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Supporting Information for

Nickel and chromium stable isotopic composition of ureilites: implications for the Earth's core formation and differentiation of the ureilite parent body

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Introduction

In Supporting Information, we mainly describe the samples and analytical methods and discuss the origin of anomalous Cr stable isotope data for NWA 2236. We also show the figures for elemental mappings of one ureilite, the Ni contents in ureilite metal and bulk samples and the correlation with Fo%, and the 1/[Ni] vs. Ni stable isotope compositions of ureilites. Finally, we reviewed and listed all the Ni stable isotope data for chondrites from literature and this study.

Text S1. Samples and Analytical Methods

For Ni stable isotope compositions, we analyzed 22 un-brecciated ureilites found in Sahara (N = 13) or Antarctica (N = 9), and 3 carbonaceous chondrites (Allende (CV3), Paris (CM2) and Orgueil (CI)). Samples were kindly provided by the NASA meteorite working group (MWG), the National Institute of Polar Research (NIPR), Université de Bretagne Occidentale (UBO), the Museum National d'Histoire Naturelle de Paris and, the Smithsonian Institution (Table 1). Compositions of the olivine cores of the ureilites were previously determined (Barrat et al., 2015 and reference therein), and the bulk compositions of the same samples of chondrites were determined by Barrat et al. (2012) and Hewins et al. (2014). Chemical mapping (Ni, Mg, Fe, S) of a section of a ureilite (NWA 6056) was performed using a Cameca SX100 electron microprobe at service commun "Microsonde Ouest", Plouzané (Figure S1).

Fragments (each of 300–1000 mg) were crushed to a homogeneous fine powder using a boron carbide mortar and pestle. Twenty mg of crushed material was digested by sequential mixtures of HF/HNO₃, HNO₃ and HCl. These procedures allow a perfect dissolution of all the phases except graphite and diamond, which are Ni-free, and have no effect for our study. Aliquots of the solutions were used for the determination of the Ni concentrations by ICP-AES (inductively coupled plasma–atomic emission spectrometry) using a Horiba Jobin Yvon Ultima 2 spectrometer in IUEM, Plouzané.

As for Ni isotope measurements, we employed the double spike technique to correct the potential Ni stable isotope fractionation on column purification and instrumental mass fractionation. Nickel in the dissolved samples was then purified by a two-step ion-exchange chromatography, including the first AG1-X8 resin and second Ni-spec resin (Eichrom). Ni isotopic compositions were determined following the procedure described by Gueguen et al. (2013) and Gueguen and Rouxel (2021), using the Thermo Electron Neptune multi collector–inductively coupled plasma–mass spectrometer (MC-ICP-MS) housed at Ifremer, Plouzané. The reader is referred to the paper by Gueguen et al. (2013) for an extensive description of the procedure. Nickel isotope values are reported relative to the standard NIST SRM 986 using the δ -notation:

$$\delta^{60/58}\text{Ni} = \left(\frac{{}^{60}\text{Ni}/{}^{58}\text{Ni}_{\text{sample}}}{{}^{60}\text{Ni}/{}^{58}\text{Ni}_{\text{NIST SRM 986}}} - 1 \right) \times 1000 \text{ (1)}.$$

The error associated with sample values is reported as both 2SD and 2SE of the mean (2se) calculated with the 50 cycles of measurements during acquisition on the MC-ICPMS. A sample-standard bracketing method was employed during the analytical protocol, and

thus, replicate measurements of NIST SRM 986 allows the calculation of a two-standard deviation (2SD) as the external error of $\sim 0.05\%$. Analyses of USGS certified reference material BHVO-2 measured along with these samples were already published in Gueguen et al. (2013).

As for Cr stable isotope measurements, we used sample aliquots (40%) from Zhu et al. (2020) that includes 11 Sahara main-group ureilites and one ureilitic trachyandesite (Bischoff et al., 2014). To estimate the data quality, we also measured the Cr isotope compositions of the NIST SRM 3112a standard and several USGS geological standard rock powders: BHVO-2 (basalt) and PCC-1 (peridotite). Ten of 11 main-group ureilites overlap with the samples in the Ni isotope list (Table 2). After addition and homogenization of a ^{50}Cr - ^{54}Cr double spike (with a spike ratio of $\sim 28\%$), the purification for Cr is similar than in previous studies (Sossi et al., 2018; Zhu et al., 2021a; Zhu et al., 2021b; Zhu et al., 2019), using a two-step cation-exchange chromatography (Trinquier et al., 2008). The yield is 60% - 90%, and the total Cr blanks are 2-5 ng that is negligible compared to the total 5-20 μg of Cr in the ureilite samples. The final Cr cut was evaporated in concentrated HNO_3 drops, 3 to 5 times, to convert the HCl medium to HNO_3 and to remove leftover organics (i.e., some resin) following which the samples were diluted to a concentration of 1 ppm Cr, in 2% (0.317 M) HNO_3 for isotope analysis.

Neptune Plus MC-ICP-MS, housed at the Institut de Physique du Globe de Paris (IPGP) was used to determine Cr isotope ratios. Analytical details are described in Zhu et al. (2021a), Sossi et al. (2018) and Zhu et al. (2019). The purified sample solution was diluted to ~ 1 ppm, which returned a signal of $\sim 15\text{V}$ of ^{52}Cr with a Thermo Scientific Stable Introduction System (SIS) with a $50\mu\text{l}/\text{min}$ probe. Every block of measurements is composed of 100 cycles, with integration time of 4.194 s for each cycle. Each sample were measured 2 to 4 times. The final Cr stable isotope data ($\delta^{53}\text{Cr}$) for ureilites were corrected using their respective mass-independent Cr isotope compositions that had previously been analyzed on the same aliquots (Zhu et al., 2020). The isotopic ratio for samples is reported in delta notation relative to NIST SRM 979:

$$\delta^{53}\text{Cr}(\text{‰}) = \left(\frac{(^{53}\text{Cr}/^{52}\text{Cr})_{\text{sample}}}{(^{53}\text{Cr}/^{52}\text{Cr})_{\text{NIST SRM 979}}} - 1 \right) \times 1000 \quad (2).$$

The uncertainties quoted are the 2SD of individual sample measurements or the 2SD reproducibility of several NIST SRM 979 measurements in the same analytical session ($\sim 0.02 \text{‰}$), whichever is the largest. The Cr contents were measured by ICP-MS housed at IPGP, and the 2σ uncertainty is 5 to 10%. (Steele et al., 2012).

The $\delta^{60/58}\text{Ni}$ values for Allende, Orgueil and BIR-1 are consistent with the literature data (Chernozhkin et al., 2016; Gall et al., 2017; Klaver et al., 2020; Steele et al., 2012; Wang et al., 2021). Furthermore, our two replicate measurements for Allende ($0.23 \pm 0.02 \text{‰}$ and $0.25 \pm 0.02 \text{‰}$) are also consistent, indicating the good external reproducibility of our Ni stable isotope data. Similarly, the Cr stable isotope data of standard samples, including NIST SRM 3112a (unpurified), PCC-1 and BHVO-2, also accord with previous studies (Bonnand et al., 2016; Liu et al., 2019; Schoenberg et al., 2016; Sossi et al., 2018; Zhu et al., 2019), confirming the accuracy of the $\delta^{53}\text{Cr}$ values to the stated uncertainties.

Text S2. Anomalous $\delta^{53}\text{Cr}$ Value for NWA 2236

NWA 2236 has an anomalous $\delta^{53}\text{Cr}$ value ($0.09 \pm 0.02\text{‰}$) that is higher than all the other ureilites. This is not caused by terrestrial weathering that results in an enrichment in the light Cr isotopes (Frei et al., 2014). Note that NWA 2236 also possesses low Cr content (2600 ppm) and high $^{55}\text{Mn}/^{52}\text{Cr}$ ratio of 1.14, relative to other main-group ureilites with $^{55}\text{Mn}/^{52}\text{Cr}$ ratios ranging from 0.5 to 0.8 (Zhu et al., 2020). However, this chemical and Cr isotopic difference may not be caused by volatile process. Zn is much more volatile than Cr (Lodders, 2003; Sossi et al., 2019), but the Zn content (216 ppm) and isotope composition for NWA 2236 is normal. NWA 2236 is not very depleted in Zn (that is much more volatile than Cr), compared to other main-group ureilites, with normal $\delta^{53}\text{Cr}$ values, e.g., NWA 7630 and NWA 7686 have Zn contents ranging from 188 and 236 ppm respectively (Brugier et al., 2019). Additionally, NWA 2236 suffered from only a low degree of shock metamorphism after its formation (Brugier et al., 2019). Alternatively, the distinct Cr stable isotope composition of NWA 2236 results from mantle heterogeneity of UPB, and this much higher $\delta^{53}\text{Cr}$ value can be caused by some mantle process through kinetic isotope fractionation, e.g., diffusion effect. We should note that NWA 2236 also possesses the highest $\epsilon^{54}\text{Cr}$ values among the main group ureilites (all the $\delta^{53}\text{Cr}$ data have corrected for their mass-independent Cr isotope fractionation, i.e., $\epsilon^{53}\text{Cr}$ and $\epsilon^{54}\text{Cr}$ values), which is likely reflecting mantle heterogeneity (Zhu et al., 2020).

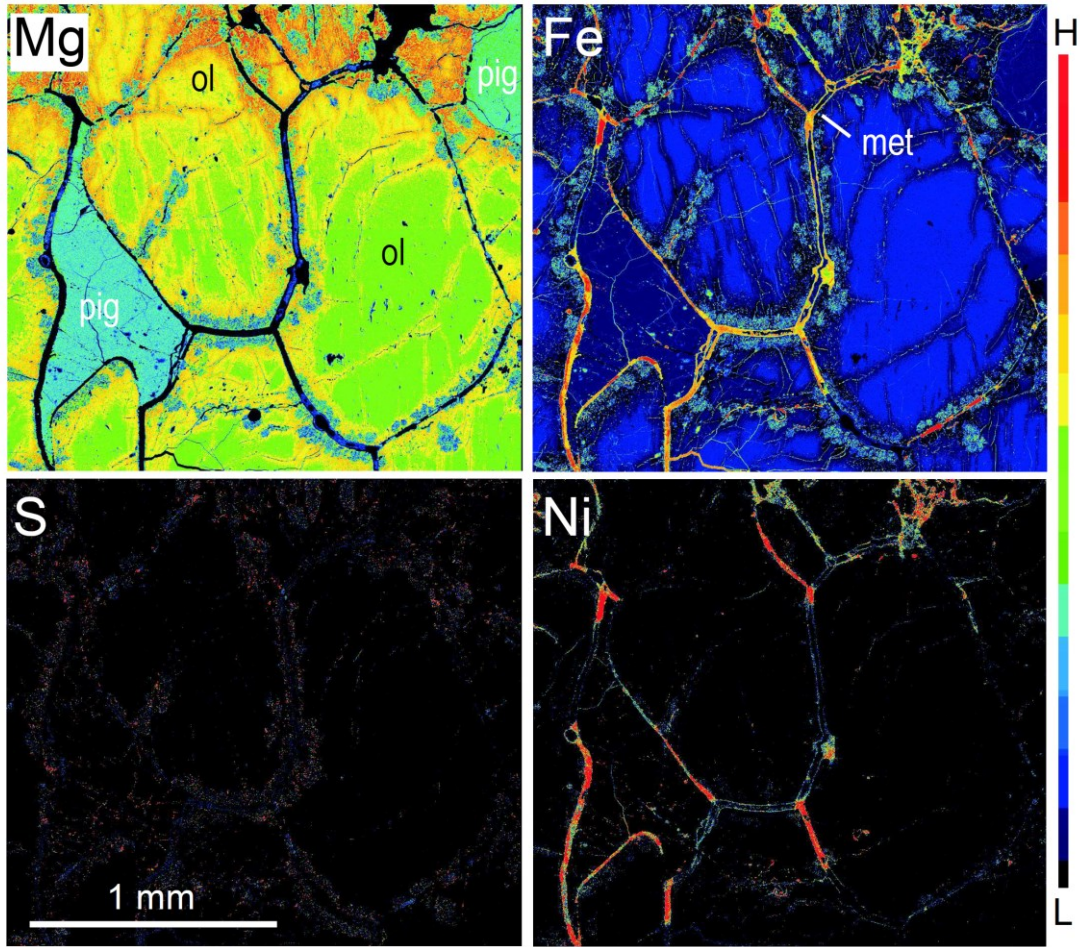


Figure S1. The elemental mapping (Mg, Fe, S and Ni) for NWA 6056. Abbreviation: Olivine-ol; Pigeonite-pig; Metal-met. Ni is mainly residing with Fe and S, and poor in silicate minerals.

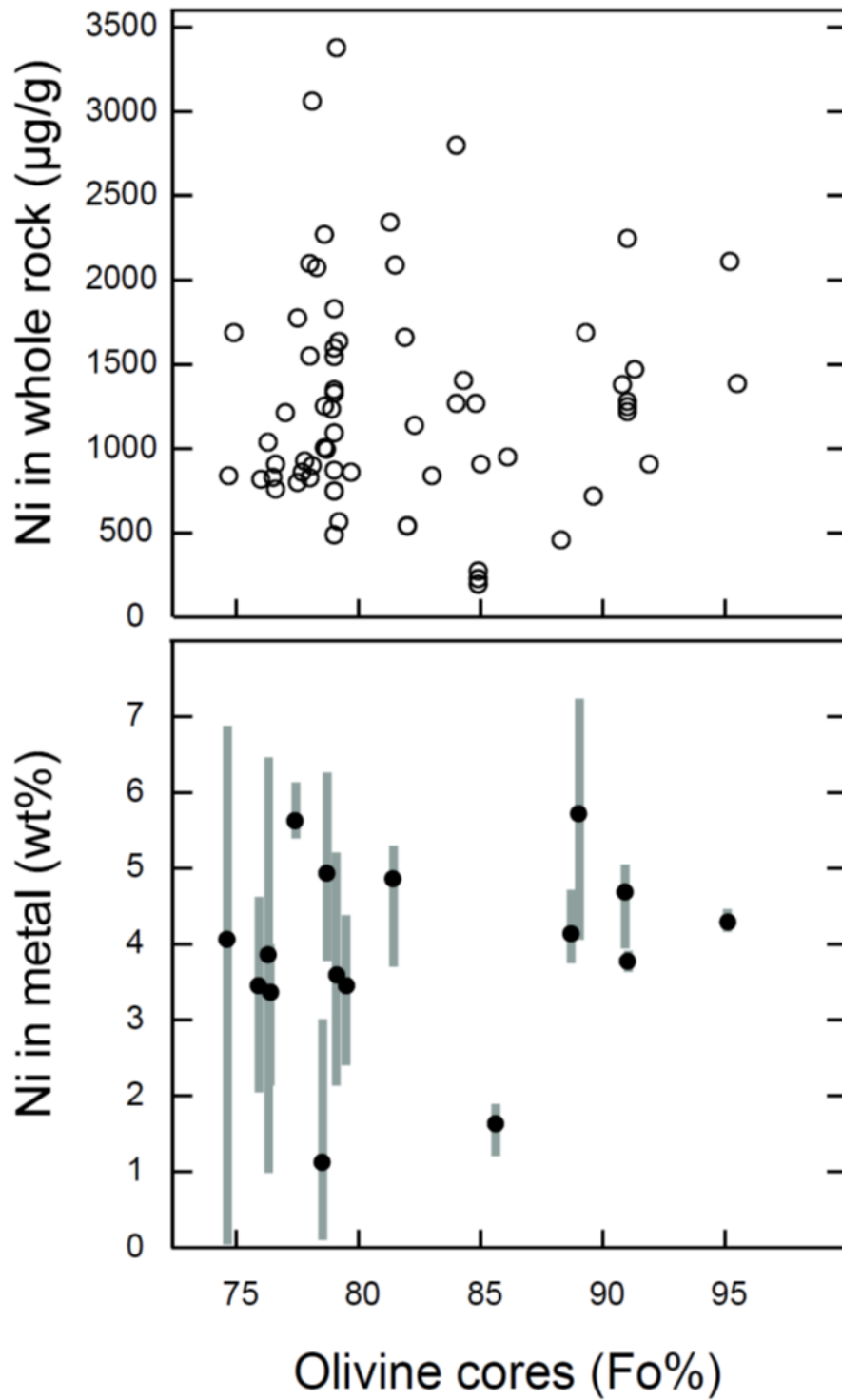


Figure S2 The relationships between Mg# in olivine cores (Fo%) and Ni contents in metals and bulk ureilites (Goodrich et al., 2004; Warren et al., 2006).

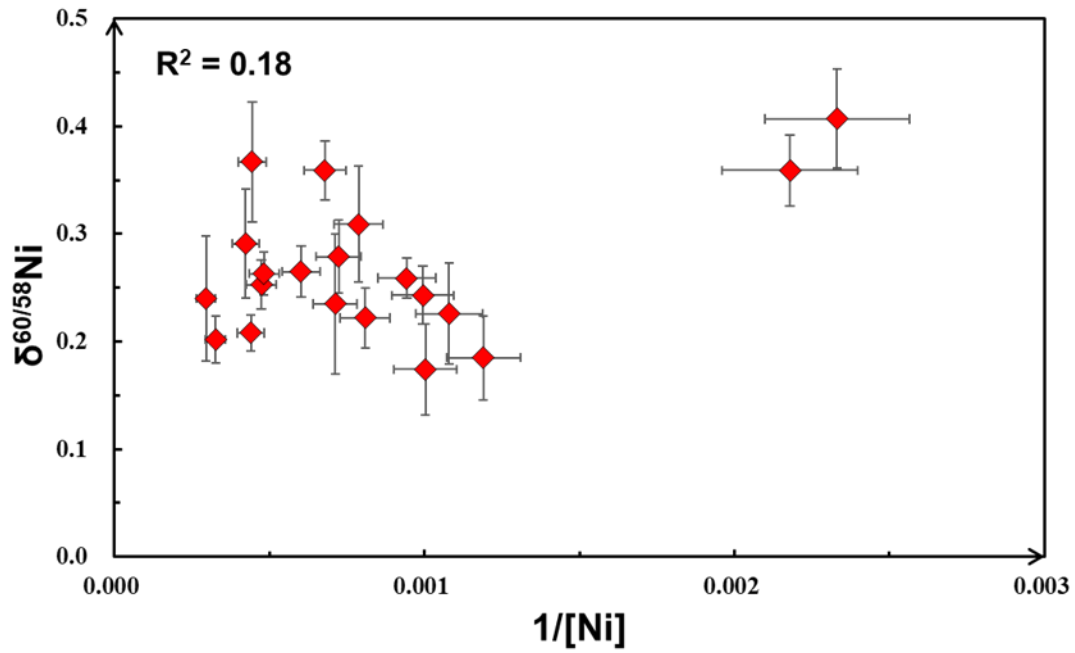


Figure S3 $1/[Ni]$ (Ni contents) versus Ni stable isotope compositions, and there is no clear relationship between them.

Table S1. Review of Ni stable isotope data for chondrites

| Sample | Group | $\delta^{60/58}\text{Ni}$ | 2SD | 2SE | N | References |
|-----------|-------|---------------------------|------|-------|---|------------------------|
| Orgueil | CI1 | 0.21 | 0.07 | | | (Cameron et al., 2009) |
| Orgueil | CI1 | 0.19 | | 0.024 | | (Steele et al., 2012) |
| Orgueil | CI1 | 0.18 | 0.04 | | 4 | (Gall et al., 2017) |
| Orgueil | CI1 | 0.02 | 0.02 | 0.008 | 8 | (Klaver et al., 2020) |
| Orgueil | CI1 | 0.12 | 0.02 | 0.006 | 8 | (Klaver et al., 2020) |
| Orgueil | CI1 | 0.14 | 0.05 | | 4 | This study |
| Aver. | | 0.14 | 0.13 | | | |
| Ivuna | CI1 | 0.11 | 0.02 | 0.006 | 8 | (Klaver et al., 2020) |
| Murchison | CM2 | 0.21 | 0.03 | | | (Cameron et al., 2009) |
| Murchison | CM2 | 0.23 | 0.07 | | 9 | (Gall et al., 2017) |
| Murchison | CM2 | 0.19 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Aver. | | 0.21 | 0.04 | | | |
| Paris | CM2 | 0.23 | 0.02 | | 4 | This study |
| Kainsaz | CO3.2 | 0.20 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Felix | CO3.3 | 0.31 | 0.07 | | | (Cameron et al., 2009) |
| Ornans | CO3.4 | 0.29 | 0.08 | | 9 | (Gall et al., 2017) |
| Ornans | CO3.4 | 0.21 | 0.02 | 0.006 | 8 | (Klaver et al., 2020) |
| Aver. | | 0.25 | 0.11 | | | |
| Renazzo | CR2 | 0.16 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Al Rais | CR2 | 0.22 | 0.02 | 0.006 | 8 | (Klaver et al., 2020) |
| Leoville | CV3.1 | 0.30 | 0.05 | | | (Cameron et al., 2009) |
| Kaba | CV3.1 | 0.22 | 0.04 | | | (Wang et al., 2021) |
| Allende | CV3.6 | 0.24 | 0.07 | | 9 | (Gall et al., 2017) |
| Allende | CV3.6 | 0.24 | 0.03 | 0.010 | 8 | (Klaver et al., 2020) |
| Allende | CV3.6 | 0.25 | 0.04 | | | (Wang et al., 2021) |
| Allende | CV3.6 | 0.23 | 0.02 | | 4 | |
| Aver. | | 0.25 | 0.01 | | | |
| Karoonda | CK4 | 0.28 | 0.02 | 0.008 | 8 | (Klaver et al., 2020) |
| Karoonda | CK4 | 0.39 | 0.04 | | | (Wang et al., 2021) |
| Aver. | | 0.33 | 0.16 | | | |
| Kota-kota | EH3 | 0.20 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Kota-kota | EH3 | 0.26 | 0.04 | | | (Wang et al., 2021) |
| Kota-kota | EH3 | 0.21 | 0.04 | | | (Wang et al., 2021) |
| Aver. | | 0.22 | 0.07 | | | |
| Abee | EH4 | 0.19 | 0.05 | | | (Cameron et al., 2009) |
| Abee | EH4 | 0.25 | 0.02 | 0.008 | 7 | (Klaver et al., 2020) |
| Aver. | | 0.22 | 0.09 | | | |

| | | | | | | |
|----------------|-------|-------------|------|-------|-----|----------------------------|
| Indarch | EH4 | 0.27 | 0.09 | | 9 | (Gall et al., 2017) |
| Indarch | EH4 | 0.19 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Aver. | | 0.23 | 0.12 | | | |
| St. Mark's | EH5 | 0.18 | 0.02 | 0.006 | 15 | (Klaver et al., 2020) |
| Khairpur | EL6 | 0.29 | 0.08 | | 9 | (Gall et al., 2017) |
| Khairpur | EL6 | 0.21 | 0.02 | 0.007 | 11 | (Klaver et al., 2020) |
| Aver. | | 0.25 | 0.12 | | | |
| Atlanata | EL6 | 0.21 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Hvittis | EL6 | 0.22 | 0.02 | 0.008 | 7 | (Klaver et al., 2020) |
| Yilmia | EL6 | 0.21 | 0.02 | 0.008 | 8 | (Klaver et al., 2020) |
| A 10224 | L3 | 0.24 | 0.01 | | 3-5 | (Chemonzhkin et al., 2016) |
| Ceniceros | L3.7 | 0.20 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Barratta | L4 | 0.31 | 0.05 | | 9 | (Gall et al., 2017) |
| Barratta | L4 | 0.19 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Aver. | | 0.25 | 0.16 | | | |
| Bruderheim | L6 | 0.51 | 0.07 | | 5 | (Gall et al., 2017) |
| Chainpur | LL3.4 | 0.28 | 0.10 | | | (Cameron et al., 2009) |
| Chainpur | LL3.4 | 0.28 | 0.04 | | | (Wang et al., 2021) |
| Aver. | | 0.28 | 0.00 | | | |
| Parnalee | LL3.6 | 0.20 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| A 09135 | LL3 | 0.15 | 0.02 | | 3-5 | (Chemonzhkin et al., 2016) |
| Parnalee | LL3 | 0.16 | 0.05 | | 4 | (Gall et al., 2017) |
| Chelyabinsk | LL5 | 0.27 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Dhurmsala | LL6 | 0.22 | 0.02 | 0.008 | 8 | (Klaver et al., 2020) |
| Kilabo | LL6 | 0.25 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| St. Severin | LL6 | 0.24 | 0.05 | | 5 | (Gall et al., 2017) |
| A 09436 | H3 | 0.20 | 0.04 | | 3-5 | (Chemonzhkin et al., 2016) |
| Bremervorde | H3 | 0.26 | 0.05 | | 5 | (Gall et al., 2017) |
| Buzzard Coulee | H4 | 0.17 | 0.03 | 0.009 | 8 | (Klaver et al., 2020) |
| Kernouve | H6 | 0.37 | 0.08 | | 4 | (Gall et al., 2017) |

Note: All chondrites show average $\delta^{60/58}\text{Ni}$ value of 0.23 ± 0.14 (2SD) ± 0.02 (2SE, N = 37). Literature sources: (Cameron et al., 2009; Chernonozhkin et al., 2016; Gall et al., 2017; Klaver et al., 2020; Steele et al., 2012; Wang et al., 2021)

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