
	c_{10}	1.211169044551	0.717396068916
	c_{11}	-1.119776245438	-1.501296438845
	c_{12}	0.579066509504	0.442826511887
B_1	c_{13}	-0.604527699539	-0.154971505863
	c_{14}	0.118371042255	0.036542289696
	c_{15}	0.008955505675	0.006784919880
	c_{16}	0.219608674529	0.401880787461
	c_{17}	0.017557536680	0.006896838546
	c_{18}	24.442309754388	24.751953435615
B_2	c_{19}	1.983490330585	1.961341923034
	c_{20}	6.781440647278	3.284009890111
	c_{21}	7.947947040974	8.379337236413
	c_{22}	-4.696499003167	-3.636259490187
	c_{23}	-0.437054238710	2.349430558787
	c_{24}	5.471252046908	5.851939658893
	c_{25}	0.639468224273	2.443227221148
	c_{26}	0.673385731705	0.301462797210
	c_{27}	3.433229044819	3.976051353364
	c_{28}	0.367036215316	1.728745711306

angle θ , or alternatively, $x = (\theta - 40)/25$. The B_0 term is
 defined as

$$278 B_0 = 10^{a_0 + a_1 v} f(a_2 v, s_0)^p \quad (\text{A2})$$

279 where

$$280 \quad f(s, s_0) = \begin{cases} (s/s_0)^\alpha g(s_0), & s < s_0 \\ g(s), & s \geq s_0 \end{cases} \quad (\text{A3})$$

281 where

$$282 \quad g(s) = 1/(1 + \exp(-s)) \quad \text{and} \quad \alpha = s_0(1 - g(s_0)). \quad (\text{A4})$$

283 The functions a_0 , a_1 , a_2 , γ , and s_0 depend on incident angle
284 only

$$285 \quad \begin{aligned} a_0 &= c_1 + c_2x + c_3x^2 + c_4x^3 \\ a_1 &= c_5 + c_6x \\ a_2 &= c_7 + c_8x \end{aligned} \quad (\text{A5})$$

$$286 \quad \begin{aligned} \gamma &= c_9 + c_{10}x + c_{11}x^2 \\ s_0 &= c_{12} + c_{13}x. \end{aligned} \quad (\text{A6})$$

290 The B_1 term is defined as follows:

$$291 \quad B_1 = \frac{c_{14}(1+x) - c_{15}v(0.5+x - \tanh[4(x+c_{16}+c_{17}v)])}{1+\exp(0.34(v-c_{18}))}. \quad (\text{A7})$$

293 The B_2 term is chosen as

$$294 \quad B_2 = (-d_1 + d_2v_2)\exp(-v_2). \quad (\text{A8})$$

295 Here, v_2 is given by

$$296 \quad v_2 = \begin{cases} a + b(y-1)^n, & y < y_0 \\ y, & y \geq y_0 \end{cases}, \quad y = \frac{v+v_0}{v_0} \quad (\text{A9})$$

297 where

$$298 \quad y_0 = c_{19}, n = c_{20} \quad (\text{A10})$$

$$299 \quad a = y_0 - (y_0 - 1)/n, b = 1/[n(y_0 - 1)^{n-1}]. \quad (\text{A11})$$

300 The quantities v_0 , d_1 , and d_2 are functions of incident angle
301 only

$$302 \quad \begin{aligned} v_0 &= c_{21} + c_{22}x + c_{23}x^2 \\ d_1 &= c_{24} + c_{25}x + c_{26}x^2 \\ d_2 &= c_{27} + c_{28}x. \end{aligned} \quad (\text{A12})$$

305 The coefficients are given in Table III.

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A Geophysical Model Function for Wind Speed Retrieval From C-Band HH-Polarized Synthetic Aperture Radar

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Abstract—Synthetic aperture radar (SAR) imagery is routinely acquired at HH-polarization in high-latitude areas for measuring surface wind over the ocean. However, in the contrary of VV-polarization, there is no HH-polarization geophysical model function (GMF) exists to directly retrieve wind speed from SAR images. In general, HH-polarized normalized radar cross section (NRCS) is thus converted into VV-polarization and then conventional CMOD functions are used with auxiliary wind direction information for wind speed retrieval. In this letter, we propose a new GMF for SAR ocean surface wind speed retrieval, called CMODH, which relates the C-band NRCS acquired at HH-polarization over the ocean, to the 10-m height wind speed, incident angle, and relative wind direction. We first use more than 220 000 ENVISAT ASAR radar backscatter measurements collocated with ASCAT winds to derive the CMODH coefficients. Subsequently, 1459 RADARSAT-2 (RS-2) and 428 Sentinel-1A/B (S1-1A/B) HH-polarized SAR acquisitions under different wind speeds are matched to *in situ* buoy observations to validate CMODH. The statistical comparisons between SAR-observed and simulated NRCS show a bias of -0.07 dB and a root-mean-square error of 1.62 dB for RS-2, and -0.01 dB and 2.48 dB for S1-1A/B. These results suggest that the proposed CMODH has the potential to directly retrieve ocean surface wind speeds using C-band SAR images acquired at HH-polarization, with no need for NRCS transformation by using various empirical and theoretical polarization ratio models.

Index Terms—HH-polarization, ocean surface wind speed, synthetic aperture radar (SAR).

Manuscript received October 21, 2018; revised February 3, 2019; accepted March 13, 2019. This work was supported in part by the National Key Research and Development Program of China under Grant 2016YFC1401001, in part by the National Science Foundation of China for Outstanding Young Scientist under Grant 41622604, in part by the Excellent Youth Science Foundation of Jiangsu Province under Grant BK2016090, in part by the Data Utilization Application Plan of Canadian Space Agency, in part by the Canadian Office of Energy Research and Development, and in part by the ESA Dragon Project. (Corresponding author: Biao Zhang.)

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Digital Object Identifier 10.1109/LGRS.2019.2905578

I. INTRODUCTION

SYNTHETIC aperture radar (SAR) is a very useful microwave sensor for remote sensing of ocean surface features with ability for high-resolution and almost all-weather observations including night and day. Spaceborne SARs are on onboard polar-orbiting satellites and thus have the capability to observe high-latitude areas with shorter revisit time than middle-and low-latitude regions. In particular, they are routinely used to monitor seasonally or annual variations in sea-ice in the Arctic, Antarctic, and other high-latitude areas such as Chukchi Sea, the Labrador Sea, or the Beaufort Sea. When only one co-polarized channel is available, HH-polarization is preferred for sea-ice application because the contrast with ocean is more pronounced than at VV-polarization [1]. As a consequence, the acquisition strategy at high latitudes gives priority to HH-polarization (and cross-polarization when available). In the contrary, VV is widely at mid-latitude for ocean applications, including wind field measurements.

Moreover, the high-resolution ocean surface wind is not only a key complementary geophysical parameter to investigate sea-ice motion and melt but also the coupling between ocean and the marginal ice zone (MIZ) in high-latitude regions. Long time-series satellite-tracked sea-ice motion observations reveal that changes in wind speed are linked to the observed increase in ice drift speeds in the Center Arctic, on climate timescales [2]. Antarctic sea-ice motions observed at large scale with satellites also suggest that significant trends in Antarctic ice drift can be statistically related to local winds in most geographic areas [3]. For example, in coastal East Antarctica, persistent katabatic winds from the ice sheet’s interior cause widespread snow erosion and enhance the melt of blue ice and firn snow [4].

SAR ocean surface wind speed retrieval is usually based on geophysical model functions (GMFs) [5]–[7], which are derived from VV-polarized scatterometer observations. Conventional scatterometer-derived GMFs have been demonstrated to be applicable for moderate wind speed retrieval with an accuracy of 1.5 m/s from ERS, ENVISAT ASAR, Sentinel-1A, or RADARSAT-2 VV-polarized SAR images [8]–[12]. Mouche and Chapron [13] pioneered the idea of relying directly and massively on SAR data to describe the relationship between NRCS and ocean surface wind in both co-polarizations and possibly derive GMF from SAR data.

Recently, a new GMF has been proposed for C-band SAR coastal wind speed mapping [14], by using SAR measurements acquired in VV-polarization collocated with buoy observations. However, no similar well-documented wind retrieval model exists to derive wind speed from the HH-polarized SAR imagery. Therefore, various empirical and theoretical polarization ratio (PR) models have been developed to transform normalized radar cross section (NRCS) from HH into VV-polarization [15]–[20], along with CMOD GMFs to retrieve wind speed. A systematic comparison of PR models for ocean surface wind retrieval using ENVISAT ASAR images shows a small wind speed underestimation or overestimation at very small or large incident angles, respectively [21].

Although the CMOD GMF-PR approach constitutes a practical hybrid model for wind speed retrieval from SAR images acquired at HH-polarization, the NRCS transformation certainly induces error in wind speed estimation. In order to directly retrieve wind speeds using HH-polarized SAR imagery, we develop a new GMF relating NRCS at HH-polarization to radar incident angle, wind speed at 10 m height, and relative wind direction. Finally, to complement the VV-polarization, the next European scatterometer mission onboard MetOp-SG will also operate at HH-polarization (and VH-polarization) [22]. The proposed GMF should, thus, directly benefit to this forthcoming satellite mission. This letter is organized as follows. Data sets are described in Section II. GMF development and validation are presented in Section III. Summary and conclusion are given in Section IV.

II. DATA SET

In this letter, more than 2700 ENVISAT/ASAR images acquired in wide swath mode (WSM) in 2009 are collocated with MetOp/ASCAT and ECMWF winds to develop the HH-polarized GMF. The incident angles of WSM are between 16° and 42° in the near and far ranges, respectively. SAR acquisitions in WSM cover about 400 km in the range direction and can be more than 1000 km in the azimuth direction. We processed all ASAR WSM products to derive the NRCS, incident angles, and azimuth angles at a spatial resolution cell of 12.5 km. These measurements are then matched to ASCAT wind speeds and directions. This results in more than 220 000 collocations. The ENVISAT ASAR and ASCAT orbits and field of view enable temporal intervals of less than 1 h, for collocation. This collocated data set has already been used to analyze wind sensitivity for different incident angles and azimuth angles [13].

In order to validate the proposed GMF, we collect 1352 RS-2 SAR images acquired in quad-polarization (HH + HV + VH + VV) and 107 dual-polarization (HH + HV) imaging modes between October 2008 and April 2013, and 428 Sentinel-1A/B (S1-1A/B) images acquired in extra wide swath (EW) mode between October 2014 and April 2018. A summary of the parameters for the RS-2 quad-and dual-polarization modes, including incident angles, spatial resolutions, swaths, and noise-equivalent sigma-zero (NESZ) values, is given in Table I. The EW mode of S1-1A/B data has the capability to observe the ocean surface with dual-polarization (HH + HV) channels. Table II contains the major parameters

TABLE I
MAJOR PARAMETERS OF THE RS-2 QUAD-AND
DUAL-POLARIZATION IMAGING MODES

Parameter	quad-polarization mode	dual-polarization ScanSAR Narrow mode
Polarization	HH+HV+VH+VV	VV+VH
Incident angles	20° – 49°	20° – 46°
Azimuth resolution	8 m	60 m
Range resolution	5.4 m	79.9–37.7 m
Swath	25 km	300 km
NESZ	-36.5 ± 3 dB	28.5 ± 2.5 dB

TABLE II
MAJOR PARAMETERS OF THE S1-1A/B IW AND EW IMAGING MODES

Parameter	Extra wide swath (EW) arrow mode
Polarization	HH+HV
Incident angles	19° – 47°
Azimuth resolution	40 m
Range resolution	20 m
Swath	400 km
NESZ	-22 dB

for EW mode. All RS-2 and S1-1A/B SAR images are collocated with 102 *in situ* National Data Buoy Center (NDBC) buoys in the Gulf of Alaska, off the East and West coasts of the USA, and the European Seas. The time interval for the collocation is smaller than 30 min. We process the NRCS at HH-polarization at a spatial resolution of 10 km to match the resolution used by Mouche and Chapron [13], where medium (instead of high) resolution was used to be in line with collocated wind field from ASCAT. This approach results in 1527 and 461 collocated data pairs for RS-2 and S1-1A/B, respectively, which are used to compare simulated and SAR-measured NRCS.

III. MODEL DEVELOPMENT AND VALIDATION

Similar to NRCS at VV-polarization, HH-polarized NRCS also depend on radar incident angle, wind speed, and direction (see [13, Fig. 3(b) and (d)]). Therefore, we describe NRCS at HH-polarization by a nonlinear mapping function of the incident angle, wind speed, and relative wind direction (the angle between the true wind direction and the radar observation direction), which is given as

$$\sigma_0^{\text{hh}}(v, \phi, \theta) = (B_0(v, \theta)[1 + B_1(v, \theta) \cos(\phi) + B_2(v, \theta) \cos(2\phi)])^p \quad (1)$$

where the isotropic B_0 , the upwind/downwind B_1 , and the upwind/crosswind B_2 amplitude terms are the functions of wind speed v and incident angle θ . The superscript p is a constant with a value of 1.6. The superscript p and the transfer functions used to define B_0 , B_1 , and B_2 are adopted from CMOD5 [6] for use in this study. The proposed form of the GMF is almost the same as that of CMOD5 excepted for the p exponent, on B_0 .

For each incident angle and wind speed bin, using the collocated ENVISAT ASAR and ASCAT data set, we first analyze the NRCS at HH-polarization as a function of azimuth angle

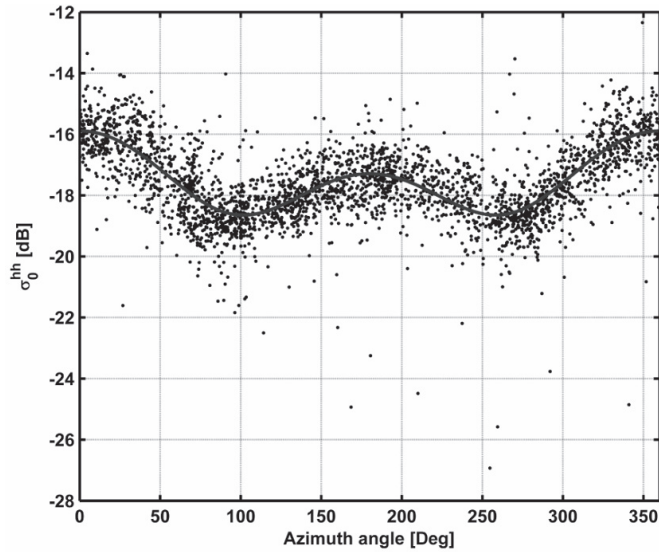


Fig. 1. HH-polarized NRCS as a function of azimuth angle, for a 7 m/s wind speed and 35° incidence angle. Black dots are SAR-observed NRCS at HH-polarization. Blue line represents regression results.

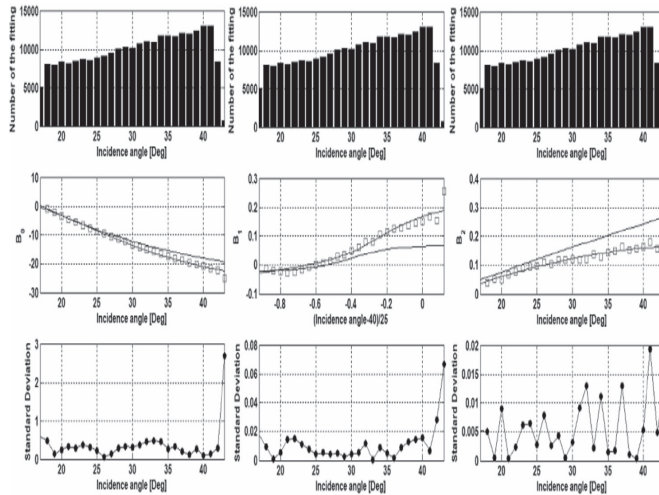


Fig. 2. Number of the data for (Top) B_i coefficients derivation, (Middle) B_i coefficients, and (Bottom) standard deviation of the regression as a function of incident angle for 7 m/s wind speed. B_i coefficients derived from (1), the regression analysis approach, and the CMOD5 are represent as red squares, and blue and black solid lines, respectively.

to determine the corresponding B_i ($i = 0, 1, 2$) coefficients by using a regression analysis approach. Fig. 1 presents the ASAR-observed NRCS (black dots) versus azimuth angles for 7 m/s wind speed and 35° incident angle. Fig. 1 clearly shows the azimuth modulation of NRCS at HH-polarization. Estimated NRCS values using (1) are indicated in the blue solid line in Fig. 1, which are in good agreement with the radar observations, suggesting that (1) is appropriate to describe the proposed GMF.

Subsequently, to further assess the rationality of the functional form, B_i coefficients are also analyzed as a function of wind speed and incident angle. The top three panels in Fig. 2 show the number of data used for B_i coefficients derivation. For 7 m/s wind speed and different incident angles, B_i coefficients derived from (1), the regression analysis

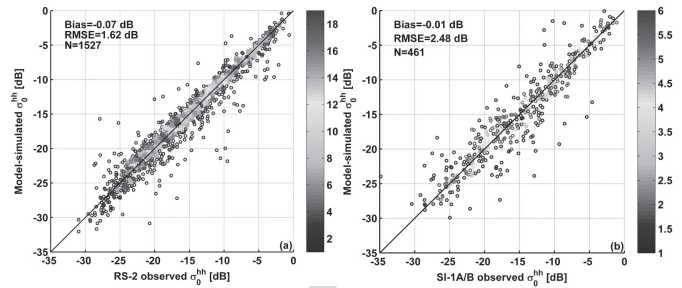


Fig. 3. Comparisons of observed NRCS in HH-polarization (x-axis) from (a) RS-2 and (b) S1-A/B with CMODH model simulations (y-axis). The color bar indicates the scatter density.

approach, and the CMOD5 function are represented as red squares, and blue and black solid lines, in the middle three panels of Fig. 2.

Overall, the trend of B_i coefficients obtained by our analysis and those provided by CMOD5 GMF are in very good agreement excepted at the largest incident angle of 43°. The B_0 values from the regression analysis and from CMOD5 estimation are almost identical at low incident angles (<23°). As incident angles increase, the discrepancies between B_i coefficients obtained for HH from our regression analysis and from CMOD5 gradually increase. This is in line with a known different sensitivity to ocean sea-surface backscattering between VV- and HH-polarizations. Indeed, at C-band, polarization effect is strongly incident angle-dependent. In particular, our analysis confirms that B_1 values at VV are smaller than at HH, while B_0 and B_2 , are larger at VV.

The standard deviations of the regression analysis for each incident angle bin are illustrated in the bottom three panels in Fig. 2. The standard deviation for the large incident angle bin of 43° is much larger than those for other incident angle bins, which is possibly caused by the very few data numbers in this bin. Consequently, we do not use NRCS for this large incident angle bin to develop our HH-polarized GMF. In this study, B_i coefficients are determined via a nonlinear weighted regression approach by minimizing the standard deviation between ENVISAT ASAR observations and regression analysis for all wind speeds and incident angle bins. The number of the data in each wind speed and incident angle bin is considered as a weighted factor for the regression analysis. The final model formulation and 28 coefficients for HH-polarization are given in the Appendix. Since our collocated data sets both involve NRCS at HH- and VV-polarization, we also use (1) to derive the coefficients of GMF for VV-polarization, which are shown in the Appendix. Therefore, one can estimate NRCS at HH-and VV-polarization using (1) and coefficients with given incident angle, wind speed, and direction.

To assess the accuracy of the proposed GMF, we make statistical validations using collocated RS-2 and S1-1A/B observations and buoy measurements. Fig. 3(a) shows the simulated and SAR-measured NRCS at HH-polarization, with a bias of -0.07 dB and a RMSE of 1.62 dB. For comparison between model simulations and S1-1A/B observations, the bias and RMSE are -0.01 and 2.48 dB, which are shown in Fig. 3(b). For NRCS comparisons, the RS-2 data have smaller

222 RMSE than the S1-1A/B data. The reason for this result may
 223 be associated with two factors: 1) RS-2 has lower NESZ than
 224 the S1-1A/B and 2) the collocations between S1-1A/B and
 225 buoy data are less accurate than that of RS-2. For S1-1A/B
 226 collocation, some buoys are outside the SAR image which
 227 leads to the large spatial interval of the collocation. However,
 228 for RS-2 collocation, each buoy locates in the SAR image and
 229 thus there is almost no spatial match error.

230 IV. CONCLUSION

231 Accurate wind field information in high latitudes and polar
 232 regions plays an important role in sea-ice motion and melt
 233 simulations and predictions. The high sensitivity of SAR
 234 backscatter at HH-polarization to sea-ice makes it more
 235 preferable for sea-ice monitoring, rather than VV-polarization.
 236 However, unlike CMOD GMFs, no HH-polarized GMF can be
 237 directly used to retrieve ocean surface wind speed. As a result,
 238 various empirical and theoretical PR models are used to trans-
 239 form NRCS from HH into VV-polarization, which induces
 240 inevitable errors in the retrieved wind speeds. In order to
 241 solve this problem, we develop a new HH-polarized GMF for
 242 C-band SAR, taking advantage of a large number of C-band
 243 SAR data and collocated scatterometer wind measurements.
 244 The proposed GMF has potential to improve SAR wind speed
 245 retrieval accuracy and can be used to directly retrieve wind
 246 speed in mixing areas of ice and water, like the MIZ.

247 The new GMF relates NRCS at HH-polarization to radar
 248 incident angle, wind speed at 10 m height, and relative wind
 249 direction. We use a nonlinear weighted regression method
 250 and the collocated data set (ENVISAT ASAR backscatter
 251 and ASCAT winds) to derive all the coefficients in the
 252 GMF. Furthermore, we use an independent matching of
 253 data (RS-2 and S1-1A/B backscatter and buoy measurements)
 254 to validate the proposed GMF. The statistical comparisons
 255 between radar observations and GMF simulations show a bias
 256 of -0.07 dB and a root-mean-square error of 1.62 dB for RS-2,
 257 and -0.01 dB and 2.48 dB for S1-1A/B, respectively. Based on
 258 this new GMF, we do not need to make a NRCS transformation
 259 with a PR model before wind speed retrieval. The scatter-
 260 omer (SCA) onboard EUMETSAT Polar System-Second
 261 Generation (EPS-SG) mission shall provide observations with
 262 VV-, HH-, or VH-polarization. The addition of HH- or
 263 VH-polarization is considered as an option for extending
 264 the upper limit of the dynamic measurement range. There-
 265 fore, the proposed GMF can be a benefit for EPS-SG SCA
 266 wind retrieval. Moreover, this also opens new perspectives
 267 for the forthcoming launch of the RADARSAT Constellation
 268 Mission (RCM) with new polarization configurations.

269 APPENDIX

270 CMODH MODEL FORMULATION AND COEFFICIENTS

271 The form of the CMODH model is

$$272 \sigma_0^{\text{hh}}(v, \phi, \theta) = (B_0(v, \theta)[1 + B_1(v, \theta) \cos(\phi) \\ 273 + B_2(v, \theta) \cos(2\phi)])^p \quad (\text{A1})$$

274 where p is constant with a value of 1.6 , B_0 , B_1 , and
 275 B_2 are functions of wind speeds v and the incidence

TABLE III
CMODH COEFFICIENTS

Function	Coefficients	HH	VV
B_0	c_1	-0.72722756511	-0.13393789593
	c_2	-1.1901195406	-0.74081314533
	c_3	0.33968637656	0.34811480603
	c_4	0.086759069544	0.019382338942
	c_5	0.003090124916	-0.008066293463
	c_6	0.011761378188	0.006426074015
	c_7	0.129158495658	0.096343783534
	c_8	0.083506931034	0.042280179737
	c_9	4.092557781322	5.007750349297
	c_{10}	1.211169044551	0.717396068916
	c_{11}	-1.119776245438	-1.501296438845
	c_{12}	0.579066509504	0.442826511887
B_1	c_{13}	-0.604527699539	-0.154971505863
	c_{14}	0.118371042255	0.036542289696
	c_{15}	0.008955505675	0.006784919880
	c_{16}	0.219608674529	0.401880787461
	c_{17}	0.017557536680	0.006896838546
	c_{18}	24.442309754388	24.751953435615
B_2	c_{19}	1.983490330585	1.961341923034
	c_{20}	6.781440647278	3.284009890111
	c_{21}	7.947947040974	8.379337236413
	c_{22}	-4.696499003167	-3.636259490187
	c_{23}	-0.437054238710	2.349430558787
	c_{24}	5.471252046908	5.851939658893
	c_{25}	0.639468224273	2.443227221148
	c_{26}	0.673385731705	0.301462797210
	c_{27}	3.433229044819	3.976051353364
	c_{28}	0.367036215316	1.728745711306

angle θ , or alternatively, $x = (\theta - 40)/25$. The B_0 term is
 defined as

$$278 B_0 = 10^{a_0 + a_1 v} f(a_2 v, s_0)^7 \quad (\text{A2})$$

279 where

$$280 \quad f(s, s_0) = \begin{cases} (s/s_0)^\alpha g(s_0), & s < s_0 \\ g(s), & s \geq s_0 \end{cases} \quad (\text{A3})$$

281 where

$$282 \quad g(s) = 1/(1 + \exp(-s)) \quad \text{and} \quad \alpha = s_0(1 - g(s_0)). \quad (\text{A4})$$

283 The functions a_0 , a_1 , a_2 , γ , and s_0 depend on incident angle
284 only

$$285 \quad \begin{aligned} a_0 &= c_1 + c_2x + c_3x^2 + c_4x^3 \\ a_1 &= c_5 + c_6x \\ a_2 &= c_7 + c_8x \end{aligned} \quad (\text{A5})$$

$$286 \quad \begin{aligned} \gamma &= c_9 + c_{10}x + c_{11}x^2 \\ s_0 &= c_{12} + c_{13}x. \end{aligned} \quad (\text{A6})$$

290 The B_1 term is defined as follows:

$$291 \quad B_1 = \frac{c_{14}(1+x) - c_{15}v(0.5+x - \tanh[4(x+c_{16}+c_{17}v)])}{1+\exp(0.34(v-c_{18}))}. \quad (\text{A7})$$

293 The B_2 term is chosen as

$$294 \quad B_2 = (-d_1 + d_2v_2)\exp(-v_2). \quad (\text{A8})$$

295 Here, v_2 is given by

$$296 \quad v_2 = \begin{cases} a + b(y-1)^n, & y < y_0 \\ y, & y \geq y_0 \end{cases}, \quad y = \frac{v+v_0}{v_0} \quad (\text{A9})$$

297 where

$$298 \quad y_0 = c_{19}, n = c_{20} \quad (\text{A10})$$

$$299 \quad a = y_0 - (y_0 - 1)/n, b = 1/[n(y_0 - 1)^{n-1}]. \quad (\text{A11})$$

300 The quantities v_0 , d_1 , and d_2 are functions of incident angle
301 only

$$302 \quad \begin{aligned} v_0 &= c_{21} + c_{22}x + c_{23}x^2 \\ d_1 &= c_{24} + c_{25}x + c_{26}x^2 \\ d_2 &= c_{27} + c_{28}x. \end{aligned} \quad (\text{A12})$$

305 The coefficients are given in Table III.

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