

## Mercury deposition in the Eastern Mediterranean: Modern fluxes in the water column and Holocene accumulation rates in abyssal sediment

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### Abstract :

Modern and past mercury (Hg) fluxes in the oceanic water column and abyssal sediments are poorly quantified. Here, we investigated the particulate transfer of Hg in the water column of the ultra-oligotrophic Ionian Sea (Eastern Mediterranean) with sediment traps during a one-year period, and its accumulation in the deep central abyssal plain using sediment cores comprising the last 10 ka. The Hg concentrations in the particles collected in the sediment traps varied from 112 to 401 ng g<sup>-1</sup> and enabled quantifying annual Hg fluxes of 2.0, 2.5, and 2.5 µg m<sup>-2</sup> a<sup>-1</sup>, for traps deployed at 250, 1440, and 2820 m deep, respectively. Hg collected in the upper trap originates from atmospheric deposition, including Saharan dust, which is scavenged by the biological pump. Higher Hg fluxes found at mid-depth and near-bottom than in the upper water layer are attributed to lateral advection under the mixed layer of Hg-rich resuspended sediments from the Adriatic continental margin. In the abyssal sediment, Hg concentrations range from 15 to 134 ng g<sup>-1</sup> with the highest levels in the Sapropel S1. Methylmercury concentrations varied from 0.06 to 0.24 ng g<sup>-1</sup> following the distribution of total Hg, with evidence of its specific accumulation at the oxidized front of the sapropel. We estimated that <1.8% of the total Hg in the sedimentary column was diagenetically reallocated. The reconstruction of historical Hg accumulation rates (HgAR) during the Holocene shows low pre-anthropogenic values (~0.3 µg m<sup>-2</sup> a<sup>-1</sup> before 4 ka BP), increasing up to ~0.9 µg m<sup>-2</sup> a<sup>-1</sup> during the late Iron Age and the Roman period (1.5–2.5 ka BP), and up to 2.9 µg m<sup>-2</sup> a<sup>-1</sup> during the Industrial Era. During the Sapropel S1 period (~6–10 ka BP), HgARs rose to 6.4 µg m<sup>-2</sup> a<sup>-1</sup> likely due to the intensity of the Hg removal by the biological pump, the organic matter

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preservation, along with high inputs of Hg-rich terrigenous matter and a possible restricted recycling in the atmosphere. Hg accumulation in the Ionian Sea deep sediment is found ~3-fold lower than those in the western Mediterranean abyssal plain.

### Highlights

► Mercury (Hg) was studied in water column particles and sediments of the Ionian Sea. ► Average Hg fluxes through the water column varied between 2.0 and 2.5  $\mu\text{g m}^{-2} \text{a}^{-1}$ . ► Sediment Hg accumulation rates (HgAR) for the Industrial Era was 2.9  $\mu\text{g m}^{-2} \text{a}^{-1}$ . ► HgAR in sediment has increased 3-fold in the last 4 ka. ► Before 6000 years ago in the Sapropel S1, the Hg AR reached up to 6.4  $\mu\text{g m}^{-2} \text{a}^{-1}$ .

**Keywords** : mercury, Mediterranean, flux, accumulation rates, sediment traps, sediment core

## **1. Introduction**

The mercury (Hg) transfer and accumulation in deep-sea sediments are still poorly quantified (e.g., Outridge et al., 2018; Hayes et al., 2021). Estimates of Hg fluxes associated with oceanic settling particles vary by more than one order of magnitude from 0.3 to 4.6  $\mu\text{g m}^{-2} \text{a}^{-1}$  (Munson et al., 2015; Cossa et al. 2021), and the global deep-sea Hg annual burial estimations fluctuate by a factor of 7 from 220 to 1540 Mg (Amos et al., 2014; Zhang et al., 2015; Outridge et al., 2018; Hayes et al., 2021). This is explained by the large variability of the few published measurements on Hg concentrations in sediment

trap material, and abyssal sediments (e.g., Gobeil et al., 1999; Ogrinc et al., 2007; Gehrke et al., 2009; Munson et al., 2015; Cossa et al., 2021). The high Hg contents and accumulation rates found recently for hadal sediments reinforce the need for more work on deep-ocean Hg fluxes to better constrain global Hg models (Sanei et al., 2021). It is, therefore, necessary to acquire new Hg measurements and further estimate the flux of sinking particles down to the deep-sea floor from different oceanographic regions presenting biogeochemical contexts as diverse as possible.

Several approaches can be used to estimate the Hg transfer and accumulation rate (HgAR): multiplying Hg concentrations in a given layer of sediment or in particles collected in a sediment trap (Rutten et al., 2000; Outridge et al., 2007; Lim et al., 2017; Aksentov et al., 2021), by sedimentation rates (SR), whilst taking into account porosity and density of the sediments; alternatively, it is feasible to access HgAR by normalizing Hg sediment concentrations to the concentration of a tracer with a known vertical flux, such as  $^{230}\text{Th}$  (e.g., Bacon, 1984). Hayes et al. (2021) assert that the  $^{230}\text{Th}$  approach has the advantage of not being affected by age model error and by post-depositional redistribution of the sediments. Indeed, the first approaches (i.e., Hg concentration multiplied by SR) are only seemingly simple. Although post-depositional mobility of Hg in sediments is small, it is not completely negligible (Gobeil and Cossa, 1993; Mercone et al., 1999; Rydberg et al., 2008; Gehrke et al., 2009; Cossa et al., 2021). On the other hand, the journey of particulate Hg through the ocean water column down to the sediments is accompanied by homogenous and heterogenous chemical transformations (e.g., Lamborg et al., 2016). The efficiency of the vertical transfer of Hg to sediments is governed by plankton productivity and its uptake rate (i.e., the intensity of the biological pump) and by the aggregation of biogenic particles together and with lithogenic material (Mason et al., 2012; Lamborg et al. 2014; Munson et al., 2015; Zaferani et al., 2018; Jiskra et al., 2021). Indeed, Hg incorporation and accumulation into settling particles depend on the scavenging capacity of the carrier phase (particulate organic matter (POM) including plankton, detrital or authigenic material), and, *a contrario*, on the ability of Hg remobilization during the solubilization-mineralization of the settling POM, occurring particularly *via* Hg methylation or reduction during the downward export of settling material. As reported above, field measurements have shown that the intensity and

composition of the downward Hg fluxes are highly variable depending on the water depth, the geographical location, and the season. In the open ocean, Hg particulate fluxes originate mainly from atmospheric deposition, including aerosol deposition and gas exchange, whereas, in areas closer to the margins or ridges, lateral advection of material from shelves and hydrothermal sources constitutes additional sources (e.g., Fitzgerald et al., 2007; Mason et al., 2012). Atmospheric dust deposition upon oceans is patchy and gas exchanges at the sea surface are highly dependent on the existing physical conditions. Hydrothermal inputs are rare and very localized. In that context, several questions concerning the scavenging and accumulation rates of Hg in open ocean sediments remain open to discussion. Among them: (i) what are the Hg sources (ii) what are the variations of the downward Hg flux in the water column, (iii) what is the time trend of the HgAR over millennia, and (iv) which proportion of Hg is involved in methylation/demethylation processes and diagenetic rearrangement after its incorporation into sediments?

Here, we present the results of Hg distributions in particles collected in sediment traps during a one-year period and in ~30 cm long sediment cores from a central deep location of the ultra-oligotrophic Ionian Sea (eastern Mediterranean). This allowed us to explore the Hg deposition during the Holocene and its possible mobility associated with the presence of the well-known S1 sapropel (e.g., De Lange et al., 1989). This article is a companion paper to an analogous study carried out in the Balearic basin of the western Mediterranean Sea (Cossa et al., 2021), which provided a Late Holocene reconstruction of HgARs.

## **2. General settings of the Ionian Sea sediments**

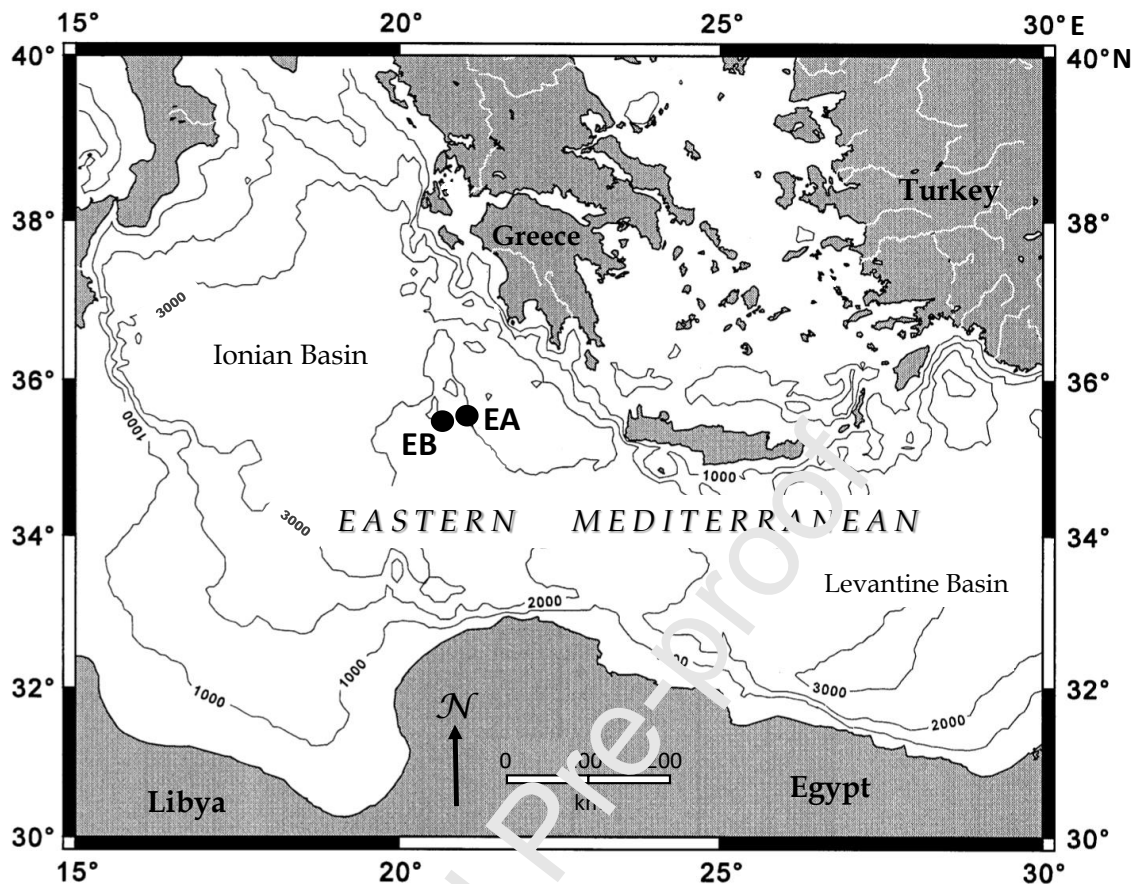
The pelagic eastern Mediterranean Sea is ultra-oligotrophic with a low sedimentation rate (SR) averaging between ~2 and 4 cm ka<sup>-1</sup> (e.g., Thomson et al., 1995; Garcia-Orellana et al., 2009; The MerMex Group, 2011). In the Ionian Sea, SRs have remained constant for at least the last past 1000 years (Castagnoli et al., 1990). The principal contributors to sedimentation are eolian deposition mainly from Sahara dust, and autochthonous organic matter, including biogenic carbonates, whereas continental inputs from coastal erosion and rivers are low (e.g., Guerzoni et al., 1997; Rutten et al., 2000). However, mud

volcanoes are common structures in the eastern Mediterranean but are mostly located along the Mediterranean ridge (e.g., Nikitas et al., 2021). Recent surficial sediments (~1.5 cm) are well-oxygenated and bio-mixed (e.g., Pruyssers et al., 1991; Ogrinc et al., 2007; Angelidis et al., 2011). In contrast, older and deeper sediments are composed of alternating sequences of organic-poor (<0.3% organic carbon) hemipelagic sediments and organic-rich (from 0.5 to 30% organic carbon) sapropelic layers (De Lange et al., 1989; Calvert et al., 1992; Bouloubassi et al., 1999). These alternations result from sharp changes in the photic layer productivity and redox conditions during sediment deposition, such as the sapropel dated to the transition between the Holocene and the Pleistocene (Gehrke et al., 2009; The MerMex Group, 2011).

### 3. Material and methods

#### 3.1. Samplings

Sediment trap moorings and collections and sediment corings have been performed in April 2001, October 2001, and April 2002 during the ADIOS cruises conducted in the Ionian basin. Automated sediment traps (Technicap® cylindrical traps, Heussner et al., 1990) with a 0.125 m<sup>2</sup> collecting area were deployed at Sta. EB in the Ionian Sea (Fig. 1). Three sediment traps with 12 collection cups were set at each line at nominal depths of 250 m and 1440 m below the surface, and 2820 m (25 m above the sea bottom). The samples were collected every two weeks for more than 11 months from the 1<sup>st</sup> of May 2001 to the 15<sup>th</sup> of April 2002. A total of 69 samples were collected for the elementary analyses. However, when the amount of material was insufficient for the analysis, samples were sometimes pooled resulting in 28 samples for Hg analyses. The samples were frozen immediately upon recovery, then freeze-dried, homogenized, and subsampled for various subsequent chemical analyses. An undisturbed 30 cm-long box core was obtained from Sta. EA in the Ionian Basin of the Eastern Mediterranean in April 2001, 25 km away from Sta. EB (Fig. 1). The core has been sliced at 0.5 cm intervals over the first 5 cm, at 1 cm intervals from 5 to 20 cm, and at 2 cm intervals from 20 cm to the bottom, on board immediately after recovery. An additional shorter box core was also obtained at Sta. EB, and was sliced only down to 15 cm.



**Figure 1.** Locations of the sampling stations in the Ionian Sea. EA (Lat.: 35°04.08'N; Long.: 2°50.94'E): coring; EB (Lat.: 35°54.01'N; Long.: 2°30.49'E): sediment traps mooring. See Suppl. Fig. SI.1 for the meteorological conditions and Saharan dust distribution prevailing during the sampling cruise.

### 3.2. Analyses and accumulation rate calculations

Total Hg analyses were performed using an automatic atomic absorption spectrophotometer (Altec®, model AMA-254), according to the protocol described by EPA (2007). Briefly, weighed aliquots of freeze-dried sediment were introduced into the analytical system where they were heated to 550°C. The volatilized Hg was carried by a stream of oxygen to a gold trap onto which Hg was concentrated by amalgamation, before being thermally dissociated and analyzed by atomic absorption spectrometry. The detection limit (DL) was 7 ng g<sup>-1</sup>, calculated as 3.3 times the standard deviation of the blanks. We used the MESS-2 marine sediment certified reference material (CRM) from the National Research Council of Canada to ensure the accuracy of the analyses; the mean value obtained (92 ± 2 ng g<sup>-1</sup>) was always within the range of the certified value: 92

$\pm 9 \text{ ng g}^{-1}$ . The reproducibility, defined as the coefficient of variation of six replicate analyses of the same CRM was 2%, whereas the analytical uncertainty (Eurachem/Citac, 2000) is  $\sim 10\%$ . Labile Hg concentrations were estimated upon partial dissolution of the sediment with diluted HCl (0.5 M) solution according to the protocol described by Kostka and Luther (1994).

Methylmercury (MeHg) concentrations were determined using a method adapted from Leermakers et al. (2001). MeHg was solubilized in 4 M HNO<sub>3</sub> from a freeze-dried sediment aliquot, then extracted by CH<sub>2</sub>Cl<sub>2</sub> and transferred into demineralized water by evaporation of the organic solvent. MeHg in the water phase was then ethylated and purged on a Tenax-packed column. The ethylmethylmercury formed was isolated from other volatile compounds by gas chromatography and quantified by cold vapor atomic fluorescence spectroscopy. The DL (DL = 3.3 times the sd of blank values) was  $0.003 \text{ ng g}^{-1}$ . The analytical precision, determined from replicate analyses ( $n = 6$ ), was  $<15\%$ . Recoveries of MeHg in the CRM 405 material from the International Atomic Energy Agency (IAEA) were  $91 \pm 8\%$  ( $n = 6$ ).

Total organic carbon (C<sub>org</sub>) content and its isotopic composition ( $\delta^{13}\text{C}_{\text{org}}$ ) were measured on freeze-dried and homogenized subsamples after decarbonation. Total carbon (C<sub>t</sub>) and nitrogen (N<sub>t</sub>) were determined with an elemental analyzer (Model CN 2000, LECO®) after acidification with 2 M HCl (overnight, at 50°C. The precision (6 replicates of a sample) for C<sub>org</sub> and C<sub>t</sub> analyses was 0.1%, and 0.3 % for N<sub>org</sub>. Concentrations are expressed as the weight percent of dry matter. Calcium carbonate content was calculated from mineral carbon (C<sub>t</sub> - C<sub>org</sub>) using the molecular mass ratio CaCO<sub>3</sub>:C = 100:12. Stable isotopes have been on a Flash 1112 EA elemental analyzer interfaced to a Delta C Finnigan® MAT isotope ratio mass spectrometer. Uncertainty was lower than 0.05‰.

Major elements were analyzed by X-ray fluorescence spectrometry and results were published earlier (Angelidis et al., 2011). Labile Fe (and associated Hg) concentrations were estimated upon partial dissolution of the sediment with dilute HCl (0.5 M) solutions according to protocols described by Kostka and Luther (1994). The HCl extraction



dissolves carbonates, crystalline Fe-oxyhydroxides, and acid-volatile sulfides. It should be noted that the 0.5 M HCl extraction does not dissolve HgS (Mikac et al., 2000).

HgARs were computed with consideration of the sediment compaction, as performed previously for the western Mediterranean Sea (Cossa et al., 2021), following the equation:

$$\text{HgARs} = 10 * [\text{THg}] * \text{SR} * (1 - \varphi) * d \quad (1)$$

where [THg] is the Hg concentration ( $\text{ng g}^{-1}$ ), SR the sedimentation rate ( $\text{cm yr}^{-1}$ ),  $\varphi$  the porosity of the sediment (Suppl. Table S1), and d the average dry bulk sediment density,  $2.65 \text{ g cm}^{-3}$  according to Hamilton (1976).

## 4. Results

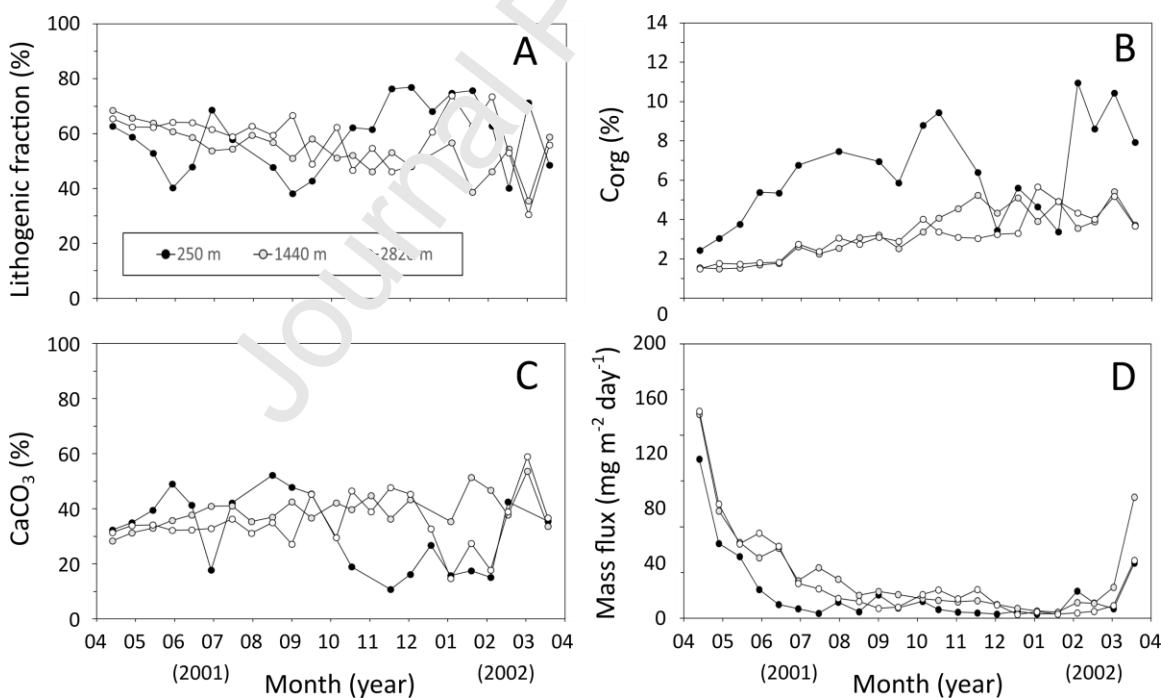
### 4.1. Sediment trap material

Microscopic examination and X-Ray diffraction (XRD) analysis of the material collected in the traps revealed the presence of fecal pellets and planktonic organisms with the dominance of foraminifera and protozoa species embedded in an amorphous organic and lithogenic material. The details of these observations are given in the ADIOS final report (ADIOS, 2004) and all the analytical results are presented in Suppl. Table SI.2. The lithogenic fraction was prevailing in terms of mass (i.e., 40 to 77%, Fig. 2A). Biogenic carbonates constituted the second major component of the collected material in the traps (i.e., 11 to 59%, Fig. 2C) and their distribution over time mirrors the lithogenic fraction. None of these fractions showed statistically significant differences with depth (Table 1).  $C_{\text{org}}$  increases from 2% at the beginning (April 2001) to 10% at the end of the experiment (March 2002), with a marked drop in winter more than likely as a result of the low productivity in the mixed layer at this season (Fig. 2B). The annual mean  $C_{\text{org}}$  concentration at 250 m was twice ( $p < 0.01$ ) the ones at other depths except during winter. This is illustrative of the dynamic of biological production in the mixed layer and of the remineralization process below. The  $\delta^{13}\text{C}_{\text{org}}$  ranged between  $-22.2\text{‰}$  and  $-24.5\text{‰}$ , and the  $\delta^{15}\text{N}$  signature between  $+0.9\text{‰}$  and  $+3.9\text{‰}$  (Suppl. Table SI.3). These tracers varied in-mirror, with  $\delta^{13}\text{C}_{\text{org}}$  decreasing and  $\delta^{15}\text{N}$  increasing, during summer and autumn at

depth suggesting an enhancing contribution of terrigenous organic material (Suppl. Fig. SI.2). The settling particle total mass flux (Fig. 2D) peaked during the spring (i.e., April 2001 and March 2002) with 10 to 20 times higher fluxes than the autumn and winter baseline values. The total mass flux during the entire collection period for the site EB was the lowest at 250 m ( $15 \text{ g m}^{-2} \text{ d}^{-1}$ ), compared to 27 and  $23 \text{ g m}^{-2} \text{ d}^{-1}$  at 1440, and 2820 m, respectively (ADIOS, 2004).

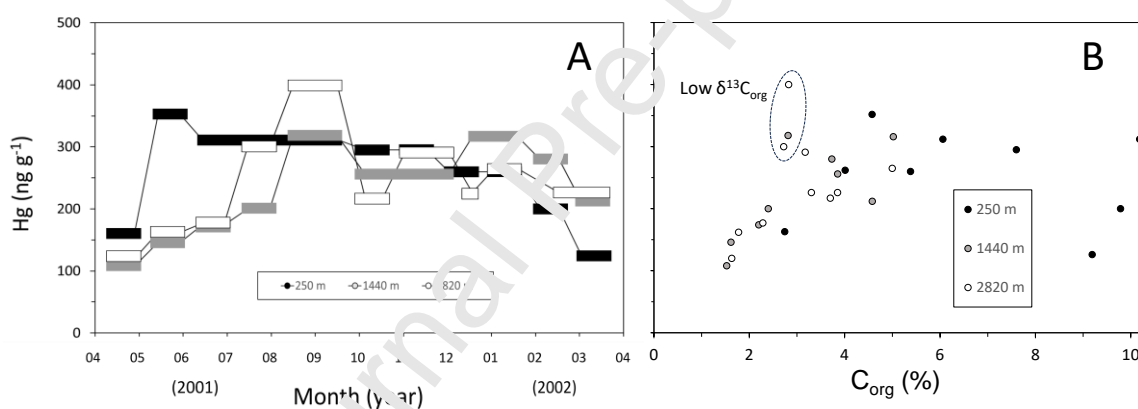
**Table 1.** Mean concentrations ( $\pm 1SD$ ) of the main components of the particles collected twice a month in the traps between May 2001 and April 2002 at Sta. EB. \* Mean concentration significantly ( $p < 0.01$ ) different from means at 1440 and 2820 m); other differences are non-significant.

Depth (m)	Lithogenic (%)	Carbonates (%)	C <sub>org</sub> (%)	Hg <sub>p</sub> (ng g <sup>-1</sup> )	Hg/C <sub>org</sub> (molar ratio 10 <sup>-6</sup> )
250	58.9 $\pm$ 12.9	31.7 $\pm$ 13.7	6.3* $\pm$ 2.5	255 $\pm$ 68	0.29* $\pm$ 0.13
1440	54.2 $\pm$ 8.0	39.3 $\pm$ 6.2	3.3 $\pm$ 1.7	236 $\pm$ 69	0.45 $\pm$ 0.12
2820	58.5 $\pm$ 9.7	35.1 $\pm$ 9.6	3.2 $\pm$ 1.1	248 $\pm$ 82	0.49 $\pm$ 0.19



**Figure 2.** (A) Lithogenic fraction, (B) organic carbon, (C) carbonates, and (D) mass flux during the year 2001-2002 in the three sediment traps at Sta. EB. Saharan dust events were mainly observed in spring.

Hg concentrations measured in the trapped particles varied from 112 to 401  $\text{ng g}^{-1}$ , with an overall mean Hg concentration of  $250 \pm 73 \text{ ng g}^{-1}$ . Mean Hg concentrations for the entire sampling period did not differ significantly with depth ( $t$ -tests,  $p > 0.10$ , Table 1). In contrast, concentrations differed over time with significantly lower values in the spring compared to other seasons (Fig. 3A). In particular, at 1440 and 2820 m, mean Hg concentrations were two-fold lower in spring compared to autumn ( $134 \pm 23 \text{ ng g}^{-1}$  versus  $310 \pm 73 \text{ ng g}^{-1}$ ,  $p < 0.01$ ,  $t$ -test). Interesting to note is the occurrence of high Hg concentrations in autumn when  $\delta^{15}\text{N}$  and  $\delta^{13}\text{C}_{\text{org}}$  values testify to the enhanced presence of terrestrial material. Despite the seasonal variation of Hg concentrations, Hg fluxes follow the total mass fluxes with the highest vertical transport in spring (Fig. 2D). The integrated Hg fluxes, calculated for the period May 2001 to April 2002, were  $2.0 \pm 0.8$ ,  $2.5 \pm 0.7$ , and  $2.5 \pm 0.9 \mu\text{g m}^{-2} \text{ a}^{-1}$ , at 250, 1440, and 2820 m, respectively.

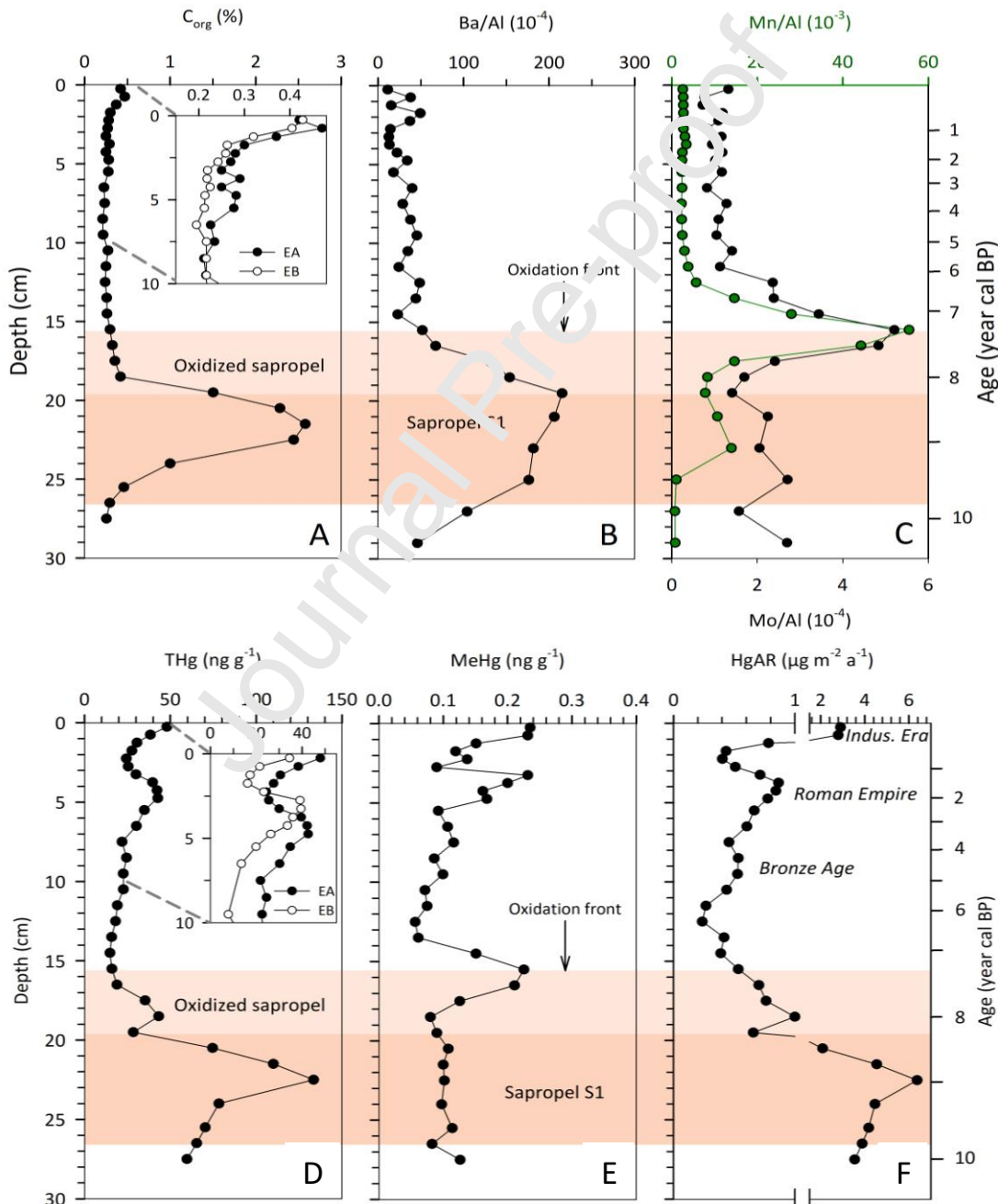


**Figure 3.** (A) Hg concentrations in the particulate material collected during the year 2001-2002 in the three sediment traps at Sta. EB. Low Hg concentrations are observed in spring; (B) Hg vs organic carbon ( $C_{\text{org}}$ ) in particulate material collected during the year 2001-2002 in the three sediment traps at Sta. EB. Low  $\delta^{13}\text{C}$  values, suggesting terrigenous OM, are associated with Hg-rich particles.  $C_{\text{org}} > 5\%$  values may involve a bio-dilution of particulate Hg

#### 4.2. Sediment core material

Core EA consists of light brown calcareous pelitic muddy sediments from surface to 18 cm. Below, a dark layer down to 24 cm corresponds to the level where Sapropel S1 has already been identified in the Eastern Mediterranean (e.g., De Lange et al., 1989; Thomson et al., 1995; Murat & Got, 2000). Core EB was similar but shorter and did not include the sapropel layer.  $C_{\text{org}}$  concentrations (Fig. 4A, insert) were  $< 0.5\%$  in most of the cores decreasing from the surface to 10 cm below (from  $\sim 0.5$  to  $\sim 0.2\%$ ) in both cores.

However, below 18 cm in core EA,  $C_{org}$  increased sharply to peak (2.6%) at 21.5 cm, testifying of the sapropel layer, which is confirmed by the highest concentration of Ba/Al, a paleoproductivity indicator, between 20 and 25 cm (Fig. 4B). The distributions Mn/Al and Mo/Al ratios (Fig. 4C) were typical of redox-sensitive elements. They exhibit well-defined peaks above the sapropel (~15-17 cm) indicating the oxidation front of the sapropel, consistent with patterns reported for the Sapropel S1 in a nearby station of the same basin (Filippidi and De Lange, 2019).



**Figure 4.** Vertical distributions of  $C_{org}$  (A), Ba/Al (B), Mn/Al with Mo/Al (C), Hg (D), MeHg (E), and HgAR (F) in sediment cores EA and EB. The onset, ending, and oxidized part of the sapropel S1 are indicated (see text). The oxidized front is defined by the presence of Mn/Al and Mo/Al peaks.

Hg concentrations in cores EA and EB ranged from 15 ng g<sup>-1</sup> to 134 ng g<sup>-1</sup> (Fig. 4D). Vertical profiles of Hg concentrations are characterized by smooth variations including (i) a decrease between 0 and 2 cm, then (ii) a bump between 2 and 7 cm, and (iii) low concentrations down to 10 cm. In addition, two Hg peaks were associated with the sapropel layer in core EA: the main Hg peak (between 21 and 23 cm) coincidental with the maximum of  $C_{org}$  (Fig. 4A), and the secondary peak (18.5 cm) in the oxidized part of the sapropel (Fig. 4D). It is important to note that the normal zea relations of Hg/Al or Hg/Si exhibit similar profiles as the Hg concentration (Suppl. Fig. SI.3), which demonstrates that the Hg variations are not due to changes in the lithology of the sediments. The MeHg measured in the EA core (0.05 to 0.24 ng g<sup>-1</sup>) represents on average 0.4% of the total Hg. Its vertical distribution (Fig. 4E) exhibits 3 peaks; one at the surface, the second between 3 and 5 cm, and the third between 14 and 17 cm. The first two peaks of the MeHg distribution are coincidental with total Hg peaks (Fig. 3A and B).

## 5. Discussion

### 5.1. Sources of particles and associated Hg in the trapped material

Particulate matter collected in the sediment traps at station EB is mainly composed of lithogenic particles and biogenic carbonates suggesting the presence of terrestrial material embedded in a biogenic matrix. Vertical mass fluxes reach their maximum in spring whatever the depth, but are higher in the deep traps compared to the upper one. This pattern supports major particulate mass fluxes during the biological productive period and lateral advection below the mixed layer (Fig. 2). In contrast, the OM fraction is higher in the upper trap material compared to deeper ones except during winter when the biological production collapses.

The composition of the lithogenic material is dominated by quartz, clay minerals, and feldspars (ADIOS, 2004), indicating various sources including the advection of river

particles, coastal erosion, resuspension of sediments, and Saharan dust deposition. Indeed, Adriatic sediments are composed of sand, silt, and clay, including significant fractions of quartz, calcareous material, feldspars, and OM (Faganeli et al 1994; Droghini et al., 2019). On the other hand, according to Kandler et al. (2007), the composition of the Saharan dust particles is dominated by phyllosilicates, with significant fractions of quartz, and carbonaceous material. Although the expected increase in kaolinite during Saharan dust events was not ascertained, 40% of our sampling period consisted of dusty days and Saharan dust events have been registered in the spring of 2001 (Suppl. Fig. SI.1; Georgopoulos et al., 2002; ADIOS 2004). The presence of biogenic carbonates formation in the mixed layer at station EB is attested by the temporal covariation of  $\text{CaCO}_3$  and  $\text{C}_{\text{org}}$  content of the material collected in the trap at 250 m (Fig. 2B and C). Our results are consistent with previous observations showing that biogenic carbonates constitute a major part of the phytoplankton production and carbonate flux between February and June in the Ionian Sea (Moulin et al., 1998; Cowley et al., 2017; Skampa et al., 2020; Pedrosa-Pamies et al., 2021). Such an increase in primary production has been shown to be triggered by the nutrient enrichment of surface water resulting from Saharan dust deposition (e.g., Gallisai et al., 2014; Pedrosa-Pamies et al., 2016). Thus, the seasonality of the mass flux appears to be governed by variations in phytoplankton production in the photic zone and Saharan dust deposition in a process similar to that described for the Northwestern Mediterranean wherein the downward particulate flux is modulated by the lithogenic contribution to settling particles ballasting the biogenic material (Ternon et al., 2010). The isotopic compositions of C and N allow us to identify the origin of the trapped POM. The highest  $\delta^{13}\text{C}_{\text{org}}$  values found in spring indicate a marine origin of the POM, attributable to phytoplanktonic production in the euphotic layer. In May 2001, the  $\delta^{13}\text{C}_{\text{org}}$  and  $\delta^{15}\text{N}$  values of sinking settling POM were very similar and covaried at 250 m, 1440 m, and 2820 m, supporting a top-down control of the organic fraction flux, from the euphotic zone to the bottom. In contrast, at the end of the summer and the beginning of autumn, lower  $\delta^{13}\text{C}_{\text{org}}$  and higher  $\delta^{15}\text{N}$  values of POM collected in the intermediate and deeper traps than at 250 m supports increased inputs of terrigenous material below the mixed layer.

The involvement of the Hg vertical transfer in the water column with the POM dynamic (biological pump and the terrigenous advection) is illustrated in Figure 3B. When  $C_{org}$  is  $< 5\%$ , the positive covariation between the Hg and  $C_{org}$  ( $R^2 = 0.39$ ,  $p < 0.05$ ) suggests that POM plays a major role in the vertical transfer of Hg to the bottom sediments. This positive relationship is even much stronger ( $R^2 = 0.73$ ,  $p < 0.01$ ) when values of the terrigenous samples (with low  $\delta^{13}C_{org}$  values) collected at the end of summer and the beginning of autumn at 1440 and 2820 m are removed from the calculation. When  $C_{org}$  concentrations are  $> 5\%$ , i.e., during the productive period of early spring 2002 (Fig. 2B), Hg concentrations do not increase with  $C_{org}$  (Fig. 3B), suggesting that the Hg available for the biological pump has been exhausted from the mixed layer. In other words, this pattern would result from the biodilution of the Hg pool of the euphotic zone during phytoplanktonic efflorescence. Besides, we assume that most of the lateral inputs of particulate matter to the EB site originate from the resuspension of Adriatic sediment with a substantial terrigenous origin (Manca et al., 2002) conveyed to the interior of the Ionian Sea by convective currents (e.g., Berline et al., 2021; Pedrosa-Pamies et al., 2021). Thus, the Hg concentrations of the Adriatic surface sediments should determine the magnitude of the Hg advected with this process. Hg concentrations in Adriatic sediments vary widely from  $\sim 50 \text{ ng g}^{-1}$  for matter originating from the Southern part of the Adriatic to  $> 1000 \text{ ng g}^{-1}$  for that originating from its Northern part (Droghini et al., 2019). The presence of anthropogenic Hg in the Northern Adriatic sediments has been evidenced and ascribed to mining activities and direct industrial inputs (e.g., Kotnik et al., 2015). Combining these observations, it can be inferred that the high  $C_{org}$ , low  $\delta^{13}C_{org}$ , and high Hg content ( $> 250 \text{ ng g}^{-1}$ ) of the settling material collected at depth in autumn at the EB site reflect the advection of Hg-enriched POM originating from the resuspension of Hg contaminated sediments likely from the Northern Adriatic. On the other hand, the Hg concentrations in the Saharan dust are less constrained than those of the Adriatic sediments. The few Hg concentrations available for the Saharan dust aerosols measured at the Cape Verde and Canaries Islands (Bailey, 2021) give a mean Hg concentration of  $73 \pm 7 \text{ ng g}^{-1}$  ( $n=5$ ) for the year 2020. In the spring of 2001, the Eastern Mediterranean atmosphere was characterized by southerlies transporting high quantities of Saharan dust (Georgopoulos et al., 2002; Sciarra et al.,

2004). Moreover, the spring-2001 cruise took place immediately after a cloud of Saharan dust passed over the Ionian region (Suppl. Fig. SI.1). At that time, the local circulation was favorable for dust transport from the surface to the deeper layers, due to the presence of a deep anticyclone. In spring 2001, Hg concentrations collected at the 250 m trap were greater than  $150 \text{ ng g}^{-1}$ , ruling out the Saharan dust as the main Hg carrier. On the contrary, Saharan dust probably dilutes the Hg-rich biogenic particles. Furthermore, the most probable role of Saharan dust in the Hg transport to the bottom is a ballasting effect upon the biological pump. The contribution of Hg deposition associated with Saharan dust can be independently assessed based on the most recent estimates of the total mass deposition rates of Saharan dust onto the eastern Mediterranean basin ( $7.4 \text{ g m}^{-2} \text{ a}^{-1}$  according to Vincent et al., 2016) and the only available mean Hg concentration in the Saharan aerosols ( $73 \text{ ng g}^{-1}$  according to Bailey, 2021). In these conditions, the mean contribution of particulate Hg fluxes from the Sahara is  $\sim 0.52 \text{ } \mu\text{g m}^{-2} \text{ a}^{-1}$  and represents one-fifth of the total Hg collected in the traps. However, this estimate is fraught with great uncertainty, since Hg concentrations in Saharan dust are still weakly documented and the total mass deposition rates of Saharan dust on the Ionian Sea are poorly constrained, ranging from 1 to  $50 \text{ g m}^{-2} \text{ a}^{-1}$  (Guerzoni et al., 1999; ADIOS, 2004; Guieu and Shevchenko, 2015; Vincent et al., 2016).

### 5.2. Sources and accumulation of Hg in sediment above the sapropel

Recently, Friedling et al. (2023) concluded that high amplitude oxygen fluctuations in the water column and shallow sediment may lead to increased variability in Hg and OM buried in sediments. Indeed, according to these authors, under oxygen depletion conditions, a part of Hg can escape from sediments after methylation and reduction. Contrarily, aerobic degradation would affect the sedimentary OM, whereas Hg would be mainly retained in the sediment. In order to explore the possible Hg mobility associated with redox changes we proceeded to an estimation of the Hg species solubilized with dilute HCl (here called “mobile Hg fraction”), which is supposed to dissolve amorphous oxy(hydr)oxides and sulfides (Kostka and Luther, 1994). The mobile Hg fraction, which constitutes  $< 1.8\%$  of the total Hg (Suppl. Fig. SI.5), presents a vertical distribution in surface sediment indicating that the vertical Hg distribution within this layer is mainly governed by external input and the relative proportion of various carrier phases.



Consistently, the first 15 cm sediment of the core EA (i.e., above the S1 sapropel) is an oxic layer without any trace of change of redox conditions, which would make a possible reallocation of the Hg in the sedimentary column; thus, the possible involvement of Hg in diagenetic homo- and heterogeneous reactions is very unlikely.

As discussed for the particulate Hg flux through the water column, the main carrier which conveys Hg to the sediment is POM from planktonic production and terrestrial detritus inputs. First, the biological pumping, initiated by phytoplankton production, scavenges Hg from atmospheric deposition. Second, the advection of coastal material enriched or not with Hg from coastal regions has been identified in sediment traps. In addition, in the Ionian Sea, the role of volcanic and hydrothermal inputs must be taken into account. The contribution of the biological pump in bringing Hg to the sediments can be explored using the Hg/C<sub>org</sub> ratios (Suppl. Fig. SI.5A). As in the trapped particles, the Hg is significantly correlated with C<sub>org</sub> in deposited sediments, but under certain conditions. In Cores EA and EB the overall relationship is highly significant ( $R^2 = 0.59$ ,  $p < 0.01$ ), but mostly driven by the sapropel content, since above it the  $R^2$  is only 0.26 ( $p < 0.10$ ) (Suppl. Fig. SI.6B). In these parts of the cores (above the sapropel), the Hg:C<sub>org</sub> molar ratios varied from 0.32 to  $1.01 \times 10^{-6}$  and from 0.30 to  $1.08 \times 10^{-6}$  for the EA and EB, respectively (Suppl. Fig. SI.6A). In the redox conditions of the EA station, the aerobic degradation of the Hg-OM appears to raise the Hg:C<sub>org</sub> atomic ratios as illustrated by the difference between the sediment trap material collected at 25 mab (0.49) and the surface of the cores (0.68). These differences suggest that during the OM consumption below the mixed layer, the Hg initially bound to the POM and released upon its solubilization or remineralization must be re-bound to the residual settling particles and that this process occurs mostly in the water column near the bottom. If this process was continuing in the sediments we should observe an increase of the Hg:C<sub>org</sub> ratio with depth which is not the case (Fig. SI.6A). There, time variations of HgAR should be the driving factor.

Besides, it has been suggested that a part of Hg concentration within the first 5 cm below the interface in Ionian Sea sediments originates from the deposition of volcanic ashes (Ogrinc et al., 2019); this hypothesis was based on the changes of  $\delta^{202}\text{Hg}$  within this layer. However, the  $\delta^{202}\text{Hg}$  values are very variable Italian volcanic emissions (-1.74

to  $-0.11\text{‰}$  according to Zambardi et al., 2009), and more data are needed to accurately identify the trace of Hg from a volcanic origin in the deep Mediterranean sediments.

### 5.3. Evidence for Hg relocation in the S1 sapropel

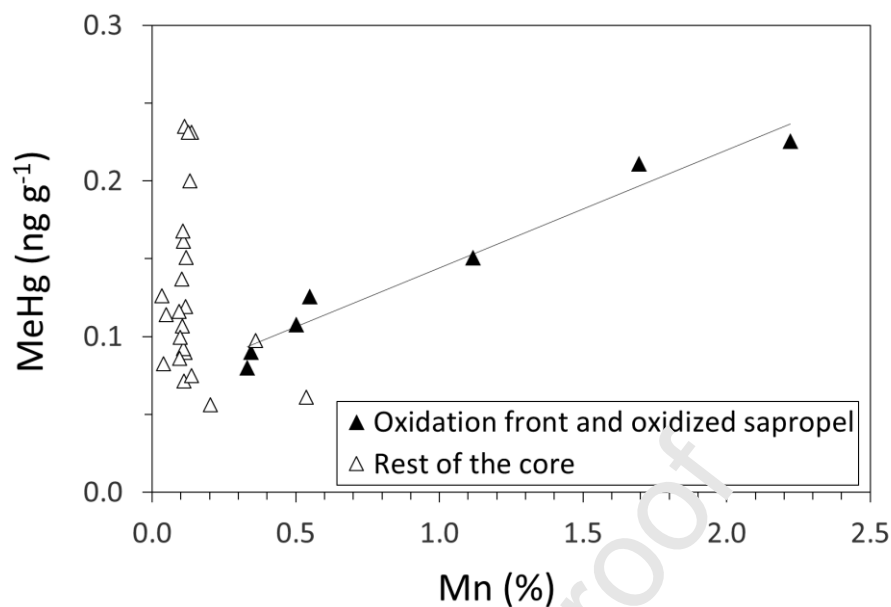
In the sediment core EA, sapropel S1 is characterized by high  $C_{\text{org}}$  concentrations (Fig. 4A) which result from OM preservation owing to oxygen-depleted conditions in the bottom waters and high primary production in surface water of the Ionian sea (e.g., Filippidi and De Lange, 2019; Murat & Got, 2000). The high primary production behind the formation of S1 is thought to be due to an increased discharge of freshwater into the Mediterranean following a change to a warmer and wetter climate during the African Humid Period (Ariztegui et al., 2000). The use of Ba/Al distribution, as a paleoproductivity indicator (due to the immobility of Ba as barite according to Thomson et al., 1995), gives more precise limits of the sediment layer concerned by the sapropel, namely 18-27 cm (Fig. 4B). The major element composition of the sedimentary column (Suppl. Fig. SI.4) indicates little changes in the source of material within the sapropel, with less biogenic carbonates and a larger lithogenic fraction. This likely results from a larger contribution of terrigenous material from rivers or Saharan dust during the sapropel deposition (Wehausen and Brumsack, 2000; Böttcher et al., 2003).

High Hg concentrations in Mediterranean sapropels have already been shown: (i) in S1 sapropel collected in the Aegean Sea (Mercone et al., 1999), and (ii) in a mi-Pleistocene sapropel collected in the Pyrenean basin of the western Mediterranean (Gehrke et al., 2009). According to these authors, this Hg enrichment is due to the intensity of the Hg removal by OM associated with oxygen depletion in the water column. In the case of the mid-Pleistocene Mediterranean sapropel, Gehrke et al. (2009) concluded even that OM quantitatively delivers the Hg present in the mixed layer to the seafloor. Other authors also argued that the high terrigenous fraction of the sapropelic OM may favor its elevated Hg: $C_{\text{org}}$  ratio (e.g., Böttcher et al., 2003). In core EA, the Hg: $C_{\text{org}}$  atomic ratio was quite variable (Fig. SI.6A) as it is in the settling particles collected at 250 m in the water column (Fig. 3B). These similar situations suggest that during the highly productive period of Sapropel S1, the biological pump was exhausting all the Hg present in the euphotic zone as observed by Gehrke et al. (2009) in the mid-Pleistocene sapropel.

In the oxidized part of the sapropel, a diagenetic mechanism may have induced the development of the secondary Hg peak observed at 18.5 cm depth (Fig. 4D) in the sediment layer defined by the dramatic decrease of Ba/Al and the rise in Mn/Al above (Fig. 4), where barite is dissolved in presence of (trace) sulfides, while Mn starts to precipitate in oxic sediments. This Hg peak should result from the reallocation of a small amount of Hg associated with sapropelic OM. In the Sapropel S1, sampled nearby in the Aegean Sea, Mercone et al. (1999) have already observed sharp Hg peaks in suboxic conditions where sediments containing trace pyrite have been re-oxidized. These authors proposed that such reallocation occurred following the release of Hg into the pore water solution during the reduction of oxide minerals. The Hg is then re-immobilized under reducing conditions associated with pyrite or in the form of H<sub>2</sub>S and/or HgSe. Here, the hypothesis of Hg incorporation in pyrite is supported by the significant relationship between Hg and Fe in the sapropel layer, i.e., between 16 and 25 cm below the sediment-water interface ( $R^2 = 0.58$ ,  $n = 9$ ,  $p < 0.01$ ). The mobile Hg fraction (i.e., Hg<sub>HCl</sub>, Suppl. Fig. SI.5), presents a main peak on the top of the oxidized sapropel coincidental with the Mn/Al and Mo/Al peaks (Fig. 4C). This suggests that a small proportion of Hg was remobilized during oxidation of OM diffusing along the concentration gradient, and readsorbed onto oxy(hydr)oxides (associated or not with remaining POM), as Cu and Zn already analyzed in this core (see figure 3 in Angelidis et al., 2011). As suggested by Mercone et al. (1999): "...[the] redox gradient across the [oxic-suboxic] boundary, imposes thermodynamic controls on elemental speciation, which leads to the development of peaks at slightly different depth levels for different elemental geochemistries". The amounts of relocalized trace elements vary in response to their relative reactivity with OM and adsorptive rate by oxy(hydr)oxides, in the order Cu < Zn << Hg (Feyte et al., 2010). Angelidis et al. (2011) have noted that, in core EA, the Pb profile (Suppl. Fig. SI.7), contrary to the cases of Cu and Zn, does not show evidence for diagenetic remobilization, and its distribution is governed by external input. During sediment sulfidization, especially in the sapropel layers, the Hg-Fe association would shift from the adsorption onto the oxy(hydr)oxides to the incorporation into sulfides.

#### 5.4. Methylmercury in sediments

MeHg pattern is consistent with the general model showing that the net MeHg abundance in sediments is a function of the inorganic Hg considered as the substrate of enzymatic reactions and that a part of the signal may remain inscribed in the oxic sediments (Cossa et al., 2014). The third MeHg peak is located above the sapropel, while the sapropel itself has low MeHg concentrations, but very high total Hg levels *a contrario* to what occurs in the two MeHg peaks described above (Fig. 4D and E). High MeHg:Hg ratios in sediments are often associated with Hg net methylation in suboxic environments resulting from sulfate or iron-reducing bacteria activities, whereas low MeHg:Hg ratios are associated with net demethylation occurring mostly under anaerobic conditions (e.g., Figueiredo et al., 2018; Bravo and Cosio, 2019; Regnell and Wat as, 2019). High MeHg levels at the oxidation front of the sapropel suggest a Hg methylation by sulfate-reducing bacteria at a relict suboxic zone. During its diffusion upward from the oxidizing zone, MeHg appears to be adsorbed onto Mn-oxy(hydr)oxides as suggested by the strong correlation between MeHg and Mn at the oxidation front (Fig. 5) and constitutes most of the mobile Hg fraction at the top of the oxidized sapropel (Suppl. Fig. SI.5). This link between MeHg and Mn-oxy(hydr)oxides should be favored by OM through the formation of stable bidentate complexes as predicted by modeling (Feyte et al., 2010). Below the oxidizing zone, it can be speculated that, if MeHg has been formed in suboxic conditions associated with OM-rich sediment of the sapropel, it has been subsequently demethylated since.



**Figure 5.** Methylmercury (MeHg) and manganese (Mn) relationship in the EA sediments. The oxidation front and oxidized sapropel cover the 14-20 cm sediment layer. In this layer, the MeHg vs Mn relationship exhibits an  $R^2$  of 0.96 ( $p < 0.01$ ).

#### 5.5. Historical reconstruction of Hg accumulation rates in Ionian abyssal sediment

To reconstruct historical Hg fluxes in the Ionian abyssal sediment, a chronological framework was established for core EA (Suppl. Fig. SI.8). An age-depth model was built through the combination of  $^{210}\text{Pb}$  activity data and radiocarbon ages of foraminifera (*Universa orbulina*) collected by Garcia-Orellana et al. (2009), at 0-0.5 cm ( $845 \pm 25$  a BP), 9-10 cm ( $4790 \pm 50$  and  $4520 \pm 40$  a BP), 13-14 cm ( $6420 \pm 35$  a BP), and 26-28 cm ( $11765 \pm 50$  a BP). The decay of  $^{210}\text{Pb}_{\text{ex}}$  activities and the distribution of other radionuclides (i.e.,  $^{137}\text{Cs}$  and  $^{239+240}\text{Pu}$ ) within the first 1.5 cm in Core EA indicates a mixing of this sediment covering the last past ~120 years (Garcia-Orellana et al., 2009). This is also consistent with the lowest  $^{206/207}\text{Pb}$  values of the EA record ( $< 1.20$ , Suppl. Fig. SI.7), which confirms that anthropogenic Pb from the industrial revolution (i.e., since ~100 BP or 1850 CE) is confined in this layer (Angelidis et al., 2011). In addition, the bottom and top of the Sapropel S1 were identified by correlating Ba/Al, Mn/Al, and  $\text{C}_{\text{org}}$  excursions (Fig. 4), and the corresponding ages (Crusius and Thomson, 2003; De Lange et al., 2008; Grant et al., 2016; Rohling et al., 2015; van Kemenade et al., 2023) were used to constrain the age model. Finally, the age-depth model was derived using the IntCal20 calibration curve (Reimer et al., 2020) with consideration of a  $390 \pm 85$  years

marine reservoir age of the Mediterranean Sea (Siani et al., 2000). All details for the construction of the age model are provided in Suppl. Fig. SI.8. Below the 1.5 cm, (i.e., before the Industrial Era), the obtained linear SRs were  $2.2 \pm 1.0 \text{ cm ka}^{-1}$ , in agreement with previously published SRs in that area (e.g., Rutten et al., 2000, De Lange et al., 2008). At the bottom of the core, linear SR increased in the Sapropel S1, up to  $\sim 4.3 \text{ cm ka}^{-1}$  consistently with reported values for the eastern Mediterranean Sea (De Lange et al., 2008; Filippidi and De Lange, 2019).

The relatively elevated Hg concentrations in the first 1.5 cm below the sediment surface (Fig. 4D) are a common feature in the abyssal sediment of the western Mediterranean (Cossa and Coquery, 2005; Ogrinc et al., 2007; Heimburger et al., 2012; Cossa et al., 2021). This Hg-enrichment has been attributed to the increased anthropogenic Hg depositions during the Industrial Era, partially bio-mixed within the 1.5 cm surface layer. The resulting Hg accumulation rate (HgAR) calculated for this period is  $2.9 \mu\text{g m}^{-2} \text{ a}^{-1}$  (Fig. 4F), which is similar to HgAR calculated from the sediment trap near the bottom ( $\sim 2.5 \mu\text{g m}^{-2} \text{ a}^{-1}$ ) but about 2 times lower than the values reported for the western basin surficial sediments ( $\sim 8.9 \mu\text{g m}^{-2} \text{ a}^{-1}$ , Fig. 6) (Cossa et al. 2021). It is worth mentioning that these superficial signals, as well as the deeper ones, have been smoothed by a high biological mixing (García-Crellana et al., 2009).

Below the superficial layer, a second Hg bump in concentration between 3 and 7 cm depth corresponds to a rise in HgAR from a baseline of  $\sim 0.3$  up to  $\sim 0.9 \mu\text{g m}^{-2} \text{ a}^{-1}$  and centered between 1500 and 1800 BP. A similar observation is made with Pb concentrations, PbARs and  $^{206/207}\text{Pb}$  profiles (Suppl. Fig. SI.7), and interpreted as the effect of anthropogenic deposition during that period (Angelidis et al., 2011). Such increases in Hg and Pb depositions for the same period have already been reported in a lagoon and the abyssal plain of the western Mediterranean, as well as in Alpine Lake sediments (Elbaz-Poulichet et al., 2011, 2020; Thevenon et al., 2011; Cossa et al., 2021) and attributed to the metallurgic development during the Iron Age and Roman periods. The climate optimum during the Roman Era may have also influenced HgAR, as this period is known as the warmest period of the last 2 ka, being  $\sim 2^\circ\text{C}$  warmer on average (Margaritelli et al., 2020). This may explain the slight increase in  $C_{\text{org}}$  observed between 3 and 6 cm in core EA (Fig. 4A insert). However, if such climatic influence may have

influenced HgAR in core EA, it is not the case for core EB where the  $C_{org}$  bump around 3-6 cm is not discernable (Fig. 4A, insert), whereas the Hg signal is distinctly present (Fig. 4D, insert). Another slight increase in HgARs (up to  $0.53 \mu\text{g m}^{-2} \text{a}^{-1}$ ) at ~8 cm depth (Fig. 4F) corresponds to the Bronze Age, a period marked by increased Hg depositions in certain Alpine lake records attributed to local mining activities (Elbaz-Poulichet et al., 2020). Finally, during the Sapropel S1 period (i.e., ~ 5.7 to 9.3 ka cal BP), the HgAR has reached up to  $6.4 \mu\text{g m}^{-2} \text{a}^{-1}$  (Fig. 4F) which is ~20 times higher than those of the preindustrial period. From the HgAR profile, the anthropogenic Hg accumulated since 6 ka in these sediments is  $4.0 \text{ mg m}^{-2}$  compared to 8.5 during the S1 period.

#### 5.6. Comparison of Hg dynamics between Eastern and Western Mediterranean

A comparison between Hg deposition, vertical transfer, and bottom accumulation between eastern and western Mediterranean abyssal plains is proposed in Fig. 6. The overall average Hg concentration of settling particles collected in the eastern Mediterranean was twice the one of those measured in the western Mediterranean during the same period (Cossa et al., 2021), whereas  $C_{org}$  was twice lower. This results in mean annual Hg: $C_{org}$  molar ratios (Table 1) being 3-5 times higher than their equivalent measured in the western Mediterranean basin. In other words, the POM is enriched in Hg in the ultra-oligotrophic eastern Mediterranean. The complex relationship between Hg and  $C_{org}$  testifies to the multiplicity of the POM and Hg sources. Hg originates from atmospheric and terrigenous sources, whereas POM mainly comes from *in situ* productivity in the mixed layer and lateral advection of terrigenous sources. The atmospheric Hg deposition consists of aerosol depositions (including Saharan dust and Hg-contaminated aerosols from Europe) and the gaseous exchange, which are removed from the water of the mixed layer by the biological pump and the embedding with marine snow. In this scavenging process, the Hg: $C_{org}$  ratio of the marine OM is limited by the amount of Hg deposited at the sea surface. On the other hand, the lateral advection of Hg-rich terrigenous OM (with high Hg: $C_{org}$  ratios due to anthropogenic inputs from riverine, erosion, and catchment sources) is conveyed from the Adriatic to the deep waters by the vertical convection (Manca et al., 2002). Broadly speaking, Hg is carried by OM but the Hg: $C_{org}$  ratios vary depending on the origin of the OM (marine vs terrigenous) and their anthropogenic components. To sum up, the decrease in Hg flux from surface to bottom

waters of the western basin is due to terrestrial contributions (anthropogenic aerosols and fluvial contributions) scavenged by the biological pump and combined with OM degradation below the upper layer where a part of Hg is involved in the microbiological loop to form mobile species such as  $\text{Hg}^0$  and MeHg. In contrast, in the eastern basin, where Hg flux tends to increase with depth, the main Hg inputs are due to Saharan aerosols ballasting the low planktonic matter and the resuspension of sediments from the continental margin. As a result, the vertical Hg transfer at mid-depth and near the bottom is not different between the eastern and the western basins ( $p < 0.01$ , t-tests). The slightly lower Hg fluxes measured in the bottom trap than in surficial sediments observed in the eastern basin, but more marked in the western basin, probably reflect the recent decrease in anthropogenic Hg emissions around the Mediterranean basin (Cossa et al., 2022). This trend is more clearly observed in the western basin, which is more subject to the influence of anthropogenic inputs, than in the eastern basin.

Western Mediterranean			Eastern Mediterranean			
Hg ( $\mu\text{g g}^{-1}$ )	Hg/ $\text{C}_{\text{org}}$ ( $10^{-6}$ )	Hg ( $\mu\text{g m}^{-2} \text{a}^{-1}$ )		Hg ( $\mu\text{g g}^{-1}$ )	Hg/ $\text{C}_{\text{org}}$ ( $10^{-6}$ )	Hg ( $\mu\text{g m}^{-2} \text{a}^{-1}$ )
$0.13 \pm 0.03$	$0.06 \pm 0.02$	$4.5 \pm 1.5$	250 m	$0.27 \pm 0.07$	$0.29 \pm 0.13$	$2.0 \pm 0.8$
$0.13 \pm 0.02$	$0.11 \pm 0.04$	$3.2 \pm 1.1$	1440 m	$0.24 \pm 0.07$	$0.45 \pm 0.12$	$2.5 \pm 0.7$
$0.13 \pm 0.02$	$0.15 \pm 0.05$	$3.1 \pm 1.1$	2820 m 25 mab	$0.25 \pm 0.08$	$0.49 \pm 0.19$	$2.5 \pm 0.9$
$0.08 \pm 0.01$	$0.82 \pm 0.1$	8.9	Industrial Era	$0.04 \pm 0.01$	$0.59 \pm 0.14$	2.9
$0.02 \pm 0.01$	$0.27 \pm 0.1$	0.9	Pre-anthropogenic*	$0.02 \pm 0.00$	$0.39 \pm 0.06$	0.3

**Figure 6.** Mercury (Hg) concentrations and fluxes in the eastern Mediterranean (Ionian abyssal plain) and western Mediterranean (Balearic abyssal plain). Data for the western Mediterranean are from Cossa et al. (2021). (mab) meter above the bottom. \* Refers to the pre-anthropogenic period, i.e., before 3 ka cal BP for the western basin and before 6 ka cal BP for the eastern basin excluding Sapropel S1.



## 6. Conclusions

This work represents the first direct estimation of the Hg transport through the water column and sedimentary accumulation in the deepest part of the eastern Mediterranean Sea. From the data acquired by sediment traps, mean annual Hg fluxes (May 2001-April 2002) of 2.0, 2.5, and 2.5  $\mu\text{g m}^{-2} \text{a}^{-1}$  have been calculated for depths of 250, 1440, and 2820 m (~25 mab), respectively. These Hg flux values are in the same order of magnitude as those estimated for the Mediterranean using the  $^{230}\text{Th}$  method ( $\sim 3 \mu\text{g m}^{-2} \text{a}^{-1}$  see figure 8c in Hayes et al., 2021), slightly lower than those calculated for the western Mediterranean Basin (4.5, 3.2, and 3.1  $\mu\text{g m}^{-2} \text{a}^{-1}$ ) at equivalent depths (Cossa et al., 2021), and ten times lower than those obtained in sediment trap moored close to the Mediterranean continental margin (Ligurian Sea), i.e., 30  $\mu\text{g m}^{-2} \text{a}^{-1}$  (Heimbürger et al., 2012). The highest rates of Hg deposition on the entire water column occur in spring associated with a marine snow consortium composed of Saharan dust and plankton detritus. In other seasons lateral advection of Hg-rich continental sediment occurs below the mixed layer; the convection of Hg-contaminated riverine and coastal sediments is the most probable source for this input.

The HgAR in the Ionian Sea sediments are characterized by the increasing Hg emissions during the Industrial Era and also reflect their changes in the Holocene period. We estimate the pre-anthropogenic HgARs  $\sim 0.3 \mu\text{g m}^{-2} \text{a}^{-1}$ , and the fluxes in the last 150 years around  $2.9 \mu\text{g m}^{-2} \text{a}^{-1}$ ; during the Roman period, HgARs have reached  $\sim 0.9 \mu\text{g m}^{-2} \text{a}^{-1}$ . During the Sapropel S1 period, in the mid-Holocene, the Hg fluxes would have been up to  $6.4 \mu\text{g m}^{-2} \text{a}^{-1}$ , as a result of the intensity of the Hg removal by the biological pump and the organic matter preservation in low oxygen environments, along with high inputs of Hg-rich terrigenous matter and a possible restricted Hg recycling in the atmosphere because of the higher stratification of the water column which would have limited the Hg reduction and its reinjection in the atmosphere. Finally, MeHg distribution in the sediment core is a function of the abundance of inorganic Hg and its accumulation in the sediment derives in part from external inputs, but Hg methylation in the transient suboxic conditions of the sapropel is suggested.

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**Data availability**

All the data are given in Suppl. Tables SI.1 and SI.2.

**Declaration of interests**

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

The authors declare the following financial interests/personal relationships which may be considered as potential competing interests:

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### Figure captions

**Figure 1.** Locations of the sampling stations in the Ionian Sea. Sta. EA: boxcore; Sta. EB: sediment trap mooring.

**Figure 2.** (A) lithogenic fraction, (B)  $C_{org}$ , (C) carbonates, and (D) total mass flux in the three sediment traps at Sta. EB during the period May 2001 to April 2002.

**Figure 3.** (A) Hg concentration in the particulate material collected during the year 2001-2002 in the three sediment traps at Sta. EB. Low Hg concentrations are observed in spring; (B) Hg vs organic carbon ( $C_{org}$ ) in particulate material collected during the year 2001-2002 in the three sediment traps at Sta. EB. Low  $\delta^{13}C$  values, suggesting terrigenous OM, are associated with Hg-rich particles.  $C_{org} > 6\%$  values may involve a bio-dilution of particulate Hg

**Figure 4.** Vertical distributions of  $C_{org}$  (A), Ba/Al (B), Mn/Al with Mo/Al (C), Hg (D), MeHg (E), and  $Hg^{AP}$  (F) in sediment cores EA and EB. The onset, ending, and oxidized part of the sapropel S1 are indicated (see text). The oxidized front is defined by the presence of Mn/Al and Mo/Al peaks.

**Figure 5.** Methylmercury (MeHg) and manganese (Mn) relationship in the EA sediments. The oxidation front and oxidized sapropel cover the 14-20 cm sediment layer. In this layer, the MeHg vs Mn relationship exhibits an  $R^2$  of 0.96 ( $p < 0.01$ ).

**Figure 6.** Mercury (Hg) concentrations and fluxes in the Eastern Mediterranean (Ionian abyssal plain) and Western Mediterranean (Balearic abyssal plain). Data for the western Mediterranean are from Cossa et al. (2021). (mab) meter above the bottom. \* Refers to

the pre-anthropogenic period, i.e., before 3 ka cal BP for the western basin and before 6 ka cal BP for the eastern basin excluding Sapropel S1.

### **Table captions**

**Table 1.** Mean concentrations ( $\pm 1SD$ ) of main components of the particles collected in the traps between twice a month May 2001 and April 2002 at Sta. EB. \* Mean concentration significantly ( $p < 0.01$ ) different from means at 1440 and 2820 m; other differences are non-significant.