
Evaluating the ecological status of cold-water coral habitats using non-invasive methods: An example from Cassidaigne canyon, northwestern Mediterranean Sea

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Abstract :

Cold-water coral ecosystems have been identified as vulnerable, but quantitative data on their conservation status is very limited. The Marine Strategy Framework Directive (MSFD) is the tool implemented by the European Union's Integrated Maritime Policy to achieve Good Environmental Status (GES) of marine waters by 2020. In this context, the aim of this study was to evaluate the Ecological Status of benthic habitats in Cassidaigne canyon, focusing in particular on cold-water coral habitats dominated by *Madrepora oculata*.

Data were collected during the Videocor1 cruise (2017). Videos and photos collected during eight dives of the H-ROV Ariane were used to reconstruct, in 3-dimensions, the areas where cnidarians have settled in the canyon. A total of 33 3D models were built, which allowed measuring the spatial and vertical distribution, surface, density and size structure of cnidarian populations at four different sites. When 3D reconstructions were not possible, GIS tools were used. The seven cnidarian species considered were the scleractinian *M. oculata*; three antipatharians: *Leiopathes glaberrima*, *Antipathella subpinnata*, *Antipathes dichotoma*; and three aclyonaceans: the precious red coral *Corallium rubrum* and the gorgonians *Callogorgia verticillata* and *Viminella flagellum*.

Using photogrammetry, we were able to reveal the size structure of the dense population of *M. oculata* in the canyon, as well as to obtain knowledge on a complex site (Cassis-200) composed of 15 knolls, and to quantify the surface occupied by *M. oculata* at a separate site (Cassis-500) influenced by industrial discharges. At the southern flank of the canyon we found a highly diverse site (SW Flank) dominated by antipatharians and gorgonians composing large forests, and finally a reservoir of *M. oculata* was identified under overhangs at a site called the Wall. The diversity of accompanying species is also reported and marine litter quantified. Images collected before 2017 were compared to the 3D models to precisely locate them on the sites, and assess temporal changes in *M. oculata* colony sizes at Cassis-200 site. We also report on the ground-truthing of predicted habitat maps produced previously, and confirm their good representation of the distribution of cold-water coral habitats. Finally, we quantified the criteria defined by the MSFD, aimed at evaluating the GES of benthic habitats for *M. oculata* ecosystems, at the scale of the

Cassidaigne canyon. Measurements showed that the extent of loss of the observed *M. oculata* habitat reached 56% according to the MFSD definition.

Highlights

► Cold-water corals distribution, density and size structure are reported. ► Size measurements were performed using 3D-models. ► A diverse site dominated by antipatharians and gorgonians forests was found. ► The extent of loss of the observed *M. oculata* habitat reached 56%. ► Ground-truthing of predicted habitat maps produced previously is reported.

Keywords : Biodiversity, Photogrammetry, Habitat mapping, Bauxite residues, Vulnerable marine ecosystems

44 1. Introduction

45 Cold-water corals encompass many species of the subclass Hexacorallia, including the order of stony
46 corals (Scleractinia) among which are framework-building species, and the order of black corals
47 (Antipatharia) forming arborescent colonies. Some species of the subclass Octocorallia, including
48 gorgonians and alcyonaceans, are also arborescent corals. The two main species known to form
49 complex three-dimensional structures in the Mediterranean Sea are *Lophelia pertusa* and *Madrepora*
50 *oculata*. These framework-building species can occur in very high densities, increasing habitat
51 heterogeneity thus providing niches for several species at different scales among their polyps,
52 branches, colonies, patches and mounds, and enhancing species density and/or diversity at a local or
53 regional scale (Henry and Roberts, 2017). Framework-building and arborescent corals are very
54 important habitat-structuring species that provide many ecosystem services, by increasing the local
55 biodiversity in the deep Mediterranean Sea (D'Onghia et al., 2010).

56 Various human activities are known to have a negative effect on cold-water coral communities.
57 Bottom trawling is the most destructive (Benn et al., 2010), and has been banned below 1000 m in
58 the Mediterranean Sea since 2005 (<http://www.fao.org/gfcm/data/maps/fras/en/>). Although having
59 much less impact than bottom trawling, bottom longlines and gill nets may cause damage to coral
60 colonies when these fishing gears get entangled with corals which are then broken during retrieval of
61 the gear or if gears are lost, trapped in the colonies and moved by water currents. Sediment
62 resuspension and accumulation, seafloor dredging, mineral activities, oil prospection, litter, waste
63 disposal, submarine cables, CO₂-induced global-warming and ocean acidification are other potential
64 threats to the environmental integrity of cold-water coral habitats caused by human activities
65 (Freiwald et al., 2004; Guinotte et al., 2006; McCulloch et al., 2012; Puig et al., 2012; Fabri et al.,
66 2014; Roberts and Cairns, 2014; Lastras et al., 2016). These impacts lead to habitat degradation or
67 destruction and therefore loss or change in biodiversity.

68 The Marine Strategy Framework Directive (MSFD), adopted in June 2008, commits European Union
69 (EU) Member States to adopt an ecosystem approach to managing the marine environment
70 (European Parliament, 2008). By implementing this directive, the member states aim to achieve good
71 environmental status (GES) of their marine waters by 2020, based on 11 descriptors. This includes
72 deep-sea waters within the European Union Exclusive Economic Zone (EEZ) and embraces, as defined
73 by the MSFD, the water column and the seabed and subsoil under the water column. Monitoring,
74 assessing the environmental status, and managing the deep sea on the basis of knowledge of
75 ecosystem biodiversity, functions and services are crucial actions to ensure its long-term
76 sustainability. Among the 11 descriptors of the MSFD framework, seafloor integrity (D6) is the
77 descriptor used to characterize the status of benthic habitats (European Commission, 2017a, b).
78 Habitat destruction is one of the main threats to environmental integrity (Claudet and Fraschetti,

79 2010). Assessing the consequences of human impacts is crucial to both predict and prevent structural
80 and functional changes of habitats.

81 In French Mediterranean waters, structure-forming scleractinians are known from two canyons:
82 Lacaze-Duthiers and Cassidaigne (Fabri et al., 2014). Our study focuses on Cassidaigne canyon, where
83 potential distribution maps of cold-water coral habitat have been produced previously (Fabri et al.,
84 2017; Bargain et al., 2018). These potential suitable areas were inspected in 2017 with a Remotely
85 Operated Vehicle (ROV) to validate the predicted habitat distribution and evaluate the status of
86 populations of different species of benthic megafauna in relation to anthropogenic impacts. For 48
87 years (1967 to 2015), Cassidaigne canyon was subjected to massive discharges from the aluminum
88 industry, which is known to have had impacts on the local fauna of the canyon (Dauvin, 2010;
89 Fontanier et al., 2012; Fabri et al., 2014; Fontanier et al., 2015; Fabri et al., 2017).

90 Video and high-resolution images acquired during the 2017 surveys were used to build three-
91 dimensional (3D) models of outcrops on which coral colonies have settled using photogrammetry.
92 The reconstructions of the real-scale 3D models were built from a series of overlapping photos and
93 snapshots of videos (Arnaubec et al., 2015). Photogrammetry provides a permanent record of the 3D
94 coral colonies, gives access to quantitative parameters, replicates measurements and allows
95 monitoring over time. This non-invasive methodology has recently been used for monitoring marine
96 habitats, mostly in shallow waters. For example, it was used to evaluate gorgonian densities and
97 abundances in the Marine Protected Area of Portofino (Palma et al., 2018) and a methodology
98 dedicated to the management and conservation of red coral populations has been formulated (Royer
99 et al., 2018). Regarding deep waters, one study was performed in the Whittard canyon (Atlantic
100 ocean), where photogrammetry was used to highlight the complexity of coral habitats living on
101 vertical walls (Robert et al., 2017). In our study we used photogrammetry to evaluate the ecological
102 status of cold-water coral habitats in Cassidaigne canyon. Our objectives were to:

103 (1) draw the spatial and vertical distribution of colonial cnidarians (the scleractinian *M. oculata*; the
104 antipatharians: *Leiopathes glaberrima*, *Anthipathella subpinnata* and *Antipathes dichotoma*; the red
105 coral *Corallium rubrum*, and the gorgonians *Callogorgia verticillata* and *Viminella flagellum*);

106 (2) measure the surface covered by, and the density of, populations of *M. oculata*, *C. rubrum* and
107 gorgonians using photogrammetry and video analyses;

108 (3) measure the size structure of *M. oculata* and *C. verticillata* populations at different locations in
109 the canyon;

110 (4) evaluate the diversity of megafauna species at the different locations in the canyon;

111 (5) quantify the seafloor integrity in areas dominated by *M. oculata* populations at the scale of the
112 Cassidaigne canyon, using the two criteria defined by the MSFD to evaluate the status of benthic
113 habitats;

114 (6) assess temporal changes in ecological status at a location in Cassidaigne canyon that had been
115 previously surveyed in 2010 and 2013;

116 (7) validate the habitat suitability model previously built for cold-water corals at the scale of the
117 Cassidaigne canyon.

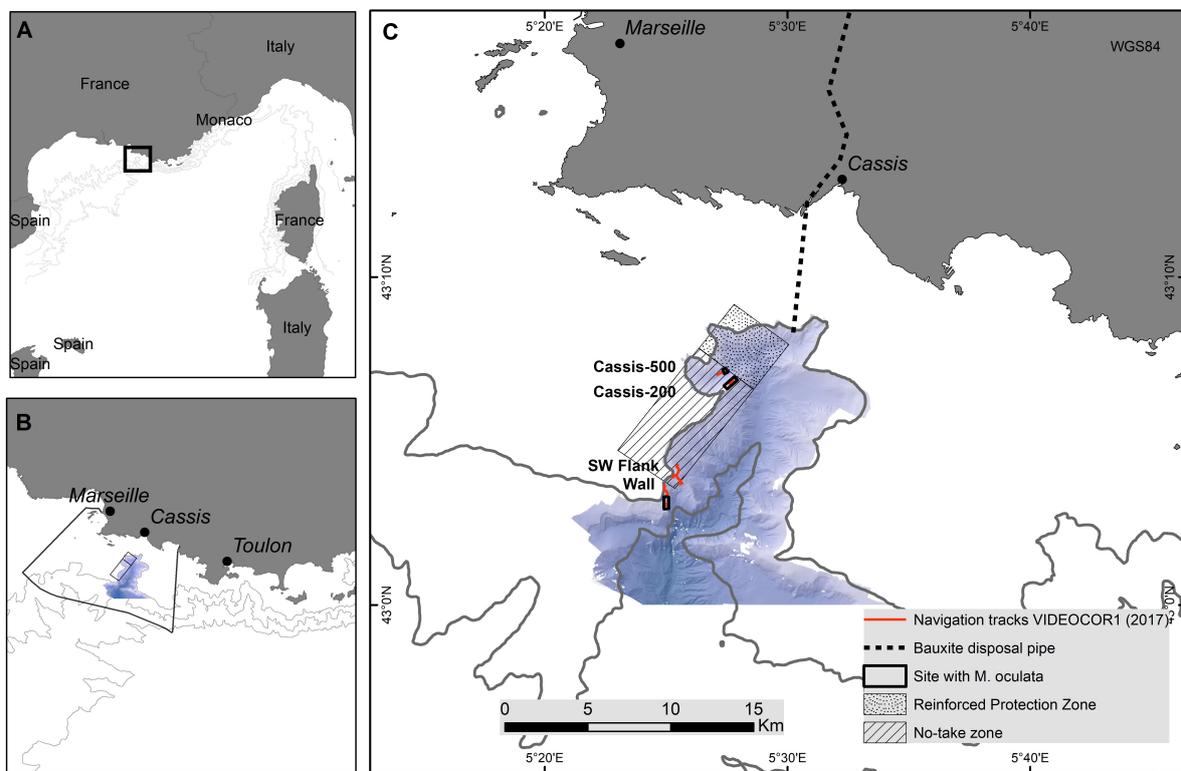
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119 2. Materials and methods

120 2.1 Study area

121 The head of the Cassidaigne canyon (Fig. 1) is located at a depth of 190 m and 8 km away from the
122 coast, in the Eastern part of the Gulf of Lion (NW Mediterranean Sea). In this region, the general
123 water circulation (the Northern Current, from 0 to 200 m depth) follows a westward direction along
124 the continental slope. In the presence of the prevailing Mistral wind (from north or north-west),
125 intense upwelling is generated in the canyon area (Millot, 1990; Alberola and Millot, 2003).

126



127

128 Figure 1: Location map of the study area. A: The global location is the eastern part of the Gulf of Lion; B: the Marine Protected Area
129 (MPA) of the "Parc National des Calanques"; C: The Cassidaigne canyon, the bauxite disposal pipe, the reinforced protection zone
130 and the no-take zone, and the four locations surveyed by the ROV Ariane during the VIDEOCORI cruise (2017).

131

132 In Cassidaigne canyon, *M. oculata* colonies were previously reported only on exposed bedrock and
133 on vertical walls of the western flank (Fabri et al., 2014; Fabri et al., 2017). Four locations were
134 explored during the Videocor1 cruise (2017): Cassis-200, Cassis-500, SW Flank and the Wall at depths
135 from 200 to 515 meters (Fabri et al., 2017).

136 Since 1967, the canyon has been impacted by the direct disposal of bauxite residues (red mud) from
137 two different aluminum factories. After 1988 the discharge of bauxite residues was reduced as one

138 factory stopped production. This disposal is known to have impacts on the local fauna of the canyon
139 (Dauvin, 2010; Fontanier et al., 2012; Fabri et al., 2014; Fontanier et al., 2015; Fabri et al., 2017).
140 Nevertheless, in an attempt to apply the protocols of the Barcelona Convention (1976), the French
141 government decided to stop the disposal of red mud in January 2016, by granting a six-year license
142 for the disposal of a fluid with no suspended matter and only chemicals, instead.

143 Additionally, to reduce impacts from fisheries, the Cassidaigne canyon is now included in the Marine
144 Protected Area (MPA) of the “Parc National des Calanques” (www.calanques-parcnational.fr –
145 Decree 2012-507). Thus, two zones were created on the western flank of the canyon: a (1)
146 ‘reinforced protection zone’, i.e. a no-take zone with the exception of local artisanal fisheries, and a
147 (2) ‘no-take zone’, where no fishing activity is allowed (Fabri et al., 2014).

148

149 2.2 Data acquisition

150 Scientific data were collected during the Videocor1 cruise in 2017. This cruise, on board the scientific
151 R/V L’Europe, was the first to deploy the new Hybrid-ROV Ariane (©Ifremer) (Raugel et al., 2019).
152 Data were obtained from a total of 8 dives (supp. mat. A), with the acquisition of 24 hours of video
153 (10h on Cassis-200, 4h on Cassis-500, 1h30 on the Wall, and 8h30 on SW Flank) and more than 8,000
154 photos of different rugged sites along the canyon. No *in situ* measurements were performed since a
155 laser measurement system was not available during the cruise. The four locations explored during
156 Videocor1 were chosen based on the results of the habitat suitability models previously built for the
157 canyon (Fabri et al., 2017; Bargain et al., 2018).

158 Ariane was equipped with an HD photo camera (Nikon D5200) and two video cameras, an HD pan
159 and tilt mounted camera with zoom capacity (Camera Principale) and an HD tilt mounted camera
160 (Camera Science). The navigation data were derived from the acoustic USBL GPS positioning system.

161 In this study, the 2-meter resolution bathymetry used was that collected locally at the same four sites
162 during the Bathycor 2 cruise (2016) by the Autonomous Underwater Vehicle (AUV) Idef’X (©Ifremer).
163 (DOI: [10.17600/16011300](https://doi.org/10.17600/16011300)).

164 2.3 Video and image processing

165 The videos recorded by Ariane were analyzed either with Matisse and 3D Metrics processing tools
166 (©Ifremer) or Adelie Video (©Ifremer) depending on the type of site (flat or with vertical parts) and
167 the ability to build a 3-dimension (3D) reconstruction or not.

168 2.3.1 Matisse and 3D Metrics processing tools (Photogrammetry)

169 The 3D models were built using the ‘Structure from motion’ technique, based on multi-view
170 geometry and image matching. A global view of the scene can be obtained by building a map using
171 multiple images. This technique uses the redundancy of successive images or within a video flux to
172 put all the pieces together and build the global view (Bruno et al., 2011; Arnaubec et al., 2015).

173 Limited visibility imposes taking images close to the sea-floor, with a small footprint for each image,
174 and requires artificial illumination of the scene as light is greatly attenuated in water. However,
175 artificial light cannot be optimized to light the scene uniformly, therefore deep-sea optical images

176 must be compensated for non-uniform illumination. Taking other disturbances such as color
177 dependent attenuation and marine snow into account are part of the preprocessing step. These
178 steps are followed by the construction of a sparse 3D cloud that is then densified. The dense point
179 cloud is used to create a surface (mesh), and then the original images are projected onto the mesh to
180 obtain the final textured model. In order to geo-reference these data and make measurements, the
181 absolute position and scale are obtained by merging the navigation system positioning and the
182 camera positions retrieved during the reconstruction process.

183 In this study, a total of 33 models were built, when possible based on HD photos collected regularly
184 by automatic acquisition during the cruise. When automatic acquisition had not been triggered,
185 video stills were used instead of photos if needed (supp. mat. A). The 3D Metrics processing tool was
186 used to view the models and measure population parameters.

187

188 2.3.2 Adelie Video for georeferencing species occurrences

189 The Ifremer post-processing software: Adelie Video© was used to analyze videos recorded along
190 linear dive transects where photogrammetry was not useful to map distribution because of the flat
191 shape of the seafloor, such as along the SW-Flank site. This processing allowed georeferencing
192 species occurrence observed on videos using the corresponding navigation file, and mapping the
193 distribution in a GIS (ArcMap ©ESRI).

194

195 2.4 Data analysis

196 2.4.1 Species distribution

197 The species considered for this study are not only the structure-forming scleractinian *M. oculata*, but
198 also other cold-water coral species providing 3D-habitat to other species: antipatharians *Leiopathes*
199 *glaberrima*, *Antipathella subpinnata* and *Antipathes dichotoma*; alcyonaceans *C. rubrum* (precious
200 red coral), gorgonians *Callogorgia verticillata* and *Viminella flagellum*.

201 The distribution of the species considered was obtained using either photogrammetry or video
202 analysis. Each colony was then manually plotted using 3D Metrics or Adelie Video, to obtain the
203 depth, latitude and longitude. In the second step, occurrence points were exported and mapped in
204 ArcMap©ESRI for spatial distribution, or processed in Excel©Microsoft for vertical distribution.
205 Species whose recognition/identification was not easy in the 3D models were plotted with the help
206 of video and HD photos to find their exact position in the reconstructions.

207

208 2.4.2 Surface and density

209 A 3D approach was used to calculate the surface and density occupied by species settled on rocky
210 outcrops with complex topography, i.e. *M. oculata* and *C. rubrum*, and a 2-dimension (2D) approach
211 was used for species seen along linear transects on smooth areas, i.e. *C. verticillata* and *V. flagellum*.

212 In the 3D approach, a maximum number of colonies of *M. oculata* and *C. rubrum* was plotted on the
213 3D models, giving the total number of colonies (nb) in each reconstruction. Colonies under 2 cm were
214 not plotted. The surface occupied by the colonies plotted was measured using the polygon function
215 of the 3D Metrics processing tool. This allowed calculating the density of species considering the
216 actual surface area of the bottom taking into account its 3D structure (nb colonies.m⁻²). The points
217 plotted with 3D Metrics could also be exported and mapped in plan view through GIS. Areas
218 occupied by colonies were then contoured by polygons and the area of each polygon was calculated
219 on a horizontal view. The number of colonies belonging to each polygon was known, which allowed
220 estimating the colony density projected onto a horizontal plane (nb colonies.m⁻²).

221 Where colonies were settled under the overhangs of the Wall, it was not possible to take photos
222 because the HD photo camera could not be directed upwards. Therefore, this site was reconstructed
223 using video images from the 'Camera Principale' located on the top of the ROV, which could be
224 directed upwards. The resulting 3D models had low resolution, allowing identification of only a
225 quarter of the individual colonies from the models. However, colonies could be counted using
226 snapshots of the video of the different parts of the huge overhangs.

227 In the 2D approach, species observed and settled along linear transects, i.e. *C. verticillata* and *V.*
228 *flagellum*, were plotted in a GIS using Adalie-Video, and a linear density was calculated (nb
229 colonies.m⁻¹). Sections of 20 m-long were calculated using the GIS tool, Adalie-Observation@Ifremer.
230 At Cassis-500, *C. verticillata* was also present but in smaller areas, and thus the sections used were
231 10 m in length.

232 No estimation of density and surface could be calculated for antipatharian species, given the
233 difficulty of reconstructing them in 3D, and the difficulty of counting, from video, the total number of
234 individuals forming forests and gardens.

235

236 2.4.3 Size

237 Colony heights were measured using 3D Metrics. Size measurements were performed on *M. oculata*,
238 *C. rubrum*, *C. verticillata* and *V. flagellum* populations, in attempts to measure as many colonies as
239 possible in each 3D model. No antipatharian size measurements could be performed since colonies
240 always move with the water flow and local currents, therefore they cannot be represented by 3D
241 reconstructions. Nevertheless, a qualitative description of antipatharian colony sizes was made
242 through video and photo analysis.

243

244 2.4.4 Diversity of accompanying species

245 The different species observed were identified to the lowest taxonomic level possible, using high
246 resolution digital photos and the book "Guide de la faune profonde de la mer Méditerranée" (Fourt
247 et al., 2017). Since no specimen was sampled, some species could not be identified (i.e. sponges), but
248 were morphologically distinguished. Abundances are given as presence and absence data based on
249 species observations at each location.

250

251 2.4.6 Statistical analysis

252 Statistical analysis was performed using XLSTAT 2019 for Excel (Addinsoft, 2019). Box-plots of the
253 vertical distribution of the different species were produced. A non-parametric Kruskal-Wallis test was
254 used to compare mean sizes of *M. oculata* colonies at three locations. Non-parametric Mann-
255 Whitney tests were computed to compare mean sizes of *C. verticillata* colonies at two locations, and
256 the same was applied for *V. flagellum*. Gaussian mixture models were run to identify the different
257 size groups at each location for the same species.

258 A hierarchical clustering of the four locations was performed based on the presence or absence of
259 species, including habitat-forming cnidarians. The Sorensen similarity coefficient was used to build
260 the similarity matrix because this coefficient is based on binary data, and it emphasizes the presence
261 of species (absence may be due to weak sampling effort).

262

263 2.5 Degradation and seafloor integrity

264 Areas where coral rubble was found were contoured and measured with 3D Metrics, allowing their
265 estimation in 3D, after which they were transferred to a GIS for positioning and measurement in 2D.

266 Marine litter observations were georeferenced and mapped in a GIS using Adelie Video. The
267 observations were classified following the guidelines of the MSFD Technical Subgroup on Marine
268 Litter for the Mediterranean Sea (Annex 5.1.) (European Commission, 2013). The three classes
269 considered were: Plastic, Metal and Glass/Ceramics. Linear density was calculated using the total
270 length of navigation tracks on each site.

271 The presence or absence of waste disposal (bauxite residues) at each location was recorded.

272

273 2.6 Temporal changes

274 A qualitative analysis of the temporal changes of *M. oculata* colonies distribution and sizes of the
275 Cassis-200 location was performed. 3D models and images taken in 2017 were compared with video
276 data recorded previously in 2010 (ESSROV 2010) and 2013 (ESSNAUT 2013) (Fabri et al., 2014; Fabri
277 et al., 2017). The videos recorded in 2010 and 2013 did not allow to build 3D-models because they
278 were not recorded for that purpose: the video coverage and image quality were not sufficient, and
279 the submarine navigations were not precise enough. But the 3D-models of 2017 made it possible to
280 precisely identify the scene observed on the videos recorded in 2010 and 2013 and to recognize
281 which knoll and which face of the knoll had been observed at the time. It was therefore possible to
282 choose the orientation of the 3D-model and to compare the same scene at 5 and 7 years apart.

283

284 2.7 Validation of the habitat suitability maps built previously

285 The *in situ* cold-water coral population distribution observed in 2017 was compared to the
286 theoretical distribution obtained from the habitat suitability models (Fabri et al., 2017; Bargain et al.,
287 2018). For this we overlapped both maps in GIS and compared the extent of the distributions. We
288 also gathered all the navigation tracks we were aware of in Cassidaigne canyon (MARUM_2009,
289 MEDSEACAN_2009, ESSROV_2010, ESSNAUT_2013, ESSTECHROV_2017, ESSTECHROVGEN_2017, VIDECCORI_2017 cruises) in order
290 to map absence of *M. oculata* colonies.

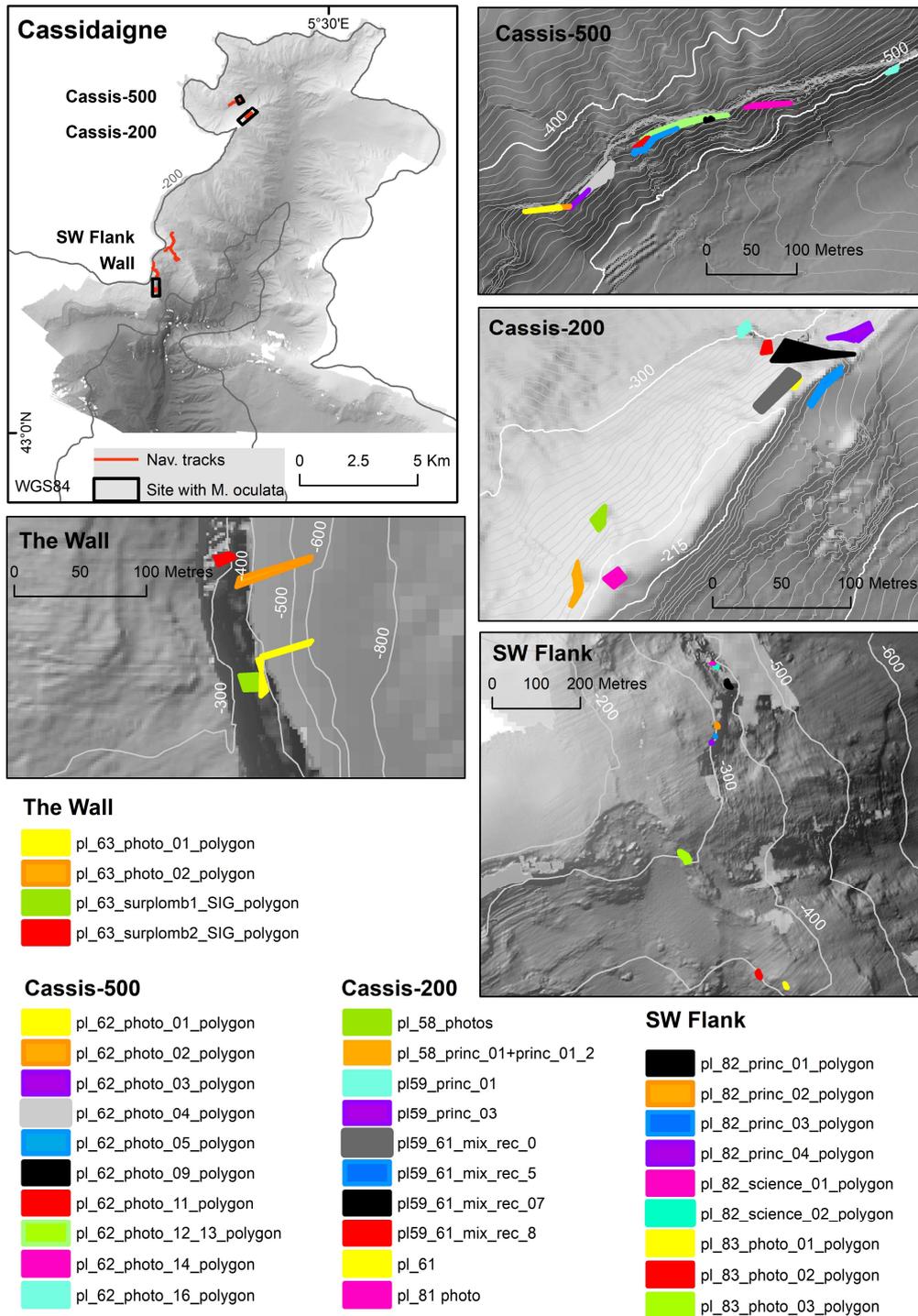
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292 3. Results

293 3.1 Spatial distribution of the 3D models

294 A total of 33 models were built in 3D; 10 at Cassis-500, 10 at Cassis-200, 9 on the SW Flank and 4 at
295 the Wall (Fig. 2).

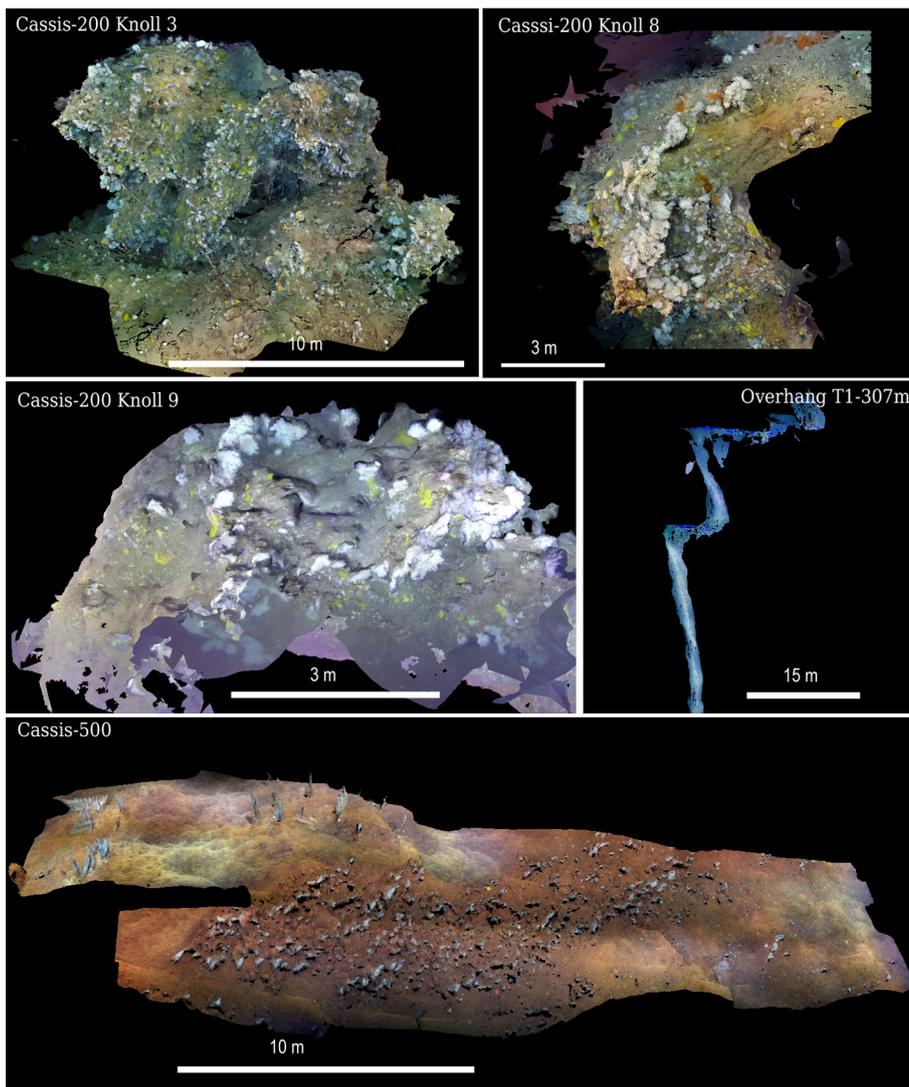
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297

298 Figure 2: Spatial distribution of the 33 models built in 3D at Cassidaigne canyon, specifically at Cassis-500 (10 models), Cassis-
 299 200 (10 models), SW Flank (9 models) and the Wall (4 models). High-resolution bathymetry at each site was collected during the
 300 Bathycor Cruise (2016). See supp.mat. B for an exhaustive view of 3D-models.

301



302

303 Figure 3: Examples of 3D-models of selected areas within Cassidaigne canyon. See supp.mat B for an exhaustive view of all the 3D-
304 models constructed.

305

306

307 3.2 Scleractinian *Madrepora oculata*

308 3.2.1 Spatial and vertical distribution of *M. oculata*

309 The presence of *M. oculata* was observed in high abundance at three locations: Cassis-500, Cassis-
310 200 and the Wall; but only three colonies were observed at the SW Flank. The vertical distribution
311 ranged from depths of 203 to 484 m.

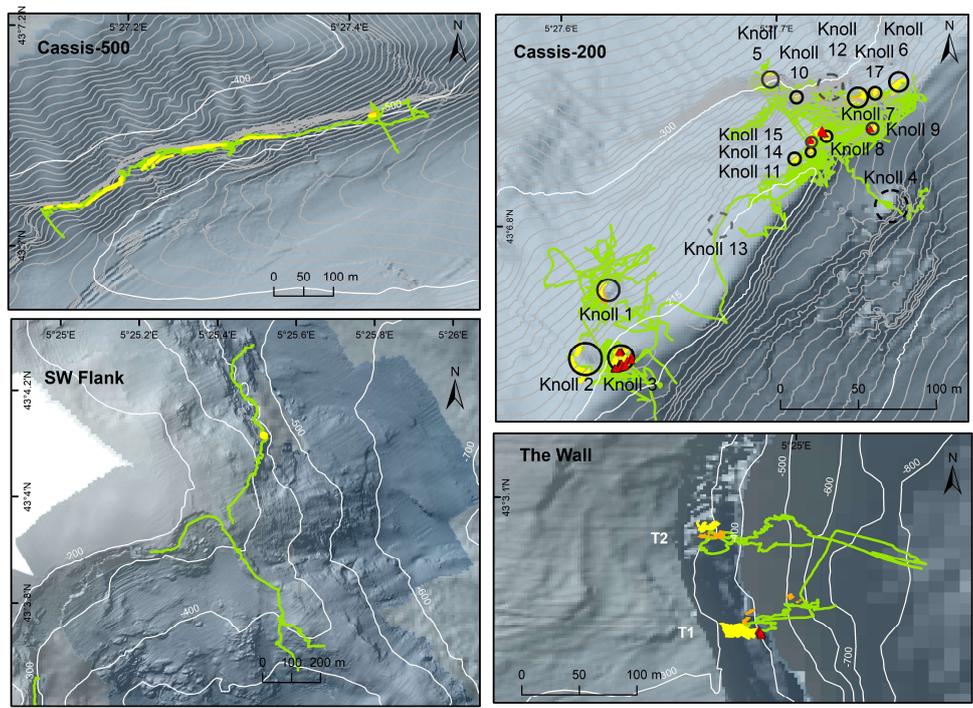
312 At the head of Cassidaigne canyon, at Cassis-200, 1294 colonies of *M. oculata* were counted from the
313 3D reconstructions. Colonies have settled on particular rocky elevations of the area that we called
314 knolls (Fig. 3). We identified a total of 13 knolls where the presence of *M. oculata* was mapped (Fig.
315 4). Only a few isolated colonies were found in areas in between the knolls. Video observations also
316 showed the presence of colonies on Knoll 13, however no 3D model could be built for this knoll
317 because the video coverage we had was not enough. One knoll was also found with no *M. oculata*
318 (Knoll 4 on the eastern side of Cassis-200), and one knoll (Knoll 12 on the northern side of Cassis-200)
319 was not explored but was identified with the help of bathymetry. This means that at Cassis-200, *M.*
320 *oculata* was present on a total of 14 knolls, different in size and structure (supp. mat.B). Colonies of
321 *M. oculata* occupied a 3-dimensional spread distribution on each knoll, presenting a south-westerly
322 orientation, facing the dominant currents of the area. Here, the mean depth for the presence of the
323 species was 217 ± 14 m. Because of the patchy distribution of *M. oculata* on the knolls, the vertical
324 distribution on each knoll is presented (Fig. 4). Knoll 3 is the shallowest one, with a mean depth of
325 206 ± 2 m, ranging from 203 to 211 m. Knoll 5 is the deepest one, with a mean depth of 271 ± 2 m,
326 ranging from 269 to 276 m.

327 At Cassis-500, also located at the canyon's head, a total of 3492 colonies of *M. oculata* were plotted.
328 The linear exploration performed at this site showed that colonies have settled on the vertical part of
329 a rocky ledge, occupying the whole structure explored (Fig. 4). The entire area was covered by red
330 sediment, and it is important to note that the basal part of tall colonies was covered by red mud.
331 Very small colonies were also seen under overhangs, with no red mud. The deepest observation for
332 the presence of *M. oculata* in the canyon was registered at this site, where colonies were found at a
333 mean depth of 432 ± 10 m, ranging from 416 m to 484 m (Fig.4).

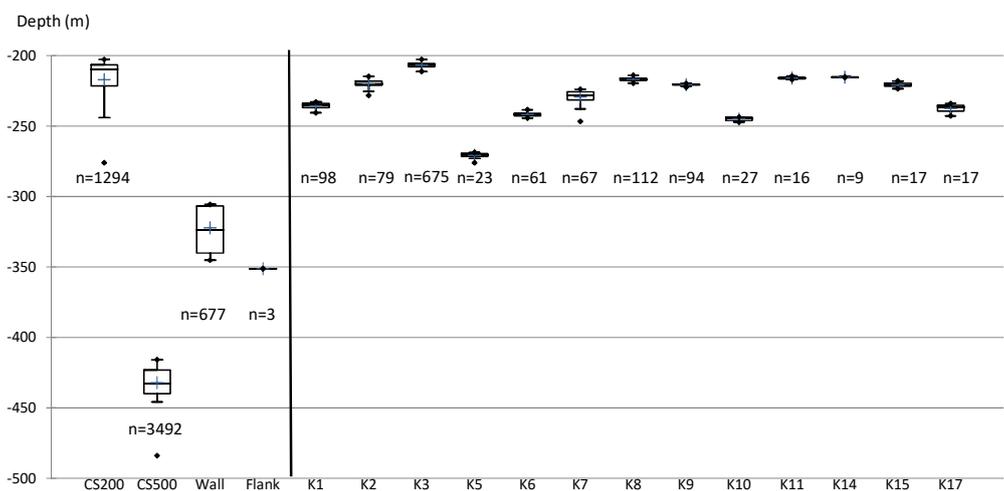
334 At SW Flank, located at the exit of Cassidaigne canyon, only three colonies of *M. oculata* were
335 observed at the top of a rocky structure, at a depth of 351 m (Fig. 4).

336 The Wall, a long escarpment also located at the exit of the canyon, was surveyed through two
337 vertical transects of approximately 300 m in height (T1 and T2) (Fig. 5; supp.mat. B). This location was
338 very difficult to explore due to the technical challenge of navigating in this environment below an
339 overhang, something that had never been done before since the tilt cameras of ROVs generally
340 cannot be directed upwards. In this study, it was possible to build 3D models based on video since
341 the part under the overhang was filmed when the H-ROV Ariane accidentally bumped the ceiling;
342 however, no photos could be taken. These 3D models allowed showing the complexity of this site
343 where colonies have settled. Along the first transect (T1), two overhangs at different depths
344 (Overhang T1-307 m and Overhang T1-320 m) were found with the presence of *M. oculata* (Fig. 5),
345 whereas a third overhang was found (Overhang T2-340m) along transect 2 (T2), at a deeper part of
346 the wall (Fig. 5). At the Wall, *M. oculata* exclusively colonized the area under the overhangs, with
347 colonies having an upside-down orientation. Around 3000 small colonies were counted under the

348 three overhangs (only 677 colonies were plotted on the 3D models) and they occupied the entire
 349 area. At the Wall (n=677), the mean depth where colonies could be found was 322 ± 15 m. Since
 350 colonies had settled under the overhangs, their vertical distribution corresponded to the depth
 351 ranges of the latter. Therefore, the shallowest overhang was found between a depth of 305 and 307
 352 m, whereas, just below, the second overhang was found at depths from 323 to 325 m. The deepest
 353 occurrence of *M. oculata* at the location was recorded under the third overhang, with a depth
 354 distribution that ranged from 339 to 345 m.



355

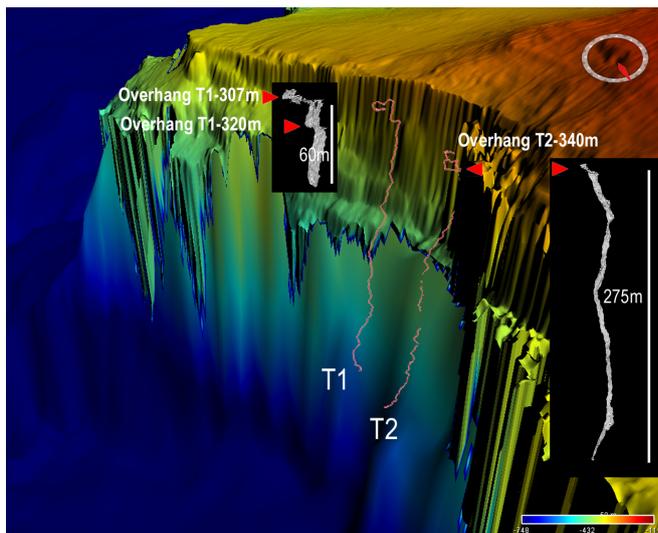


356

357 Figure 4: Spatial and vertical distribution of *M. oculata* and *C. rubrum*. Maps show the distribution of *M. oculata* colonies in yellow,
 358 rubble in orange, *C. rubrum* colonies in red. Dive navigation tracks are represented as green lines. The graph shows depth-range
 359 box-plots of *M. oculata* in Cassidaigne canyon, for each location (Cassis-200, Cassis-500, the Wall and SW Flank) and the detail for
 360 each knoll at Cassis-200 (K1 to K17). The number of plotted colonies is indicated on the graph (n) for each location and knoll.

361

362



363

364 Figure 5: 3D-Representation of the Wall, with the two vertical navigation tracks T1 and T2 and the position of the 3 overhangs
365 located at different depths, where 3D models were made.

366

367 3.2.2 Surface coverage and colony density of *M. oculata*

368 At Cassis-200, the total surface occupied by colonies was 409 m² on the 3D models, and was half as
369 much when transferred into GIS and seen in plan view (Table I). Colonies of *M. oculata* occupied
370 different surfaces and presented different densities depending on the knoll. The largest surface
371 occupied by colonies was seen for Knolls 2 and 3 (Table I). The highest density was found on Knoll 8,
372 with 12 colonies.m⁻² calculated on the 3D model. Overall, at Cassis-200, the mean density of *M.*
373 *oculata* calculated in 3D was 4 colonies.m⁻².

374 At Cassis-500, the surface occupied by *M. oculata* was 924 m² on the 3D models, and less than half
375 when transferred into a GIS. The mean density was 4 colonies.m⁻² in 3D.

376 Under the overhangs of the Wall, colonies occupied a total surface area of 317 m² measured on the
377 3D models, which was almost the same when transferred into a GIS because the colonies occupied a
378 horizontal plane under the overhangs, allowing good representation in 2D. At the Wall, the mean
379 density was 9 colonies.m⁻², the highest mean density of *M. oculata* in Cassidaigne canyon.

380 At SW Flank, no estimation of surface/density was calculated since only three colonies were
381 observed.

382

383

384 **Table I.** Surface coverage and colony density, in 2D and in 3D, for *M. oculata* at Cassis-200, Cassis-500 and the Wall. For the
 385 location Cassis-200, the surface and density are presented for each knoll. Total surface and mean densities were calculated for
 386 the 3 locations.
 387

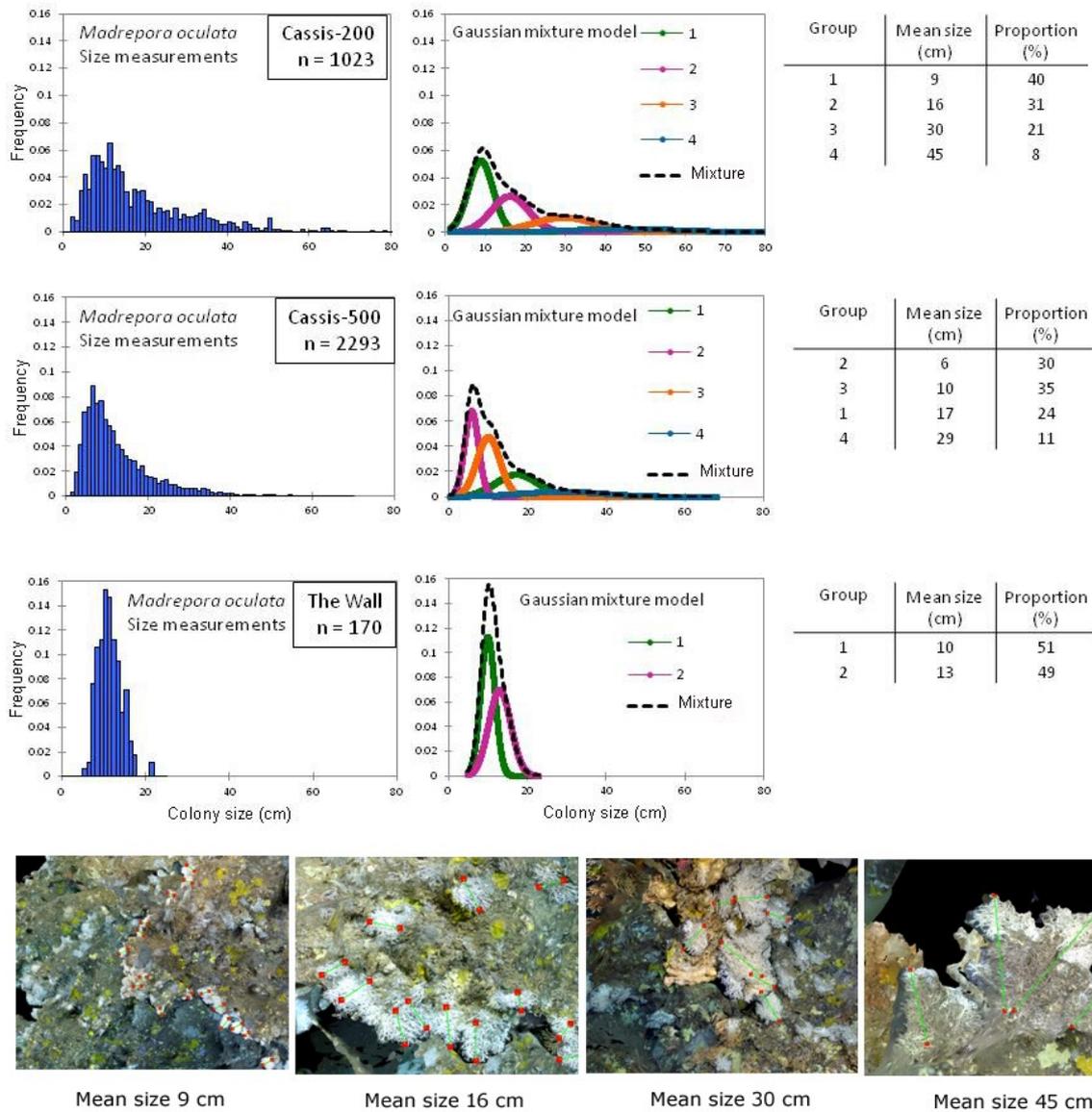
Location	Knoll	Surface 2D (m ²)	Surface 3D (m ²)	n	Density 2D (nb colonies.m ⁻²)	Density 3D (nb colonies.m ⁻²)
Cassis-200	Knoll 1	19	38	98	5	3
Cassis-200	Knoll 2	44	104	79	2	1
Cassis-200	Knoll 3	49	101	98	14	7
Cassis-200	Knoll 5	8	42	23	3	1
Cassis-200	Knoll 6	14	32	60	4	2
Cassis-200	Knoll 7	17	30	67	4	2
Cassis-200	Knoll 8	8	9	98	14	12
Cassis-200	Knoll 9	11	14	94	8	7
Cassis-200	Knoll 10	6	6	27	5	5
Cassis-200	Knoll 11	5	5	16	3	3
Cassis-200	Knoll 14	2	2	9	4	4
Cassis-200	Knoll 15	9	11	17	2	2
Cassis-200	Knoll 17	2	3	17	8	7
Cassis-200		200	409	1388	6	4
Cassis-500		409	924	3492	9	4
SW Flank				2		
Wall		322	317	677	9	9

388

389

390

391 3.2.3 Size structure of *M. oculata* populations



392 Mean size 9 cm Mean size 16 cm Mean size 30 cm Mean size 45 cm

393 Figure 6: Histograms of size distribution for *M. oculata* populations at Cassis-200, Cassis-500 and the Wall, and the Gaussian
 394 mixture models implemented to identify size groups. Illustration of size measurement on the 3D models for *M. oculata* at Cassis-
 395 200 (green line between two red dots).

396 At Cassis-200, the mean height of *M. oculata* was 18 ± 12 cm ($n=1023$). Size measurements show
 397 colonies as small as 2 cm (smaller colonies were not plotted), while the tallest colony measured 78
 398 cm. The Gaussian mixture model (Fig. 6) shows a size structure of colonies composed of 4 distinct
 399 size groups. At Cassis-200, 40% of the colonies were included in the group with the smallest mean
 400 size of 9 cm (group 1), whereas 8% of the colonies were included in the largest size group of 45 cm
 401 (group 4). At this location, 52% of the colonies were therefore included in intermediate size groups,
 402 with 31% of the colonies included in group 2, with a mean size of 16 cm, and 21% included in group
 403 3, with a mean size of 30 cm.

404 At Cassis-500, the size measurements (n=2293) show that the smallest colony measured 1 cm and
405 the tallest 67 cm, the mean size was 12±8 cm. The Gaussian mixture model for Cassis-500 (Fig. 6)
406 shows that colonies of *M. oculata* belong to 4 different size groups. At Cassis-500, there are
407 therefore 2 groups (groups 2 and 3) in which the smallest colonies were included, with a mean size of
408 6 and 10 cm, which accounted for 30% and 35% of the colonies, respectively. The model also shows
409 that 24% of colonies of *M. oculata* belong to a group in which the mean size was 17 cm and 11% of
410 the colonies fit in the group in which the mean size observed was 29 cm. Therefore, the results
411 presented by the model indicated that at Cassis-500 a larger part of the population is included in the
412 smaller size groups (with 65% in group 2 and 3) than in the taller size groups (35% in group 1 and 4).

413 At the Wall, the size measurements (n=170) show colonies with a minimum size of 6 cm and a
414 maximum of 22 cm, and with a mean size of 11±3 cm. The low number of size measurements is due
415 to the low image resolution of the 3D reconstructions made from video for this location. However,
416 the Gaussian mixture model analysis could be performed (Fig. 6). The model suggests that the
417 colonies of *M. oculata* belong to two size groups, with 51% of the colonies included in a group with a
418 mean size of 10 cm and the other 49% included in a group with a mean size of 13 cm.

419 The Kruskal-Wallis test performed to compare the mean size of *M. oculata* populations on the three
420 locations shows that they are all different. The mean size of *M. oculata* in Cassidaigne canyon was
421 significantly different between Cassis-200 (18±12 cm), Cassis-500 (12±8 cm) and the Wall (11±3 cm)
422 (p-value ≤ 0.05). The global mean size of *M. oculata* in the Cassidaigne canyon was 14±10 cm.

423

424 3.3 Forests and gardens of antipatharians

425 3.3.1 Spatial and vertical distribution of antipatharians

426 Three antipatharian species, *A. dichotoma*, *A. subpinnata* and *L. glaberrima*, were observed at
427 Cassidaigne canyon. The species *A. subpinnata* was found only at Cassis-200. The two species, *A.*
428 *dichotoma* and *L. glaberrima*, were found at both Cassis-200 and at SW Flank, while at Cassis-500
429 only one antipatharian colony was observed, *A. dichotoma*, and no antipatharian was observed at
430 the Wall (Fig. 7).

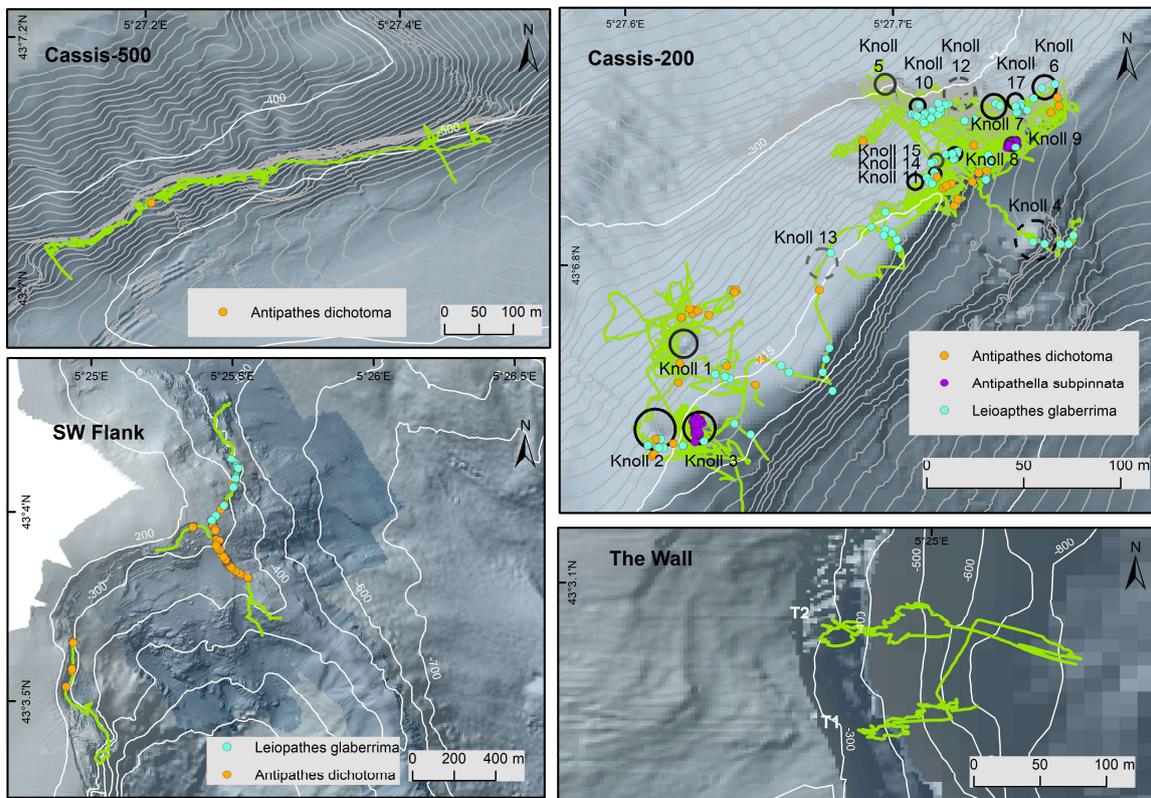
431 The distribution of *A. subpinnata* was limited to Cassis-200 at a mean depth of 210±8 m (Fig. 7),
432 where two dense forests of the species were seen at the tops of knoll 3 and knoll 9.

433 Gardens of *L. glaberrima*, on the contrary, presented a much more widespread distribution, with the
434 presence of the species at both Cassis-200 and the SW Flank, with a depth distribution ranging from
435 200 to 376 m (Fig. 7). At Cassis-200 (n=124), *L. glaberrima* was observed settled on the side of knolls,
436 as well as between knolls as part of coral gardens at a mean depth of 225±14 m. At the SW Flank
437 (n=21), colonies of this species were observed only in the northern part of the site where *L.*
438 *glaberrima* occupied the sides of a rocky structure at a mean depth of 322±37 m and dense colonies
439 of *V. flagellum* were observed at the top. Colonies of *L. glaberrima* could be observed either as white
440 or orange. Both morphs could be observed with no link with the size of the colony or its location.

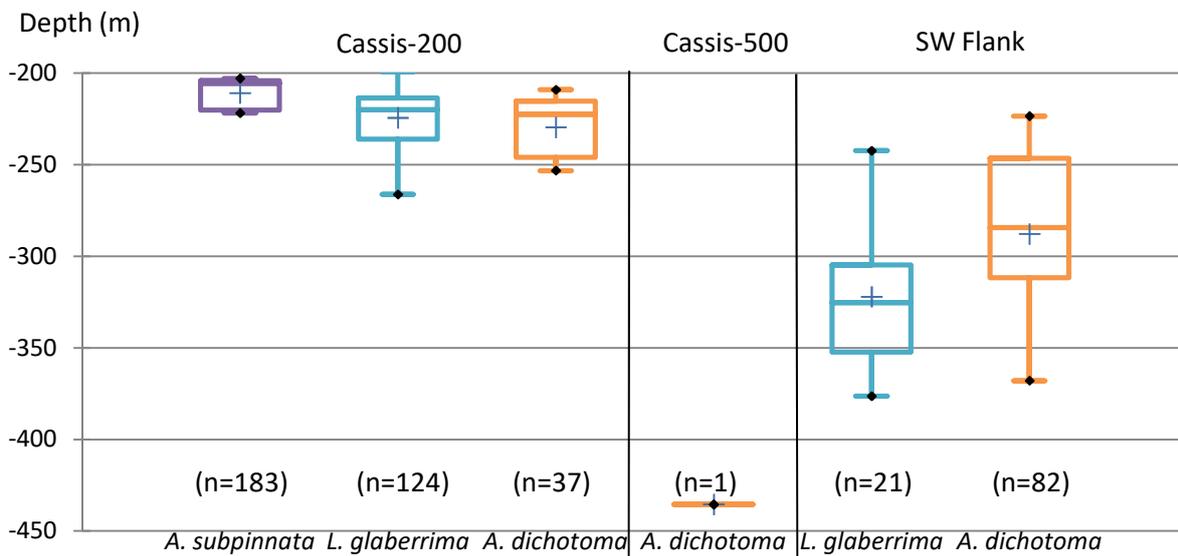
441 Colonies of *A. dichotoma* was always observed as isolated colonies from 209 to 435 m (Fig. 7) at
442 Cassis-200, SW Flank and Cassis-500.

443 Colonies of *L. glaberrima* and *A. dichotoma* presented a more extended vertical distribution than the
 444 forest forming species *A. subpinnata* (Fig. 7).

445



446



447

448 Figure 7: Spatial and vertical distribution of antipatharian species in the Cassidaigne canyon. Maps show the distribution of forests
 449 of *A. subpinnata* (violin), gardens of *L. glaberrima* (blue), and isolated colonies of *A. dichotoma* (orange). The graph shows depth
 450 range box-plots of the three species observed at each location during the Videocorl cruise. No antipatharian was found at the Wall.
 451 The number of plotted colonies is indicated on the graph (n) for each species.

452

453

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456 3.3.2 Size measurements of antipatharian populations

457 It was not possible to identify antipatharians on the 3D models due to their movement with the
458 water flow. Hence, it was not possible to perform any size measurements for this group of corals.
459 Nevertheless, a qualitative description of the size structure of the populations was possible based on
460 video and photo analysis (Fig. 8).

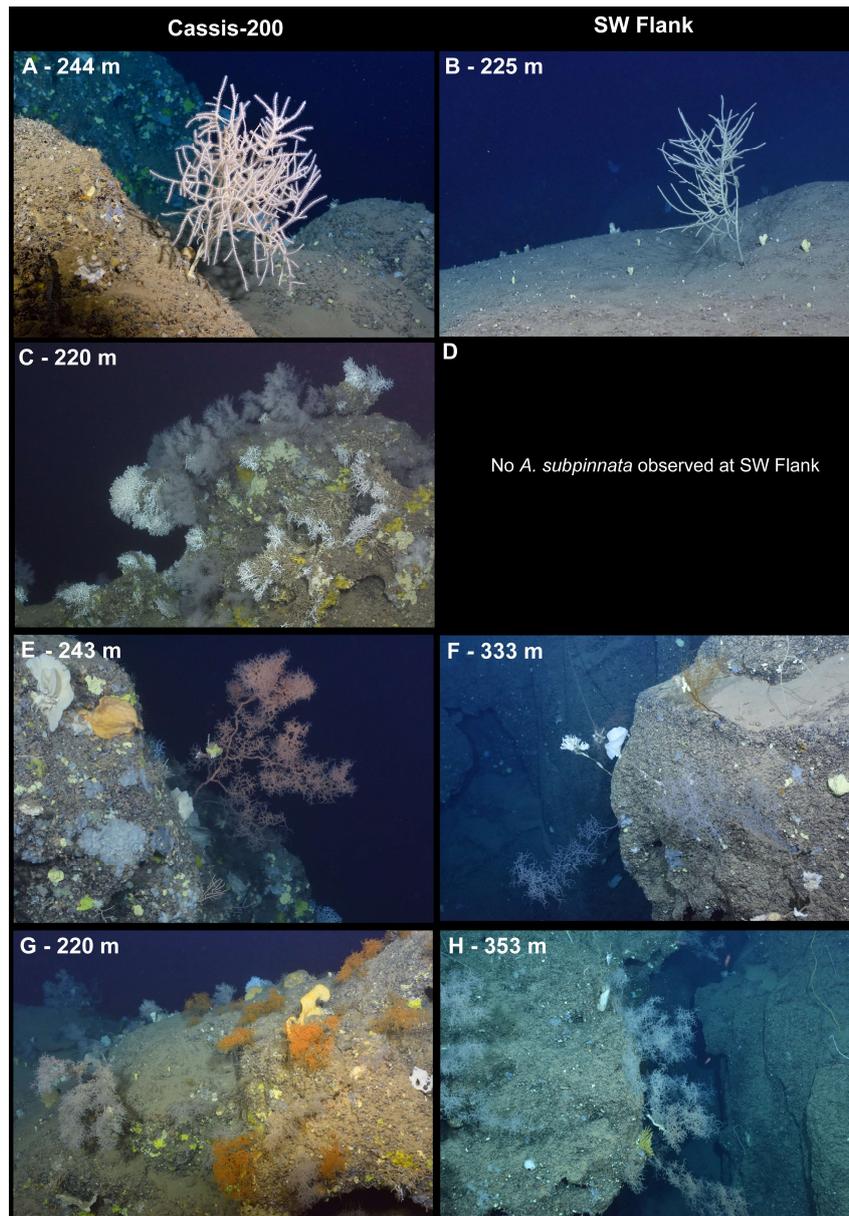
461 At both locations, Cassis-200 and SW Flank, size seemed to be similar among colonies of *A.*
462 *dichotoma* and no small colonies could be observed (Fig. 8A & 8B).

463 At Cassis-200, forests of *A. subpinnata* included large colonies occupying the top of two knolls but
464 small colonies had also settled on the sides of knoll 3 (Fig. 8C), whereas no *A. subpinnata* was
465 observed at SW Flank (Fig. 8D).

466 Colonies of *L. glaberrima* of different sizes were observed at both Cassis-200 and SW Flank. Large
467 colonies of the species could be observed, in particular occupying the sides of knolls at Cassis-200
468 (Fig. 9E) and vertical walls at SW Flank (Fig. 8F). However, small and intermediate size colonies were
469 also observed in high density, forming coral gardens (Fig. 8G). A mix of small and large colonies was
470 also found at the site SW Flank (Fig. 8H).

471

472



473

474 Figure 8: Qualitative size description of antipatharians from Cassis-200 and SW Flank sites. A and B: Large isolated colonies of *A.*
 475 *dichotoma*; C and D: Forest of large colonies of *A. subpinnata* with small colonies around; E and F: Large colonies of *L. glaberrima*;
 476 G: Gardens of *L. glaberrima* with colonies of different sizes; H: Colonies of *L. glaberrima* of different sizes.

477

478

479 3.4 The Octocorallia Scleraxonia *Corallium rubrum*

480 3.4.1 Spatial and vertical distribution of *C. rubrum*

481 The red coral *C. rubrum* was found at two locations: Cassis-200 and the Wall (Fig. 4).

482 At Cassis-200 (n=123), the presence of *C. rubrum* could be observed on Knolls 3, 8, 9 and 15 (Fig. 4),
483 which are amongst the shallowest. Here, *C. rubrum* was found at a mean depth of 209 ± 5 m, with a
484 vertical distribution ranging from 203 to 220 m depth (Fig. 9). The use of videos and HD photos was
485 essential to locate small patches of colonies, always underhanging the knolls. On Knoll 3, 94 colonies
486 of *C. rubrum* were counted and plotted, while only 16 colonies were counted on Knoll 8, 9 on Knoll
487 15 and 4 on Knoll 9. Colonies of *C. rubrum* presented a wide distribution along the entire rocky
488 structure of Knoll 3, while on other knolls colonies were concentrated in a small area (Fig. 4).

489 At the Wall, a total of 42 colonies of *C. rubrum* were counted at a mean depth of 306 ± 1 m (Fig. 9).
490 Colonies were found to have settled all along the outer edge of the overhang T1-307m. The length of
491 the explored edge was around 10m long. Since the occurrence of the species was limited to the outer
492 edge of the overhang, the observations of red coral varied only between 305 and 307 m depth.

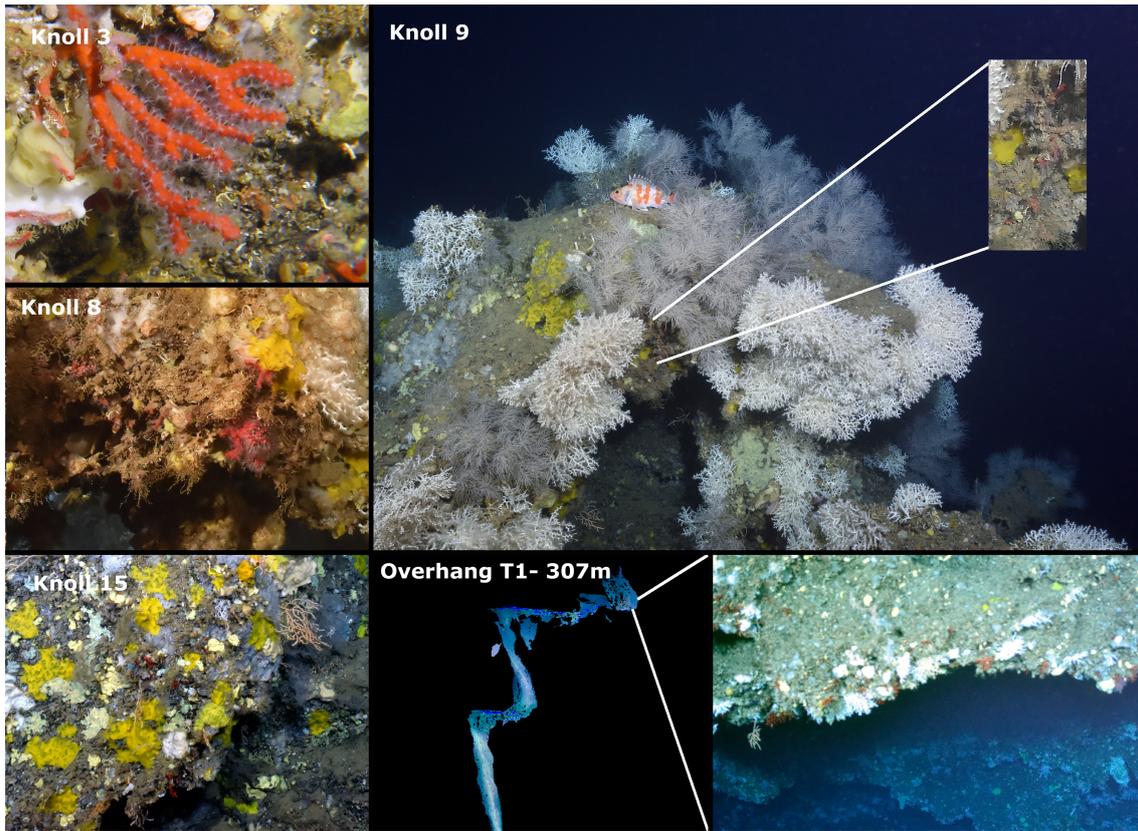
493

494 3.4.2 Surface coverage and colony density of *C. rubrum*

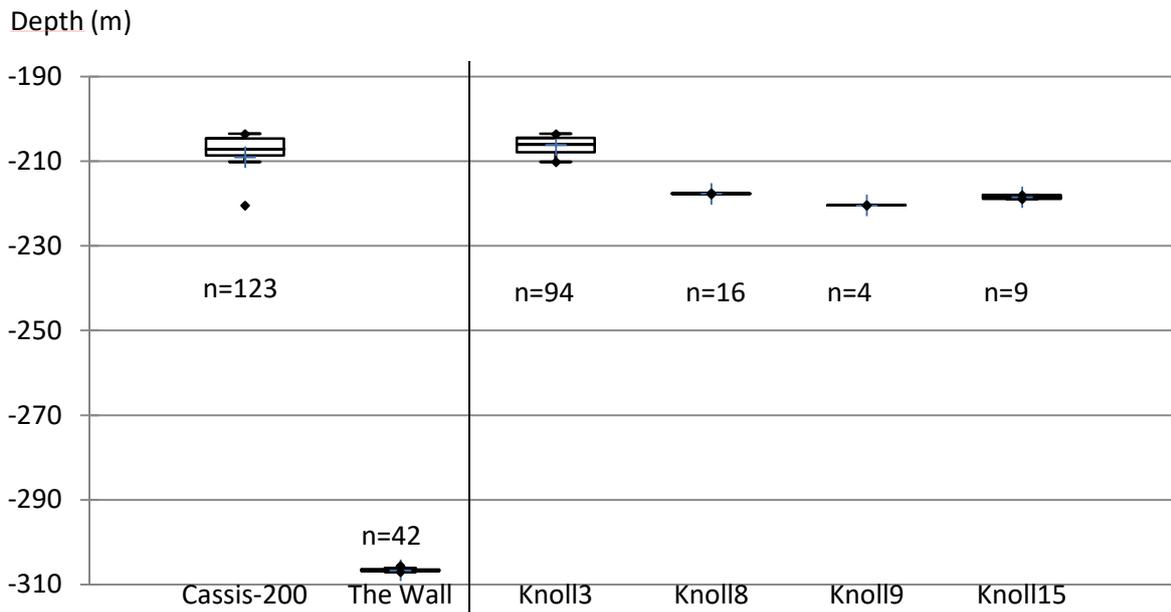
495 At Cassis-200, *C. rubrum* colonies were counted on the 3D models when possible, and this was
496 complemented by counting small colonies on HD photos. The total surface area measured on the 3D
497 models was 11 m^2 and we counted 123 colonies, which gives a density of $11\text{ colonies.m}^{-2}$. Colonies of
498 *C. rubrum* covered the largest surface (10 m^2) on Knoll 3, and showed the highest density on Knoll 8,
499 where the colonies were all gathered in an area smaller than 1 m^2 .

500 At the Wall, colonies of *C. rubrum* settled on the horizontal edge of the shallowest overhang. We
501 counted 42 colonies on a surface area of 2 m^2 , corresponding to a density of $20\text{ colonies.m}^{-2}$. Since
502 the colonies were all aligned along the 10 m long edge, a linear density was calculated,
503 corresponding to 4 colonies.m^{-1} .

504



505



506

507 Figure 9. Spatial and vertical distribution of *C. rubrum* in Cassidaigne canyon for each location. Photos show *C. rubrum* in
 508 Cassidaigne canyon, at Cassis-200 and the Wall, with close-ups on Knolls 3, 8, 9, 15 and on the ledge of Overhang T1-307m. The
 509 graph shows depth-range box-plots of *C. rubrum* in Cassidaigne canyon for each location. The number of plotted colonies is
 510 indicated on the graph (n) for each location and knoll.

511

512

513 3.4.3 Size measurements of the *C. rubrum* population

514 At Cassis-200 on Knoll 3, several size measurements of *C. rubrum* colonies were performed (n=26) on
515 the 3D reconstructions, showing a mean size of 6 ± 1 cm, with the smallest colony measuring 3 cm and
516 the tallest one measuring 16 cm (Fig. 9).

517 No measurement could be performed on Knolls 8, 9, and 15 because the colonies were too small, or
518 at the Wall, because the 3D models were made with video snapshots, therefore the resolution was
519 not good enough to measure small colonies, although they could be seen on the HD images (Fig. 9).

520

521 3.5 Gorgonians *Callogorgia verticillata* and *Viminella flagellum*

522 3.5.1 Spatial and vertical distribution of gorgonian populations

523 Many tall colonies of both gorgonians species were observed at SW Flank while only a few isolated
524 colonies were observed at Cassis-200, and only one species (*C. verticillata*) was observed at Cassis-
525 500 (Fig. 10). Neither of the two species was observed at the Wall.

526 At SW Flank, dense and tall colonies of *C. verticillata* (n=84) had settled on patches of rocky substrate
527 covered by a thin layer of sediment at a mean depth of 371 ± 19 m (Fig. 10). At this location, dense
528 aggregations of *V. flagellum* (n=234) were also found at a mean depth of 297 ± 45 m (Fig. 10). Linear
529 transects carried out along the edge of the continental shelf have shown its presence almost
530 everywhere. White and yellow-colored *V. flagellum* colonies formed vast forests, settled on rocky
531 substrate.

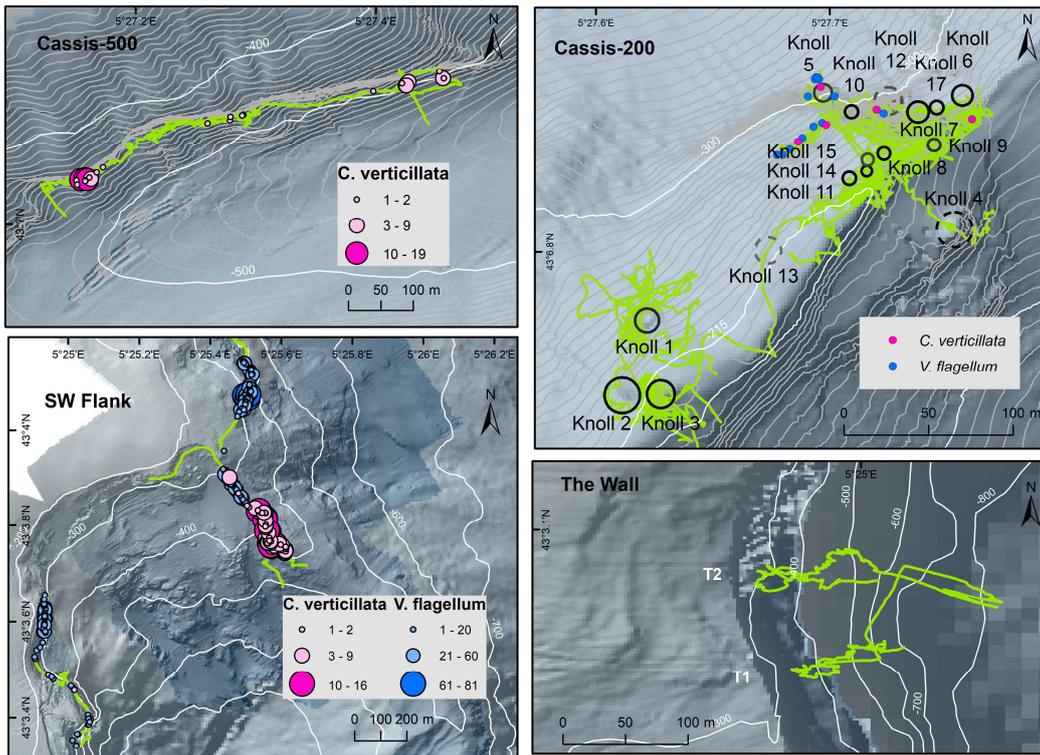
532 At Cassis-200, only a few gorgonian colonies were seen at knoll 5 and around knolls in the deepest
533 area. Cassis-200 was the shallowest location where *C. verticillata* (n=6) was observed at a mean
534 depth of 254 ± 17 m, together with *V. flagellum* at a mean depth of 267 ± 17 m (Fig. 10).

535 At Cassis-500, tall colonies of *C. verticillata* have settled on the top of the ledge previously described,
536 but no *V. flagellum* were observed. At this site, *C. verticillata* (n=57) reached a deeper part of the
537 canyon, at a mean depth of 450 ± 18 m (Fig. 10). The depth range of *C. verticillata* at Cassis-500 varied
538 between 424 and 483 m, being the widest depth range of all the locations, with the highest density
539 recorded at a depth around 430 m.

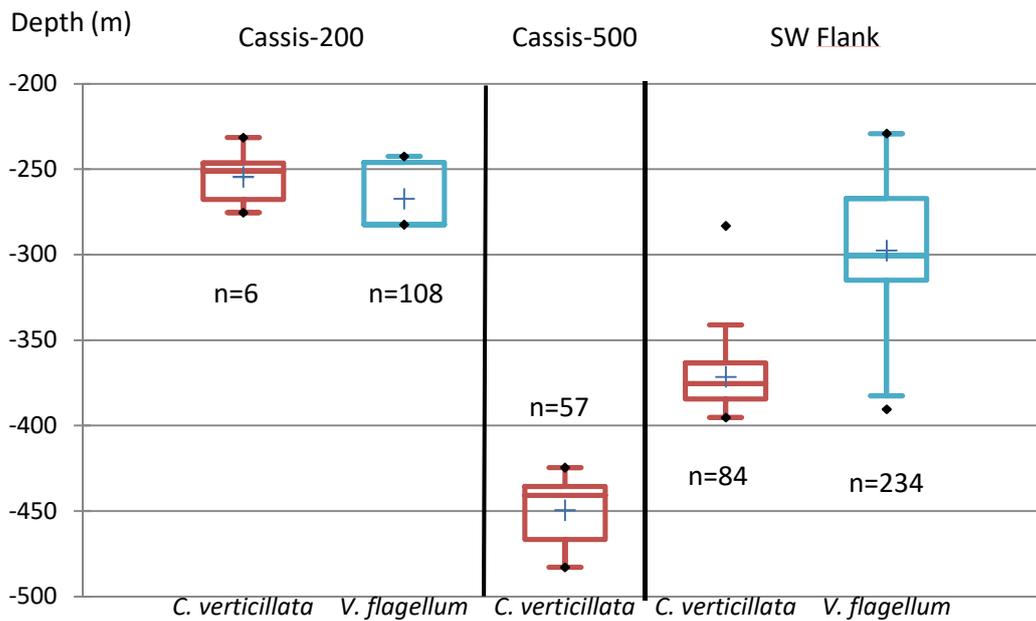
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545 Figure. 10. Spatial and vertical distribution of the two gorgonian species (*C. verticillata* in red and *V. flagellum* in blue) in
 546 Cassidaigne canyon. The number of plotted colonies is indicated on the graph (n) for each species.

547

548

549

550 3.5.2 Surface coverage and density of gorgonian populations

551 At SW Flank, colonies of *C. verticillata* (437 colonies) were irregularly distributed along a linear
552 transect of 1.2 km. The maximum density estimated was 47 colonies observed along a 20 meter long
553 video transect, corresponding to 2 colonies.m⁻². Many more colonies of *V. flagellum* were counted,
554 approximately 3870 colonies irregularly distributed along a linear transect of around 2.2 km in
555 length. The highest density estimated for *V. flagellum* at SW Flank was 316 colonies along a 20 meter
556 long video transect, corresponding to 16 colonies.m⁻¹.

557 At Cassis-200 on Knoll 5, 24 colonies of *V. flagellum* could be plotted, covering a surface of 7 m²,
558 corresponding to a density of 4 colonies.m⁻², while only 2 small colonies of *C. verticillata* were seen
559 on the same knoll.

560 At Cassis-500 a total of 104 colonies of *C. verticillata* were counted on videos at the beginning,
561 middle and end of the linear transect of 440 m length. The highest density measured was 19 colonies
562 along two 10-meter video transects, corresponding to 2 colonies.m⁻¹.

563

564 3.5.3 Size structure of gorgonian populations

565 Gorgonian *C. verticillata*

566 The global mean size of *C. verticillata* in the Cassidaigne canyon was 69 ± 29 cm.

567 At the SW Flank (n=37), the mean size was 80±27 cm, the smallest colony of *C. verticillata* measured
568 22 cm and the tallest one measured 144 cm (Fig. 11). The size histogram is bimodal and the result of
569 the Gaussian mixture model shows one group with a mean size of 65 cm (27% of the population) and
570 another group with a mean size of 86 m (73% of the population).

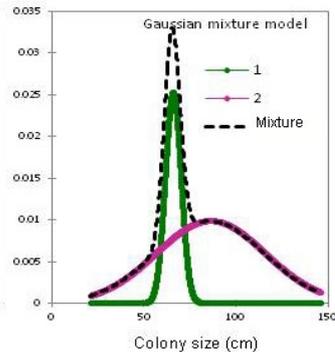
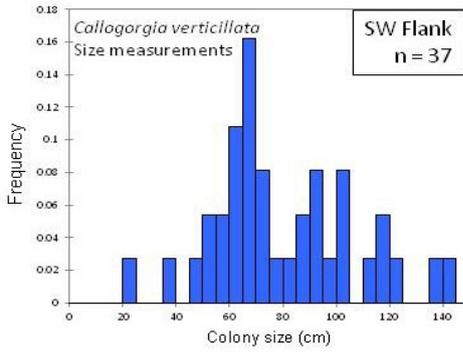
571 At Cassis-500, size measurements (n=44) show a mean size of 60±28 cm, a minimum size of 18 cm
572 and a maximum size of 109 cm (Fig. 11). The Gaussian mixture model shows three different size
573 groups. The smaller one contains 23% of the population with a mean size of 27 cm, an intermediate
574 group contains 32% of the population with a mean size of 52 cm and the larger group contains
575 almost half of the population (45%) with a mean size of 87 cm.

576 At Cassis-200, the size of the very few small colonies of *C. verticillata* observed at the periphery of
577 knolls was not measured.

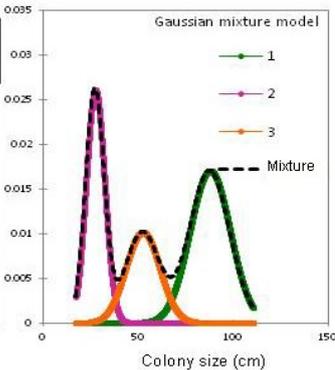
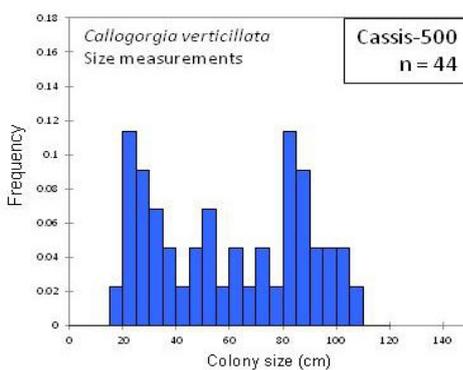
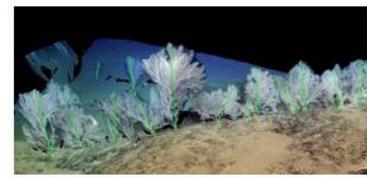
578 The Mann-Whitney test performed to compare mean sizes of *C. verticillata* colonies at SW Flank and
579 Cassis-500 evidence a difference between the two populations (p-value≤0.05).

580

581



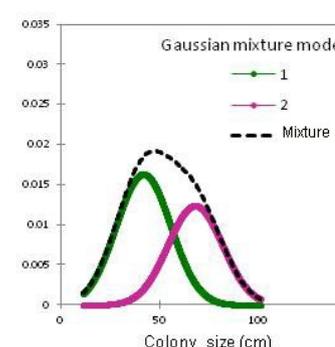
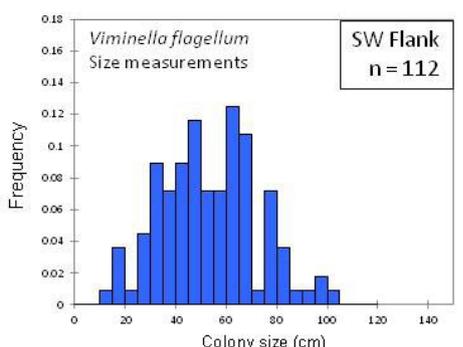
Group	Mean size (cm)	Proportion (%)
1	65	27
2	86	73



Group	Mean size (cm)	Proportion (%)
2	27	23
3	52	32
1	87	45



582



Group	Mean size (cm)	Proportion (%)
1	42	57
2	69	43



583

584

585 Figure. II. Histograms of gorgonian size distribution, and gaussian mixture models implemented to identify size groups for *C.*
 586 *verticillata* populations at SW Flank and Cassis-500, and for *V. flagellum* population at SW Flank.

587

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589

590 Gorgonian *V. flagellum*

591 The global mean size of *V. flagellum* in the Cassidaigne canyon was 53 ± 15 cm. Observations showed
592 that most colonies of *V. flagellum* were unbranched, however, when in the presence of one or two
593 branches, the entire colonies measured around 20 cm at most.

594 At SW Flank (n=112), the smallest colony measured 12 cm, while the tallest measured 100 cm. The
595 Gaussian mixture model shows a population composed of two size groups with mean heights of 42
596 cm and 69 cm, with a respective distribution of 57% and 43% (Fig. 11). The overall mean size was 53
597 ± 19 cm.

598 At Cassis-200, only 11 colonies could be measured, and no Gaussian mixture model was computed.
599 The smallest colony measured 39 cm, while the tallest one measured 72 cm; the mean size was $52 \pm$
600 11 cm.

601 The Mann-Whitney test performed to compare the mean sizes of *V. flagellum* colonies at SW Flank
602 and Cassis-200 locations showed that there is no difference between the two groups. No significant
603 difference (p -value ≥ 0.05) was observed between the mean sizes at Cassis-200 and SW Flank.

604

605 3.6 The diversity of accompanying species

606 In Cassidaigne canyon, the highest species richness was observed at Cassis-200 and SW Flank, with a
607 total of 63 and 52 species of megafauna observed (supp. mat. C). At Cassis-500, we could observe 34
608 different species on photos and videos, while at the Wall only 23 species were observed.

609 Colonies of *M. oculata* were seen at the four locations, and some accompanying species were also
610 present at these locations: the poriferans *Haliclona (Haliclona) magna*, an encrusting bright yellow
611 sponge (*Hexadella* cf. *crypta*) (Reveillaud et al., 2012) and an encrusting blue gray sponge
612 (*Hamacantha* cf. *falcula*); the scleractinian *Desmophyllum dianthus*; serpulids with calcareous tubes;
613 the decapod *Munida* sp.; the brachiopod *Gryphus vitreus*; the echinoid *Cidaris cidaris*; and the fish
614 *Pagellus bogaraveo*.

615 A minimum of 14 visually distinct sponges could be seen on the photos and videos, but not named
616 because no sampling was performed. Certain others are known from the literature to live in the
617 Cassidaigne canyon and were easily recognized. Therefore, the taxonomic and morphologic list of
618 sponges provided is probably underestimated (supp. mat. C).

619 At Cassis-200, species from the continental shelf could be observed mixed with species from the
620 deep sea, i.e. several colonies of the gorgonian *Paramuricea clavata* (tall and small ones), the
621 decapods *Palinurus elephas* and fish species *Anthias anthias* and *Lapanella fasciata*. The soft coral
622 *Daniela koreni* (pers. Comm. P. Lopez-Gonzales) was also present in high abundance between knolls.

623 The SW Flank is characterized by the dominance of the structuring species *C. verticillata* and *V.*
624 *flagellum*. However, the species composition at this location is similar to what was seen at Cassis-
625 200, in terms of presence/absence but is very different in terms of species abundance and
626 distribution. The analysis of videos and photos showed that *C. verticillata* at SW Flank is a host for
627 several species of epifauna. Species from groups such as zoantharians, sea urchins, hydrozoa,

628 nudibranchs, crustaceans and ophiuroids were seen occupying the canopy of the alcyonaceans. In
629 most cases, more than one group of epifaunal species was found on a single colony of *C. verticillata*.
630 A total of 44 colonies of *C. verticillata* (out of 437) providing habitat for species of epifauna were
631 observed (10%). Two colonies were found to host the zoantharian *Zibrowius primnoidus* (Carreiro-
632 Silva et al., 2017), 8 Tritoniidae were observed in 8 colonies, and juveniles of Ophiuroidea were found
633 in 17 colonies of *C. verticillata*. Additionally, hydrozoans were found on dead structures. Several
634 individuals (one to four) of the sea urchin *Cidaris cidaris* were found in a total of 7 colonies, most of
635 which were damaged.

636 The Wall is composed of three different 'habitats' presenting a low diversity of megafauna: the
637 vertical cliff, terraces with coral rubble, and overhangs. Under the overhangs, *M. oculata* was found
638 covering the entire ceiling with a bright blue encrusting sponge between coral colonies. An
639 encrusting pale yellow sponge was also seen on protrusions that resembled fossilized oysters
640 hanging from the ceiling. The outer ledge of the overhang was very rich and harbored many small
641 colonies of the red coral *C. rubrum*, together with *M. oculata*. The vertical cliff was covered by many
642 sponges with *Munida* sp. hidden in small holes of the granular cliff. Terraces serve as repositories for
643 coral rubble falling from the overhangs above. Many skeletons of the scleractinians *Dendrophyllia*
644 *cornigera* and *M. oculata* were seen on these shelves. Some colonies were still alive as were
645 echinoids probably attracted by the fauna close by.

646 At Cassis-500, two glass sponges *Farrea bowerbanki* (Boury-Esnault et al., 2017) and cf. *Oospacas*
647 *minuta* (Bakran-Petrcicoli et al., 2007) were observed exclusively at this location, in very high
648 abundances. Another glass sponge *Tretodictyum reiswigi* (Boury-Esnault et al., 2017) was also seen in
649 high abundance but was not restricted to Cassis-500. This location is also characterized by the
650 presence of a mixture of small and tall colonies of *C. verticillata*, and only 3 colonies out of 104 (3%)
651 were observed with the presence of epifauna. At this location, the epifaunal groups included one
652 individual of the nudibranch Tritoniidae, a juvenile of the crustacean *Anamathia rissoana*, the
653 zoantharian *Zibrowius primnoidus* (Carreiro-Silva et al., 2017) and a high abundance of juvenile
654 Ophiuroidea. Eggs of an unknown species were also found in the canopy of *C. verticillata* at Cassis-
655 500.

656 The number of necroses of *C. verticillata* was also registered during video analysis. At Cassis-500, no
657 necroses were found on the 104 colonies of *C. verticillata* however, at the SW Flank, 20 (out of 437)
658 colonies with necroses were observed (4.5 %).

659 The cluster analysis computed to explore the similarity between the four locations, based on their
660 species composition in presence or absence, grouped Cassis-200 and SW Flank with a Sorensen
661 similarity index above 70%. The Wall was similar to Cassis-200 and SW Flank at 50%, and finally
662 Cassis-500 presented only a 48% similarity to the three other sites (supp. mat. D). The Wall and
663 Cassis-500 harbor different assemblages of species.

664

665

666

667 3.7 State of degradation and seafloor integrity

668 3.7.1 Coral rubble of *M. oculata*

669 Coral rubble was found at the three locations where *M. oculata* was present in high abundance,
670 Cassis-200, Cassis-500 and the Wall. The surface occupied by rubble differed between locations
671 (Table II).

672

673 **Table II.** Surface covered by coral rubble, calculated using either a 2-dimension or a 3-dimension approach, for each location
674 where *M. oculata* was observed. The surface covered by coral rubble is also represented in % in relation to the total area
675 occupied by living *M. oculata* at each location (Table I).

676

Location	Rubble 2D (m ²)	Rubble 3D (m ²)	% rubble 2D	% rubble 3D
Cassis-200	68	90	34	22
Cassis-500	9	10	2	1
the Wall	74	154	23	49

677

678

679 At Cassis-200, coral rubble was observed on 9 of the 13 knolls reconstructed in 3D. Rubble was found
680 on the floor under places where colonies have settled. Knoll 3 had the largest surface area covered
681 by coral rubble. Overall, at Cassis-200, coral rubble was measured to cover a surface area of 90 m²,
682 representing 22% in relation to the total coverage occupied by living colonies.

683 At Cassis-500, coral rubble was only observed in a small area of the site, occupying a surface area of 9
684 m² according to the calculations in 2D, and of 10 m² according to the calculations in 3D. Coral rubble
685 was found on the floor, under an area covered by living colonies of *M. oculata*, corresponding to 1-
686 2% of their total surface area at Cassis-500. The area where coral rubble was found at Cassis-500 was
687 the only site of the location where images of the seafloor were recorded beneath the ledge colonized
688 by *M. oculata*. Despite the small surface area covered by rubble, Cassis-500 was the only location
689 where dead colonies of *M. oculata* were observed standing in place (see section 3.7.3).

690 The Wall was the site where coral rubble occupied the largest surface area. Coral rubble covered the
691 rugged sections of the vertical wall, just below the overhangs where colonies have settled. The total
692 area occupied by rubble along the Wall was 154 m² in 3D, representing 49% of the total surface area
693 observed with living *M. oculata*. When considering the 2D calculations of the surface area, the rubble
694 surface area was 74 m², representing 23% of the total area occupied by living *M. oculata*.

695

696 3.7.2 Marine litter

697 A total of 112 litter items were found at the four locations, of which 46 at Cassis-200, 16 at Cassis-
698 500, 41 at SW Flank and 9 at the Wall. The spatial distribution of marine litter observed during video
699 analysis shows the presence of three categories of litter at the four locations in the Cassidaigne

700 canyon: plastic, metal and glass/ceramics, with a density ranging from 4.9 to 10.6 items.km⁻¹ (Table
701 III).

702

703 **Table III.** Distribution of marine litter observation and linear density (item/km) at each location. Sub-categories of plastic have
704 been defined according to the criteria established by the MSFD for the Mediterranean Sea.

	Cassis-200	Cassis-500	SW Flank	Wall
A2 : bottles	1			
A5 : other plastic objects	1			
A - PLASTIC A6 : Fishing nets	0	1	3	2
A7 : Fishing longlines	44	13	35	6
A8 : Other fishing related (beverage can)		1	1	1
C -METAL D - GLASS (bottle)		1		
			1	
Total nb of item	46	16	40	9
Total navigation length (km)	4.838	2.035	3.775	1.825
Item/km	9.5	7.8	10.6	4.9

705

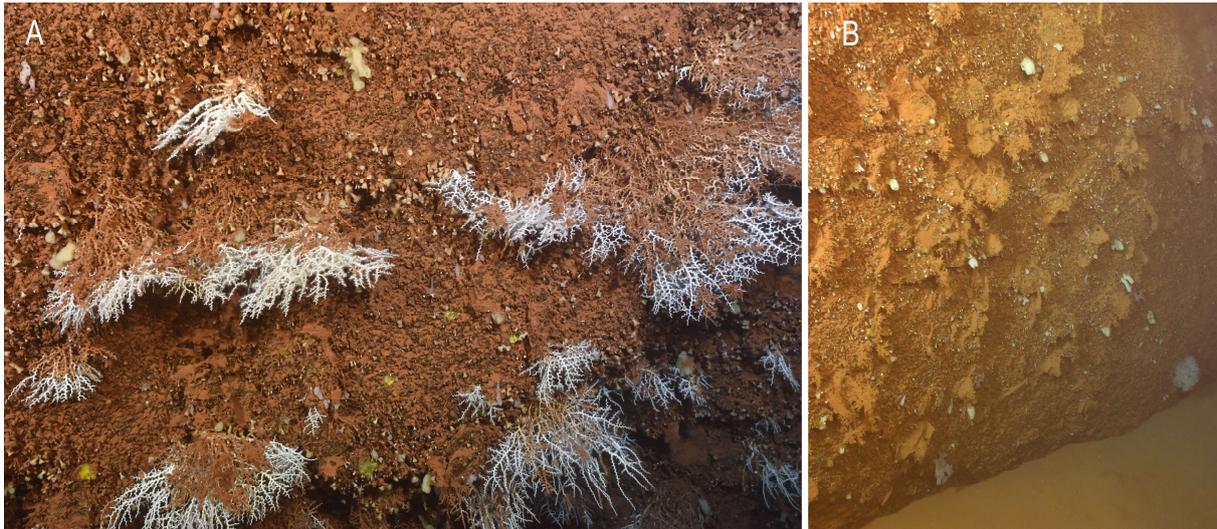
706 Derelict fishing gear (98 longlines and fishing nets) belonging to the plastic category (105 items), was
707 the type of marine litter observed most, as it was found at all four locations (Table III). Entanglement
708 of coral colonies is one of the principal impacts observed at Cassis-200 (see Knoll 3 and 5 in supp.
709 mat. A). Plastic, including fishing gears, was the most abundant category of marine litter found at the
710 Cassidaigne canyon, accounting for 93 % of all litter items. A total of 5 sub-categories of plastic, out
711 of the 10 defined by the MSFD, were found at the Cassidaigne canyon (Table III). Derelict fishing
712 gears were the sub-categories that contributed to the high abundance of plastic along the canyon,
713 with fishing lines (category A7) representing between 67% and 98% of the abundance of plastic
714 marine litter.

715

716 3.7.3 Waste disposal

717 Cassis-500 was the only location where video images showed the presence of bauxite residues. The
718 entire ledge inspected was covered by red sediment (Fig. 12) with living colonies of *M. oculata*,
719 having their basal part covered by red mud. In the deepest part of Cassis-500, colonies of *M. oculata*
720 were seen standing in place but dead and completely covered by red mud (Fig. 12B), showing no
721 existence of a living part at the apex of the colony. Cassis-500 corresponds to a surface area of 924
722 m² of a total of 1650 m² occupied by *M. oculata* (Table I), which means 56% of the explored *M.*
723 *oculata* habitat is impacted by bauxite residues in the Cassidaigne canyon.

724



725

726 Figure 12. Colonies of *M. oculata* impacted by the disposal of bauxite residues at Cassis-500: (A) The basal part of living colonies
727 covered by red mud and (B) a vertical wall covered with dead colonies.

728

729

730 3.8 Temporal changes at Cassis-200

731 Two knolls (Knoll 3 and 9), easily identifiable at the location Cassis-200 on the 3D reconstructions,
732 were recognized in the video footage recorded previously in 2010 and 2013.

733 Photos of Knoll 3 were taken with an interval of seven years (2010 and 2017), while photos of Knoll 9
734 were taken with one of five years (2013 and 2017) (Fig. 13). The corresponding 3D reconstructions
735 were built and helped the recognition of both knolls (Fig. 13).

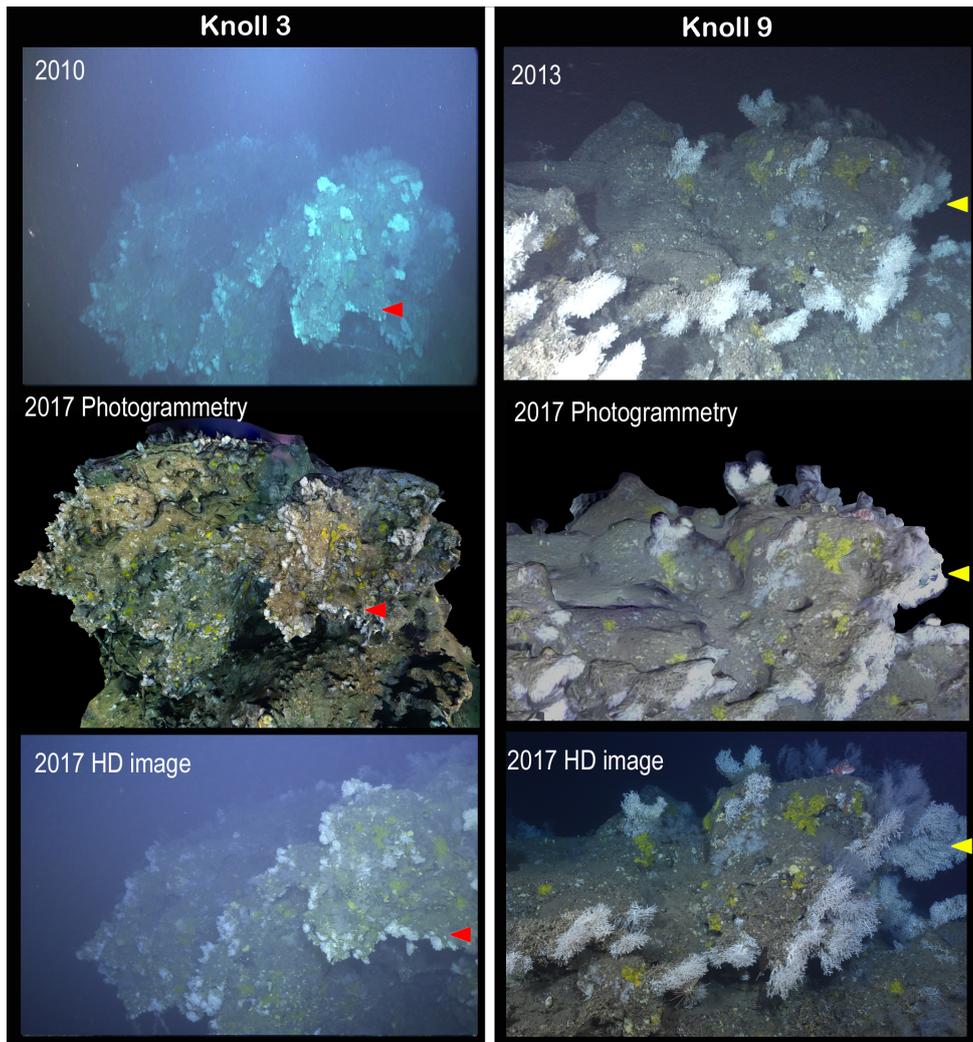
736 At Knoll 3, no major differences could be seen between 2010 and 2017. However, possible growth in
737 the size of colonies and the establishment of two new colonies could be observed, although the
738 videos from 2010 were of poor quality and no HD images were taken (Fig. 13). At Knoll 3, a fishing
739 long line seen in the 3D reconstruction of 2017 was also seen during the analysis of the video of
740 2010, showing its persistence over seven years.

741 At Knoll 9, no major differences could be seen between 2013 and 2017. Only apparent growth was
742 noted for *M. oculata* colonies (Fig. 13). However, no big differences in the sizes of the rest of the *M.*
743 *oculata* colonies were observed.

744 However, videos from 2010 and 2013 were not dedicated to *M. oculata* observation, and no HD
745 images were taken and no 3D models were built at those times, therefore differences could not be
746 quantified and these results should be interpreted with caution.

747

748



749

750 Figure 13: Illustration of the temporal changes at Cassis-200. Photos and 3D reconstruction of Knoll 3 at seven years interval, the
 751 arrows indicate the place where two new colonies could be observed in 2017 compared to 2010, and Knoll 9 at five years interval,
 752 yellow arrows indicate the colonies that seem to have grown in 2017 compared to 2013.

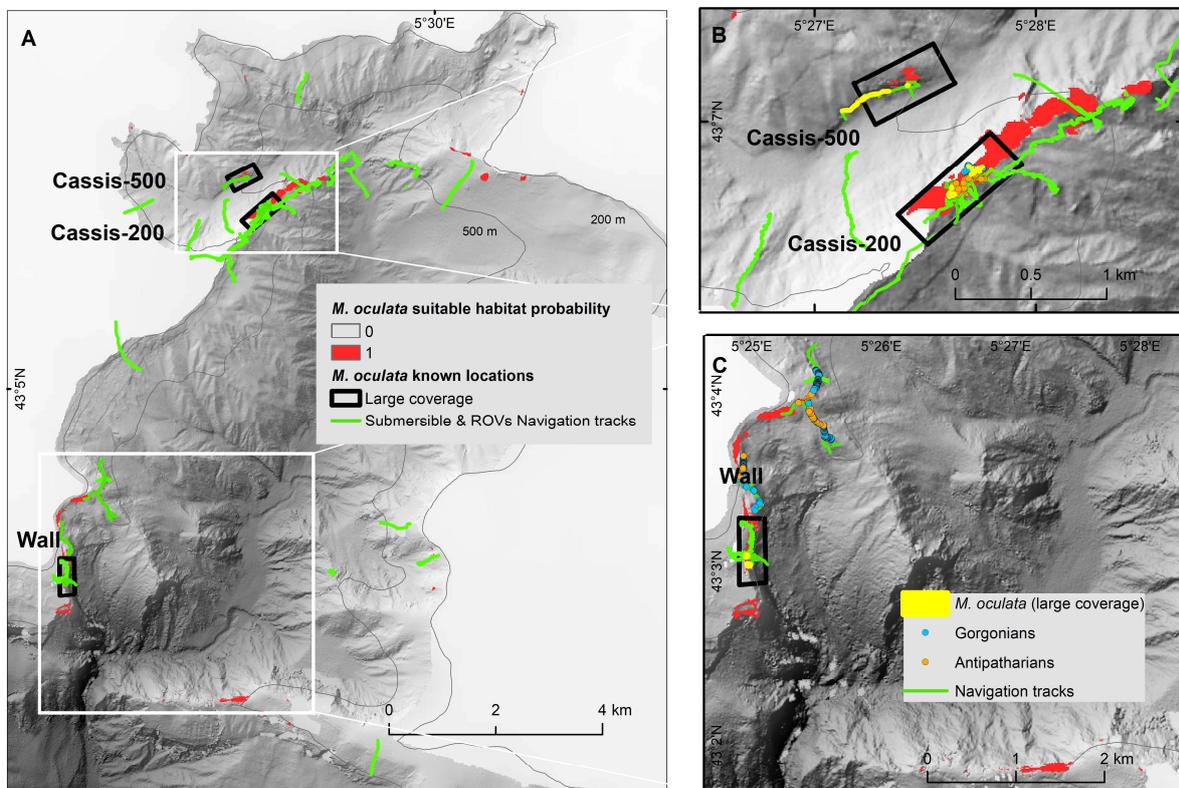
753

754

755 3.9 Validation of the habitat suitability model built previously

756 The inspection of Cassidaigne canyon in 2017 was driven by the habitat suitability maps built
757 previously (Fabri et al., 2017; Bargain et al., 2018). However, not all the locations highlighted as
758 suitable by the model could be explored, especially that located on the eastern flank (Fig. 14). But the
759 exploration of Cassis-200, Cassis-500 and the Wall was a success and revealed a very high
760 concentration of *M. oculata* colonies, as described in this paper. The exploration of the SW Flank did
761 not specifically show *M. oculata* populations, but other cold-water corals, such as gorgonians and
762 antipatharians were observed in high concentrations (Fig. 7, 11 & 15). Distribution of the absence of
763 *M. oculata* colonies is shown by mapping all the navigation tracks in Cassidaigne canyon (Fig. 14).
764 Despite the effort made for the exploration of the eastern flank, no *M. oculata* colony could be
765 found.

766



767

768 Figure 14. A - Map displaying the probability of the presence of *M. oculata* in Cassidaigne canyon generated with a model built using
769 MaxEnt, from (Fabri et al., 2017), and all submersible and ROVs navigation tracks we could gather (cruises MARUM_2009,
770 MEDSEACAN_2009, ESSROV_2010, ESSNAUT_2013, ESSTECHROV_2017, ESSTECHROVGEN_2017, VIDEOCORI_2017). B and C: Probability
771 of presence of *M. oculata* in red, Navigation tracks in green and *in situ* observation of different species of cold-water corals from
772 Videocorl, *M. oculata* in yellow, gorgonians in blue, antipatharians in orange.

773

774

775 4. Discussion

776 4.1 Spatial and vertical distribution

777 In our study *M. oculata* was seen to occupy a vertical distribution at depths ranging from 203 to 483
778 m. At Cassis-200, colonies of *M. oculata* were found on 14 knolls (including Knoll 13 where no 3D
779 model was built). The mean depth of occurrence of *M. oculata* was 215 ± 14 m. In other parts of the
780 Gulf of Lion they were found as shallow as 129 m depth in the La Fonera canyon, but most frequently
781 between 180 and 300 m (Lastras et al., 2016). At Cassis-500, colonies of *M. oculata* settled on the
782 vertical part of a deep (432 ± 10 m) and long (circa 400 m) rocky ledge, with the lower limit of the
783 observed distribution at 483 m. However, during a previous cruise in 2013, *M. oculata* colonies were
784 observed further east along the rocky ledge, reaching depths of 515 m (Fabri et al., 2017). At the
785 Wall, a vertical cliff more than 300 m high, colonies of *M. oculata* had settled in an upside-down
786 position and in high density under overhangs encountered at different depths (307, 320 and 340 m).
787 Due to the challenging navigation conditions in this complex location, the ROV navigated at a safe
788 distance from the cliff and the footage of the cameras therefore covered a 10-meter wide band. But
789 the overhangs are probably larger than the surveyed area, and colonies of *M. oculata* probably cover
790 a surface area larger than that measured at the Wall during the present work. On the SW flank only
791 three colonies could be plotted at a depth of around 350 m. Two other colonies had already been
792 seen in this area in 2009 but they were located shallower at 324 and 244 m (Fabri et al., 2014).
793 Therefore, we can assume that at this location either colonies are sparsely settled, or that we have
794 not yet found the area of high concentration. The wide depth distribution of *M. oculata* seen in
795 Cassidaigne canyon is probably related to its high tolerance to environmental conditions (Wienberg
796 et al., 2009), making it the most dominant structuring species in the Mediterranean Sea (Freiwald et
797 al., 2009; Gori et al., 2013; Fanelli et al., 2017), whereas in the North Atlantic scleractinian structuring
798 ecosystems are mainly composed of *L. pertusa* (Roberts et al., 2003; Costello et al., 2005; van den
799 Beld et al., 2017).

800 Forests of *A. subpinnata* were observed on the top of two knolls at Cassis-200, at depths of 203 m
801 (Knoll 3) and 220 m (Knoll 9). Colonies were grouped together, surrounded by *M. oculata* colonies. In
802 the Mediterranean Sea, the depth distribution of *A. subpinnata* is often reported to be between 60
803 to 100 m deep (Bo et al., 2009); however, around Malta, this species was observed between 250 to
804 400 m (Deidun et al., 2015; Evans et al., 2016). Our observations between these two depth ranges
805 provide new information on the depth distribution of the species, insofar as relatively little is known
806 about this distribution.

807 | Gardens of *L. glaberrima* presented a wider vertical distribution, between 214 and 355 m depth. This
808 is consistent with previous observations commonly reporting the species at depths exceeding 600 m
809 in the Mediterranean Sea (Carlier et al., 2009; Angeletti et al., 2014; Bo et al., 2014; Mytilineou et al.,
810 2014; Angeletti et al., 2015; Bo et al., 2015; Ingrassia et al., 2016). In the Cassidaigne canyon, gardens
811 of *L. glaberrima* were seen in the vicinity of *M. oculata* habitats and also in the vicinity of dense
812 forests of gorgonians *V. flagellum*, something that was also described in Maltese waters (Deidun et
813 al., 2015). Spatial distribution of *L. glaberrima* in the Mediterranean Sea has mainly been highlighted
814 in the eastern part of the western basin (Massi et al., 2018), therefore our records provide new
815 information on its distribution in the western basin.

816 The antipatharian *A. dichotoma* was observed as isolated colonies at depths from 209 to 435 m,
817 which is consistent with previous records in the Mediterranean Sea reporting the species at depths
818 from 100 to 650 m (Mastrototaro et al., 2010; Bo et al., 2011; Mytilineou et al., 2014; Fourt et al.,
819 2017; Bo et al., 2018).

820 The precious red coral *C. rubrum* was found at depths between 203 and 307 m in the Cassidaigne
821 canyon. In the NW Mediterranean Sea, red coral presents a very wide vertical distribution, ranging
822 from 16 m (Tsounis et al., 2006) to the deepest record at 1016 m off Malta (Knittweis et al., 2016).
823 We observed *C. rubrum* in association with *M. oculata* ecosystems at two locations (Cassis-200 and
824 the Wall), an association that has already been described in other canyons of the NW Mediterranean
825 Sea such as in the Spanish canyons, Cap de Creus between 50 and 100 m (Rossi et al., 2008) and La
826 Fonera between 100 and 160 m (Lastras et al., 2016).

827 The fan-forming alcyonacean *C. verticillata* was found at depths between 377 and 483 m at SW Flank,
828 and isolated colonies were also found at a depth of 232 m at Cassis-200. The upper limit of the *C.*
829 *verticillata* population in the Mediterranean Sea was determined as 120 m (Bo et al., 2012b; Bo et al.,
830 2014; Bo et al., 2015), but its distribution is normally reported to be between 300 and 1000 m deep
831 (Mastrototaro et al., 2010; Deidun et al., 2015), which corresponds to our observations.

832 The whip gorgonian *V. flagellum* was found forming large forests at depths between 279 and 390 m.
833 This species has often been observed in the Atlantic Ocean (OCEANA, 2011), presenting a vertical
834 distribution from 90 to 1000 m. In the Mediterranean Sea, the species has been documented in
835 Italian waters (Giusti et al., 2012; Angiolillo et al., 2014; Cau et al., 2017; Giusti et al., 2017), and its
836 distribution is known between 90 to 200 m. Our results extend knowledge on the vertical distribution
837 of *V. flagellum* with the report of large forests at water depths down to 390 m.

838

839 4. 2 Surface and density

840 In this study *M. oculata* was seen to cover a global surface of around 1650 m² in the canyon, with
841 56% of the surface area at Cassis-500, 25% at Cassis-200 and 19% at the Wall. However, we can
842 assume that the surface area is probably larger at the Wall since the overhangs are probably larger
843 than what could be seen. The surface area is also probably larger at Cassis-500 because we had to
844 stop the dive before reaching the end of the ledge, and we know that further east, colonies of *M.*
845 *oculata* have settled to a depth of 515 m (Fabri et al., 2017). Surface areas are rarely mentioned in
846 the literature because navigation tracks are not usually dedicated to photogrammetry. The highest
847 density of *M. oculata* was 12 colonies.m⁻² on Knoll 8 at Cassis-200, but colonies were irregularly
848 distributed among knolls and the density was 7 colonies.m⁻² on Knolls 3 and 9, and even lower on
849 Knolls 2 and 5 (1 colony.m⁻²). On the contrary, at the Wall colonies were regularly distributed and
850 entirely covered the ceiling of the overhangs, but they were of smaller sizes. The mean densities
851 were 4 colonies.m⁻² at Cassis-200 and Cassis-500, and 9 colonies.m⁻² at the Wall. Our results are
852 similar to those found previously in the Cap de Creus canyon where densities of *M. oculata* were
853 calculated to reach a maximum of 11 col.m⁻², with a mean density lower than 1 col. m⁻² (Orejas et al.,
854 2009).

855 The precious red coral *C. rubrum* was found on areas of 11 m² at Cassis-200 (11 colonies.m⁻²) and
856 only 2 m², but in higher density (20 colonies.m⁻²), along the edge on the top of the cliff at the Wall.
857 Red coral was reported in other canyons of the Gulf of Lion, mainly on the upper walls and near
858 vertical walls (Lastras et al., 2016), but no surface areas or densities were calculated.

859 Forests of gorgonians were not mixed and at SW Flank they were either composed of *C. verticillata* (2
860 colonies.m⁻¹) or *V. flagellum* (16 colonies.m⁻¹) distributed in patches each around 20 meters long. At
861 Cassis-500 a population of *C. verticillata* had settled on the ledge (2 colonies.m⁻¹) in patches of
862 around 10 meters long, while at Cassis-200 very few colonies were seen and only in the deepest part
863 of the site. On the Montenegrin margin, the gorgonian *C. verticillata* was reported in densities of up
864 to 5 colonies.m⁻² (Angeletti et al., 2014), which is higher than what we could observe in the
865 Cassidaigne canyon. Around the island of Sardinia and in the Strait of Sicily, *V. flagellum* was reported
866 in densities lower than 0.5 colonies.m⁻² (Giusti et al., 2017) and up to 60 colonies.m⁻² (Angiolillo et al.,
867 2014). We only calculated linear densities, but images of *V. flagellum* forests show similar densities in
868 Italian waters as in the Cassidaigne canyon.

869

870 4.3 Size structure of populations

871 At Cassis-200 we could observe the tallest population of *M. oculata*, with the highest colony being
872 78.3 cm. The structure of the population was composed of four different size groups (mean sizes 9
873 cm, 16 cm, 30 cm and 45 cm) in which 41% of the colonies are in the group of small specimens, and
874 only 8% are in that of the tallest. A population structure composed of similar size groups (<10cm; 10-
875 20 cm; 20-40 cm and higher than 40 cm) was also described in the Cap de Creus canyon (Gori et al.,
876 2013). The presence of small colonies is described to indicate active recruitment (Gori et al., 2013).

877 At Cassis-500 the size structure of *M. oculata* population is composed of four groups (mean sizes = 6
878 cm, 10 cm, 17 cm, 29 cm), in which 65% of the population belong to the two smallest size groups.
879 These mean sizes are smaller than those measured at Cassis-200. We hypothesize that the small size
880 of the population is a consequence of anthropogenic impacts occurring during the last 50 years
881 (Dauvin, 2010; Fontanier et al., 2012; Fontanier et al., 2015; Fabri et al., 2017). The high
882 sedimentation rate of bauxite residues may have prevented *M. oculata*'s growth by suffocation, by
883 starvation or by preventing settlement of new colonies. Colonies were seen only on vertical parts or
884 under the rocky outcrops but not on the top of the ledge where there is probably too much
885 sedimentation for larvae to cope with and settle. According to the minimal growth rate of *M.*
886 *oculata*, shown to be 3 mm/year (Orejas et al., 2008), the smallest size groups of colonies (mean size
887 = 6 and 10 cm) observed in our study are around 20 and 30 years old. The measurement of the size of
888 the colonies performed in this study supports what was initially hypothesized for this location (Fabri
889 et al., 2017), that is the recruitment of *M. oculata* at Cassis-500 probably occurred after the
890 reduction of red mud discharges in 1988, 30 years ago, due to the termination of production by one
891 of the two main aluminum factories. Since the discharge of particles stopped in January 2016, our
892 measurements performed one year after the end of this pressure can be used as a reference for
893 future measurements dedicated to monitoring this site's recovery.

894 At the Wall the population of *M. oculata* was composed of small colonies (two groups of mean sizes
895 10 and 13 cm), in very high density under three overhangs at depths between 307 and 340 m. No

896 large colonies could be observed, all of them were upside down, and coral rubble was seen to cover
897 terraces below the three overhangs. We can thus assume that colonies fell due to their own weight
898 and because of their downward orientation. Colonies are thus naturally unable to reach large sizes at
899 this location, nor can they be physically impacted by human activities because they are protected by
900 the overhang. This makes the Wall a location that can be considered as a reservoir for the *M. oculata*
901 population in the Cassidaigne canyon, ensuring the persistence of the species, since it can spread to
902 other areas given their reproduction processes based on external fertilization (Freiwald et al., 2009;
903 Pires et al., 2014). However, if colonies keep falling off once they reach a certain size, it is also
904 possible that the population at the Wall itself also relies on recruitment of larvae from elsewhere.

905 Colonies of *C. rubrum* were very small, with a mean size of 6 cm height and with the tallest colony
906 found measuring 16 cm in Cassidaigne canyon. In other canyons, heights of *C. rubrum* populations
907 were also measured to be small, i.e. in Cap de Creus canyon they hardly reached 6 cm (Rossi et al.,
908 2008) and in La Fonera canyon, *C. rubrum* was reported as small, centimetric colonies (Lastras et al.,
909 2016). The small size of *C. rubrum* colonies has been shown not to be related to depth, but to
910 disturbances occurring in the environment (Cau et al., 2016). Therefore, the small size observed for
911 the *C. rubrum* population in Cassidaigne canyon is probably due to intense hydrodynamic processes
912 occurring in the canyon (Fabri et al., 2017). The water flow generated by upwelling and downwelling
913 phenomena in the canyon could be a disturbance slowing down the colonies growth. Whereas larger
914 colonies can be observed in deeper waters where there are less constraints from the environment,
915 and a more stable temperature (Galli et al., 2016). For example, *C. rubrum* colonies of good size, in
916 view of their many branches, were seen at a depth of 1000 m off Malta (Knittweis et al., 2016).

917 The height of antipatharian colonies could not be measured because of their flexibility and their
918 movement in the water flow, which prevented photogrammetry. However, it was possible to observe
919 different sizes of both species *A. subpinnata* and *L. glaberrima*, suggesting active recruitment and
920 healthy populations. *Antipathes dichotoma* was always seen in tall colonies and no small ones were
921 reported. This might be because small colonies could not be seen on the videos as they are very thin
922 and flattened against the substrate. Antipatharians are amongst the oldest known marine organisms,
923 with very slow growth and millennial longevities (Roark et al., 2009; Carreiro-Silva et al., 2013).
924 *Leiopathes glaberrima* and accompanying benthic invertebrates have been reported as Vulnerable
925 Marine Ecosystem indicators and protection measures should be adopted as a precautionary
926 approach (Massi et al., 2018).

927 The population of *C. verticillata* at the SW Flank was composed of two size groups, 65 cm and 86 cm
928 divided respectively into 27% and 73%, both corresponding to tall sizes, with no record of small
929 specimens. Small colonies may be located at another place that we may not have explored. A bi-
930 modal size structure has already been described for *C. verticillata* in the Gulf of S. Eufemia
931 (Tyrrhenian Sea) where colonies of *C. verticillata* presented two groups of colonies of smaller sizes
932 (15-20 cm or 55-60 cm) (Bo et al., 2012b). At Cassis-500, the size structure of the *C. verticillata*
933 population was divided into three groups of 27, 52 and 87 cm (23%, 32% and 45 %). The presence of
934 small colonies supports the idea of a population with recruitment. Since the medium size class is less
935 abundant than the two others, it may have suffered from the high sedimentation rate of bauxite
936 residues expelled at the site, but now that the discharge of these residues has stopped a new
937 generation may have settled.

938 The whip gorgonian *V. flagellum* formed large forests at SW Flank and we measured a mean size of
939 53±19 cm, divided into two groups of 42 and 69 cm (57% and 43%). The measurements may be
940 inaccurate because *V. flagellum* is flexible and the 3D models may not have properly reconstructed
941 the apex of the colonies. No measurement for *V. flagellum* is reported in the literature.

942

943 4.4 Diversity of accompanying species

944 At Cassis-200, a higher megafauna species richness was observed compared to the three other
945 locations in the canyon, and exceptional diversity patterns were associated with the highest
946 topographic elevations (knolls) covered by *M. oculata*, i.e. forests of *A. subpinnata* were found at
947 Knolls 9 and 3; *Paramuricea clavata* was found at Knoll 3; and *C. rubrum* was observed at Knolls 3, 8,
948 15 and 16. These knolls are high topographic elevations on which species can be exposed to currents.
949 This factor was reported to be a driver of the distribution of deep-water corals, for example *M.*
950 *oculata* (Fabri et al., 2017; Bargain et al., 2018) and *C. rubrum* (Lastras et al., 2016) in the
951 Mediterranean Sea, but also *Lophelia pertusa* in the Atlantic Ocean (Mohn et al., 2014). The
952 association of the precious red coral *C. rubrum* with the scleractinian *M. oculata* has often been
953 reported in the Mediterranean Sea (Rossi et al., 2008; Costantini et al., 2010; Lastras et al., 2016) and
954 deep occurrences have been discovered recently (Evans et al., 2016; Knittweis et al., 2016). Two
955 scleractinian species, the yellow coral *Dendrophyllia cornigera* and the solitary coral *Dendrophyllium*
956 *dianthus*, are widely described in association with *M. oculata* in other canyons, such as in the
957 Sardinian canyons (Taviani et al., 2017) and the Levante canyon (Fanelli et al., 2017). The presence
958 of tall colonies of *M. oculata* at Cassis-200, as well as the multi-modal size structure of the
959 population, suggests a healthy population. This exuberant development of colonies is known to
960 provide a habitat structure for many other species (D'Onghia et al., 2012) and may explain the high
961 number of accompanying species of megafauna we recognized (63 species or morphotypes)
962 following Fourt et al. (2017). The presence of *M. oculata* has already been shown to enhance the
963 local diversity of species at a site (D'Onghia et al., 2010). In comparison with lists of species observed
964 at other *M. oculata* locations in the Mediterranean Sea, we observe that fewer than 50 megafauna
965 species are commonly reported in cold-water ecosystems using images only, i.e. 24 megafauna
966 species were reported in the Bari canyon (Angeletti et al., 2014); 30 megafauna species in La Fonera
967 canyon (Lastras et al., 2016). But when using sampling, including megafauna and macrofauna, the
968 number of species increases: e.g. 222 species were identified at Santa Maria di Leuca (Mastrototaro
969 et al., 2010).

970 At Cassis-500, colonies of *M. oculata* were spaced from each other, growing only on the vertical part
971 of a ledge and not on the top. Accompanying species were less numerous (34 species or
972 morphotypes) than at Cassis-200, and composed of different species (only 48% similarity). Cassis-500
973 has been impacted by waste disposal for 50 years (Dauvin, 2010; Fabri et al., 2017), and the effects
974 are visible in every video and photo. Therefore, species growing at Cassis-500 are those able to live
975 without being buried and able to cope with sedimentation. Since this site is also the deepest in our
976 study, we cannot exclude the hypothesis that depth could also be a factor explaining a change in
977 species diversity. However, the exceptional diversity found recently deep in the Mediterranean Sea,
978 for example at depths of 400 m in Sardinian canyons (Moccia et al., 2019) or at depths of 1000 m off

979 Malta (Evans et al., 2016), suggest that depth is not the main factor that prevented accompanying
980 fauna to develop at Cassis-500.

981 We observed the lowest number of accompanying species at the Wall (23 species or morphotypes)
982 and a similarity of up to 50% with Cassis-200 and SW Flank. We could see *M. oculata* and a blue
983 sponge entirely covering the ceiling of overhangs, and the red coral *C. rubrum* living on the border of
984 the overhang between small colonies of *M. oculata*. A few other species, mainly sponges, have
985 settled on the vertical part of the cliff, but do not specifically accompany *M. oculata*. The difficult
986 navigation conditions encountered along the vertical cliff and the difficulties in observing the ceiling
987 under the overhangs prevented good observation of species, which may explain the low number of
988 species reported in this study. But this location is also the most distant from the coast and from the
989 organic matter input, which may also explain the low number of species.

990 Forests of *A. subpinata* and gardens of *L. glaberrima* were seen in the vicinity of either *M. oculata* at
991 Cassis-200, or *V. flagellum* at SW Flank. Those two locations presented the highest numbers of
992 megafauna species (respectively 63 and 52 species or morphotypes) and presented a species
993 composition with more than 70% similarity. Antipatharians have already been described as true
994 oases of biodiversity (Bo et al., 2012b). Given their large tree-like structure, antipatharians can form
995 gardens or forests that have been compared to terrestrial forests given their longevity and the
996 amount of high biodiversity they support (Bo et al., 2015). Antipatharians also have a nursery role in
997 the ecosystem (Bo et al., 2015). The fragility of antipatharians has been recognized and the FAO now
998 considers coral gardens as VMEs (FAO, 2016). *Leiopathes glaberrima* appears as an endangered
999 species on the IUCN Red list of species (Bo et al., 2015). The three species of antipatharians observed
1000 in the Cassidaigne canyon are protected under three different conventions: CITES (Convention on
1001 International Trade in Endangered species), the Bern convention (Treaty no. 104; Bern, 19.IX.1979)
1002 and the Barcelona convention (IG.23/10/EC) where antipatharians are listed in Annex II (*L.*
1003 *glaberrima*) and Annex III (*A. subpinnata* and *A. dichthoma*).

1004 Both gorgonian species, *C. verticillata* and *V. flagellum*, were present as deep forests at SW Flank and
1005 may play important habitat structuring roles. We observed gorgonians in association with
1006 antipatharians (above paragraph), scleractinians (*D. dianthus* and *D. cornigea*) and commercially
1007 important species, such as the seabream *Pagellus bogaraveo* and the lobster *Palinurus mauritanicus*.
1008 The presence of epifauna species in the canopy of *C. verticillata* was often seen. A high number of
1009 colonies of *C. verticillata* were observed acting as hosts for nudibranchs. A previously identified
1010 zoantharian species, *Zibrowius primnoidus*, was reported as a parasite of *C. verticillata* (Carreiro-Silva
1011 et al., 2017). This zoantharian is known to use the colony as a host and as a food resource by eating
1012 the tissue. Observations of the high abundance of epifauna showing negative effects on colonies of *C.*
1013 *verticillata* have also been demonstrated in Sicilian waters (Bo et al., 2014). In our study, we also
1014 observed the presence of the echinoid *C. cidaris* together with necroses in the canopy of *C. verticillata*
1015 at SW Flank, but neither the echinoid nor the gorgonian necroses were seen at Cassis-500. We
1016 assume that a relationship exists between the high number of sea urchins and the high number of
1017 necroses in the canopies. At Cassis-500 it seems that bauxite residues and sedimentation rate
1018 prevent the settlement of *C. cidaris* and therefore its feeding on *C. verticillata*. Bauxite residues could
1019 thus protect *C. verticillata* from grazing.

1020 With regards to diversity, caution is necessary when considering the number of megafauna species
1021 presented in this study as it was based on photos and video analysis, which can lead to bias and the
1022 misidentification of species. Moreover, macrofauna and meiofauna are not considered in our study
1023 since no sampling was performed. Quantitative evaluation of diversity was not possible because the
1024 species matrix was not quantitative and the only data available was the presence and absence of
1025 species. Although the number of species identified at Cassis-200 appears to illustrate high diversity
1026 when compared to the three other locations explored in the canyon, the number of species at this
1027 location nonetheless remains underestimated. For example, video observations have shown a high
1028 abundance and diversity of sponges at Cassis-200, but the species could not be identified, since the
1029 identification of sponges requires physical sampling (Bo et al., 2012a), and sometimes genetic
1030 analysis (Bakran-Petricioli et al., 2007).

1031

1032 4.5 Seafloor integrity

1033 4.5.1 Marine litter and waste disposal

1034 Marine Litter was mainly composed of plastics at the four locations, reflecting the general trend
1035 previously described in French Mediterranean canyons (Fabri et al., 2014; Gerigny et al., 2019).
1036 Plastic items were identified as derelict fishing gears such as fishing long lines. They were seen in high
1037 abundance in all the coral habitats explored, together with coral rubble. Overall, in Cassidaigne
1038 canyon, the surface corresponding to *M. oculata* rubble was measured as 15% of the total surface
1039 occupied by the species. It is known that fishing longlines can have impacts on colonies due to
1040 entanglement and physical breakage (Lastras et al., 2016; D'Onghia et al., 2017; Fanelli et al., 2017;
1041 Galgani et al., 2018). Yet, in Cassidaigne canyon, the locations Cassis-200 and Cassis-500 have been
1042 included within regulated fisheries zones since 2012. This indicates that the fishing gears observed
1043 during our study, entangled in coral colonies, were lost before that time. Our study took place on the
1044 upper part of Cassidaigne canyon's flank, therefore we could not observe heavy litter items as those
1045 described in three large submarine canyon off Spain (Tubau et al., 2015), but we found the same
1046 distribution of marine debris, mainly composed of longlines, observed between 30 to 300 m in the
1047 Tyrrhenian Sea (Angiolillo et al., 2015). These fishing gears were seen colonized by *M. oculata* and
1048 similar observations have also been described in the Levante canyon (Fanelli et al., 2017). This shows
1049 the persistence of the abandoned fishing gears and the resilience capacities of the ecosystem. The
1050 largest surface of coral rubble was found at the Wall, but at this location of coral rubble is more likely
1051 caused by the natural fall of the colonies under their own weight than by fishery-related impacts.

1052 Another impact on the *M. oculata* ecosystem in Cassidaigne canyon is the presence of bauxite
1053 residues at Cassis-500, covering 56% of the total area where *M. oculata* colonies are known to have
1054 settled. It is possible that a non-evaluated part of the habitat was completely buried by bauxite
1055 residues, and this would increase the proportion of the ecosystem impacted. At Cassis-500 the local
1056 species composition is affected, evidenced by the low number of species observed and the lowest
1057 similarity in species composition compared to the three other sites. Moreover, at Cassis-500, the size
1058 structure of *M. oculata* population is different from that measured at Cassis-200, and the mean size
1059 of colonies is smaller. The gorgonian *C. verticillata* has a size structure with a very low number of
1060 colonies in the middle size group, which may correspond to a group of colonies that had settled
1061 when industrial wastes were expelled in large quantities. Even though the disposal of waste in solid

1062 particle form was stopped in January 2016 (Fabri et al., 2017), it will take many years before
1063 noticeable changes occur.

1064 4.5.2 Seafloor Integrity criteria for the European MSFD

1065 To consider the Good Ecological Status of benthic habitats, the MSFD stipulates the use of descriptor
1066 6, Seafloor Integrity, and more specifically two criteria: the extent of loss of the habitat type (D6C4),
1067 and the extent of adverse effects on the condition of the habitat type (D6C5). These criteria measure
1068 the degree of anthropogenic effects on biological parameters that should not exceed a specified
1069 proportion of the natural extent of the habitat (EU Decision 2017/848). According to the MSFD, the
1070 physical loss of the natural seabed due to permanent change of the seabed substrate or morphology
1071 should be considered as a permanent change to the seabed that has lasted or is expected to last for a
1072 period of two reporting cycles (12 years) or more.

1073 In this study, we considered the habitat type to be the *M. oculata* ecosystem at the scale of the
1074 Cassidaigne canyon, and we used photos, videos and photogrammetry to evaluate the status of the
1075 ecosystem with the two criteria suggested for benthic habitats. Our results show that the proportion
1076 of the habitat impacted by bauxite residues (waste disposal) was 56% of the total extent of the *M.*
1077 *oculata* habitat. This proportion represents the total surface occupied by *M. oculata* at Cassis-500,
1078 where we can attest that in the deepest part, *M. oculata* colonies were completely dead; the density
1079 of the structuring species *M. oculata* is lower compared to the other sites in the canyon; the mean
1080 sizes of the structuring species *M. oculata* and *C. verticillata* are smaller than the size at visually non-
1081 impacted sites; the size structure of these populations (*M. oculata* and *C. verticillata*) is different
1082 compared to the size structure at visually non-impacted sites; the accompanying species presented a
1083 particular and different composition, compared to visually non-impacted sites in the canyon.

1084 It is noteworthy that a 12-year interval in the deep sea is very short and the impact on the deep
1085 seafloor is of much greater concern than on the continental shelf because of the low resilience of
1086 deep-sea benthic ecosystems (Fabri et al., 2018). We can therefore conclude that in the light of
1087 criterion D6C4, the extent of loss of the *M. oculata* ecosystem equals 56% in the Cassidaigne canyon
1088 and will need more than 12 years to recover despite the fact that the pressure has stopped.

1089

1090 4.6 Temporal changes

1091 The MSFD advises that special conservation areas should be monitored every 6 years to evaluate the
1092 Ecological Status of European marine waters (EC 734/2008). In our study, we presented an attempt
1093 to perform a temporal analysis at 7 and 5 year intervals on two luxuriant *M. oculata* knolls at Cassis-
1094 200. Using 3D models of 2017 made it possible to recognise which knolls had been observed in 2010
1095 and 2013 and to compare them with what we observed in 2017. It was difficult to quantify changes in
1096 the general size and distribution of *M. oculata* colonies, and sizes could not be quantified because we
1097 did not possess good resolution images for the 2010 and 2013 cruises, and because videos were not
1098 recorded with the objective to build 3D models, we could not quantify sizes. In addition, the low
1099 growth rate and low fecundity of *M. oculata*, explain the small changes seen between years.
1100 Therefore, based on these results we suggest that if legal conservation instruments are applied in the
1101 area, instead of the 6-year monitoring cycles considered by the MSFD, a 12-year monitoring cycle

1102 should be enough for *M. oculata* ecosystems at Cassis-200. Monitoring should be strengthened at
1103 Cassis-500, where recovery should be observed in terms of species diversity since the disposal of
1104 bauxite residues was stopped in 2016. However, although the necessary time for the recovery is not
1105 known, this site could be a good observatory. Obviously, in the deepest part of the site, where hard
1106 bottoms are covered by bauxite residues, sessile benthic species will probably never be able to settle
1107 again.

1108

1109 4.7 Validation of potential habitat maps

1110 The suitable habitats predicted previously in Cassidaigne canyon with statistical models (Fabri et al.,
1111 2017; Bargain et al., 2018) were visually explored during the Videocor1 2017 cruise described in our
1112 study. At every location predicted to be a probable habitat for *M. oculata* we found large cnidarian
1113 species able to play a structural role in the ecosystems, i.e. *M. oculata* or antipatharians or large
1114 gorgonians. At SW Flank we did not find the exact species we expected to find, therefore we can
1115 hypothesize that: (1) the parameters used as drivers for the distribution of *M. oculata* are the same
1116 for the different species of cnidarians and we may not have elucidated the parameters strictly
1117 related to the scleractinians; (2) and if the parameters are the same, competition between species
1118 may have led to one species dominating instead of another. For example, at SW Flank where we
1119 expected to find *M. oculata* populations but where we saw large forests of antipatharians and
1120 gorgonians, it may be that every hard substrate available for the establishment of *M. oculata* was
1121 already occupied by antipatharians and gorgonians, therefore we saw only sparse colonies of *M.*
1122 *oculata*. Whatever the case we found habitat-forming species of cnidarians sheltering a diverse
1123 ecosystem where potential areas for cold-water corals were predicted.

1124

1125 5. Conclusion

1126 Photogrammetry provided much information on the complex locations in Cassidaigne canyon. Every
1127 site/location can now be easily recognized by its particular shape, and 3D models helped to better
1128 understand where vulnerable ecosystems are located. Moreover, the detailed maps produced in this
1129 work will be of great help for future monitoring of the area. Also, the precise navigation now possible
1130 with ROVs and submersibles allows exact positioning and returning to study locations very easily,
1131 ensuring repeatability in the methodology. In a world in which the number of impacts on marine
1132 environments is increasing, and in which assessments sometimes use methodologies like sampling
1133 that damage marine ecosystems, innovative non-intrusive methods such as imagery and 3D models
1134 can be used to better understand protected habitats. Photogrammetry is therefore a good means of
1135 obtaining quantitative data in the deep sea and of better calculating surfaces, densities and size
1136 parameters. However, it is noteworthy that density results correspond to an estimation limited by
1137 several methodological factors. First, the number of colonies counted and the measurement of
1138 surface areas were limited to the areas reconstructed in 3-dimensions. Second, the number of
1139 colonies plotted in 3D corresponded to an approximation of their total number, and this number is
1140 underestimated, especially in 3D models built from video data where the low image resolution
1141 makes it difficult to identify every colony. Third, in sites with dense agglomerations of colonies it was
1142 difficult to distinguish and count individual ones. Additionally, there is a substantial risk of not

1143 plotting colonies of small size due to misidentification. Nonetheless, this methodology allows
1144 comparing quantitative data between sites and monitoring their evolution through time. 3D models
1145 built at regular intervals will allow monitoring vulnerable ecosystems, giving access to distribution,
1146 density and size parameters.

1147

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1158

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