Paleomagnetic recording efficiency of sedimentary magnetic mineral inclusions: implications for relative paleointensity determinations

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Abstract :

Sedimentary relative paleointensity (RPI) records are often carried by complex magnetic mineral mixtures, including detrital and biogenic magnetic minerals. Recent studies have demonstrated that magnetic inclusions within larger detrital silicate particles can make significant contributions to sedimentary paleomagnetic records. However, little is known about the role such inclusions play in sedimentary paleomagnetic signal recording. We analyzed paleomagnetic and mineral magnetic data for marine sediment core MD01-2421 from the North Pacific Ocean, offshore of central Japan, to assess how magnetic inclusions and other detrital magnetic minerals record sedimentary paleomagnetic signals. Stratigraphic intervals in which abundant magnetic inclusions dominate the magnetic signal are compared with other intervals to assess guantitatively their contribution to sedimentary RPI signals. The normalized remanence record from core MD01-2421 does not correlate clearly with global RPI stacks, which we attribute to a demonstrated lower paleomagnetic recording efficiency of magnetic inclusions compared to other detrital magnetic minerals. We also carried out the first laboratory redeposition experiments under controlled Earth-like magnetic fields for particles with magnetic inclusions using material from core MD01-2421. Our results confirm that such particles can be aligned by ambient magnetic fields, but with a lower magnetic recording efficiency compared to other detrital magnetic minerals, which is consistent with normalized remanence data from core MD01-2421. Our demonstration of the role of sedimentary magnetic inclusions should have wide applicability for understanding sedimentary paleomagnetic recording.

Keywords : magnetic mineral inclusions, relative paleointensity, detrital remanent magnetization

1. Introduction

Understanding recording processes associated with detrital remanent magnetization (DRM) acquisition is important in sedimentary paleomagnetism because records of changing geomagnetic directions and relative paleointensity (RPI) have been widely used to understand geomagnetic field variations and deep-Earth geodynamo processes (e.g., Guyodo & Valet, 1999; Valet et al., 2005). Such records are also used widely to date sedimentary sequences (e.g., Roberts et al., 2013). As well as being controlled by the ambient field strength, RPI signals can be affected by various factors associated with the magnetic minerals in a sedimentary sequence, such as grain size distributions and magnetic properties, and a range of external factors, such as salinity, bioturbation, and post-depositional compaction during sediment deposition and burial (e.g., Katari & Tauxe, 2000; Zhao et al., 2016). Our understanding of DRM acquisition mechanisms has improved as a result of numerical simulations of individual particles and flocs (Heslop et al., 2006; Heslop, 2007; Shcherbakov & Sycheva, 2008, 2010; Heslop et al., 2014), and laboratory-controlled redeposition experiments (Tauxe et al., 2006; Mitra & Tauxe, 2009). These experimental studies include time/field dependence (Katari & Tauxe, 2000; Quidelleur et al., 1995), different materials and water contents (Levi & Banerjee, 1990; Lu et al., 1990; Jackson et al., 1991; Carter-Stiglitz et al., 2006; Paterson et al., 2013; Valet et al., 2017), and flocculation (Van Vreumingen, 1993a, 1993b; Katari & Tauxe, 2000; Mitra & Tauxe, 2009; Spassov & Valet, 2012). However, understanding of RPI signal recording remains poor due to the complex effects of the abovementioned factors (Roberts et al., 2013).

Previous studies have investigated the effects of sediment composition on DRM acquisition, including, for example, clay (King, 1955; Blow & Hamilton, 1978), inorganic magnetite (Lu et al., 1990; Jackson et al., 1991; Van Vreumingen, 1993a, 1993b), and

biogenic magnetite (Paterson et al., 2013; Valet et al., 2017). These studies demonstrate that inorganic and biogenic magnetite can be important sedimentary paleomagnetic signal carriers. Recently, Chang et al. (2016a, 2016b) suggested that magnetic mineral inclusions hosted within detrital silicate particles occur widely in marine sediments and that they represent a potentially important DRM recorder. Compared to other detrital magnetic mineral forms, which are generally treated as isolated particles, magnetic inclusions consist of fine ferrimagnetic mineral grains embedded within a coarser non-magnetic host mineral. Magnetic inclusions occur abundantly in igneous rocks and are well known to be paleomagnetically important in igneous rocks (e.g., Evans et al., 1968; Feinberg et al., 2005, 2006; Tarduno et al., 2006; Usui et al., 2015). Magnetic inclusions are predominantly distributed over the fine-grained stable single-domain (SSD) or vortex state size range (e.g., Harrison et al., 2002; Tarduno & Cottrell, 2005; Chang et al., 2016a, 2016b). Also, inclusions are surrounded and protected against surrounding environmental changes by their aluminosilicate hosts (e.g., plagioclase and pyroxene, which are abundant in igneous and metamorphic rocks). These characteristics mean that magnetic inclusions can potentially preserve stable remanent magnetizations over billions of years (Selkin et al., 2000; Renne et al., 2002; Tarduno et al., 2006, 2010; Muxworthy et al., 2013). Silicates can protect embedded magnetic inclusions from chemical alteration induced, for example, by hydrothermal fluids or oxidation in igneous rocks (Tarduno et al., 2006) or during reductive diagenesis in sediments (Chang et al., 2016a). Magnetic inclusions within silicates in sediments could, therefore, be a potential recorder of sedimentary paleomagnetic signals, even for sediments that have been subjected to extensive chemical or mechanical alterations (e.g., post-deposition compaction and bioturbation) (Chen et al., 2017).

Few studies have contributed to understanding how magnetic inclusions contribute to sedimentary paleomagnetic signals (e.g., Chang et al., 2016a; Chen et al., 2017). In this

study, we explore signals carried by magnetic inclusions through paleomagnetic analysis of marine sediment core MD01-2421 from offshore of Japan. This sediment core contains abundant magnetic inclusions, and the magnetic properties of some intervals are dominated by inclusions (Chang et al., 2016a). We divide the sediment core into intervals with abundant magnetic inclusions and those with mixed detrital magnetic mineral assemblages (which we refer to as inclusion-rich and inclusion-poor intervals, respectively), based on their distinctive magnetic properties. Such separation enables assessment of paleomagnetic recording behavior associated with sedimentary magnetic mineral inclusions. In addition, we present results of preliminary laboratory redeposition experiments using particles with magnetic inclusions from different intervals of the studied sediment core. Together, these results are used to assess DRM recording associated with magnetic inclusions.

2. Marine sediment core MD01-2421

The marine sediment core studied here, MD01-2421, is from the North Pacific Ocean, ~100 km offshore of central Japan (36°01.40 N, 141°46.80 E; 2,224 m water depth; 45.82 m long, Oba et al., 2006; Chang et al., 2016a) (Fig.1). The core was recovered from a location where the warm and saline subtropical Kuroshio Current and the cold and less saline subarctic Oyashio Current meet. The core was collected during IMAGES (International Marine Global Change Study) cruise VII-WEPAMA (Western Pacific Margin) Leg 2 of R/V *Marion Dufresne* in 2001 (Oba et al., 2006). The recovered sediments are homogenous olive-gray silty clays with calcareous and siliceous microfossils and high total organic carbon (TOC) contents (0.5-2.1 wt %; Ueshima et al., 2006). The sediments were deposited continuously with a basal age of 144 ka and an average sedimentation rate of 30 cm/kyr in a hemipelagic environment (Oba et al., 2006). A detailed age model for the core is based on

oxygen isotope stratigraphy from benthic foraminifera (Oba et al., 2006). The magnetic domain state of samples from core MD01-2421 spans a spectrum from SSD to vortex state with high coercivity (i.e., $B_c > 30$ mT) to multi-domain (MD) grains with low coercivity (i.e., $B_c < 20$ mT) (Chang et al., 2016a). Detailed magnetic measurements, including hysteresis properties, first-order reversal curve (FORC) diagrams, isothermal remanent magnetization (IRM), and thermomagnetic curves indicate consistently that the constituent magnetic minerals are dominantly magnetite and titanomagnetite (Chang et al., 2016a). Scanning electron microscope (SEM) and transmission electron microscope (TEM) observations on magnetic extracts from the sediments reveal the presence of fine magnetic particles, including the widespread occurrence of magnetite/titanomagnetite nanoparticles hosted in silicates (Chang et al., 2016a), in addition to coarse-grained and unprotected detrital magnetite/titanomagnetite grains.

3. Methods

3.1 Paleomagnetic and rock magnetic measurements of u-channels

Paleomagnetic data were obtained from the entire core from u-channel samples, which typically have 1.5 m length and 2×2 cm square cross-section. The natural remanent magnetization (NRM) of u-channels was measured at 1 cm intervals using a 2-G Enterprises 755R long-core cryogenic magnetometer with stepwise alternating field (AF) demagnetization carried out at 5 mT intervals from 0 to 20 mT at Doshisha University, Japan. Some u-channels were AF demagnetized up to 30 mT. Data from the top and bottom 5 cm of each u-channel were discarded to avoid edge effects. After AF demagnetization, an anhysteretic remanent magnetization (ARM) was imparted using a 100 μ T direct current (DC) bias field superimposed on a decaying 80 mT peak AF. ARM was measured first at Doshisha University, and was then re-measured at Kochi Core Centre (KCC), Kochi University, Japan, with a 2-G Enterprises 755R long-core cryogenic magnetometer after a few years. Nearly identical results were found between the two measurement systems. ARM data from KCC was used in this paper. The translation speed used to impart an ARM to the u-channels at KCC was 10 cm/sec. After ARM measurements, an IRM was imparted at 1 T using a pulse magnetizer, which was then measured with a 2-G Enterprises 755R long-core cryogenic magnetometer at KCC.

3.2 Magnetic measurements on discrete samples and SEM observations

In order to validate demagnetization results from u-channel samples, discrete samples were taken with paleomagnetic cubes for every section from the studied core (e.g., ~1.5 m sampling interval). AF demagnetization was carried out in 12 steps for these samples (0, 2.5, 5, 10, 15, 20, 25, 30, 35, 40, 45, 50, 55, and 60 mT) and magnetizations were measured using a flux-gate spinner magnetometer (Natsuhara-Giken SMD-88) at Doshisha University. To understand paleomagnetic recording of magnetic inclusions, we carried out redeposition experiments in a controlled laboratory environment. Experimental DRM results for two samples from inclusion-rich (MD01-2421-10-100, 14.47 m, 42.91 ka) and inclusion-poor (MD01-2421-16-100, 23.3 m, 71.43 ka) intervals are compared: one sample contains mainly particles with magnetic inclusions (where other unprotected magnetic grains were dissolved diagenetically; red star in Fig. 2), while the other sample contains a mixture of large, unprotected detrital grains and minor magnetic mineral inclusions within silicates (blue star in Fig. 2). The two samples were characterized magnetically and with SEM observations. Hysteresis and FORC measurements were made with a Princeton Measurements Corporation

vibrating sample magnetometer (model 3900) at the Paleomagnetism and Geochronology Laboratory (PGL), Institute of Geology and Geophysics, Chinese Academy of Sciences (IGGCAS), Beijing. Hysteresis loops were measured between -1 and +1 T or between -500 and +500 mT with a field step of 5 mT and an averaging time of 250-300 ms. FORC diagrams (Pike et al., 1999; Roberts et al., 2000) were obtained with a 1 T maximum applied field, 2.7 mT field increments and 350 ms averaging time. FORC diagrams were produced with the FORCinel 3.0 software of Harrison and Feinberg (2008). Electron microscope observations were made using a FEI Quanta 650 field emission gun (FEG) SEM at the School of Earth and Space Sciences, Peking University, China.

3.3 DRM redeposition experiments

Six air-dried sister specimens with similar mass were used for DRM experiments. For laboratory-controlled redeposition experiments, we used two orthogonal sets of Helmholtz coils. First, we mixed ~0.05 g of sediment with pure water and filled standard $2 \times 2 \times 2$ cm paleomagnetic plastic cubes with the slurry. After shaking to randomize particle orientation, the cubes were put inside the Helmholtz coils. The sediment settled in a 60 µT field over 6 days. To prevent dust contamination, samples were covered during settling and drying. After drying completely, the redeposited sediment formed a thin layer at the base of each cube. Remanent magnetization measurements were made with a 2-G Enterprises DC superconducting quantum interference device (SQUID) rock magnetometer (model 755; noise level 3×10^{-12} Am²) housed in a magnetically shielded room at Peking University. Magnetic susceptibility was measured with an AGICO MFK1-FA Kappabridge system. An ARM was imparted with a 100 mT peak AF and a 50 µT DC bias field. An IRM was imparted with an ASC IM-10-30 pulse magnetizer. Remanence was measured with the

SQUID magnetometer. All DRM experiments and subsequent paleomagnetic measurements, except for ARM measurements, were carried out at Peking University. ARM measurements were made at the PGL, IGGCAS, Beijing.

4. Results

4.1 Paleomagnetic behavior of magnetic mineral inclusions in core MD01-2421

4.1.1 Sedimentary intervals dominated by magnetic mineral inclusions

Magnetic and geochemical property variations for core MD01-2421 (Fig. 2a-c), including coercivity (B_c) , saturation magnetization (M_s) , and elemental S abundance (Chang et al., 2016a), have large down-core variations. Down-core trends of X-ray fluorescence (XRF) S abundance and B_c are similar, but both correlate inversely with M_s (Fig. 2). These correlations indicate a strong relationship between the magnetic mineral assemblage and geochemical processes through reductive diagenesis (Chang et al., 2016a). Magnetic minerals (both concentration and mineralogy) are influenced by variations in sulfidic diagenesis intensity through time (indicated by S variations), which control down-core B_c and M_s fluctuations. Unprotected magnetic particles, including detrital and biogenic magnetite, were dissolved, but fine-grained Fe-Ti oxide inclusions were preserved because they were protected by host silicate grains even under strongly sulfidic conditions (indicated by high sedimentary S contents) because silicates are largely unreactive to sulfide (Canfield et al., 1992; Poulton et al., 2004; Roberts, 2015; Chang et al., 2016b). The dominant contribution of inclusions to the magnetic signal within strongly diagenetically altered sediment intervals is confirmed by TEM observations and magnetic property variations (Chang et al., 2016a, 2016b). Based on these relationships, we isolated sediment intervals with abundant magnetic

inclusions (gray bar in Fig. 2). Criteria for defining these gray shaded intervals (Fig. 2) are based on absolute values of B_c , M_s , and S, and also on down-core data trends (i.e., increasing or decreasing), where 'inclusion-rich' intervals (marked by gray shading) have relatively high B_c , low M_s , and high S values. These intervals are distinct from other intervals in which more diverse detrital magnetic mineral mixtures are present, including inclusions. The presence of silicate grains that contain magnetic mineral inclusions in core MD01-2421 has been validated by further SEM imaging (section 4.2) of magnetic mineral extracts and handpicking of silicate single crystals in addition to the observations presented by Chang et al. (2016a). It is not feasible to quantify the relative abundance of magnetic mineral inclusions throughout the core using direct electron microscopic observations; our distinction is based on down-core magnetic and geochemical profiles as proxies that can only provide an estimate of relatively 'inclusion-rich' layers. Nevertheless, this approach enables a first-order separation of magnetic mineral inclusion-rich intervals from inclusion-poor intervals, so that their paleomagnetic recording behavior can be compared.

AF demagnetization results of discrete (Fig. 3a, c) and u-channel samples (Fig. 3b, d) for core MD01-2421 indicate that a viscous remanent magnetization (VRM) component can be eliminated below 10 mT for both inclusion-rich (Fig. 3a, b) and inclusion-poor samples (Fig. 3c, d). This indicates that remanence after AF demagnetization at 10 mT is primary and should reflect different DRM behavior between magnetic inclusion-rich and inclusion-poor samples. The difference in remanence is calculated from the NRM at 0 mT subtracted from the NRM after AF demagnetization at 20 mT. The difference in remanence between inclusion-rich and inclusion-poor samples is smaller after AF demagnetization at 20 mT (Fig. 2d). The average differences between inclusion-rich and inclusion-poor samples are 1.914 and 4.999 A/m with standard deviations of 1.076 and 3.868 A/m, respectively. We plotted paleomagnetic data after AF demagnetization at 10 and 20 mT. Magnetic inclusions often

occur within the tens of nanometer size range, and have stable remanent magnetizations and high coercivities (e.g., Harrison et al., 2002; Feinberg et al., 2005; Tarduno & Cottrell, 2005; Chang et al., 2016b). Intervals with significant magnetic inclusion contents (gray bars in Fig. 2d), therefore, have relatively small differences between NRM and remanence after AF demagnetization at 20 mT compared to inclusion-poor intervals (Fig. 2d), which is consistent with their more stable remanent magnetization. This confirms that our approach for isolating magnetic inclusion-rich intervals is appropriate.

4.1.2 Paleomagnetic record of MD01-2421 compared to global RPI stacks

Global paleointensity stacks, including the Sint-800 (Guyodo & Valet, 1999), PISO-1500 (Channell et al., 2009), and PADM2M stacks (Ziegler et al., 2011), are compared with the normalized remanence record of core MD01-2401 in Fig. 4b-d. Global stacks outline the general global trend of relative paleomagnetic intensity variations. Good correlation between the paleomagnetic record for core MD01-2421 and these global stacks is not expected because of large down-core changes in magnetic mineralogy, concentration, and grain size primarily due to magnetic mineral diagenesis (Chang et al., 2016a). The purpose of our comparison is to identify potential paleomagnetic recording biases due to the complex magnetic mineralogy, including those due to magnetic inclusions. Normalized remanence records were obtained by normalizing NRM (after AF demagnetization at 10 and 20 mT) by susceptibility (*k*), ARM, and IRM (Fig. 4). To minimize effects due to changing sedimentation rates between global stacks and the studied core, data for core MD01-2421 were computed using a 100-year moving average (Fig. 4b-d, darker colors). Normalized remanence records for core MD01-2421 are similar for three normalizers (Fig. 4b-d).

There is poor correlation between the global RPI stacks and normalized remanence

records for core MD01-2421 (Fig. 4). Importantly, the normalized remanences generally have minimum values within magnetic inclusion-rich intervals (Fig. 4). The weaker signals indicate that magnetic inclusions potentially induce a recording bias, although we cannot make definite conclusions from these data because of possible effects due to regional geomagnetic field differences with respect to the global field. We plot NRM (after AF demagnetization at 20 mT) versus ARM, IRM, and k separately for inclusion-rich and inclusion-poor intervals (Fig. 5a-c). Distinctly different slopes for linear fits are observed for the respective intervals, where the normalized NRM for inclusion-rich intervals has a lower slope compared to inclusion-poor intervals (Fig. 5a-c). These data support a lower recording efficiency of magnetic inclusions compared to other detrital magnetic minerals. An ARM versus k plot (Banerjee et al., 1981) indicates a dominance of finer particles for inclusion-rich intervals compared to inclusion-poor intervals (Fig. 5d), but these finer particles have relatively lower remanence recording efficiency. This is likely caused by the lower responsiveness of the relatively heavy non-magnetic silicate host particles to alignment by geomagnetic torques. The grain size of magnetic inclusions is smaller than unprotected detrital particles in the studied sediment core, so ARM and IRM are expected to be more sensitive to magnetic inclusions than to k. This probably explains why the ARM and IRM normalized records are more similar to each other than to the k-normalized record (Fig. 4). Results obtained in a 60 μ T applied field in our redeposition experiments and normalized by IRM are compared to results at the same applied field for artificial samples containing magnetosomes (Paterson et al., 2013) (dark green, Fig. 8f). IRM acquisition conditions and sediment concentrations have been ignored for our comparison (Fig. 8f). It can be seen that our DRM/IRM results from inclusion-poor intervals are consistent with values from magnetosomes (Fig. 8f), whereas inclusion-rich intervals are weaker than magnetosome results. These findings indicate different paleomagnetic recording efficiency for various

constituent magnetic particle types, where magnetic inclusions within larger silicate particles have lower paleomagnetic recording efficiency. This confirms the results of Chen et al. (2017), which indicate that biogenic magnetite has a higher recording efficiency compared to detrital magnetic inclusions. More routine diagnosis of detailed relationships between detrital magnetic minerals and DRM recording is clearly needed to improve our understanding of sedimentary paleomagnetic signal recording.

Our results enable comparison of magnetic recording efficiency for sediment intervals that are dominated by magnetic inclusions and other forms of magnetic minerals, and confirm that normalized remanence records for inclusion-rich intervals are affected by a recording bias toward weaker remanences compared to inclusion-poor intervals (Fig. 4). Such inclusion-rich intervals can distort RPI by recording weaker remanences (Fig. 4). As indicated by Roberts et al. (2012), such recording differences may not be constant throughout a sedimentary sequence because stratigraphically varying concentrations of different magnetic mineral types will produce variable DRM/*k*, DRM/ARM, and DRM/IRM slopes (Fig. 5). Thus, DRM records should be interpreted carefully, particularly considering the different magnetic mineral forms that occur in sediments and their contrasting magnetic recording efficiencies (e.g., Roberts et al., 2012; Chen et al., 2017). For example, biogenic magnetite has greater recording efficiency compared to detrital magnetite (Ouyang et al., 2014; Chen et al., 2017).

Recent studies demonstrate the widespread occurrence of different magnetic mineral types in sedimentary records. These different magnetic minerals can have different magnetic recording efficiencies (e.g., Ouyang et al., 2014; Chen et al., 2017). Further work is needed to recognize and characterize variable remanence acquisition efficiency associated with different magnetic mineral types, such as magnetic inclusions, and biogenic and detrital

magnetic minerals. It is rare to find sediment sequences with a single and constant magnetic mineral assemblage, so suitable strategies for RPI normalization are needed for mixed magnetic mineral assemblages where the studied core fails the criteria for reliable RPI records (Tauxe, 1993).

6. Conclusions

Paleomagnetic data from marine sediment core MD01-2421 were analyzed by comparing the behavior of two types of distinct sediment intervals, where the magnetic signal in one is dominated by magnetic inclusions and the other is relatively poor in inclusions. Normalized remanence for the magnetic mineral inclusion-rich intervals generally records minima that are not observed in global RPI stacks. This is likely caused by the overall lower magnetization of silicate grains that contain magnetic inclusions. Such silicate hosts are not magnetic and their larger size means that their response to a geomagnetic aligning torque will be counteracted more strongly by the hydrodynamic settling force. Our results indicate different paleomagnetic recording efficiencies for different detrital magnetic mineral types, where magnetic inclusions have lower recording efficiency. This conclusion is confirmed by redeposition experiments for natural sediments from magnetic inclusion-rich intervals and other intervals dominated by larger unprotected detrital magnetic grains. Contrasting DRM behavior is observed for these two sample types. In all cases, inclusion-poor intervals have stronger remanence than inclusion-rich intervals. Our results demonstrate that different detrital magnetic mineral types in sediments have different recording efficiency. These observations have important implications for understanding the mechanisms of paleomagnetic recording and relative paleointensity normalization.

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Figure 1. Location map of the studied marine sediment core MD01-2421 (red solid circle) in the North Pacific Ocean. The environment near the core location is affected by the cold Oyashio current (yellow arrow) and the warm Kuroshio current (orange arrow).



Figure 2. Down-core geochemical, magnetic, and paleomagnetic variations for core MD01-2421: (a) elemental S abundance, (b) coercivity (B_c), (c) saturation magnetization (M_s), and (d) paleomagnetic remanence after 0 (dark brown line) and 20 mT (brown line) and the difference between these steps (bright orange bar). Geochemical and magnetic data are from Chang et al. (2016a). Gray bars mark magnetic inclusion-rich intervals with harder magnetic minerals, as indicated by high coercivity and high S abundance, and weaker magnetization (e.g., lower M_s). Red and blue stars indicate stratigraphic positions of "inclusion-rich" and "inclusion-poor" samples used for laboratory redeposition experiments, respectively. Green (sample MD01-2421-7-110, 10.06 m, 21.21 ka; Chang et al., 2016a) and red (MD01-2421-10-100, 14.47 m, 42.91 ka; section 3.2) stars indicate the stratigraphic positions of samples from which silicate grains have been observed from electron microscopic observations to contain magnetic mineral inclusions.



Figure 3. Representative AF-demagnetization results for discrete and u-channel samples for core MD01-2421. Solid and open circles represent projections onto the horizontal and vertical planes, respectively. (a, b) Inclusion-rich, (c, d) inclusion-poor intervals, (a, c) discrete samples, and (b, d) u-channel results.



Figure 4. Comparison of (a) the Sint-800 (orange), PISO-1500 (purple), and PADM2M (pink) RPI stacks to the normalized remanence record of core MD01-2421 for (b) NRM₂₀/IRM (blue), (c) NRM₂₀/ARM (green), and (d) NRM₂₀/susceptibility (red). Solid (dotted) lines are AF-demagnetization in 20 mT (10 mT) and the light color is the original data before calculating a 100-yr moving average. Gray bars are the same as in Figure 2.

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Figure 5. Comparison of paleomagnetic recording efficiency for "inclusion-rich" and "inclusion-poor" sediment intervals from core MD01-2421. NRM_{20mT} versus (a) susceptibility, (b) ARM, and (c) IRM, and (d) ARM versus susceptibility (Banerjee et al., 1981). Lines are means for "inclusion-rich" (red) and "inclusion-poor" (blue) samples. All data are from u-channel samples. Darker (lighter) colors represent 100-yr moving average (original) data.



Figure 6. (a, b) Hysteresis loops and (c, d) FORC diagrams for the two samples from (a, c) inclusion-rich (MD01-2421-10-100, 14.47 m, 42.91 ka) and (b, d) inclusion-poor (MD01-2421-16-100, 23.3 m, 71.43 ka) intervals, respectively, which were used for the redeposition experiments. Note that the loop for the inclusion-rich sample was measured to ± 1 T, but only a partial field range is shown. Hysteresis parameters (B_c and M_{rs}/M_s) are indicated. FORC diagrams were processed with VARIFORC smoothing parameters (Egli, 2013): {sc0, sc1, sb0, sb1, λ_c , λ_b } = {6, 6, 4, 6, 0.1, 0.1}.



Figure 7. SEM images of magnetic extracts from an inclusion-rich sample (MD01-2421-10-100, 14.47 m, 42.91 ka). Observed morphologies of magnetic mineral inclusions include (a) nanoparticle clusters, and (b) dendrites within silicates.

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Figure 8. Results of laboratory redeposition experiments in a 60 μ T ambient field with set declination (0, 45, and 90°) and inclination (45°) for an inclusion-rich sediment sample (MD01-2421-10-100, 14.47 m, 42.91 ka, red) and an inclusion-poor sediment sample (MD01-2421-16-100, 23.3 m, 71.43 ka, blue) from core MD01-2421: sub-samples from the magnetic inclusion-rich interval (red) and the inclusion-poor interval (blue). Stratigraphic positions of the selected samples are marked in Figure 2. The solid line is the expected declination trend. Acquisition of DRM (a) declination, (b) inclination, and intensity normalized by (c) mass, (d) susceptibility, (e) ARM, and (f) IRM compared with results from redeposition experiments for artificial sediments using magnetosomes (dark green: Paterson et al., 2013) and natural marine sediments (this study).