

Mud volcano growth by radial expansion: Examples from onshore Azerbaijan

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Abstract :

We investigated the mode of growth of flat-topped mud volcanoes, through the study of three active edifices onshore Azerbaijan: they are the Bozdag-Guzdek, the Ayazakhtarma and the Akhtarma-Pashaly mud volcanoes. The three edifices are up to 80 m high and 3 km wide, and the recurrence time between eruptions from a few months to a few years. Surface changes during and between eruptions were documented by a combination of mapping from satellite pictures, repeated direct observation over five years and structural analysis. In addition, resistivity profiles and microgravity measurements were used to decipher their subsurface geometries. We interpret the flat-topped character of the mud volcanoes as the result of rapid "isostatic" readjustment of a brittle surface crust, < 1 m to tens of meters thick, overlying a ductile layer. The surface structure typically shows a concentric transition from an extensional regime in the vicinity of the emission center, to strike slip movements in a median ring, to a compressional regime with thrusts, and pop-up blocks or folds in the outer part of the plateau. Both the flat surface of these mud volcanoes and observed radial displacements of the surface, combined with the very low Bouguer anomaly of the Ayazakhtarma mud volcano, give arguments to propose that km-diameter, flat-topped mud volcanoes likely reflect the presence at a shallow depth (a few tens of meters?) of a large volume of soft mud. Rapid compaction at the surface forms a crust that moves away from a central feeding area, thereby transferring mud added at the center into lateral spreading, building a Coulomb prism all around the mud volcano and strongly limiting vertical buildup.

Highlights

► We propose a new model of growth for flat-topped mud volcanoes fed from the center. ► Flat-topped MVs grow by radial expansion, making a concentric structural pattern. ► Fractures evolve from extensional in center to compressional at periphery. ► Gravity inversion shows a ca 300 m thick low-density lens below the surface. ► Soft mud in the lens limits vertical growth and flattens the MV by isostasy.

Keywords : Flat-topped mud-volcanoes, Azerbaijan, Growth of flat mud volcanoes, Resistivity-gravity data, Low-density mud at depth, Mudflows stacking, Soft mud.

38 1. Introduction

39 Mud volcanoes form in sedimentary basins onshore and offshore by localized expulsion of mud
40 remobilized from the subsurface. Sediment mobilization is driven by the migration of gas and liquids resulting
41 from compaction, diagenetic transformation and thermal maturation of sediments (Brown, 1990; Le Pichon et
42 al., 1990; Dimitrov, 2002; Deville et al., 2003). Mud emission is an episodic phenomenon with eruptive phases
43 lasting less than a few days separated by periods of quiescence that typically last several years or decades (Kopf,
44 2002; Deville and Guerlais, 2009; Mazzini et al., 2009). Mud volcanoes, built over hundreds of years or more,
45 have to be distinguished from individual mudflows emplaced during a single eruption. The central part has been
46 determined as the mud expulsion area of many mud volcanoes (Le Pichon et al., 1990; Perez-Garcia et al., 2009).
47 The boundaries of the mud volcanoes have steep slopes dipping up to 20° (Henry et al., 1990). A recent review
48 of mud volcanism by Mazzini and Etiope (2017) identifies a wide range of morphologies, from conical mud
49 volcanoes to flat mud pies; the authors interpreted these morphologies as depending on the type and viscosity of
50 the extruded mud, the activity of the volcano, and the frequency of the eruptions. Conical mud volcanoes grow
51 by emission of successive radial mudflows from a central edifice (Dimitrov, 2002; Deville et al., 2006; Mazzini
52 and Etiope, 2017) whereas the mode of formation of mud pies or plateau-like mud volcanoes is less well
53 constrained (Le Pichon et al., 1990, Dupré et al., 2007). Plateau-like mud volcanoes differ from mud pies by
54 their flat top surfaces and relatively low elevations (Mazzini and Etiope, 2017), the central part being the most
55 active in terms of mud expulsion, as is the case for most types of mud volcanoes (Perez-Garcia et al., 2009). For
56 example, on the Nile deep sea fan (Egyptian continental margin), fresh mud has been recognised in the central
57 part of the flat-topped Amon mud volcano whereas concentric ridges, identified as compressional structures, are
58 observed at the rim of the mud volcano (Dupré et al., 2008). In many plateau-like mud volcanoes, the central
59 area is surrounded by a deformed unit marked by concentric ridges and fractures (Lance et al., 1998; Dupré et
60 al., 2008).

61 In all cases, mud volcanoes expel a mixture of water, gas, homogeneous mud and numerous embedded
62 clasts. The deposited material consists of clasts up to few meters in size (Kopf, 2002; Deville et al., 2003)
63 embedded in a fine-grained matrix and is known as mud breccia (Cita et al., 1981). Each deposit results from one
64 event that is generally referred to as (individual) mudflow (Mazzini and Etiope, 2017). On a mud volcano with a
65 surface slope, extruded mud flows downslope to make a tongue-like radial mudflow, whereas on a flat surface,
66 mud typically spreads away from the emission center to generate sub-circular pie-like mudflows. A feature of
67 interest that can be observed on some pie-like mudflows is the presence of mud bomb impact features, large
68 bowls, 0.3-3m in diameter, that result from the impact of mud clots ejected during the eruption on the already
69 deposited soft mud (Dimitrov, 2002; Aliev et al., 2015). The vertical growth of plateau-like mud volcanoes
70 appears to be limited but the surfaces of many of them show signs of regular activity, such as emissions of fresh
71 mud, active displacements along fractures and active surface deformation (Henry et al., 1990; Dupré et al., 2010;
72 Antonielli et al., 2014).

73 The aim of this study is to document the deformation and growth of plateau like mud volcanoes by
74 studying some particularly flat examples onshore Azerbaijan. We use structural analyses from field and remote
75 mapping, supplemented by resistivity sounding and gravity measurements and analogue modelling to
76 characterize the morphology and subsurface structure of flat-topped mud volcanoes. The results support a new
77 model of flat-topped MVs in which they are permanently fed by fresh mud. Our findings are applicable to other
78 plateau-like MVs in onshore and offshore settings and suggest they are highly active features.

79

80 **2. Geological setting of mud volcanoes in Azerbaijan**

81

82 The Caspian Sea originated as one of a chain of Mesozoic basins that formed along the southern
83 continental margin of Eurasia (Dercourt et al., 1986). From Middle Jurassic to Early Cretaceous time, extension
84 resulted in rifting and formation of an early South Caspian Basin. From Eocene time to the present, convergence
85 between the Arabian and Eurasian plates caused the uplift of the Caucasus region and the separation of the
86 Caspian Sea from the Black Sea (Smith-Rouch, 2006). A large part of the displacement of Arabia has been
87 accommodated by thrusting and folding in the Zagros Mountains and in the Caucasus region since Miocene time
88 (Forte et al., 2010). Starting in the early Pliocene, a compressional regime has generated fold and thrust
89 structures within the basin (Roberts et al., 2010). Large anticlines parallel to the Lower Kura fold-and-thrust belt
90 have an E-W trend in north Azerbaijan (Fig. 1), then curve to N-S in the eastern Kura Basin, offshore Azerbaijan
91 along the west coast of the Caspian Sea (Jackson et al., 2002). The South Caspian Sea is an aseismic block

92 surrounded by belts of intense seismic activity (Jackson et al., 2002), and present-day displacement vectors from
93 GPS measurements indicate a N-to-NE verging displacement rate of about 6-10 mm yr⁻¹ that decreases
94 northwards (Kadirov et al., 2012, 2014).

95 One result of the collisional regime is a significant deepening of the South Caspian Basin, which
96 progressively became isolated from the world ocean during the Miocene (Stewart and Davies, 2006). The total
97 sedimentary thickness reaches up to 26 km in the northern part of the basin, which is marked by a strong
98 negative gravity anomaly (Brunet et al., 2003). The sediment accumulation rate is estimated to be as high as
99 2 km/My in the offshore area during the Plio-Pleistocene (Yusifov and Rabinowitz, 2004; Vincent et al., 2005).
100 The high subsidence rate in the South Caspian Basin led to high sediment accumulation rates (Katz et al., 2000),
101 which in turn cause sediment undercompaction and high fluid pressures, both favoring the formation of mud
102 volcanoes (Dimitrov, 2002). Hundreds of mud volcanoes have been observed in the basin (Jakubov et al., 1971;
103 Stewart and Davies, 2006; Aliev et al., 2015), their formation generally interpreted as the result of rapid loading
104 of Paleogene-Miocene sediments during the Plio-Pleistocene (Jackson et al., 2002). The constituents of the
105 expelled mud (*i.e.*, mud matrix, xenoliths, liquids and gas) typically originate from the Maykop formation (Inan
106 et al., 1997) that has been buried to depths of 4-10 km. In many cases, mud volcanoes align along anticline crests
107 (Roberts et al., 2011, Bonini et al., 2013). Anticlines are interpreted to constitute a favorable location for mud
108 volcanism by acting as focusing points for hydrocarbon migration and as loci of minimum overburden thickness
109 for overpressured reservoirs because of lateral fluid-pressure transfer (Yardley and Swarbrick, 2000; Javanshir et
110 al., 2015). In addition, the high sediment accumulation rates contributed to a low temperature gradient, 14-16
111 °C/km and the top of the gas window is estimated to be at about 13-14 km (Devlin et al., 1999). The same
112 conditions are observed onshore Azerbaijan, where the temperature gradient is 10-18°C/km (Planke et al., 2003).
113 Such a large extent of the gas window in the sedimentary series is a favorable condition for the formation of
114 large volume of gas leading to the formation of mud volcanoes (Fowler et al., 2000).

115

116 3. Data and methods

117

118 We selected three mud volcanoes in Azerbaijan that contain a flat, plateau-like emission zone, and
119 whose recent activity makes it possible to observe structural changes related to successive mud emissions over
120 the past 15 years. These are the top of the Bozdag-Guzdek mud volcano (hereafter referred to as BGMV), the
121 Ayazakhtarma mud volcano (hereafter referred to as AAMV) and the Akhtarma-Pashaly mud volcano (hereafter
122 referred to as APMV). We investigated their structure using a combination of: - structural analyses from remote

123 imaging and field mapping; - resistivity sounding; - gravimetric measurements and modelling; - analogue
124 modelling.

125

126 *3.1. Structural analyses*

127 Our structural analysis consisted of mapping fault traces using drone and satellite pictures combined
128 with field measurements of fault planes and associated small-scale striae. We acquired drone pictures over the
129 mud volcanoes at different heights, from 30 to 500 m above the surface using a DJI-Mavic-Pro® drone aircraft.
130 We constructed 2-D pictures and topographic 3-D models of the studied features from the picture mosaic using
131 the Agisoft® software. We also mapped the displacement fields of the surface of the mud volcanoes by
132 comparing drone pictures acquired in 2018 and satellite pictures available on Google Earth® for the BGMV and
133 APMV volcanoes, and from satellite images of SPOT7 (1.5 m resolution, shot from 2011 to 2017), Pleiades®
134 (0.5 m resolution), as well as drone pictures for the AAMV. On the BGMV, we mapped faults and fractures from
135 satellite and drone images, and their small-scale morphological characteristics were observed directly in the
136 field. On the AAMV, we studied the morphological evolution of extensional fractures in the central area by
137 comparing successive field observations made over the years 2014-2018, and their overall organization from
138 satellite and drone pictures. On both the AAMV and the APMV, pop-up structures were identified on their
139 plateau from topography. Strike-slip faults were identified on the plateau of the AAMV by horizontal striae
140 along vertical planes, in the field. We analyzed these using the geometric-construction method of McCoss (1986)
141 which consists of drawing the displacement vectors from the intersection of the main fracture direction and the
142 direction of minor Riedel T extension fractures. In the field on both the AAMV and the APMV, small-scale
143 reverse faults were identified in the gullies that dissect their flanks from striation and vertical cleavage along
144 fault planes.

145

146 *3.2. Resistivity profiling*

147 We carried out resistivity profiling data using a “LS 2 Terrameter”® following the experimental
148 procedure of Pazdirek and Blaha (1996). The results are provided as synthetic models of electric resistivity.
149 Complete profiles across the AAMV and the crater of the BGMV were acquired with an electrode spacing of 5
150 m. A shorter spacing of 1 m for acquisition of shorter profiles over local structures of the AAMV and APMV
151 volcanoes.

152

153 3.3. Gravimetric data and modelling

154 Inspired by the results of the early work of Feyzullayev et al. (2005), gravity was measured using a
 155 SCINTREX CG5 relative gravimeter. For each mud volcano we revisited some sites in order to control and
 156 remove the instrumental drift. The positions of the gravity stations were determined-by GPS measurements using
 157 dual-frequency Topcon® GB-1000 receiver. The GPS data were analyzed with the GINS software package
 158 developed by the French Centre National d'Etudes Spatiales (CNES), using a Precise Point Positioning (PPP)
 159 technique (Marty et al., 2011). We used the CG5TOOL (Gabalda et al., 2003) and Geosoft Oasis Montaj®
 160 softwares in order to calculate the Free Air and the Complete Bouguer anomalies. We used 0.3086 mGal/m for
 161 the free-air correction and we assumed a density value of 2.5 for the terrain corrections. We estimate a RMS on
 162 the Bouguer anomalies of 0.05 mGal. Results of Bouguer anomalies were compared with a 3-D inversion using
 163 the L_p norm 'compact' inversion scheme described in detail in Fournier et al. (2016) and in Miller et al. (2017)
 164 and implemented in the open source SimPeg framework (Cockett et al., 2015).

165 In a first step, as a first-order approximation, we estimated the mass deficit ΔM required to produce the
 166 observed gravity anomaly using Gauss's theorem:

$$167 \quad \Delta M = \frac{1}{2\pi G} \iint \Delta g dS \quad (1),$$

168 where G is the gravitational constant and Δg the observed gravity signal. Next, we built a 3-D mesh
 169 composed of 55.175 cells 200 x 200 m in the horizontal plane and whose thicknesses increased from 20 to-100
 170 m at the base of the grid set at 1200 m bsl. We consider the following data misfit to measure the goodness of the
 171 fit between the observed and the modeled data:

$$172 \quad \phi_d(\mathbf{m}) = \|W_d (G(\mathbf{m}) - \mathbf{d}_{obs})\|_2^2 \quad (2),$$

173 where W_d is a diagonal matrix with $W_{d_{ii}}$ elements being equal to the inverse of the standard deviation of the i th
 174 gravity observations. Here, we considered a constant standard deviation of 0.05 mGal for all the gravity
 175 measurements. Next, we minimized the following global objective function using the scaled iterative reweighted
 176 least square optimization approach proposed by Fournier et al. (2016) and Miller et al. (2017):

$$177 \quad \phi(\mathbf{m}) = \phi_d(\mathbf{m}) + \beta \phi_m(\mathbf{m}) \quad (3),$$

178 where $\phi_d(\mathbf{m})$ is the data misfit, $\phi_m(\mathbf{m})$ is the stabilizing regularization and β the Tikhonov parameter. We
 179 performed an optimization using a two-step inversion approach, first by inverting the data with a "smooth"
 180 constraint (*i.e.*, β not too small) and next by compacting the resulting smooth model using a norm L_p , where
 181 $0 << L_p << 2$. We did not incorporate depth weighting as we search for solutions that are close to the ground
 182 surface, but we provided strong constraints by fixing a lower density contrast bound to 500 kg/m³. Our data are

183 preliminary results that have to be completed in 2019, but they already provide significant insights to illustrate
184 the 3-D architecture of the AAMV and BGMV mud volcanoes.

185

186 3.4. Analogue modelling

187 In addition, we designed and realized several analogue experiments in order to simulate the injection of
188 soft, viscous material beneath a brittle surface layer. The deeper, soft and viscous mud was modelled by a
189 silicone polymer (XIAMETER[®] manufactured by Dow Corning). The brittle part (corresponding to the
190 uppermost, dried part of the mud) was modelled by dry quartz sand (GA39, sold by SIBELCO[®], France), whose
191 grains had a mean size of about 114 μm in diameter, and a PVC powder (manufactured by SOLVIN[®], France),
192 and whose grains had a mean size of about 100 μm in diameter. The technical reason behind our choice of using
193 a combination of fine sand and PVC was to avoid introducing a density inversion between the upper brittle part
194 and the lower viscous part and thus to avoid any unwanted diapirism effect. The silicone polymer had a density
195 close to 0.95, the sand had a density of about 1.5, whereas the PVC powder had a density close to 0.5. By
196 alternating layers of sand and PVC powder, we were able to avoid a density inversion. The sand and PVC
197 powder had similar angles of internal friction, ranging between 35 and 40°.

198 The purpose of these experiments was not to model exactly the formation and evolution of the
199 Ayazaktharma mud volcano, but rather, as a test of concept on the hypothesis that the mud coming from below
200 would flow sideways, rather than rise above the older mud to build a tall edifice. Because the mud's viscosity is
201 so low, about 10^2 Pa.s, according to Rudolph and Manga (2010), and the silicone we used had a viscosity value
202 about 10^4 Pa.s, in terms of scaling, the strain-rate ratio, and hence the time ratio were very different than those in
203 usual models (*i.e.*, where the silicone simulates rock salt or the lower continental crust). A viscous layer made of
204 either halite or the lower continental crust, or even the asthenosphere, have a much higher viscosity than that of
205 the silicone, about 10^{17} Pa.s to 10^{21} Pa.s. In technical terms, this would have required us to run our experiments
206 for several months, rather than days. Mechanically, the result would have not been much different because the
207 strength contrast between the upper, brittle, dry mud and the lower, viscous mud would have remained extremely
208 strong. We can make a quick estimate of this ratio in nature and in experiments. Assuming an angle of internal
209 friction of about 30 degrees for both the natural and experimental examples, the ratio between the shear stress
210 within the lower viscous lower layer and the differential stress within the upper brittle layer would be about 300
211 in nature, and 67 in the experimental models. In both cases, the viscous stresses within the lower viscous layer
212 would be much lower than the stresses within the upper brittle part. As a result, the main effective parameter that

213 controls the deformation process would be the pressure within the lower layer, rather than its viscous shear
214 stress.

215

216 **4. Results**

217

218 We present results from the summit crater of the BGMV, and from the plateau and flank of both the
219 AAMV and the APMV. These mud volcanoes are located 20 km, 75 km and 90 km west of Baku, respectively
220 (Fig. 1). The crater of the BGMV allows us to show on a relatively small scale (≈ 100 m) the effect of loading a
221 recent, still deformable, mudflow by a younger one. In the light of this observation, we describe and interpret the
222 2-3 km diameter AAMV in terms of the episodic emission of mud from a single site over at least the last 15
223 years. Finally, we show how mud emission from three centers on the APMV can affect the structure of the mud
224 volcano.

225

226 *4.1. The Bozdag-Guzdek mud volcano crater*

227 We focus on the summit crater of The BGMV which is the active top part of a mud volcano set over an
228 eroded irregular substratum, which occasionally crops out between mudflows almost up to the top of the edifice
229 (Fig. 2). The elevation at the top is 300 m above msl but the plateau-like top part of the mud volcano does not
230 exceed just stands a few tens of meters above the nearest substratum outcrop. The elevation at the distal edge of
231 the radial mudflows varies from 50 m in the S part to 150 m NE away from the crater. The mud volcano has a
232 roughly concentric geometry in map view, with a 350-m diameter disc depressed by a few meters located at the
233 top of the edifice, and radial mudflows, the least weathered and most visible of which have moved downhill
234 towards the NE over more than 1.5 km through a breach in the crater. Below we focus on the central part of the
235 crater and its recent evolution.

236

237 *4.1.1. Surface structure*

238 The crater is partly filled by stacked pie-like mudflows (Fig. 3). Comparison of successive satellite
239 photographs of the crater available under 'archive images' in Google Earth® (GE) provides evidence of a
240 succession of mud effusion episodes over the 15 last years, each mud flow overlying the previous one(s), partly
241 or totally (Figs. 4a, b, c). We first visited the BGMV in June, 2014 before extrusion of the flow that appears on
242 the GE image of 25 August 2014. Comparing satellite images available in Google Earth indicates that it was

243 emplaced between 7/8/2013 and 9/10/2013, which means that the 2014 eruption took place 10-13 months after
244 the 2013 one. During a field trip in May 2016, we examined in detail the overprint of the 2014 event on the 2013
245 mudflow (Figs. 4b, b').

246 The 2014 effusion formed a 110-m diameter pie-like mudflow riddled with mud-bomb impacts (Fig.
247 4b). The contact of the 2014 pie-like mudflow and the underlying 2013 mudflow is clearly visible on the ground
248 and on satellite pictures because of changes in color and texture, more than in terms of topographic relief. The
249 underlying 2013 mudflow (Fig. 4a) was deformed by the excess load of the younger mud pie. It is described
250 taking as examples the W and NE sectors that illustrate contrasting types of deformation (Figs. 4d and 4e,
251 respectively). All across the deformed part (dashed red outline in Fig. 4b'), radial extension fractures have
252 accommodated the radial expansion and correlative circumferential stretching of the mudflow. In the western
253 sector, deformation and radial displacement did not reach the edge of the 2013 mudflow; the limit of the
254 deformed zone consisted of small-scale reverse faults indicating radial compression, verging alternatively
255 inwards and outwards (Figs. 4d, d'). In the NE sector (Fig. 4e), the edge of the 2013 mudflow was pushed
256 radially outwards by 2 to 5 m, whereas the older, underlying pie-like mudflow remained non-deformed, as
257 attested by the absence of any displacement of natural passive markers. The outer boundary of the 2013 mudflow
258 is segmented by conjugate, left-lateral and right-lateral strike-slip faults that transfer the compressional
259 movement in the radial direction (Fig. 4e'). The 2013 mudflow in that sector is pushed above the underlying one
260 and the increase in radius was accommodated by both strike-slip faults and radial extensional fractures.

261

262 *4.1.2. Subsurface architecture*

263 We acquired a 640-m-long resistivity profile across the crater of the BGMV (Fig. 5). The resistivity
264 profile shows a generally low contrast. Relatively high resistivity values were measured in the upper 10 m of the
265 SE part of the profile, which corresponds to the outer slope of the mud volcano located beyond the crater's edge
266 (Fig. 5a). Across most of the crater, in the upper 20 m below surface, resistivity is in the range of 1-3 Ω *m. The
267 highest value was reached in the middle part of the bottom of the profile ($>10 \Omega$ *m) and is surrounded by two
268 zones of low resistivity ($< 1.5 \text{ Ohm}^*m$), corresponding to areas of active gryphons in the SE and old structures
269 in the NW. The old structures consist of elongate sub-circular depressions, which could correspond to the
270 boundary of old mud pies or to the limit of old depressed rings. However, because of their eroded morphology it
271 is difficult to be sure.

272 We recorded gravity measurements along the line of the resistivity profile. The Free-Air Anomaly
273 classically follows the trace of the topography section, but the Bouguer anomaly (Fig. 5b) shows negative

274 gravity values in two areas: one in the NW in the area of old structures, and the second in the area of active
275 gryphons, where the anomaly reaches 1 mGal indicating an area of mass deficit under the gryphons.

276

277 4. 2. *The Ayazakhtarma mud volcano*

278

279 4.2.1. *Surface structure and deformation*

280 The Ayazakhtarma mud volcano is a large plateau-like mud volcano that lies at the southeastern
281 termination of the Gidjakiakhtarma anticline (Fig. 6). The regional topography is characterized by NW-SE-
282 trending crests and thalwegs, themselves related to the underlying geology as shown by the geologic map of
283 Azerbaijan (Alizadeh, 2008). The AAMV itself is emplaced between two NW-SE-trending crests located 5km
284 NE and 2 km SW of the mud volcano. The AAMV comprises a 2 x 3 km elliptical plateau surrounded by a flank
285 dipping about 6° outward; the upper part of the flank is dissected by shallow gullies with a 20-to-30 m spacing
286 that merge downslope to reach a 100-120 m spacing where the flank downlaps onto the surrounding topography
287 (Fig. 7). The elevation of the plateau ranges from 549 to 553 m above mean sea level, with the highest point
288 located at the center, giving it an average slope gradient of 4%, i.e. 0.25°. The plateau can be subdivided into
289 two concentric zones: a central area, subcircular and about 500 m in diameter, and a median ring extending from
290 the central area to the edge of the plateau. Eleven eruptions of the AAMV have been reported since 1926, 10 of
291 which occurred after 1969. These eruptions were characterized by the release of large volumes of mud breccia
292 and reported mudflow thicknesses vary from 0.3 to 1 m, and cover 1000 to 70 000 m², the volumes ranging from
293 5 000 to about 400 000 m³ (Aliev et al., 2009). Since we started fieldwork on the AAMV, we observed 6 new
294 mudflow deposits, 30 to 300 m in diameter; the largest of which covered 85 000 m² in 2015. In the autumn of
295 2018, the youngest eruption covered 30000 m², and we estimate from depth soundings of the soft mud
296 underlying it that it produced at least 80 000 m³ of mud breccia.

297 Antonielli et al. (2014) used radar interferometry over the period from October 2003 to November 2005
298 to show that the structure was actively deforming over time scales of a few months. They showed vertical
299 movement of tens of cm over the two year period and observed signs of strike-slip faulting at the surface of the
300 plateau. Lio and Furuya (2018) determined from SAR data over 2006 to 2011 and 2014 to 2017 periods that
301 horizontal displacements are greater than vertical ones.

302

303 4.2.2. *Central area*

304 The central area of the AAMV is characterized by rectilinear-to-sinuuous furrows and a pervasive
305 polygonal network of desiccation structures, a few tens of cm in width (Fig. 8a, 8b) and few small-sized
306 gryphons. The furrows are up to 1.5 m deep and 1-to-2 m wide, and propagate through the desiccation structures;
307 these appear as vertical columns bent on each side of the furrow (Fig. 8 c). The furrows present an overall radial
308 pattern, with the overprint of a N 40 – N 50° trend in the outer part of the area (Fig. 8). The central area is very
309 dynamic with three mudflows between October 2014 and May 2016, with diameters of about 300, 80 and 40 m;
310 successive observations over the course of four years showed that the polygonal fracture pattern and radial
311 furrows record the progressive growth of initially smooth mud flows. Because circular impact craters have never
312 been observed at the surface of the AAMV, soft mud bombs have probably not been ejected from this mud
313 volcano in recent times. The episode of mud extrusion that we observed in May 2016 was particularly quiet with
314 neither mud bombs nor gas combustion, which supports the hypothesis of continuous mud seepage, rather than
315 violent expulsion.

316 Map views from a July 2017 satellite photograph, and from our own drone data acquired in May 2018,
317 were superimposed in order to quantify the displacements on the large mudflow emitted in 2015 in the central
318 area (Fig. 9). The cross point of the furrows was chosen as a fixed reference. The mudflow extruded in May
319 2016 and that was simply translated but not deformed in May 2018, as we have checked from field observations,
320 give assurance that the two pictures had the same scale. From 2017 to 2018, the traces of the furrows widened,
321 and some bifurcation points were displaced (Fig. 9). The expansion of the 300-m-diameter 2015 mudflow
322 reached 17.8 m/yr in the NW-SE direction, versus only 8.6 m/yr in the NE-SW direction (Fig. 9). The trend of
323 the maximum extension direction was perpendicular to that of the main fracture / furrow direction, as shown by
324 the rose diagram of figure 9, derived from 340 measurements.

325

326 4.2.3. Median ring

327 Two main patterns of lineaments formed within the median ring: dominant oblique segments, and
328 subordinate concentric segments (Figs. 7 & 8). Oblique segments vary in length from a few meters to more than
329 500 m. Antonielli et al. (2014) identified from field observations one of these trends as a strike-slip fault. Our
330 observations show that most of them are arranged as conjugate pairs symmetric with respect to the radial line at
331 their intersection with horizontal striae in the muddy material of many of these lineaments (Fig. 8d). Conjugate
332 left-lateral and right-lateral pairs define triangular wedges that moved away from the central zone in a radial,
333 centrifugal fashion. Most of these oblique segments show *en échelon* T-Riedel shear zones (Fig. 8e). Applying
334 the simple geometric construction of McCoss (1986) to these oblique segments reveals a combination of strike-

335 slip motion and extensional component, i.e. transtension along both the left-lateral and right-lateral fractures.
336 Such an extensional component along the strike-slip fault is an indication of the widespread circumferential
337 tensile strain in the median ring (Fig. 8e).

338 Concentric segments are commonly organized as parallel pairs topographically expressed on the ground
339 as ridges flanked by two lows 8 to 12 m apart (Fig. 8f). The relief of the ridges with respect to the average
340 altitude of the plateau around is up to 2 m; the lows flanking them are asymmetric, steeper on the flank of the
341 ridge (25 to 30°) and gentler on the opposite side (5 to 10°). The depressions are commonly deep and regular
342 enough to retain rainwater, making pairs of “twin lakes”. These concentric segments are more abundant in the
343 two crescent-shaped areas located at the SE and WNW edges of the elliptical plateau, beyond the subcircular
344 median ring.

345 At the scale of the median ring, displacement vectors drawn by superimposing SPOT images from 2011
346 to 2017 show centrifugal displacements (Fig. 10) dominated by trends toward the SE, with magnitudes up to
347 32 m away from the center of the volcano at the edge of the central area.

348

349 *4.2.4. Outer flank of the mud volcano*

350 The outer flank is dissected by a set of gullies radiating away from the mud volcano (Fig. 8g). These
351 gullies merge downslope, preferentially at the level of two intermediate terraces at elevations of 530 and 540 m,
352 respectively. Most of the W-verging gully flanks are vegetated, whereas the E-verging flanks are bare, poorly
353 vegetated, and most clearly subjected to erosion. In some instances, freshly rejuvenated incision surfaces expose
354 at outcrop mud breccia with indications of reverse faulting (Fig. 8h). The best-preserved fault planes show striae
355 and vertical cleavage that indicate radial compression both on forethrusts dipping towards the plateau, and on
356 backthrusts dipping away from the plateau (Fig. 8h). Such small-scale reverse faults are present all around the
357 AAMV (Fig. 8, map) with strikes of fault plane and cleavage oriented orthoradially and of striae radially.

358

359 *4.2.5. Subsurface architecture*

360 We acquired a 2150-m-long resistivity profile across the NE-SW diameter of the volcano in order to
361 characterize the subsurface architecture, and a 250-m long one in order to characterize the subsurface
362 architecture of the twin lakes (Fig. 11). The long resistivity profile shows a contrast between a less than 10-m-
363 thick surface layer, whose resistivity ranges from 3 to 5 Ω *m, and a deeper layer whose base is not detectable on
364 the inverted profile which extends to 40 m depth (Fig. 11a). The thickness of the resistive layer ranges from
365 about 5 m in the central area to 8 or 9 meters in the median ring, the lowest value of 3 m recorded near the small

366 gryphons located at the center of the structure. The resistivity of the underlying low-resistivity body varies from
367 $2 \Omega \cdot \text{m}$ in the central part, to $3 \Omega \cdot \text{m}$ at the edge of the plateau. The resistivity profile of the surface crust is
368 segmented with more resistive slabs ($4\text{-}5 \Omega \cdot \text{m}$), about 50 m wide, alternating with narrower (10-25 m wide)
369 zones whose resistivity varies from 1.5 to $3 \Omega \cdot \text{m}$. The low-resistivity domains generally correspond to
370 topographic lows. The resistivity section along the flank is less regular than that on the plateau, with patches of
371 low resistivity near the bottom of the slope, in close proximity to permanent water bodies at the surface.

372 The shorter resistivity profile acquired across two pairs of twin lakes (Fig. 11b) shows a 7-to-10-m-
373 thick interval having a resistivity of $3.1\text{-}4 \Omega \cdot \text{m}$ overlying a $2.5 \Omega \cdot \text{m}$ material. The resistivity contrast between
374 the surface and deeper intervals decreases at the location of the twin lakes, with a value below $3 \Omega \cdot \text{m}$ in the
375 shallow part, smoothly decreasing downward over 25 m to the normal background value of $2.5 \Omega \cdot \text{m}$. The long
376 profile confirms that each twin lake is an area of low resistivity.

377 We recorded gravity measurements along two lines: one lies along the long resistivity profile and the
378 other is perpendicular to it, extending from the center to the NW edge of the mud volcano. In both cases, the
379 Free Air Anomaly does not follow the profile of the topography section, and the Bouguer anomaly shows
380 negative gravity values, down to -6 mGal/km , all across the AAMV (Fig. 12a). As a first approximation, we
381 estimate the mass deficit through Gauss's theorem by integrating the gravity signal across an area of 5.5 km^2
382 yielding a mass deficit of about $35 \times 10^{10} \text{ kg}$. This value must be treated with caution because of the limited
383 spatial extent of the gravity measurements; however, it provides a lower bound for the mass anomaly required to
384 produce the 6 mGal anomaly observed on the AAMV.

385 Gravity inversion was constrained to find a compact body near the surface by incorporating no depth
386 weighting and fixing a density contrast ranging from 0 to 500 kg/m^3 . The presence of a compact body near the
387 surface is supported by observations of a mobile and rapidly changing geometry of the surface. Figure 12b
388 shows a conceptual model explaining the gravity anomaly observed at the AAMV, with a body volume of
389 $11 \times 10^8 \text{ m}^3$ and a mass deficit of $35 \times 10^{10} \text{ kg}$. The Root Mean Square (RMS) of the residuals between the observed
390 and the modeled data is 0.018 mGal , and the distribution of these residuals is consistent with the stochastic error.

391

392 4. 3. *The Akhtarma-Pashaly plateau and flank*

393 The crater of the APMV is located on the Kalamaddin anticline (Aliiev et al., 2015), which is located
394 near the Caucasus thrust front (Forte et al., 2010). It is composed of three concentric zones: a $2.6 \times 2 \text{ km}$
395 diameter plateau (Figs. 13 and 14); a steep upper flank dipping $9\text{-}12^\circ$ over distances of 600 to 1000 m from the

396 edge of the plateau and in the SW quadrant of the APMV, a 2°-dipping lower flank made of vegetated radial
397 mudflows that extends 6 km away from the edge of the plateau into the Kura river floodplain. The volume of the
398 mud breccia of the Akhtarma-Pashaly is estimated to be $16 \times 10^9 \text{ m}^3$ (Aliiev et al., 2015). Our study focused on the
399 plateau and the steep upper flank.

400

401 4.3.1. Surface structure

402 The elevation of the plateau ranges from 300 m along the edge to 306 m in the center, giving the plateau
403 a maximum slope gradient of 6‰, i.e. an average slope $<0.35^\circ$. The plateau includes three active mud emission
404 zones: at the center (Central Emission Zone, CEZ), 450 m to the NW (NWEZ), and 600 m to the E (EEZ); in
405 addition, a vegetated, inactive emission center is present 350 m SW from the center (SWEZ, Fig. 14). Both the
406 NW and E emission zones show evidence of recent activity (2016 and 2017); the CEZ erupted in early 2018 and
407 we visited it the following May, allowing observations of the pristine deformation structures described below.

408

409 4.3.2. Central emission zone

410 The CEZ is made of a stack of pie-like mudflows. The most recent mud flow, emitted on 27 April 2018,
411 partly covers two older mudflows (Figs. 14 and 15a). Its surface is flat and marked by gentle concentric ridges
412 (“growth rings”) and by 0.3-to-2 m-diameter, subcircular, impact craters (ascribed to soft “mud bombs”), but
413 without evidence for gas ignition such as terracotta-like ground or sinter cones (Hovland et al., 1997). In
414 contrast, the older mudflows have been strongly deformed. To the NW, NE and SE a combination of peripheral
415 thrusts and conjugate strike-slip fractures are present, all of them indicating radial compression (Fig. 15b and d
416 to f). To the SW regular undulations are present around the most recent mudflow (Fig. 15c). The fold axes have a
417 circular trend with a wavelength about 5 m. Their morphology is similar to that of a folded cover detaching
418 above a shallow *décollement*, but the frontal thrust does not crop out in a distal position.

419

420 4.3.3. Plateau

421 The plateau of the APMV has a concentric structure, with a central 1300-m diameter vegetated structure
422 marked by large strike-slip fractures around the group of mud pies associated with the CEZ, the NWEZ and the
423 SWEZ (Fig. 16). The eastern emission zone lies beyond this central structure. The shortening direction is
424 everywhere radial with respect to the center of the CEZ, which is compatible with the strike-slip displacement of
425 fractures. An outer area with multiple lakes extends beyond the zone of strike-slip fractures. All the lakes are
426 elongate and set in a concentric pattern roughly centered on the CEZ, but it has not been possible to clearly

427 recognize twin lakes like the ones on the AAMV. Mudflows of the EEZ cover the eastern part of the lake-
428 bearing zone.

429

430 4.3.4. Steep upper flank

431 The outer flank is dissected by a few steep-sided gullies radiating away from the mud volcano (Figs. 14
432 and 16a). Some freshly incised gullies show outcrops of mud breccia affected by small scale reverse faulting.
433 The best preserved shear planes show striae and cleavage (Figs. 16b to 16d) that record the direction and sense
434 of shortening. A radial compression direction can be recognised all around the APMV from sub-vertical
435 cleavage and small-scale reverse faults. In some places, both radial and E-W compression can be observed (Fig.
436 16 a). Then, the pattern of these compressional structures is slightly different from that of the AAMV where only
437 radial direction is observed.

438

439 4.3.5. Subsurface architecture

440 We acquired a 460-m-long resistivity profile in May, 2017 oriented N-S across the EEZ (Figs. 14 and
441 17), where a pie-like mudflow emitted on 26 January 2016 covered two adjacent older mudflows visible on a
442 2004 Google Earth satellite image. Except for the north part of the profile acquired in an area of small lakes
443 having a very low resistivity (less than $1.4 \Omega \cdot \text{m}$), the resistivity profile shows contrasts between a less-than-10-
444 m-thick surface layer, whose resistivity exceeds $2 \Omega \cdot \text{m}$, a mid-depth zone, whose resistivity is less than $1.8 \Omega \cdot \text{m}$
445 below the 2016 mudflow and a deeper part, more than 20 m below the surface, where the 2016 emission center is
446 underlain by a 75-m-diameter low-resistivity body (Fig. 17).

447

448 5. Summary and kinematic models

449 We summarize the results for each mud volcano in terms of a kinematic model of their functioning, as
450 shown in the sketches of Figs 18-20. These are presented from the simplest, the BGMV, to the most complex,
451 the APMV.

452

453 5.1. The BGMV: kinematic model based on the deformation pattern due to the 2014 mud pie effusion

454 The depressed crater of the BGMV has been filled over recent years by a succession of pie-like
455 mudflows resulting from mud emission accompanied by ejection of “mud bombs” from the active center. When
456 the area of a mudflow exceeded that of its predecessor (*e.g.*, the 2013 mudflow on top of the pre-2010 one, Fig.
457 4a), it blanketed most, if not all, of the older one; in contrast, a mudflow smaller than its predecessor generally

458 overlies the center of the latter (Fig. 4b'). When the recurrence time between eruptions was short (e.g. less than 1
459 year between the 2013 and 2014 eruptions), the absence of topographic relief at the edge of the younger flow
460 indicates that the mud load was compensated by lateral viscous flow below, beneath a relatively thin upper crust.
461 Thus the successive deposition of mudflows on the BGMV did not form a true conical mud volcano because
462 older mudflows were not completely consolidated when younger ones were emplaced. We interpret this behavior
463 as resulting from the high water content of mud of low density and viscosity, so that the mud volcano is not
464 mechanically supported below and spreads out laterally, rather than growing up vertically.

465 The contrast between the W and NE sectors of the crater of the BGMV reflects the length of the radius
466 over which the 2013 mudflow was squeezed and deformed by the 2014 one. In contrast, the 2013 mudflow,
467 which was emplaced about 4 years after the previous one (2009), has a 30-to-50-cm-high frontal step all around,
468 probably because the 2009 mudflow had completely dried out and was mechanically stronger.

469 As a result (Fig. 18), a mudflow can deform the previous one by radially pushing its edges, creating
470 small-scale reverse faults or conjugate strike-slip faults, depending on the potential for *décollement* and shearing
471 deformation within the lower mudflow. Sometimes a larger mud extrusion bypasses the edges of the crater and
472 forms a tongue-shaped mud flow along the sides of the mud volcano, to the NE for the more recent mud flows.

473

474 5.2. The AAMV: kinematic model based on strike-slip faults, twin lakes and flank thrusts

475 The surface structural pattern of the AAMV appears to result from radial spreading, corresponding to a
476 compressional outward component previously recognised by Dupré et al. (2008) on offshore mud volcanoes. We
477 interpret the driving force of all displacements outward from the center of the AAMV to be the regular injection
478 of new mud beneath the central area, as attested by three mud flows extruded to the surface since October 2014.
479 The driving force for radial spreading may be purely gravitational, as the local increase in height of the plateau
480 observed by Antonielli et al. (2014) and Lio and Furuya (2018) is transformed into lateral flow. A consequence
481 is an increase through time of both the surface area and periphery of the mud volcano. The extensional
482 component associated with the *en échelon* T-Riedel shear zones visible in the median ring of the AAMV is the
483 expression of this peripheral extensional strain.

484 The morphology of the twin lakes can be explained by a process similar to isostasy, but at a small scale,
485 via a 10-m thick crust overlying a weak, mobile, low-viscosity material. The twin lakes formed by centrifugal
486 compression and, in sedimentary formations, this kind of deformation is known to generate critical Coulomb
487 wedges (Dahlen, 1984). The cross-section of the AAMV is long and flat. Such an elongate wedge shape is
488 compatible with a very low basal shear stress (Dahlen, 1984). The deformation pattern of the AAMV is argued

489 to be driven by continuous injection of mud at depth which turns into a large spreading mud volcano. The
490 similarity between the expected depth of the low-viscosity material obtained from the geometry of the pop-up
491 blocks and the thickness of the higher-resistivity layer observed on modeled resistivity profiles leads us to
492 interpret this upper, more resistive layer as the brittle layer and the more conductive underburden as the low-
493 viscosity material.

494 The brittle behavior observed at the surface, together with the evidence from resistivity and gravity data,
495 lead us to interpret the AAMV as comprising a thick (several hundred meters) flat-topped lens of ductile, mud
496 capped by a 10-m-thick brittle crust all across the plateau. The brittle crust moves away from the central area,
497 currently mainly towards the SE. On the basis of these observations, we propose a model of lateral spreading of
498 the AAMV (Fig. 19a), starting in the central area with the regular injection of mud from a depth of at least
499 several hundred meters below the surface. The central part of the AAMV is regularly fed by fresh mud that
500 causes the extension of the central part in the NW-SE direction under rates that can reach 17 m/yr (Figs. 9 and
501 19b). This radial displacement is transmitted to the median ring along conjugate strike-slip faults, with
502 displacement rates of about 5 m/yr (Figs. 10 and 19c). In the distal part, the displacement is transmitted to the
503 twin lakes (Fig. 19 d). The forethrusts and backthrusts present in the gullies that dissect the flank, along with the
504 vertical cleavage locally observed in the mud breccia, all indicate radial compression.

505 From the center to the flank of the AAMV, surface structures appear to be controlled by the elliptical
506 geometry of the mud volcano, which can be related to a progressive increase in mud strength (difference
507 between σ_1 and σ_3). The result is a switch of the principal vertical stresses around the lithostatic pressure, ρgz ,
508 representing gravity (Fig. 19e). At the center $\rho gz = \sigma_1$, and the deformation is that of mode I extensional
509 fractures. In the median ring $\rho gz = \sigma_2$, and strike-slip faults form. Toward the outer plateau and the flank $\rho gz =$
510 σ_3 , and the deformation is compressional (σ_1 horizontal) with reverse faults, hence there is an increase in lateral
511 confinement toward the periphery. The fault pattern is Andersonian (Anderson, 1951), and appears to be
512 controlled by the progressive porosity reduction of the surface mud from the center towards the flank due to
513 evaporation.

514

515 *5.3. The APMV: kinematic model based on emplacement of successive mudflows*

516 Our observations lead us to consider the Akhtarma-Pashaly mud volcano as composed of individual
517 mud pies nested within a larger spreading structure (Fig. 20). The three recently active emission zones (NW,
518 Center, E), and an older one to the SW, each corresponding to simple mud pies, comparable to the one that fills
519 the crater of the BGMV. The mud pies formed by emission of successive pie-like flows in their central part, with

520 ejection of some mud bombs. Each pie-like mudflow has deformed the underlying one. The low relief (Fig. 14a)
521 suggests that there is a continuous isostatic-like vertical readjustment above a weaker, ductile layer that underlies
522 the whole plateau. The position of this mud body must be very shallow, less than a few tens of meters, i.e. much
523 closer to the surface than inferred in many other known mud volcanoes (Praeg et al., 2009; Kirkham et al., 2017;
524 Dupuis et al., 2019).

525 At larger scale, the volume added at each of the emission zones causes radial compression, as observed
526 for the AAMV. In the AAPV, evidence of radial compression is provided by the concentric rings affected by
527 strike-slip faults and bounded by narrow troughs where concentric lakes have settled. The peripheral flanks are
528 structured by small-scale reverse faults indicating compression, as along the flank of the AAMV. The main
529 difference is that, whereas the AAMV is fed from a single emission area at the center, the APMV is fed from
530 multiple emission centers. Thus spreading of the APMV could have had different directions over time, resulting
531 in more complex structures due to the superimposition of the deformation pattern of each mud pie combined
532 with their changing locations over time. In addition, some indications of a westward compression are sometimes
533 superimposed on the centrifugal compression.

534

535 *5. 4. Comparison with a simple analogue model*

536 We carried out a series of simple analogue experiments, two of which are illustrated in Figure 21, by
537 injecting viscous silicone beneath a brittle cover made of layers of dry quartz sand and PVC powder (Dupuis,
538 2017). Figures 21A1 and 21A2 are a schematic cross section and a sketch of the top view of the experimental
539 setup, respectively. The rate of silicone injection was 2 cm/h in the center of the table. Results show that the
540 center of the model was uplifted by no more than a few mm, with the highest point located right above the
541 injection tube. The two models differed in their degree of lateral confinement by a pack of dry sand, which
542 represents the dried mud at the periphery of the flat mud volcano. The first model (Figs 21 B & C) was only
543 partly confined (free boundaries to the NW and S) and deformed essentially by radial extensional fractures
544 around the injection point. The second model was entirely confined laterally (Figures 21 D & E) and first
545 underwent extreme extension right above the injection point (the brittle cover thinned to oblivion so that silicone
546 outcropped at the surface), as well as radial shortening marked by concentric folds around the injection point.
547 During a second stage, once the silicone had inflated and after several episodes of sand sedimentation, the outer
548 wall of the pie was breached by peripheral reverse faults, thus allowing blocks of the cover to move radially
549 away from the center. Because these blocks moved independently from each other, strike-slip formed between
550 them. This succession of events corresponds to the one we propose in nature for large, flat mud volcanoes, such

551 as AAMV and APMV. In these plateau-like mud volcanoes, the vertical growth, up to few tens of meters, is
552 significantly lower than horizontal spreading which is few kilometers.

553

554 **6. Discussion**

555

556 *6. 1. Origin of the elliptical shape of the Ayazakhtarma MV*

557 The elliptical shape of the AAMV, which is 2 x 3 km plateau with a long axis in the NW-SE direction,
558 is interpreted to result from the main centrifugal spreading of the mud in the central area (Fig. 9) progressively
559 evolving to a NW-SE direction in the median ring structures (Fig. 10). The NE-verging, present-day
560 displacement measured by GPS (Kadirov et al., 2014) is three orders of magnitude lower than extensional
561 displacements at the surface of the mud volcano. The median ring structures of the AAMV are displaced at
562 velocities up to several m/yr in a NW-SE direction (our results, and those of Lio and Furuya, 2018) whereas GPS
563 displacements in the perpendicular direction does not exceed 15 mm/yr (Kadirov et al., 2014). This means that
564 the elliptical geometry of the AAMV appears to be, to a first approximation, the result of an anisotropic
565 orientation of displacements. Nevertheless, the NE direction of the furrows is parallel to the present day GPS
566 displacements and to the compression direction recorded by numerous anticlines in the area. In the same way, all
567 recent mud extrusion zones trend SW-NE. This is in good agreement with the analysis of Viola et al. (2005),
568 Bonini and Mazzarini (2010) and Bonini (2012) showing that mud volcanoes are potential indicators of regional
569 stress. Furthermore, the AAMV is emplaced between two NW-SE-trending crests and it lies in trend with
570 thalwegs on either side of the mud volcano close to its long axis, suggesting that the AAMV filled part of an
571 initially straight thalweg that now circumvents the edifice and the downstream flow direction is to the NW. The
572 trends of the landscape do result from the active compression toward the NE, combined with differential erosion
573 reflecting lithology contrasts. However, we consider the ellipticity of the edifice as a more likely consequence of
574 the interaction between mud emission and the preexisting topography, with the long axis roughly parallel to the
575 local structural trend. On the other hand, vertical drawdown has never been observed at the periphery of the
576 AAMV, nor is there any evidence for rim faults; as a result it is difficult to either confirm or refute the
577 assumption that the elliptical shape of the AAMV is related to a caldera rim proposed by Bonini and Mazzarini
578 (2010).

579

580 *6. 2. Recurrence time of mud extrusion*

581 Mud pies build up over time through stacking of individual mudflows from the same emission center.
582 Typical recurrence times for mud volcanoes are estimated to range from a few months to over 10 years, when
583 periods of quiescence account for 95% of the life of a mud volcano (Kopf, 2002; Mazzini et al., 2009). A major
584 uncertainty is the lifetime of a given emission center. Have the four emission centers visible on the APMV been
585 active since the onset of mud volcano activity? If this is the case, the “single large edifice” would result from the
586 coalescence over time of mud emitted through closely-spaced vents. Or did emission centers shift randomly
587 within the larger edifice, creating a homogeneous crust overlying a body of very soft mud? In other words, is the
588 currently inactive SW emission zone of the APMV just a “lazy” center, or is it now abandoned for good, while
589 another one is ready to form in another area of the plateau? Changes over time of the location of the expansion
590 center of the mud volcano may better account for the spatial variability of directions of compression than a fixed
591 framework of emission centers. The recurrence time between eruptions also has consequences for geometry,
592 frequent eruptions favoring deformation under the increase of load, whereas the mud volcanoes is more likely to
593 remain rigid and non-deformed in the case of less frequent extrusions.

594

595 *6. 3. Mud injections, extrusions and bombs*

596 Little information is available about the style of the eruptions. One of the most obvious features is the
597 presence or absence of mud bomb impact craters. This is an important issue because a lack of small craters may
598 indicate the quiet extrusion of soft mud like that we observed on a few days after 24 May 2016 at the surface of
599 the AAMV. The presence of mud bomb craters is an indication of violent ejection of mud fragments and reveals
600 the action of gas during the eruption, which in some cases is also indicated by a brown area in the central part of
601 the mud pie due to gas burning. Such eruptions may be the result of mud injection at depth and extrusion of mud
602 and gas at the surface. In some very violent eruptions, mud bombs are thrown far beyond the boundary of the
603 mud pie, examples are known in recent times, e.g. the Lokbatan (Kadirov and Mukhtarov, 2004; Bonini, 2012).

604

605 *6. 4. Pop-up structures of the twin lakes*

606 Except for the morphological symmetry observed Fig. 8f, which leads us to interpret these isolated
607 compressional structures as pop-ups (Butler, 1982) generated by the radial expansion, we have no direct
608 evidence for faults bounding the inter-lake ridge. But we suspect that they do exist because the observed
609 geometry fits that of the pop-up structures described by Treloar et al. (1992) that comprise conjugate reverse
610 faults formed in a brittle layer above a viscous layer. This is confirmed by the resistivity profiles where low

611 values at depth are restricted to twin lakes areas. There, the resistivity value is decreased by the presence of
612 water in the lakes, and we argue it to be extended at depth by downward flow of water along the faults affecting
613 the brittle surface layer.

614

615 *6. 5. Geometry of low density body (Gravity data)*

616 The gravity model proposed in figure 12b is one of the possible configurations based on field data. In all
617 cases the gravity model provides an estimation of the mass deficit of 35×10^{10} kg. On the basis of a density
618 contrast of 0.5, corresponding to measurements of Dupuis (2017), this mass deficit corresponds to a body
619 volume of 11×10^8 m³. Reducing the density contrast would lead to an increase in the volume. The geometry of
620 the low-density body is unknown. The decrease in the Bouguer anomaly of Fig. 12 could be modelled by a
621 density contrast concentrated in the middle of the mud volcano and with a larger vertical extent. However,
622 because the AAMV deforms following mud emissions from the central area to the outer flank, we theorize that it
623 is supported by a flat mud body in a surface position. A more concentrated body in the center of the mud
624 volcano, or a mud chamber at shallow depth, would lead to a different deformation pattern. Nevertheless, the
625 progressive variation of the density contrast of the figure 12b is assumed rather than supported by data. Another
626 limitation of the model is the elliptical geometry of the proposed volume, probably due to the 50m measurement
627 steps of the data of our profile, whereas a shorter measurement step would have given a more accurate geometry
628 of the mud body.

629

630 *6.6. vertical vs. lateral growth*

631 The 2013-2014 pair of pie-like mudflows in the crater of the BGMV exemplifies how the local addition
632 of material deforms a soft-cored mudflow and distributes the added volume over a wider area. This means that
633 volume addition is split between limited lifting of the mudflow surface (faulted western sector, fig. 4d-d') and
634 lateral displacement (NE sector, Figs 4e-e'). At the scale of the complete mud volcano, vertical growth is limited
635 by the emission of radial mudflows that occasionally breach the rim of the crater. The crater is a record of the
636 stability of the mud supply somewhere deep under the mud volcano.

637 The flat and continually active AAMV can be interpreted in light of the deformations observed on the
638 BGMV. A brittle crust develops immediately after emission, and is rapidly deformed by the development of
639 furrows and desiccation-like structures, thereby causing spreading of the central area. Within the median ring,
640 conjugate strike-slip faults allow for outward spreading, creating the pop-up structures bounding the twin lakes
641 located near the peripheral rim. Finally, the outer flank is affected by conjugate reverse faults. Regular spreading

642 of the mud volcano forces an increase in its surface area and circumference. The extensional component of the
643 *en échelon* T-Riedel shear zones in the median ring is the expression of the increase in perimeter. The AAMV is
644 interpreted to build up and spread out the low mechanical resistance of the mud at depth. The lateral
645 displacement rate measured from superimposing satellite and drone photographs is up to 17 m/yr in the center,
646 about 5m/yr in the median ring and decreases towards the flank. Gravity data support the interpretation that the
647 surface of the mud volcano lies above a large and thick (up to 500m) volume of low-density mud. Lateral
648 spreading of this soft mud facilitates lateral movements of the rigid surface of the mud volcano in a radial
649 direction.

650 The AAMV is composed of individual mud pies imbricated within a larger spreading structure. The
651 individual mud pies are built by the successive emissions of mud in the central part of each mud pie, with
652 ejection of some mud bombs. At a larger scale, the flank of the APMV is structured by small-scale reverse faults
653 that record both radial compression caused by lateral spreading and local westward compression probably due to
654 the Caucasus thrust front. Changes in location of different expansion centers through time and their interaction
655 with local tectonics may have led to the more complex geometry of the structure.

656 These three mud volcanoes are flat but continually active. They have grown by the successive stacking
657 of mud pies pushed by gas, like in the progressive growth of the LUSI mud volcano (Istadi et al., 2009). The role
658 of gas on the mud volcanoes studied here is documented by the occurrence in places of mud bomb impacts and
659 reported historical occurrence of gas combustion during eruptions (see list in Aliyev et al., 2015). Our results of
660 resistivity and gravity measurements, and the geometries of structures such as the twin lakes, all lead us to
661 conclude that flat mud volcanoes build up primarily by radial expansion rather than significant vertical growth.
662 The centrifugal displacement observed in all of these mud volcanoes is possible because a brittle upper part can
663 move over a thick layer of soft mud (some hundreds of meters). The thickness of the brittle mud crust varies
664 from a few meters in the central areas to a few tens of meters when approaching the edge of the features.

665

666 6. 7. *Submarine versus subaerial formation of flat-topped mud volcanoes*

667 Mud volcano morphologies with conjugate logarithmic spirals similar to those of the AAMV have been
668 reported in a number of publications: Dupré et al. (2008); Feseker et al. (2010); Pierre et al. (2014). These
669 examples are located in marine environments, where desiccation does not take place, and where the internal
670 friction angle is reduced to near 0° . Another difference from marine environments is that subaerial environments
671 are much more susceptible to erosion. Most submarine mud volcanoes record a competition between local mud
672 emission and distributed sedimentation (Loncke et al., 2004; Blouin et al., 2019). In contrast, onshore mud

673 volcanoes undergo continuous erosion punctuated by occasional emission of mud, but their flat tops appear to be
674 confined by more or less rigid boundaries. The BGMV, the AAMV and the APMV are among the few onshore
675 examples of radially expanding mud volcanoes that have not been degraded by erosion. Despite the differences
676 in setting, these onshore edifices seem comparable to offshore examples. Regardless of the location of the mud
677 volcano, the water content of the mud combined with the short period between successive mud extrusions appear
678 to be key parameters that control the behavior of the mud volcanoes as a whole, and their patterns of structural
679 organization.

680 Among the submarine mud volcanoes studied by Dupré et al. (2008), Huguen et al., 2009, Feseker et al.
681 (2010) and Pierre et al. (2014) in the Eastern Mediterranean, the Cheops MV shows a cohesive behavior in the
682 top part, evidenced by central relief (Huguen et al., 2009) and ductile deformation (Dupré et al., 2014). A
683 temperature log recorded over 449 m evidences a two-layer structure with an upper 18-m-thick thermally
684 stratified interval overlying a 431-m-thick thermally homogeneous interval indicating active convection of brine
685 (Dupré et al., 2014). The top is colonized in many places by microbial filaments (Huguen et al., 2009; Pierre et
686 al., 2014), which may provide some cohesion to the surface. Other possible causes of brittle or cohesive vs.
687 fluidal behavior in submarine settings would include hydrate crystallization (not applicable in the case of the
688 Mediterranean due to the thermal regime) and precipitation of seep carbonates at the sulfate-methane interface as
689 it migrates from top down after a methane-saturated mud body is emplaced at the seabed (Pierre et al., 2014).

690

691 **8. Conclusions**

692 We have studied the top part of three examples of flat-topped mud volcanoes. Mud volcanoes of this
693 type are characterized by a plateau-like topography (overall slope gradient $< 2 \text{ ‰}$, i.e. angle ca. 0.1°). We
694 observe these mud volcanoes to have a brittle surface layer whose deformation by faulting records outward
695 expansion and compression near the edge of the plateau. They typically show a threefold zonation, consisting of:
696 1) a central part surrounding the main mud emission area, corresponding to the most recent pie-like mudflows,
697 which may show evidence of extension; 2) a median zone characterized by strike-slip faults making conjugate
698 logarithmic spirals showing axial symmetry with respect to the emission center; and 3) a distal zone, covering
699 the outer part of the plateau and possibly the upper flanks, marked by the presence of thrusts indicating radial
700 compression.

701 The three mud volcanoes studied here highlight the mode of growth of very flat mud volcanoes, as follows:

- 702 • They grow principally by radial expansion; vertical growth due to the emplacement of an
703 individual mudflow is distributed over the whole plateau, irrespective of the surface actually
704 covered by the mudflow.
- 705 • The presence of a ductile subsurface body below the plateau is a prerequisite; it ensures that
706 overloading by a mudflow is accommodated by local subsidence through an isostasy-like
707 process, so that the surface remains flat whatever is added at the top over time.
- 708 • Radial expansion is accommodated by the development of a Coulomb wedge corresponding
709 to the flanks of the mud volcano and some limited vertical growth distributed over the entire
710 surface overlying the ductile lens
- 711 • In subaerial conditions, the formation of the brittle upper layer is likely to result from
712 dewatering and porosity reduction by surface evaporation. Published examples of submarine
713 mud volcanoes showing a similar flat geometry and surface lineament pattern suggest that
714 other processes may lead in submarine settings to the formation of a cohesive material
715 overlying a ductile mud body.

716

717

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923
924
925
926 **FIGURE CAPTIONS**

927
928 Figure 1. Schematic geological map of the Absheron Peninsula in Azerbaijan showing the location (marked by
929 stars) of the three studied mud volcanoes. Most of the structures strike NW-SE, in trend with Greater Caucasus
930 and Absheron Peninsula . Redrawn from Alizadeh (2008).

931
932 Figure 2. Geological map of the BGMV. The mud pies and mudflows overlie an eroded substratum that
933 occasionally crops out between mudflows almost up to the top of the edifice. Redrawn from Aliev et al. (2015).

934
935 Figure 3. Drone picture of the BGMV crater (May 2018). The white line is the trace of both the 640 m-long
936 resistivity profile acquired in 2017 and the 2018 gravity profile (illustrated in Fig. 5). The infill of the crater is a

937 stack of pie-like mudflows. “B” indicates the breach in the crater rim through which large tongue-like mud flows
938 spilled over towards the N and NE.

939

940

941 Figure 4. Aerial views of the BGMV crater and line drawing of successive pie-like mudflows; a) outline of the
942 2013 mudflow (Google Earth satellite photograph dated 9/10/2013); b) Google Earth satellite photograph of the
943 2014 mudflow; b') outline of the 2014 mudflow (solid red line), and deformed part of the 2013 mudflow (dotted
944 red line); c) our drone picture of the superimposed mudflows from 2004 to 2014; d) our drone picture of the
945 West side of the mudflow stack; d') line drawing of the outline of d, the internal part of the 2013 mudflow has
946 been pushed and deformed by small-scale thrusts (orange wavy line) under the action of the 2014 mudflow; e)
947 drone picture of the NE part of the mudflows; e') line drawing of e showing how the 2014 mud flow (red) that
948 has deformed and pushed the 2013 one (blue) all the way out to the edge. Note the conjugate left-lateral and
949 right-lateral shear displacements.

950

951

952 Figure 5. a) 640-m-long resistivity profile across the BGMV crater (see location in Figure 3); b) Bouguer
953 anomaly overlain onto the topographic map of the BGMV, triangles are measurement points and topography is
954 from NASA: 1 arcsec Shuttle Radar Topographic Mission digital elevation model. Note the mass deficit in the
955 SE of the mud volcano, where the Bouguer anomaly decreases right at the location of the 2014 mud pie.

956

957 Figure 6. Geological map of the surroundings of the AAMV. Structures of the Paleogene to Neogene series
958 strike NW-SE, parallel to the Gidjakiakhtarma fold axis. Redrawn from Aliev et al. (2015).

959

960 Figure 7. a) General view of the AAMV (shot from a viewpoint located 3.2 km N of the center) showing the very
961 flat morphology of the mud volcano; b) satellite view (Google Earth image dated July 7th 2017) showing the
962 elliptical shape of the Ayazakhtarma, the central area, the median ring and the outer flank dissected by numerous
963 gullies radiating away from the mud volcano.

964

965 Figure 8. Morphostructural map and photographs of the AAMV showing the central area, the median ring with
966 conjugate strike-slip faults dominant in the inner part and twin lakes in the outer part; the outer flank is marked

967 by the thalweg lines of the numerous gullies that incise it. a) Overhead photograph of the center showing
968 desiccating polygonal blocks (field notebook for scale); b and c) columnar blocks are 1.5-to-2 m high and, in
969 some places, are separated by long NE-SW furrows; d) in the median ring, strike-slip fractures showing
970 horizontal striae; e) vertical view taken with GoPro camera mounted on a 10-m fishing rod and showing *en*
971 *echelon* fractures used in the McCoss construction and proving that the displacement is transtensional in most
972 cases; f) 10-m-wide ridge (R) flanked by elongate depressions (D), the one to the right with some residual
973 rainwater; such parallel depressions, commonly filled with water, make the “twin lakes” visible in the outer part
974 of the plateau (cyan lines in the line drawing map); g) the flank is incised by numerous active gullies; h) reverse
975 faults are observed locally in the gullies, vertical cleavage and striae indicate radial compression.

976

977 Figure 9. Overlay of two line drawings of a mudflow emitted in 2015 in the central area of the AAMV, one from
978 the Google Earth satellite picture dated July 23rd 2017, and the other from our drone pictures we acquired in May
979 2018. The two pictures have been calibrated to the same scale using as a reference the small NE mudflow that
980 erupted in May 2016 and that has been displaced but not deformed since. Between 2017 and 2018, the 2015
981 mudflow expanded in the NW-SE direction by up to 17 m/yr, perpendicular to the furrows that show a dominant
982 NE-SW trend (rose diagram of fractures).

983

984 Figure 10. Displacement vectors traced on morphological features of AAMV and correlated between SPOT7
985 satellite pictures acquired in October 2011 and December 2017. The vectors trend radially away from the central
986 area, with a dominant displacement, towards the SE; the maximum displacement is 32 m just south of the central
987 area, i.e. an average 5m/yr.

988

989 Figure 11. a) Modeled resistivity cross section in the AAMV from NE to SW (location in Fig. 7): a *ca.* 10-m-
990 thick resistive upper body overlies more conductive material; b) Resistivity profile across two pairs of twin lakes
991 (location in Fig. 7). The upper few meters of the profile show a high conductivity level at the surface and a
992 medium conductivity at depth, down to 10 to 15 m under the surface. Conductivity increases in twin lakes area.

993

994 Figure 12. a) Gravity profiles across the AAMV, one trending NE along the line of the resistivity section, and the
995 other radial towards the NW, triangles are measurement points and topography is from NASA: 1 arcsec Shuttle
996 Radar Topographic Mission digital elevation model; the Bouguer anomalies are strongly negative at the center of

997 the mud volcano (-6 mGal with respect to background); b) 3-D inversion model of the gravity (details in text),
998 According to this model, the volume of the body is estimated around $11 \times 10^8 \text{ m}^3$, with a mass deficit of 35×10^{10}
999 kg.

1000

1001 Figure 13. Geological map of the surroundings of the APMV. Redrawn from Voronin et al. (1958) and Aliev et
1002 al. (2015).

1003

1004 Figure 14. a) Oblique view of the APMV shot from a viewpoint located three km N of the center. The top of the
1005 mud volcano is an almost flat plateau-like edifice (300 to 306 m) comprising four visible emission centers; b)
1006 aerial view of the APMV with drone pictures we acquired in May 2018 overlain on top of the Google Earth
1007 satellite photograph dated March 24th 2004, local irregularities are observed at the emission zones and deformed
1008 substratum around (up to 2 m positive relief), small thrusts (up to 1 m) and depressions along strike-slip faults
1009 (from -2 to +2 m). The white line marks the location of the 460-m-long resistivity profile acquired in May 2017
1010 across the N-S diameter of the east mudflow (Fig. 17). NWEZ, SWEZ, CEZ and EEZ, respectively, indicate the
1011 northwestern, southwestern, central and eastern emission zones

1012

1013 Figure 15. a and b) Drone photograph and line drawing of the central emission zone of the APMV; c) regular
1014 undulations just outboard of the most recent mudflow (April 27th, 2018), the axes of the undulations are always
1015 concentric, the wavelength is about 5 meters, and they look like a folded cover above a *décollement* surface; d, e
1016 and f) associated structural features: combination of peripheral thrust and flexure towards the NW, NE and SE
1017 and conjugate strike-slip fractures;

1018

1019 Figure 16. a) Line drawing of the plateau of the APMV from the Google Earth image dated March 24th, 2004
1020 showing a concentric zonation centered near the Central Emission Zone (CEZ); NWEZ, SWEZ, and EEZ
1021 indicate the northwestern, southwestern and eastern emission zones, respectively. The inner part of the Plateau
1022 around CEZ, NWEZ and SWEZ is affected by numerous conjugate strike-slip faults that recorded radial
1023 compression, whereas the outer part is covered by concentric elongate lakes. The EEZ belongs to this outer zone;
1024 b, c, d) are photographs of the northern outer flank that is incised by numerous gullies, where some reverse faults
1025 are present, vertical cleavage and striae, all indicate a combination of radial and westward compression (pens for
1026 scale).

1027

1028 Figure 17. Modeled resistivity profile across the eastern emission zone of the APMV; see explanations in the
1029 text.

1030

1031 Figure 18. Synthetic conceptual model of the architecture of the BGMV. The top of the mud volcano is
1032 structured by the crater ring that has been filled by successive pie-like mudflows. Some larger mud flows have
1033 overflowed the edge of the crater to the NE through a breach in the rim (curved arrow). At the center the youngest
1034 mudflow has partly covered and squeezed the previous ones under the effect of its weight. Mud bombs impact
1035 features up to one meter in diameter are present on the surface of the young mudflow.

1036

1037 Figure 19. a) Synthetic conceptual model of the architecture of the AAMV as an example of a plateau-like mud
1038 volcano, soft mud is in grey; b) the central area is the youngest part and corresponds to the feeding area of the
1039 mud volcano, with both columnar mud prisms and wide extensional fractures; c) conjugate strike-slip faults
1040 showing the radial outward compression direction in the median ring; d) strike-slip faults and pop-up-like “twin
1041 lake” structures that line up with the plateau edge, like small-scale reverse faults of the outer flank;
1042 circumferential extension is made possible by the independent movements of the deformed blocks. e) From the
1043 center to the flank, the structures observed at the surface correspond to Anderson-type faults (Anderson, 1951)
1044 and they appear to reflect a switch of the principal stress representing the gravity “ $\rho g z$ ”. At the center $\rho g z = \sigma_1$ is
1045 vertical, and the deformation is extensional; in the median ring $\rho g z = \sigma_2$, and strike-slip faults are observed
1046 whereas in the outer ring and the flank $\rho g z = \sigma_3$, and the deformation is compressional with reverse faults.

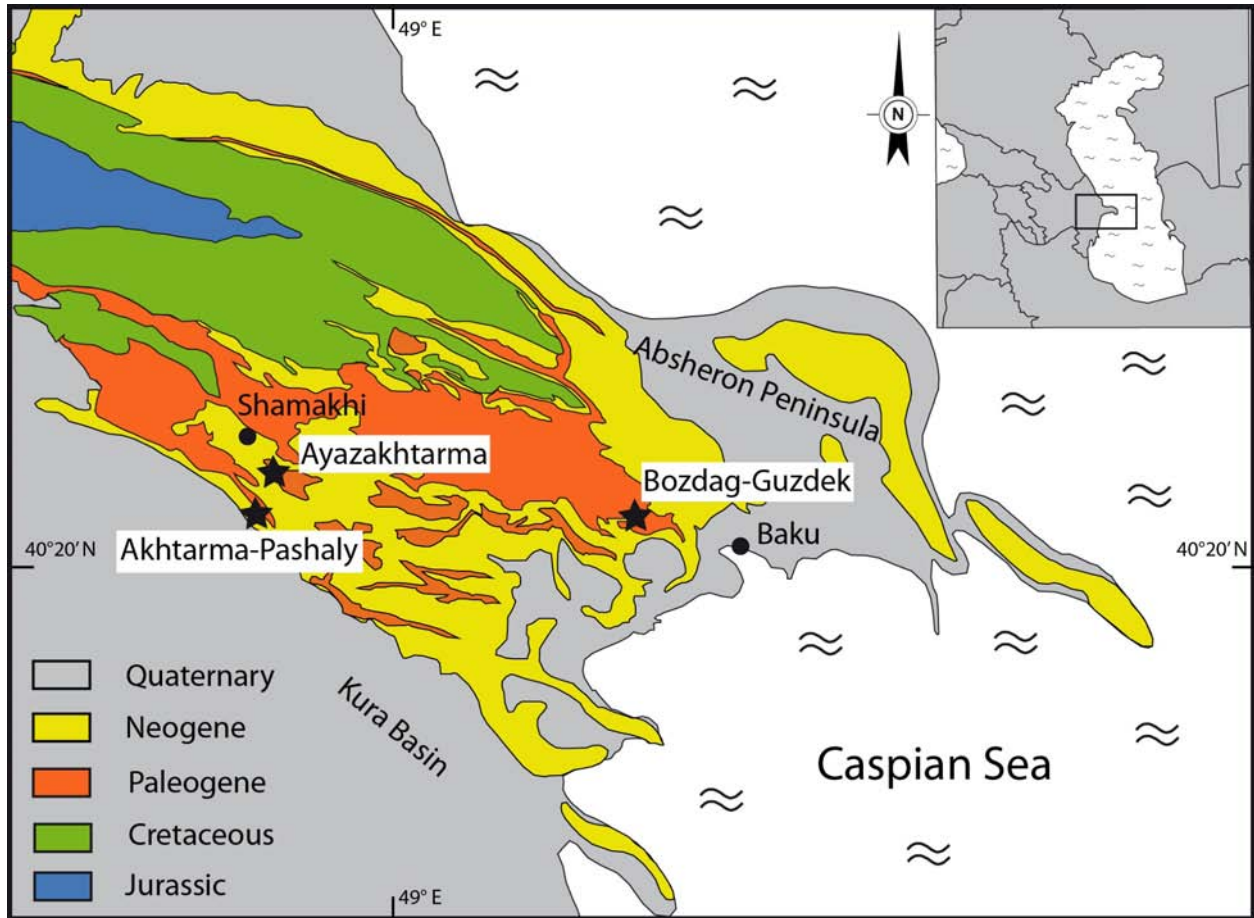
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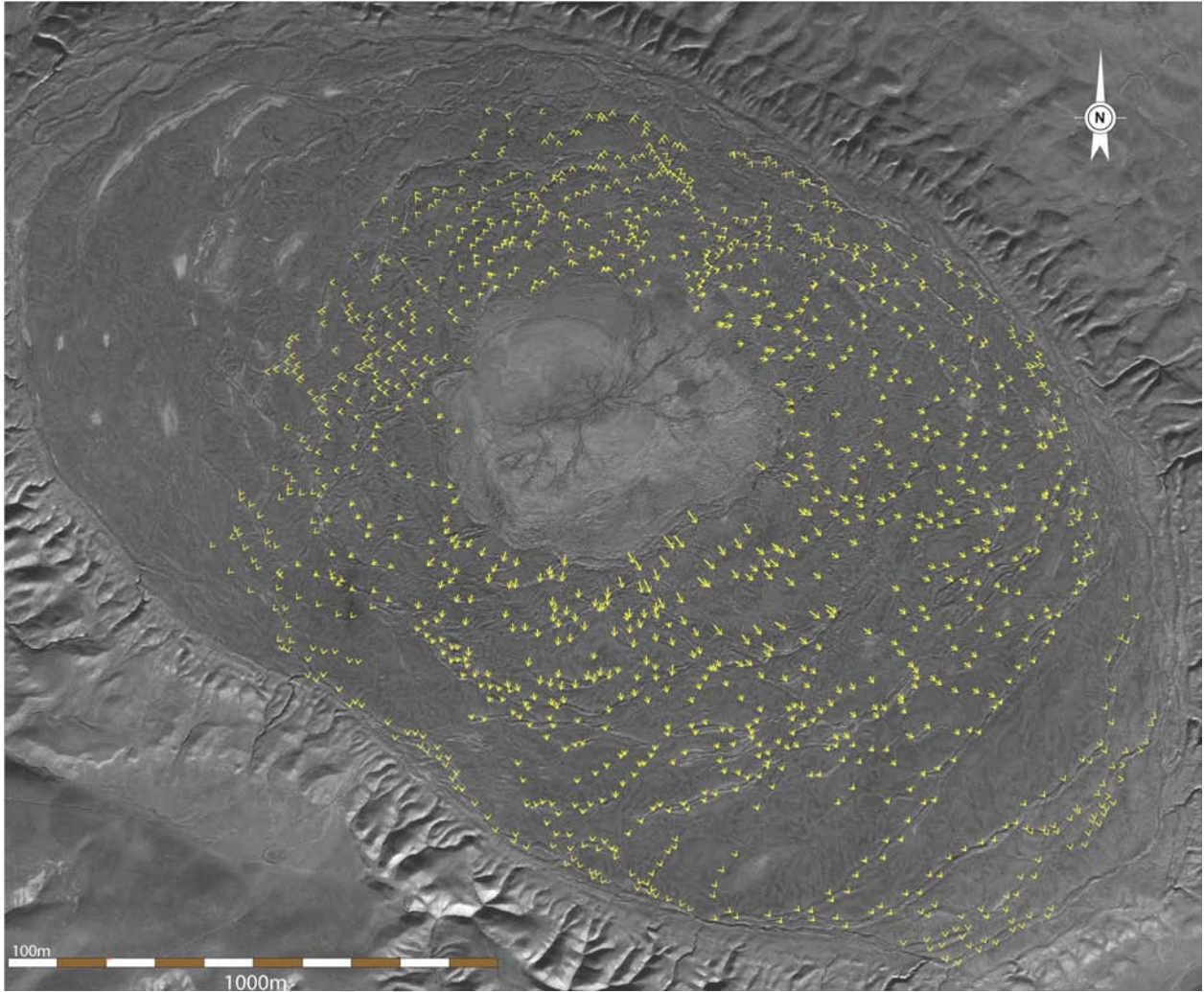
1048 Figure 20. Synthetic conceptual model of the architecture of the APMV, soft mud is in grey. The structure is
1049 composed of four mud pies, each made of a stack of individual subcircular mudflows, imbricated in a larger
1050 spreading structure. Three of the mud pies have been recently active with emission of mud and mud bombs. In
1051 the gullies that erode the outer flank, both the thrust faults and vertical cleavage result from a combination of a
1052 dominantly radial and westward compression.

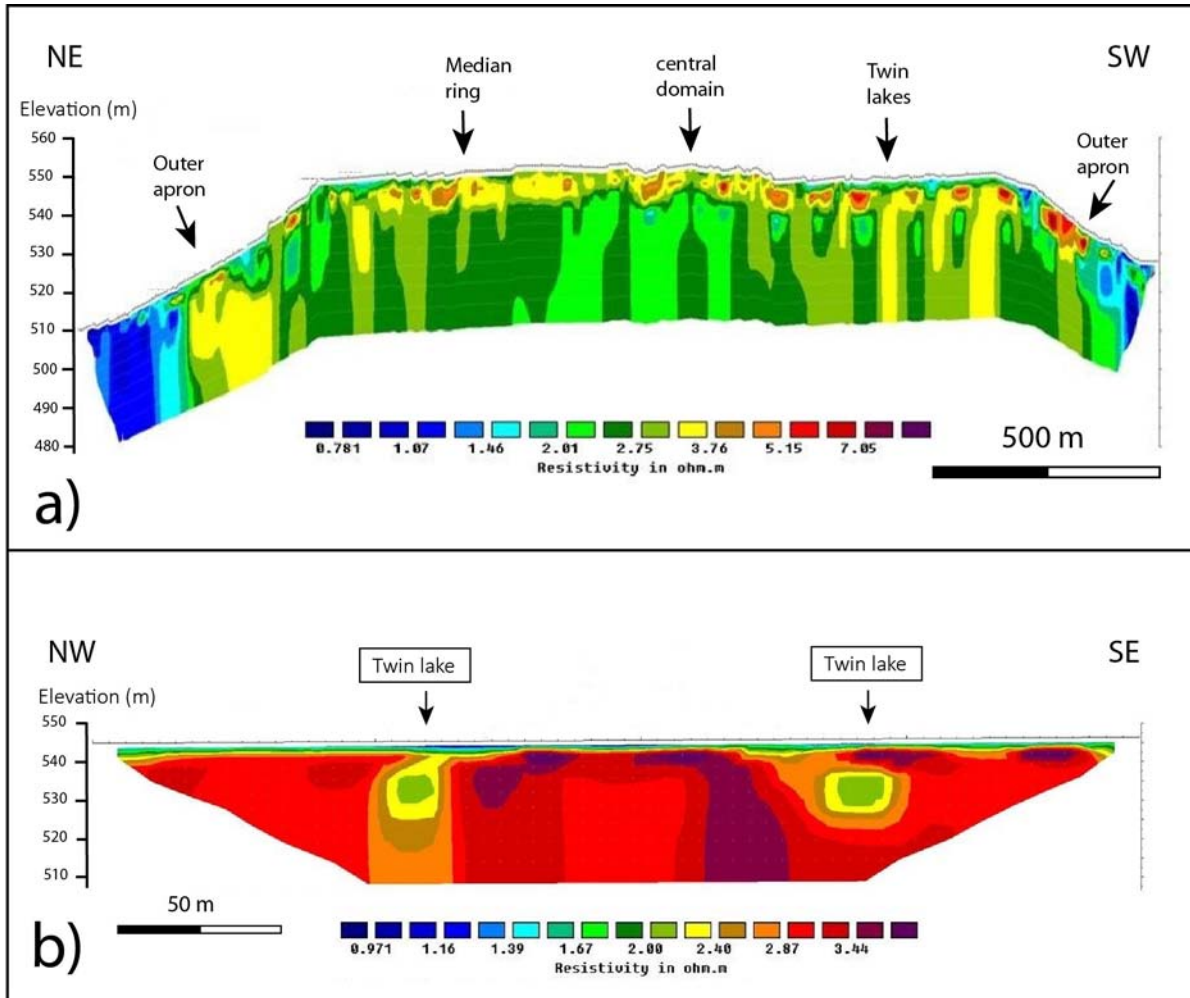
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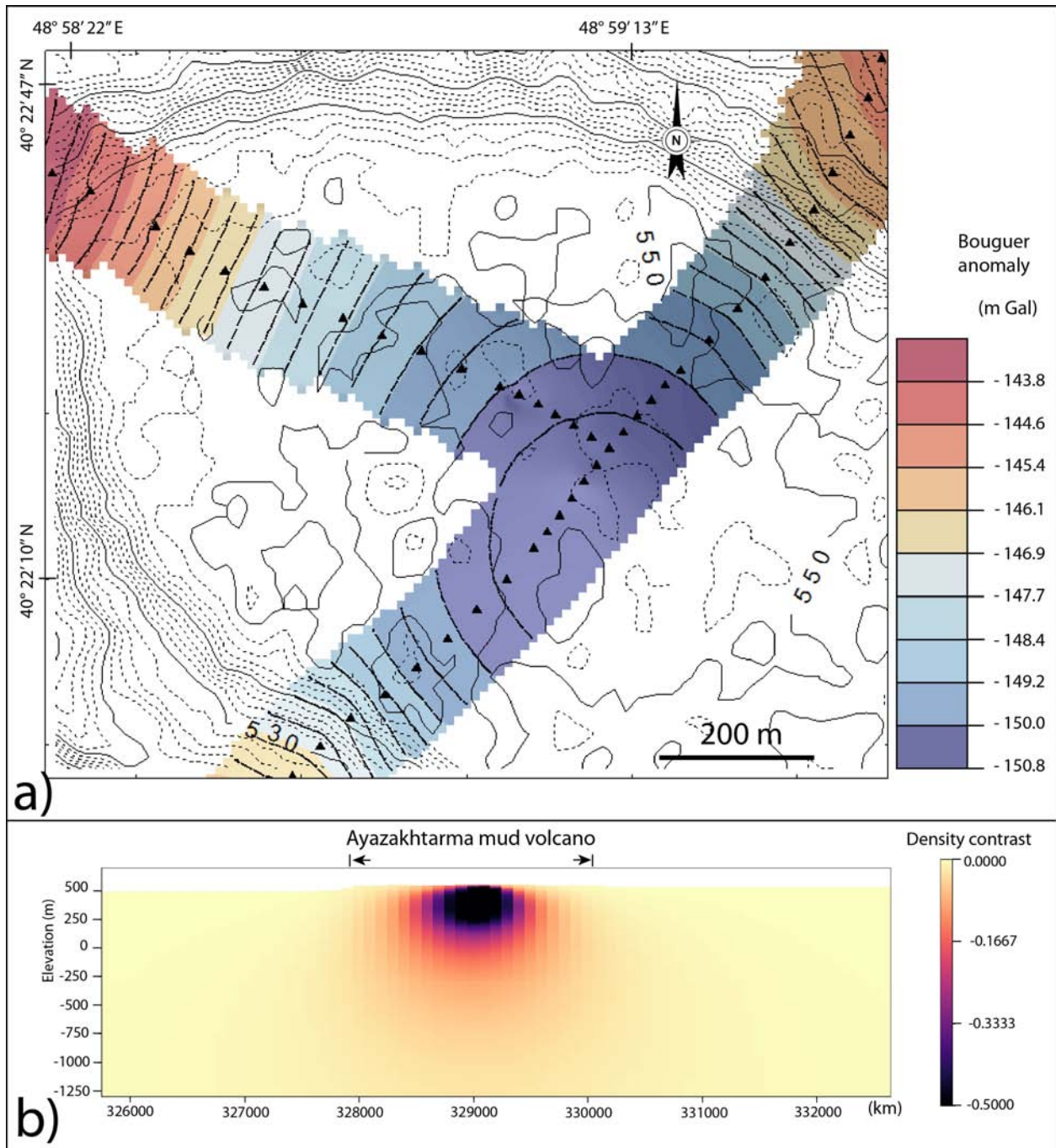
1054 Figure 21. Analogue experiment of viscous silicone injection beneath a brittle sand layer (Dupuis, 2017). A1)
1055 Schematic cross section and overhead view of the experimental setup: 1 is the rigid, horizontal base, pierced in
1056 its center to allow upward flow of the viscous silicone. 2: viscous silicone (black), analogue of the very weak,

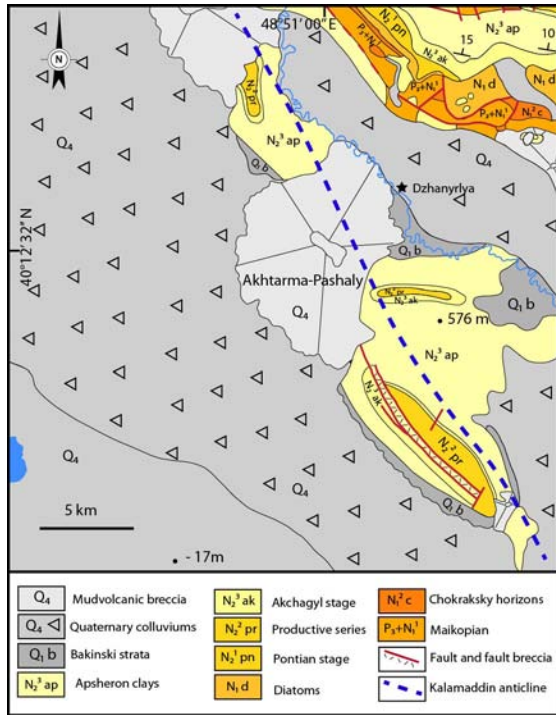
1057 viscous, basal layer of soft mud in nature. White arrows indicate the direction and sense of flow within the
1058 silicone pie. 3 is a peripheral sand mound acting as a buttress around the silicone pie. 4: thin, brittle sand cover.
1059 On the overhead view, A2) 1 is the fixed deformation table, 2 is the hole through which silicone was injected, 3
1060 is the peripheral sand mound, 5 is the mobile wall sliding beneath the fixed table and producing injection of
1061 silicone in 2, 6 is the initial area covered by a thin layer of silicone, and 7 are the fixed lateral walls; B & C)
1062 Overhead photographs of the first model mostly affected by radial extensional faults located around the center of
1063 the pie. Note that above the silicone injection point, the sand cover has been completely thinned out and the
1064 silicone crops out. The white dashed line indicates the initial position of the silicone pie (before thickening and
1065 lateral spreading). D & E) Overhead photographs of the second model that was initially entirely confined
1066 laterally by a sand mound built across the entire perimeter of the pie. The white dashed line indicates the initial
1067 position of the silicone pie (before thickening and lateral spreading). The early stages (D) were characterized by
1068 intense thinning of the sand cover above the injection point, and by concentric folds around it. Later stages, after
1069 silicone thickening and deposition of additional sand layers, show that the peripheral sand mound has been
1070 breached and the pie's cover has spread radially. Note that the different blocs are separated by strike-slip faults,
1071 and that some silicone extruded in the NE part of the model.
1072
1073

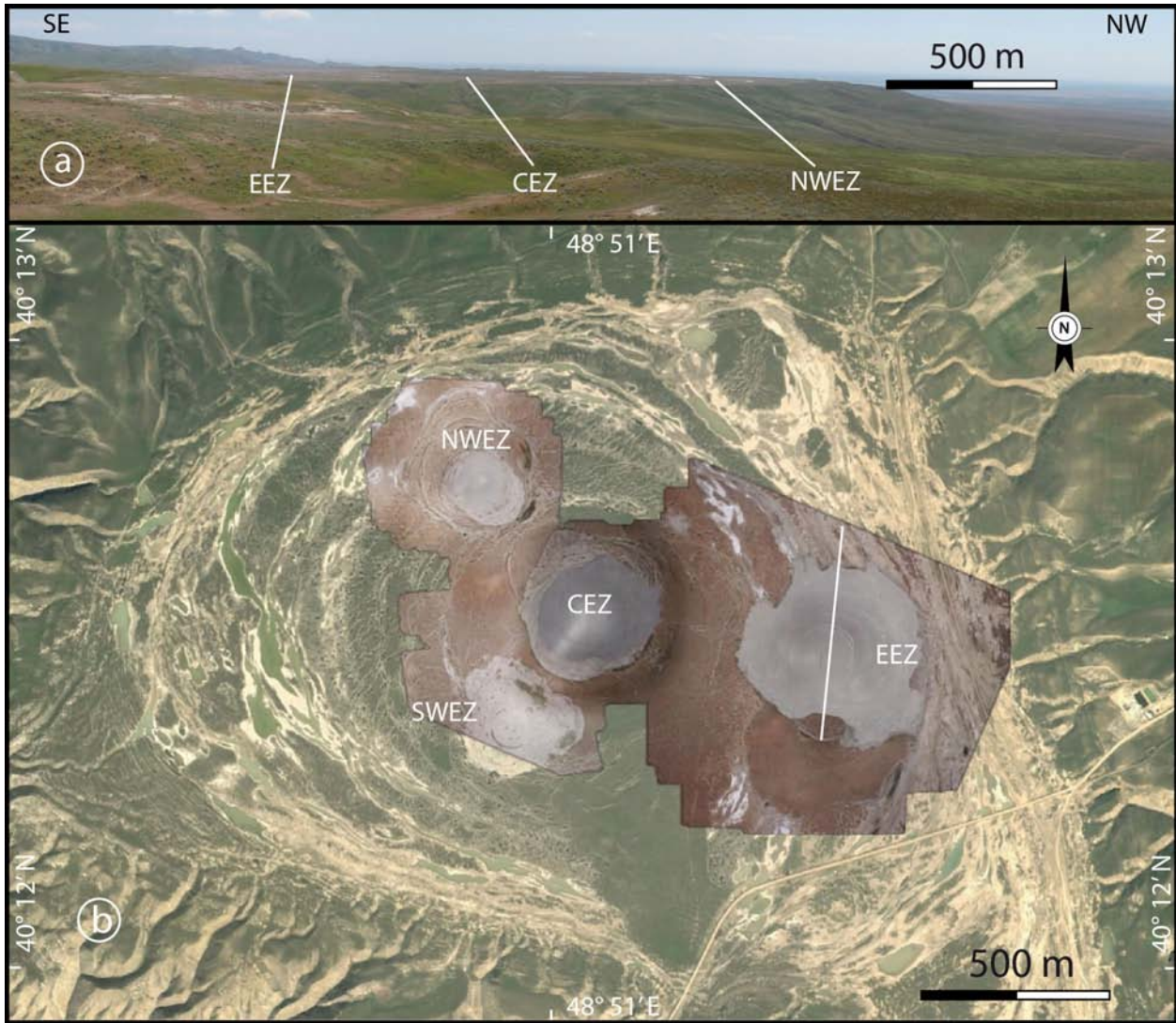


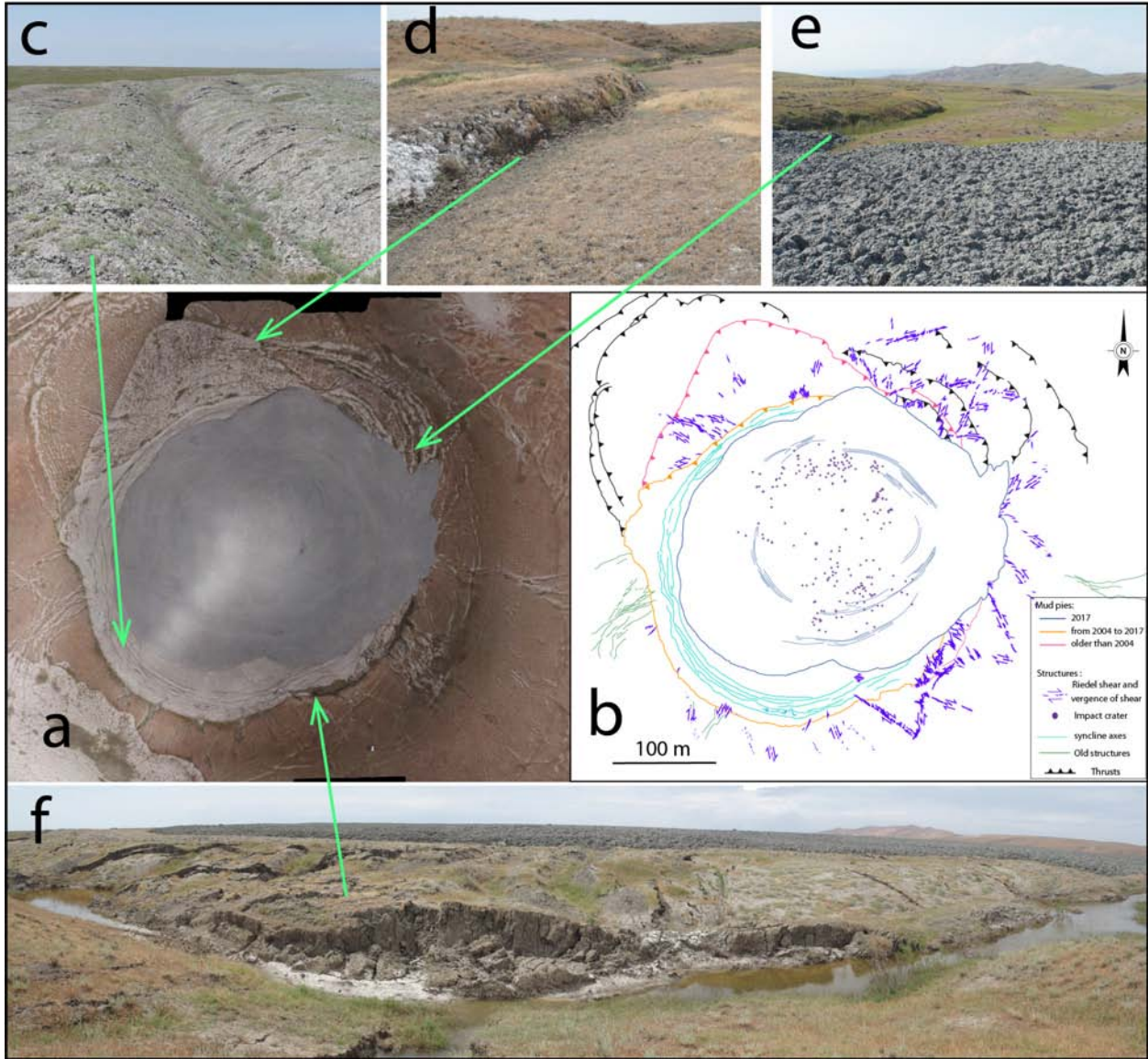


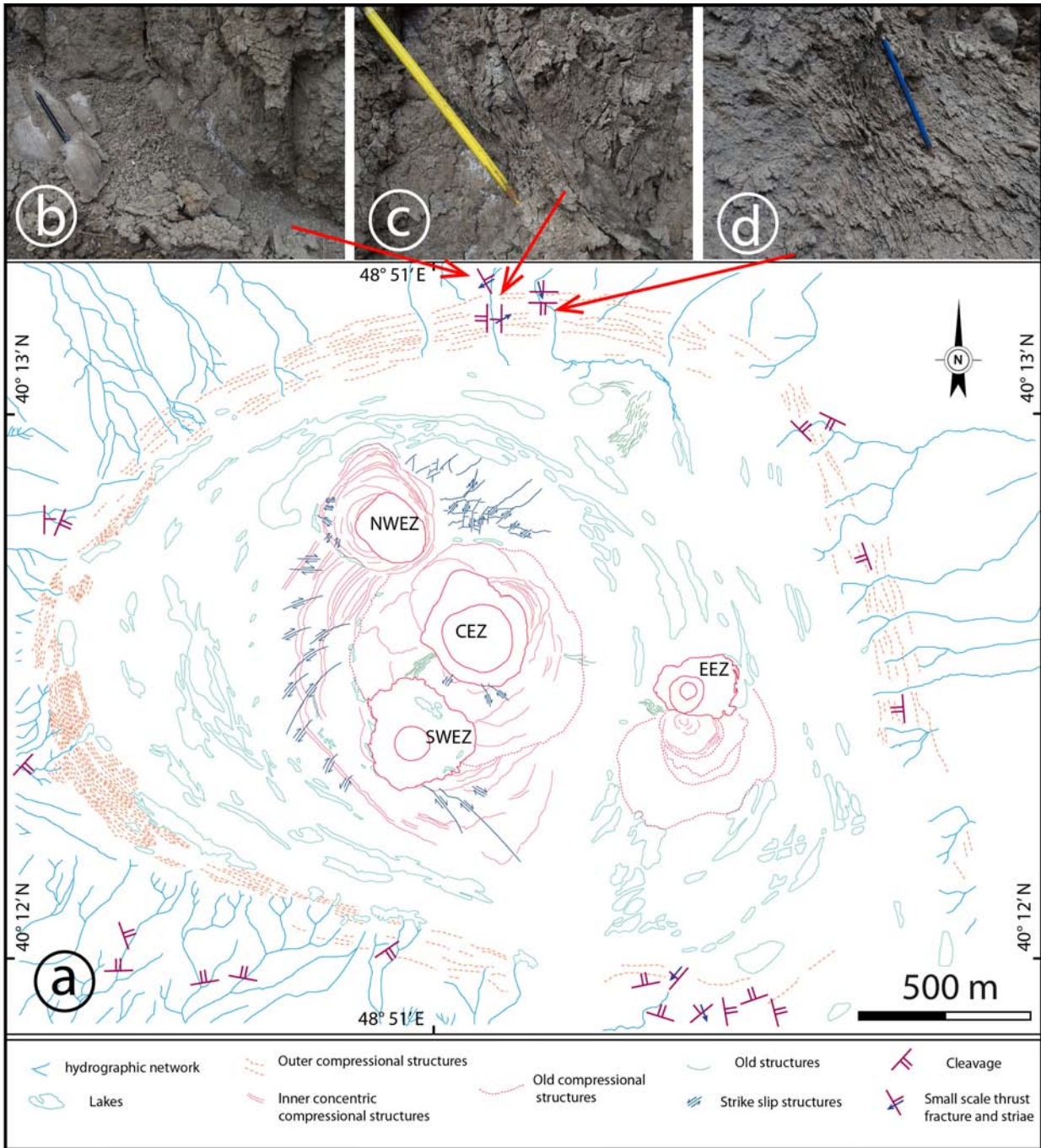


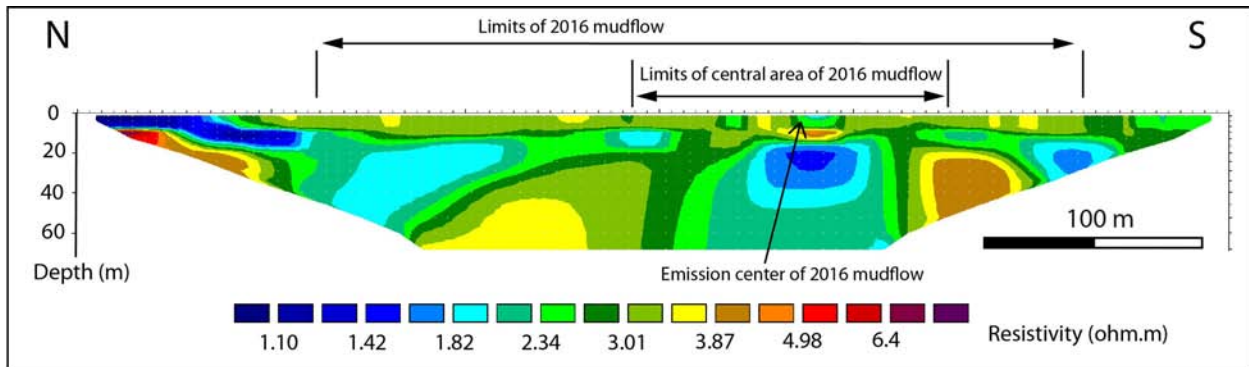




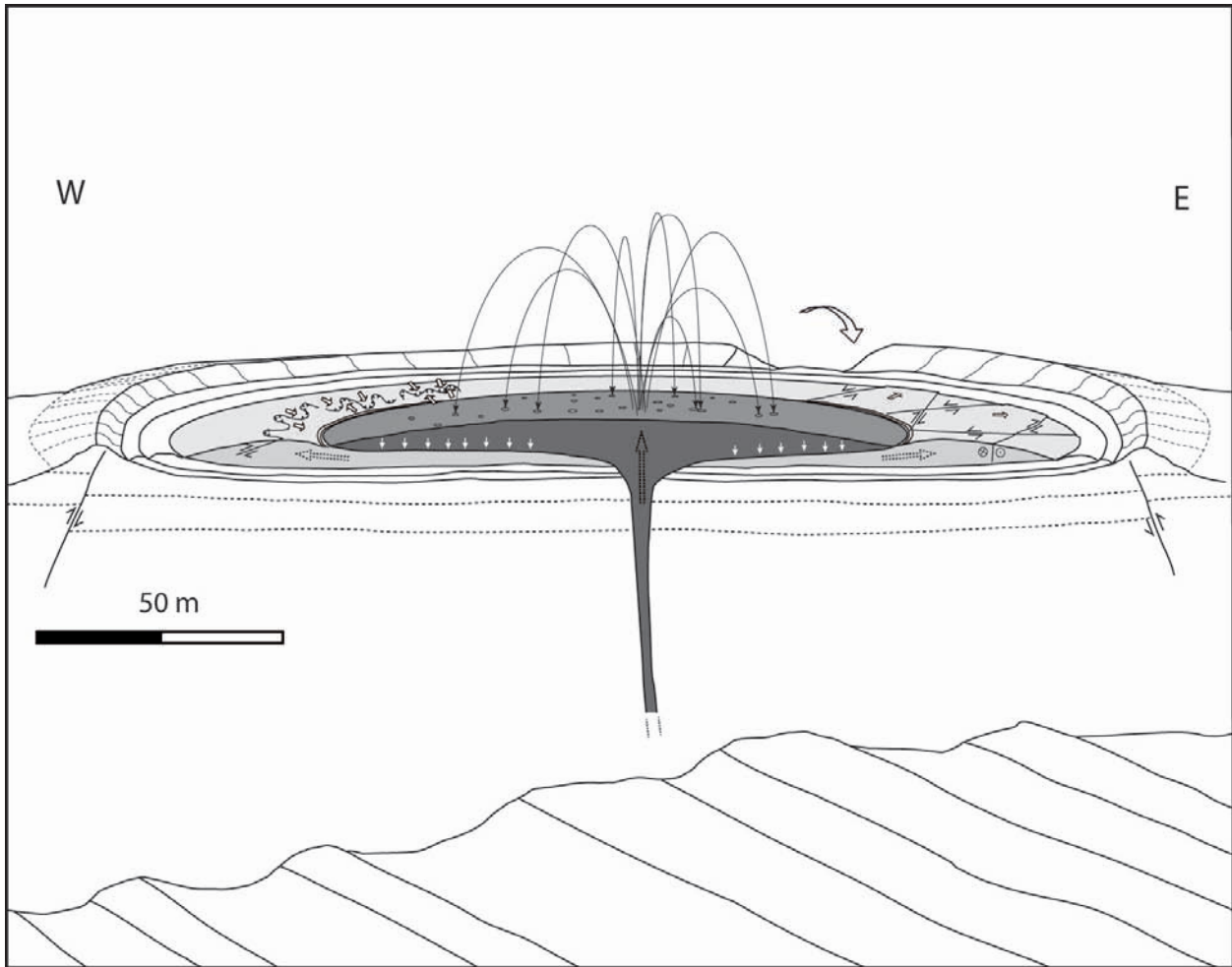


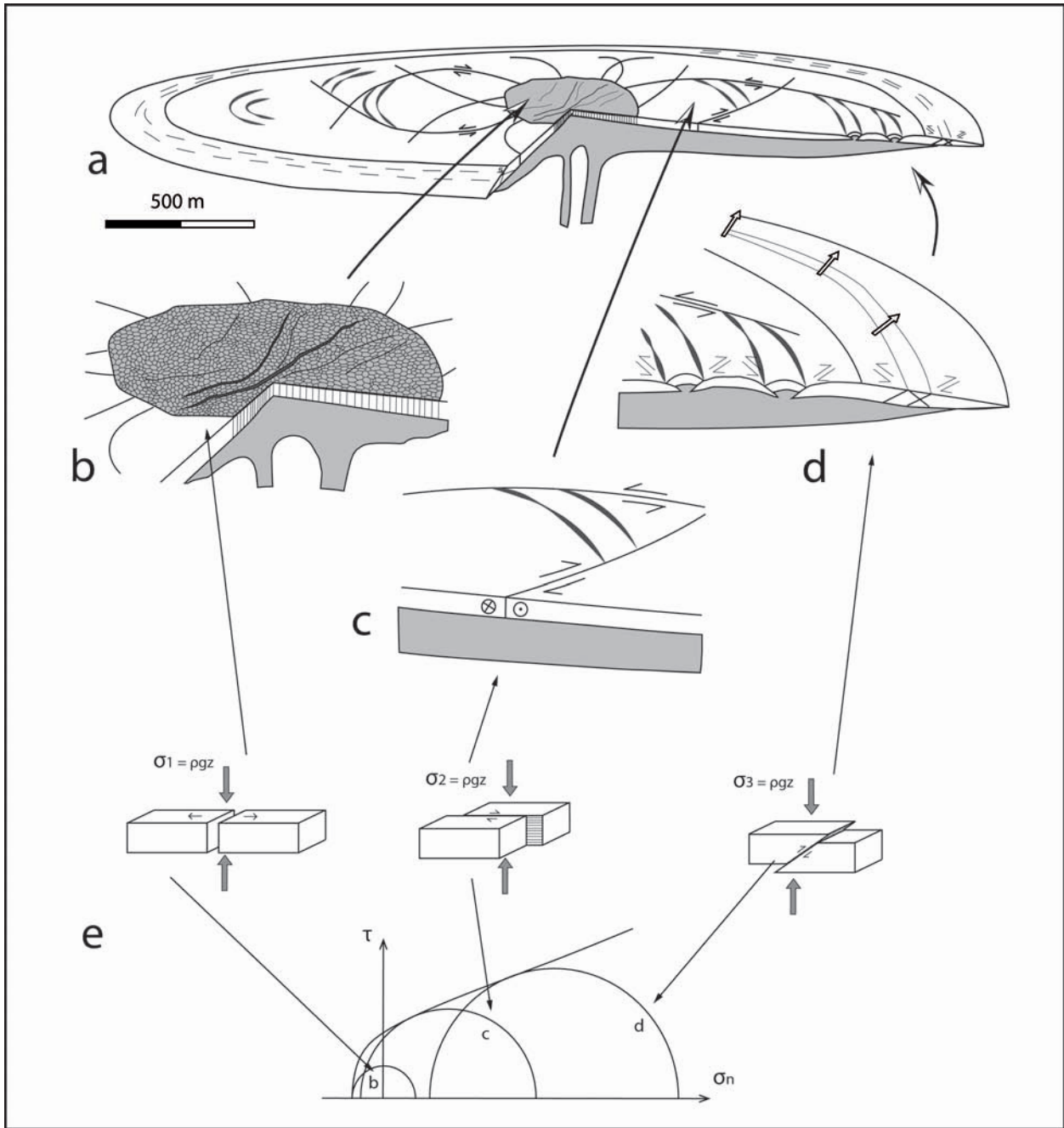


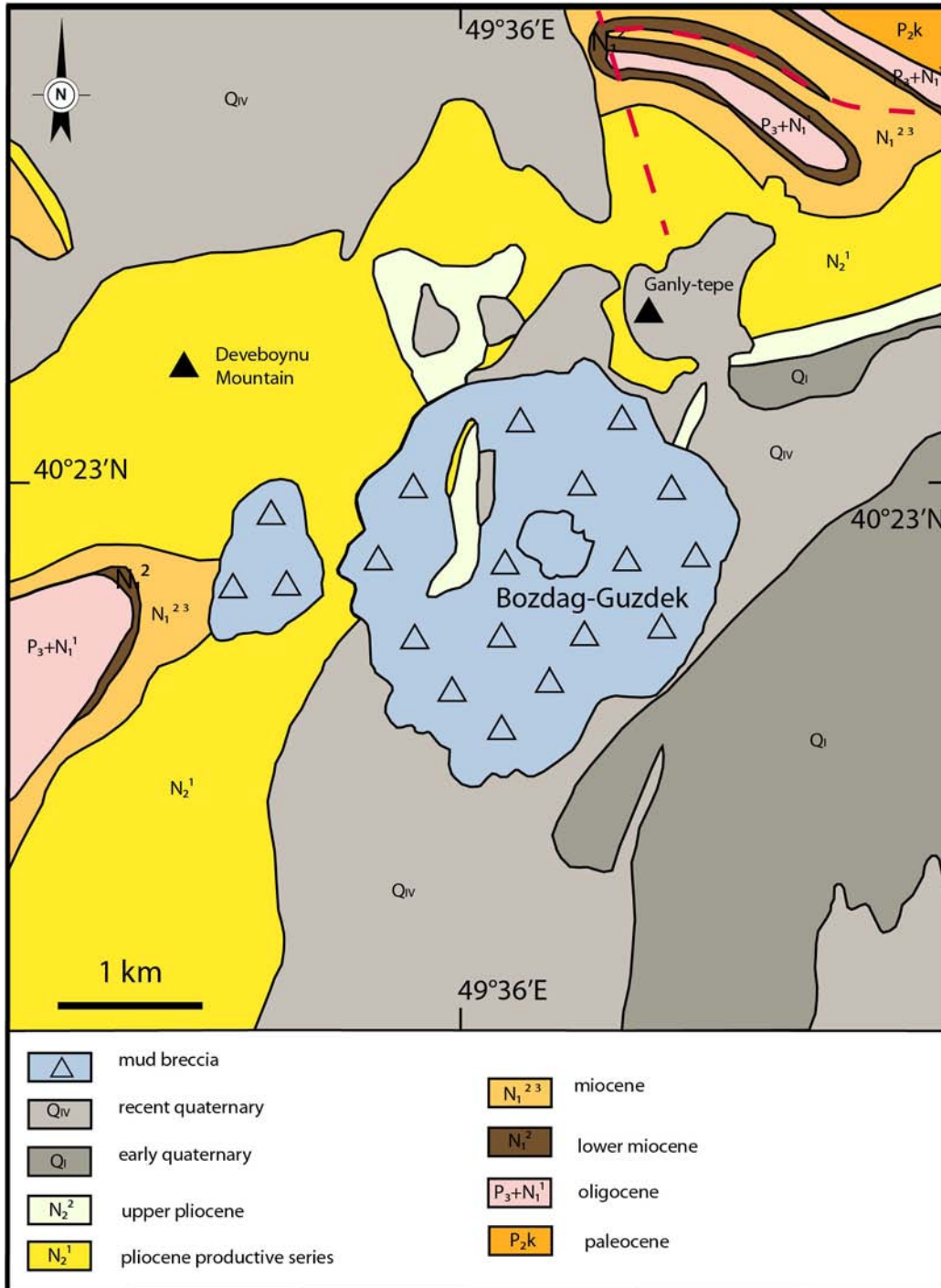


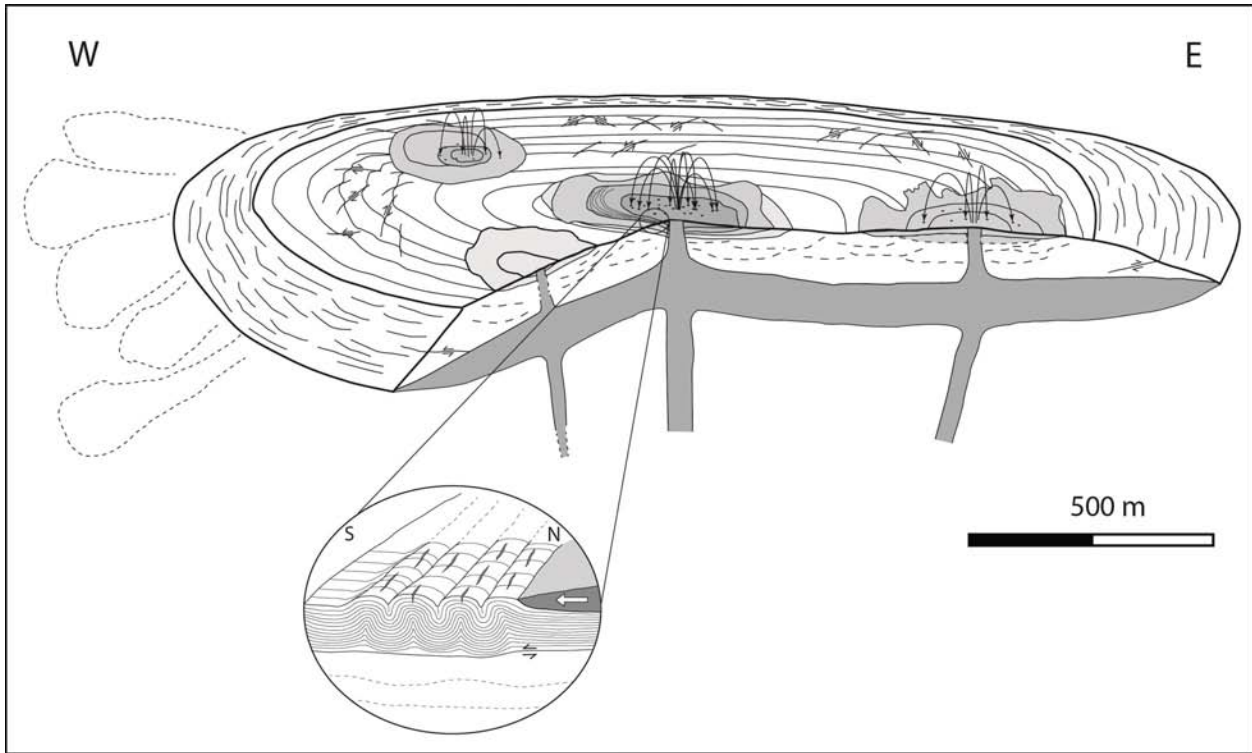


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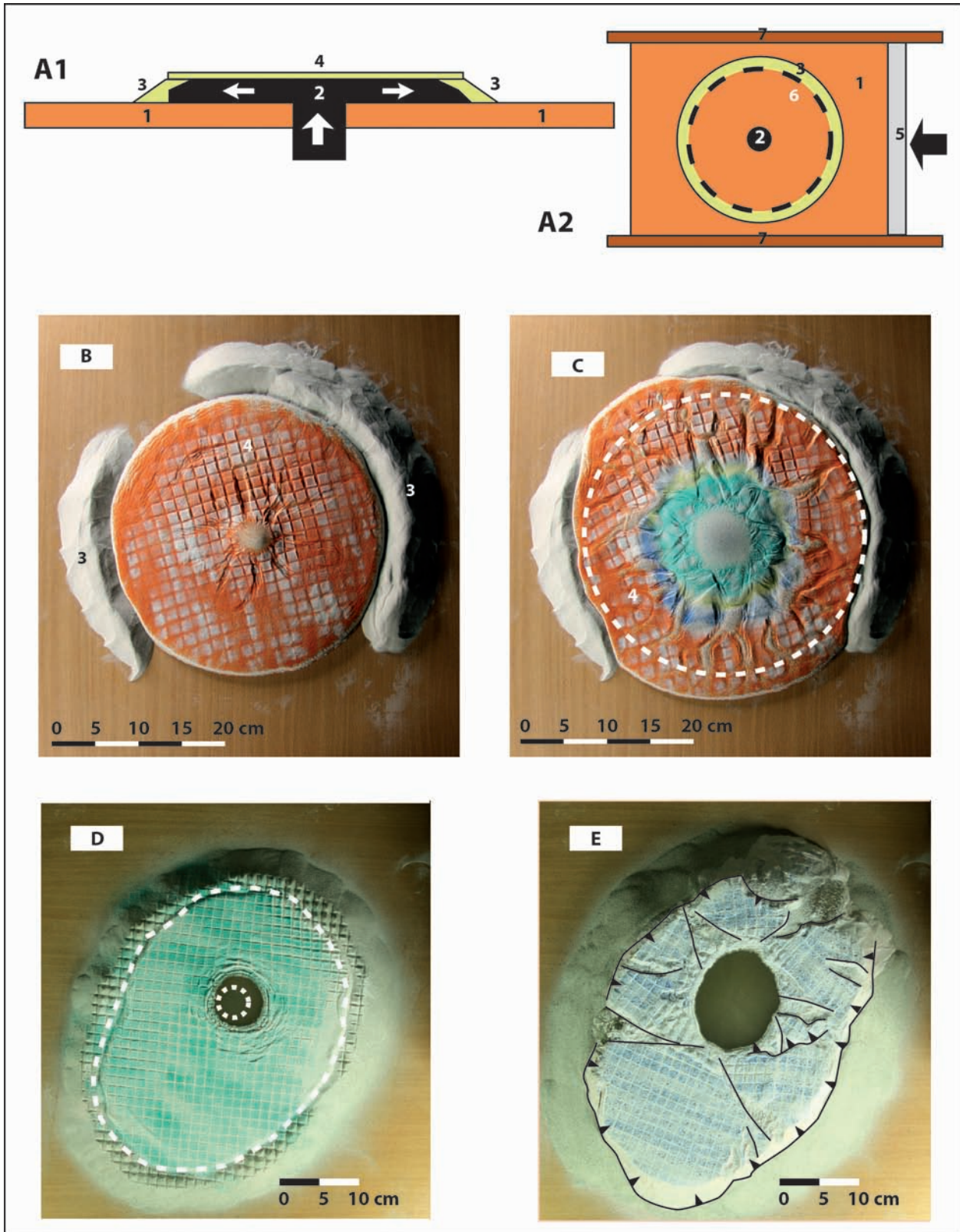


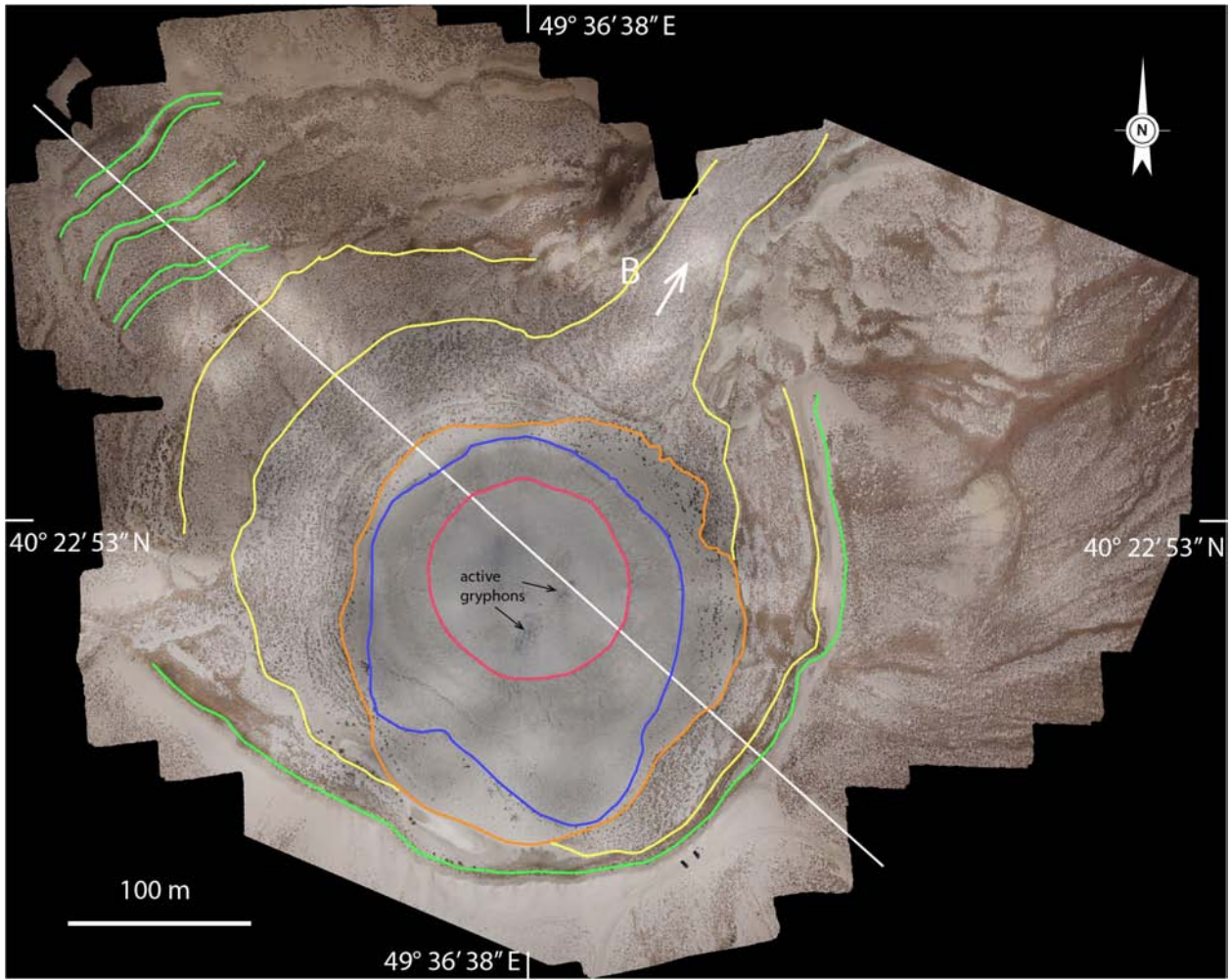


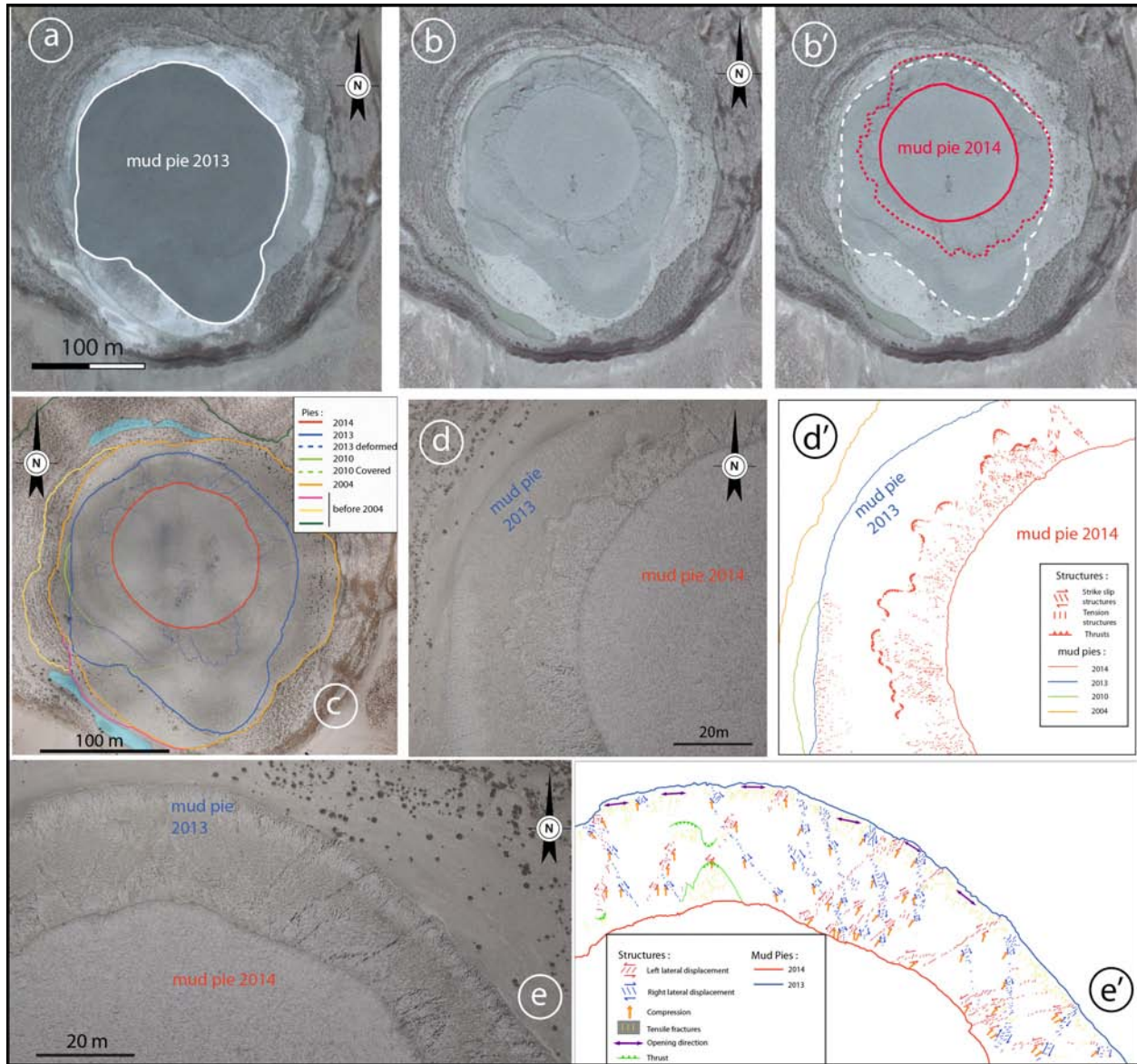


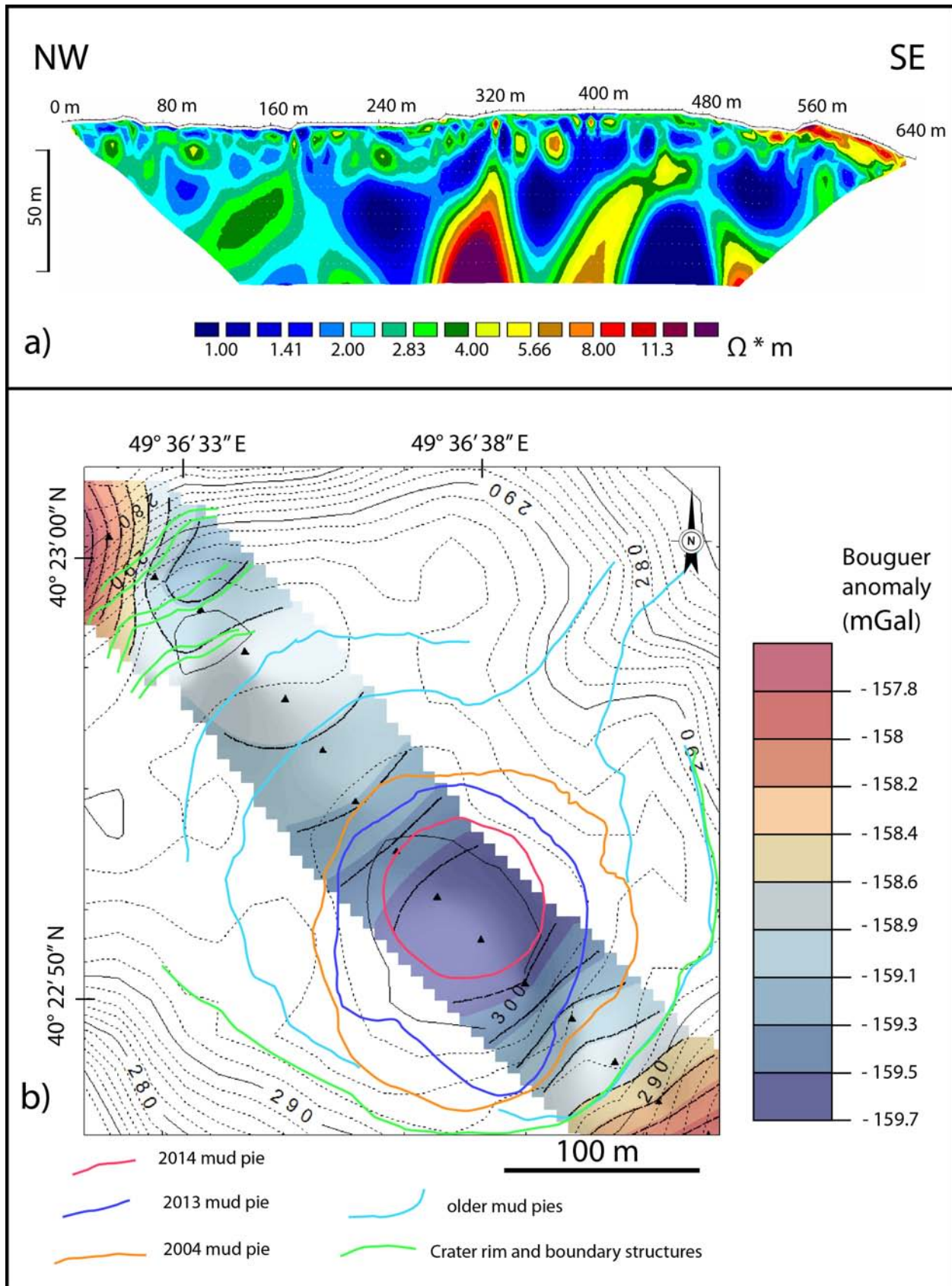


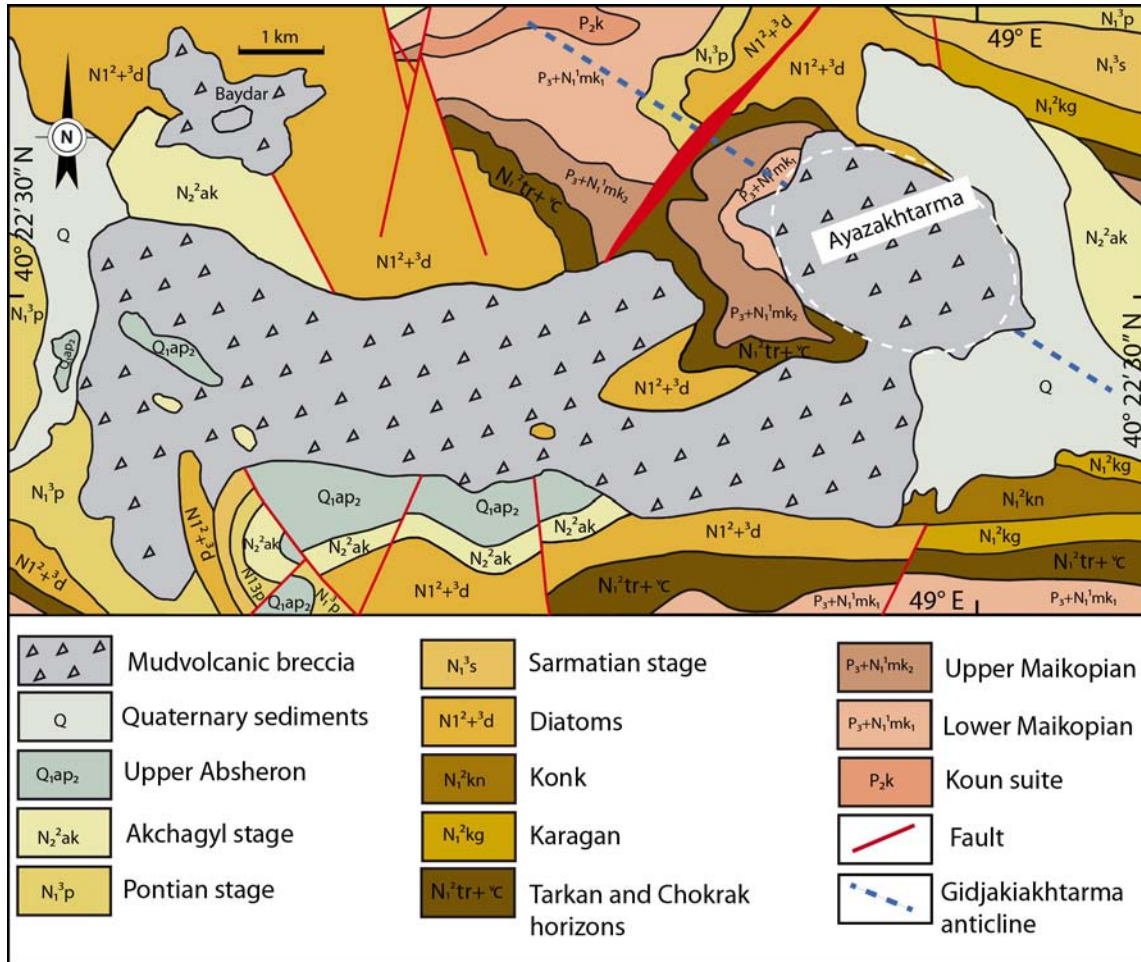
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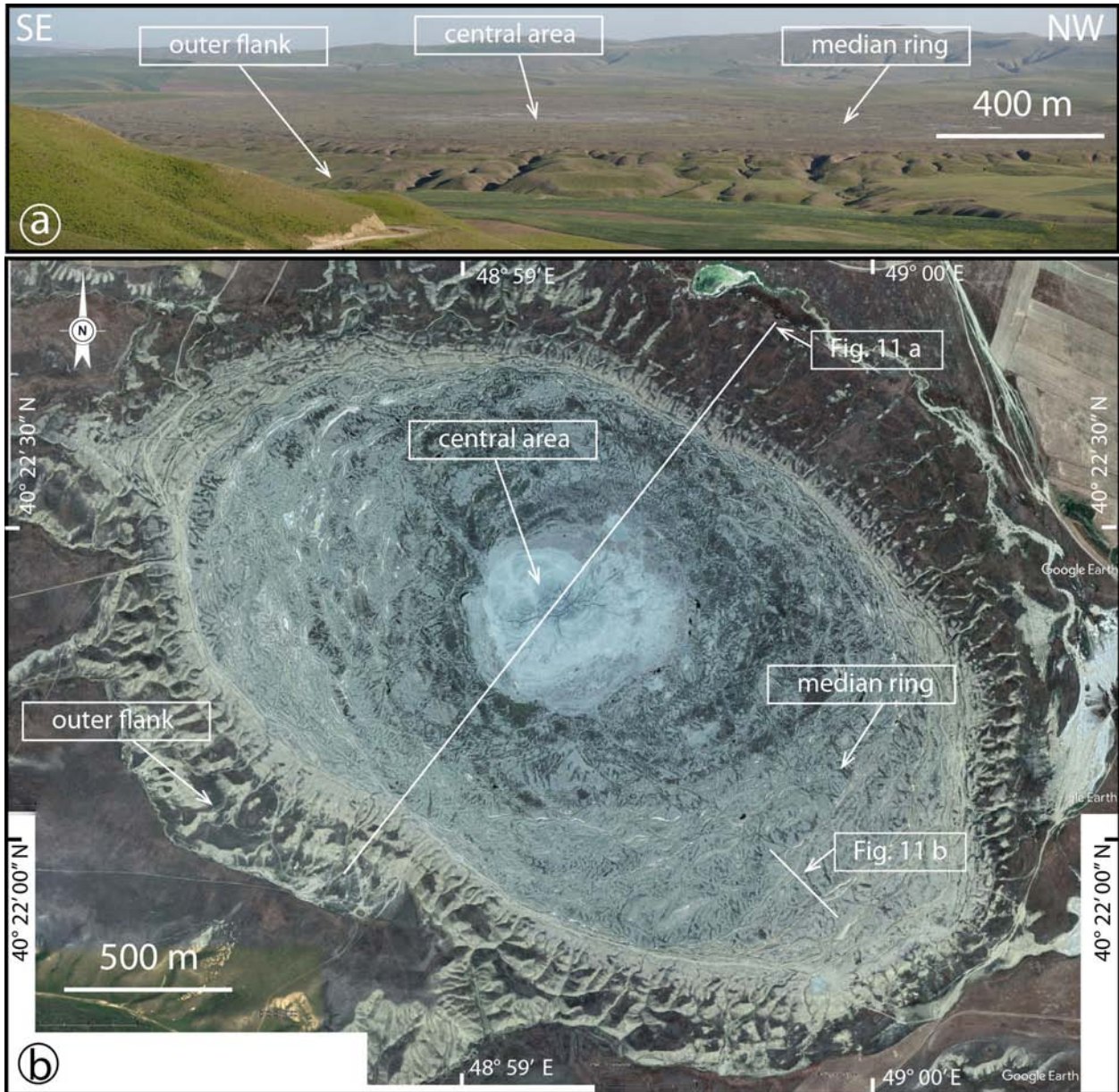


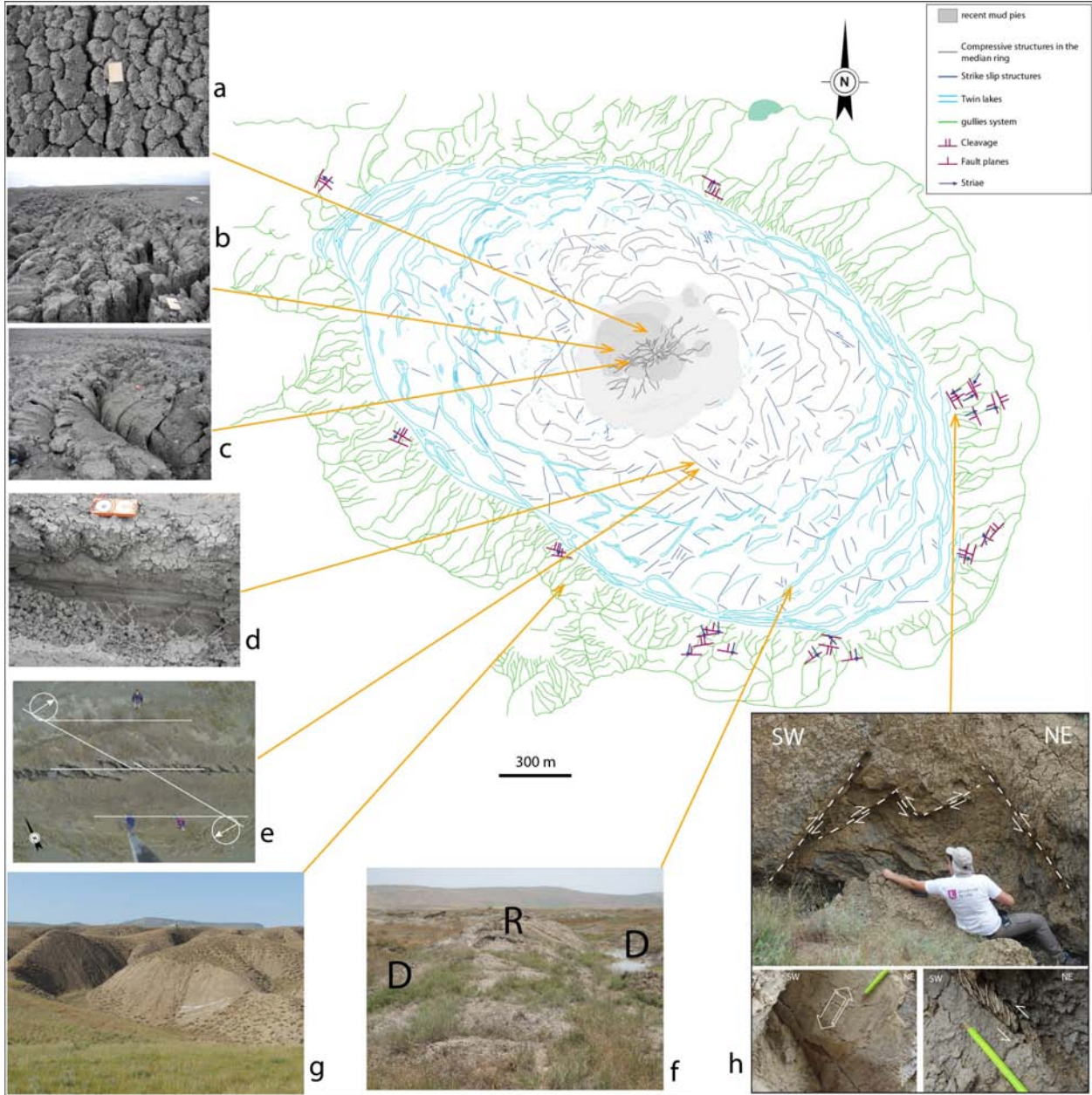


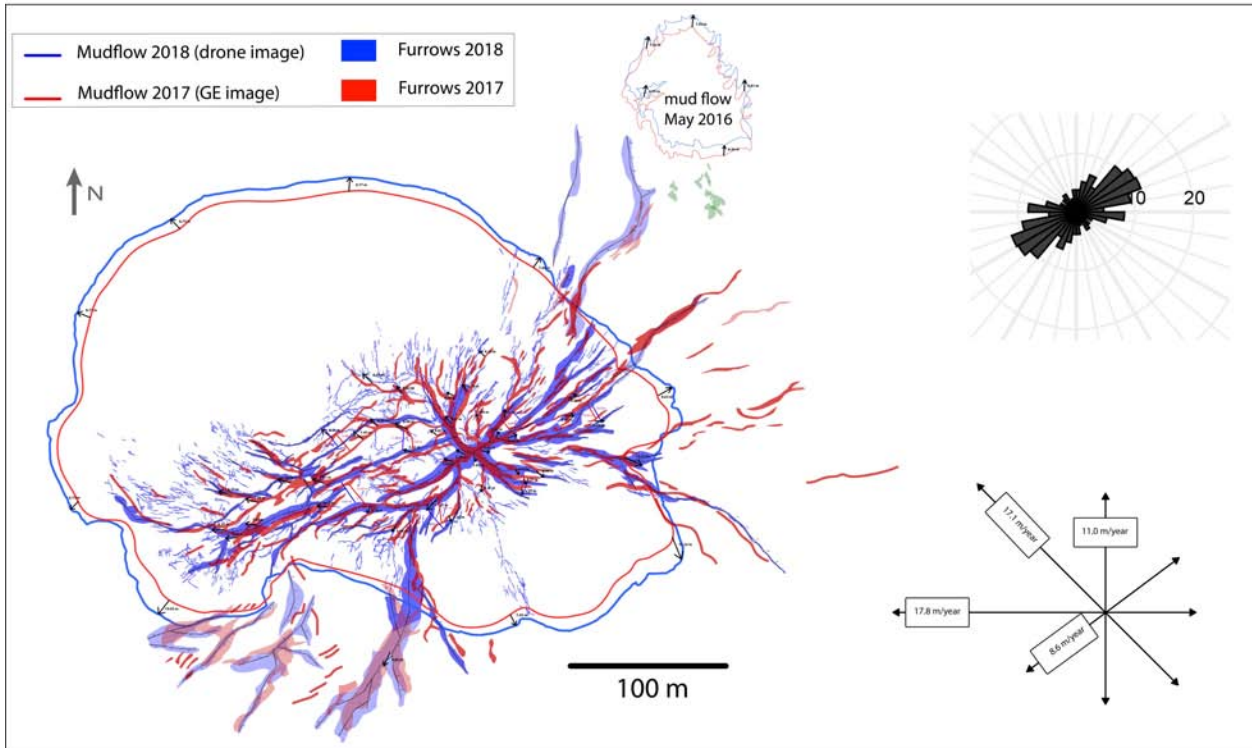












- We propose a new model of growth for flat-topped mud volcanoes fed from the center
- Flat-topped MVs grow by radial expansion, making a concentric structural pattern
- Fractures evolve from extensional in center to compressional at periphery
- Gravity inversion shows a ca 300 m thick low-density lens below the surface
- Soft mud in the lens limits vertical growth and flattens the MV by isostasy

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