1	Enhanced silicate chemical weathering of the South African Plateau and its
2	contribution to the global climate cooling trend during the late
3	Cretaceous
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17	Abstract
18	Climate cooling during the late Cretaceous (ca. 100-66 Ma) has been linked to CO <sub>2</sub> withdrawal
19	due to enhanced denudation episodes recorded in the East-South American and West-South
20	African margins. However, the onset of these episodes is recorded after the main cooling stage
21	(i.e., Turonian-Santonian — ca. 93-83 Ma). Here, we investigate the clay fraction (< 2 $\mu$ m) of
22	sediments from the borehole O-A1 located in the continental slope of the Cape Basin, adjacent
23	to the South African Margin. We combine these results with previous records from the basin

24 (DSDP 361) to investigate the controlling factors in the denudation record along the South

25 African Plateau and their relationship with the climate cooling during the late Cretaceous. The  $\Delta \varepsilon_{\rm Hf}$  (proxy for silicate weathering intensity) record from site O-A1 shows an increase in 26 27 silicate chemical weathering during the Cenomanian-Maastrichtian (ca. 95-68 Ma) interval, 28 predating the Campanian-Danian interval shown by the DSDP 361 record, which offset to the 29 site O-A1 is primarily linked to age model uncertainties from DSDP 361. This increase in silicate chemical weathering is concomitant to an enhanced phase of physical erosion and 30 31 tectonic uplift of the South African Plateau as well as global climate cooling, suggesting this 32 enhanced silicate chemical weathering phase is tectonically driven. Our  $\Delta \varepsilon_{\text{Hf(80)}}$  data from site 33 O-A1 shows for the first time an increase in silicate chemical weathering as early as the 34 Cenomanian, suggesting not only that it could have played a role maintaining the cool global 35 climate trend during the late Cretaceous, but that this episode could have contributed to trigger 36 the  $CO_2$  withdrawn responsible for this cooling that started by the Cenomanian-Turonian. 37 Additionally, we suggest that this enhanced silicate chemical weathering episode along the 38 South African Plateau might have influenced atmospheric CO<sub>2</sub> concentrations and climate 39 through two different mechanisms, initially via silicate reactions, but also through increased 40 nutrient riverine input into the basin, which could have contributed to enhanced primary productivity and organic carbon storage as shown by the increase in Baexcess during the 41 42 Santonian-Maastrichtian (ca. 85-65 Ma).

43 Keywords: Isotope geochemistry, Tectonics-Climate, Silicate weathering, South
44 African Plateau, Cretaceous cooling, Marine productivity.

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Sedimentary basins represent the best archive to track the long-term evolution of 47 48 denudation processes in ancient environments. Recent studies conducted in basins adjacent to 49 the East-South American and West-South African margins relate tectonic activity with 50 enhanced denudation episodes, subsequent CO<sub>2</sub> withdrawal, and climate cooling (Corentin et al., 2024, 2023, 2022; Gaitan et al., 2023). These studies have shown through clay mineralogy 51 52 and isotopic records (Hf-Nd isotopes on clays) increases in silicate chemical weathering during the late Cretaceous, which is recognised as one of the most important process for CO<sub>2</sub> 53 withdrawal (Berner et al., 1983; Gaillardet et al., 1999; Hilton and West, 2020). While these 54 weathering episodes have been associated with the long-term climate cooling trend observed 55 in the late Cretaceous (Friedrich et al., 2012; O'Brien et al., 2017), Hf-Nd isotopic records 56 reveal increases between the Santonian-Campanian interval in the South American margin 57 58 (Corentin et al., 2023, 2022), and between the Santonian-Maastrichtian interval in the West-59 South African margin (Corentin et al., 2024; Gaitan et al., 2023), both later to the suggested 60 onset of the cooling trend around the Turonian (Friedrich et al., 2012; O'Brien et al., 2017). In 61 the context of the South African Margin, the proposed increase in chemical weathering relies 62 only on the geochemical record from DSDP core 361. Yet this site presents relatively large 63 uncertainties in the age model of the studied sediments (Gaitan et al., 2023), hindering the identification of a precise timing for the onset of the enhanced weathering episode. 64 Additionally, no major changes in chemical weathering along the South African margin are 65 66 detected at DSDP core 361 during the Turonian-Campanian interval, which is the main cooling 67 phase reported by global isotopic records during the late Cretaceous ( $\delta^{18}$ O benthic) (Friedrich et al., 2012; O'Brien et al., 2017). Therefore, although the recorded weathering episodes along 68 69 the East-South American and West-South African margins could have contributed to the

cooling conditions during the late Cretaceous, the mechanisms responsible for the cooling onset
remain elusive.

72 The evolution of the chemical weathering signal from DSDP core 361 relies on the analysis 73 of detrital clay minerals coming predominantly from the South African Plateau and depositing 74 in the abyssal plain of the Cape Basin (Gaitan et al., 2023). However, due to the distal location of the sedimentary record, the incorporation of allogenic sediments by deep-bottom currents is 75 76 not entirely discarded (Murphy and Thomas, 2013; Natland, 1978), and differential settling 77 processes on clay minerals might also be involved introducing some bias into the paleoclimatic 78 interpretation (Gibbs, 1977; Petschick et al., 1996). Multiple thermochronological data — Apatite-Fission Track (AFT) and Apatite (U-Th)/He (AHe) — have shown that regional 79 80 structures across the Plateau are important in the rock cooling events during the late Cretaceous 81 (Green et al., 2017; Kounov et al., 2009, 2013; Raab et al., 2002; Stanley and Flowers, 2020; 82 Tinker et al., 2008; Wildman et al., 2015, 2021). These regional structures could have modified 83 the source areas of the denudation episodes (i.e., erosion and weathering) inferred from these 84 rock cooling events. These denudation episodes have been further linked to tectonic uplift involving continental tilting (Baby et al., 2018a, 2019; Braun et al., 2014; Stanley et al., 2021), 85 responsible for drainage reorganization (De Wit et al., 2009; De Wit, 1999; Goudie, 2005; 86 Stevenson and Mcmillan, 2004), and by extension for variations in the sediments source areas. 87 88 Estimations of sediment fluxes and volumes have also shown that the tilting component of the 89 uplift appears to have affected the sediment distribution along the basins located adjacent to 90 the South African margin (Baby et al., 2019). Altogether, these processes affecting the source 91 and distribution of the sediments deposited in the Cape Basin could have impacted the 92 denudation record, particularly the chemical weathering, and hinder any interpretation of the impact of these processes in the climate during the late Cretaceous. For this reason, a process-93 94 response comparison of the denudation record along the Cape Basin appears of prime

95 importance to understand the dynamics controlling this record and the implications for climatic96 interpretations.

97 The aim of this study is to investigate the sedimentary record at the continental slope of the 98 Cape Basin (borehole O-A1— Fig. 1) and compare it with previous records from the abyssal 99 plain (DSDP core 361— Fig. 1; Gaitan et al., 2023). Through this comparison we study the 100 impacts of changes in sediment distribution and sources in the erosional and chemical weathering signal recorded along the Cape Basin, also exploring the links of this record to the 101 102 global climate cooling during the late Cretaceous. To do so, we combine X-ray diffraction 103 (XRD) analyses, major and trace elements, Transmission Electron Microscopy (TEM), and Hf 104 - Nd isotopic compositions on clay fractions (< 2  $\mu$ m - Bayon et al., 2016). Through this suite 105 of techniques, we seek to understand the dynamics affecting geochemical signals, specifically 106 on the clay fraction across the sediment routing system.

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# 108 **2. Geological Setting**

#### 109 **2.1 Evolution of the southwestern African margin**

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111 The South African margin evolution began with the rifting between the South American and African Plates and the subsequent opening of the Atlantic Ocean during the late Jurassic -112 early Cretaceous (Moulin et al., 2010; Torsvik et al., 2009). It has been suggested that this 113 114 opening started at the southern segment and moved northwards (Koopmann et al., 2016; Rabinowitz and LaBrecque, 1979), but the rifting onset in the southern segment is not yet fully 115 116 understood due to age uncertainties on the magnetic anomalies (e.g., Baby et al., 2018b; Moulin 117 et al., 2010). However, it is acknowledged that magnetic anomaly M3 at 130 Ma marks the 118 rifting-drifting transition in this segment (Collier et al., 2017). Deposition during this period is

marked by large volcanic activity linked to the Paraná-Etendeka Large Igneous Province
Province (ca. 135-133 Ma; Gomes and Vasconcelos, 2021; Renne, 2015), this is supported by
the presence of Seaward-Dipping Reflectors (SDRs) overlain graben and half graben basement
structures (Gladczenko et al., 1997). Regional Valanginian-Hauterivian fault-bounded syn-rift
sequences are present across the southwestern African basins (Séranne and Anka, 2005). These
sequences comprise lava flows with continental sediments, essentially sandstones and
claystones (Hirsch et al., 2010; Jungslager, 1999; Van Der Spuy, 2003).

126 Post-rift deposits consist of thick siliciclastic sequences of about 3 to 5 km distributed along 127 the 4 basins adjacent to the margin: Walvis, Lüderitz, Orange, and Cape (Aizawa et al., 2000). 128 (Baby et al., 2018b) defined three main first-order depositional phases for these post-rift 129 deposits. The first phase is a retrogradational stage (131-93.5 Ma) marked by an increase rate 130 of accomodation space compared to sediment supply. The creation of this accomodation space 131 is associated to the thermal subsidence of the margin. During the second stage (93.5-66 Ma), 132 deposition shifts to a predominant progradational-aggradational phase, whereas the margin is 133 tilted seawards (Aizawa et al., 2000; Baby et al., 2019). A pronounced increase in sediment supply related to the tectonic uplift of the South African Plateau is linked to this depositonal 134 135 change. The final stage (66-0 Ma) represents primarily a retrogradational phase, although there is a progradational stage within this period at ca. 27 Ma (Baby et al., 2018b). The 136 137 retrogradational stage is likely associated with the stabilization of the inland Plateau and 138 climate aridification, which led to a decrease in sediment supply.

The stratigraphy of the margin is divided in 8 units bounded by stratigraphic markers (e.g., unconformities or maximum flooding surfaces) that define relevant regional deposition cycles (Baby et al., 2019, 2018b). Lithology of this post-rifting units comprises mainly siltstones, although limestones are also an important component in the Unit 1 (130-113 Ma) from the Walvis and Lüderitz basins (Baby et al., 2019).Cape Basin 144

#### 2.2 Cape Basin

The Cape Basin is divided in 8 units and bounded by the regional stratigraphic surfaces 145 defined in the basins along the margin. The Cretaceous interval in the basin spans between Unit 146 147 1 and 5 (Baby et al., 2018b). Unit 1 is bounded by the unconformities 6At1 (130 Ma) and 13At1 148 (113 Ma), the former representing the breakup unconformity whereas the latter a depositional 149 hiatus (Paton et al., 2008, 2007). This unit comprises a sedimentary package evidencing the 150 rifting-drifting transition of the margin. At the base, aeolian sandstones with basalts are covered 151 by carbonate cemented marine sandstones. On top, shelly sandstones with some local lagoonal 152 limestones are topped by marine shales (Baby et al., 2018b). The upper boundary of Unit 2 is the seismic marker 14At1 (100 Ma). This unit is dominated by silty-claystones overlain by 153 154 shales that define marker 14At1 (Baby et al., 2018b). This unit comprises a predominant 155 retrogradational trend (Baby et al., 2019), but most importantly evidences sediment supply 156 from two disconnected deltaic systems located southern of the modern Orange river system 157 (Baby et al., 2018b). Unit 3 consists wholly of deep marine shales (Baby et al., 2018b; Paton 158 et al., 2008), with marker 15At1 (93.5 Ma) as the upper boundary. Depositional trend in this 159 unit still is mostly retrogradational, marked by wedges backstepping (Baby et al., 2019). The 160 upper boundary of Unit 4 is marked by the unconformity 17At1 (81 Ma), which has been interpreted as a major sea-level fall. Claystones dominate in this unit in a predominantly 161 162 aggradational depositional trend. The top of Unit 5 is marker 22At1 at the Maastrichtian-163 Paleogene boundary (66 Ma). Depositional trend changes from aggradational to 164 progradational, linked to variations in the ratio of subsidence and sediment supply (Baby et al., 165 2018b; de Vera et al., 2010). Lithology at this unit is dominated by clayey siltstones (Baby et 166 al., 2018b). Unit 6 lies on top of the Cretaceous sequence and it is bounded on top by an Oligocene unconformity (Baby et al., 2018b; de Vera et al., 2010). This unit displays deltaic 167 progradational units (Baby et al., 2018b; de Vera et al., 2010). 168

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## **3. Material and methods**

# 171 **3.1 Borehole description**

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173 In this study we analysed 91 cutting samples from the borehole O-A1 provided by the South 174 African Petroleum Agency. This borehole is located at the continental slope of the Cape Basin 175 (Fig. 1; 33.161539°S 16.822406°E), around 100 km off the southwestern African margin. The 176 borehole comprises a stratigraphic succession of about 3650 m, spanning from the Aptian -177 Barremian to the middle Eocene period (Valicenti et al., 1993), but the samples analysed in 178 this study span between the Albian and Ypresian periods (i.e., 115-55 Ma), representing Unit 1 to 5 from those described by Baby et al. (2019) (Fig. 2). The basement of the borehole 179 180 consists of volcanic rocks presumably Hauterivian in age (ca. 132 Ma; Valicenti et al., 1993). 181 On top, there is an Aptian-Albian sequence of argillaceous sandstones (ca. 125-100 Ma), with 182 increased clay abundance at the top and the base. At the late-early Cretaceous boundary (100 183 Ma), the predominant lithology are claystones interbedded with sandstone layers (Dudus et al., 184 1992), some minor dolomite stringers are also present in this interval. Throughout the late 185 Cretaceous (100-66 Ma) the succession consists of claystones with occasional pyrite, and 186 stringers of limestones, dolomites, and siltstones (Dudus et al., 1992). Siltstones stringers are 187 predominantly argillaceous and glauconitic. The interval between the Paleocene and the middle 188 Eocene (ca. 66-50 Ma) consists of claystones with sporadic fossiliferous limestones and siltstone stringers (Dudus et al., 1992). At the base of this interval, stringers are scarce and 189 190 claystones are more calcareous, dolomite stringers are present as well. Sporadic pyrite and 191 glauconite occur within the claystones along the whole interval (Dudus et al., 1992).

The selected samples were targeted considering even distribution within the studied time interval (i.e., 115-55 Ma) and clay content. To ensure an even distribution of the analysed samples an age model (Supplementary material FS1) was built based on the planktonic foraminifera report by Valicenti et al. (1993).

There were no elements to determine neither the Aptian-Albian boundary (113 Ma) or the 196 197 Albian stage (113-100.5 Ma), which could be missing. Only the occurrence of *Globuligerina* 198 hoterivica around 4170 mbsf indicates late Aptian ages, so the Aptian-Albian boundary was 199 tentatively placed at this depth (113 Ma). A late Albian-Mid Cenomanian (ca. 105-97 Ma) 200 interval was defined around 3624 mbfs based on the last occurrence of Rotalipora appenninica. The Cenomanian-Turonian boundary (93.9 Ma) was placed at 3570 mbsf based on the 201 202 appearance of Helvetoglobotruncana helvetica. The last occurrence of this assemblage 203 indicated a Mid-Turonian age (ca. 91 Ma) at 3408 mbsf, but there were no elements to define the Turonian-Coniancian boundary (89.8 Ma). An increase in the occurrence of 204 205 Marginotruncana coronata and the appeareance of Globotruncana species enabled to place an 206 interval Coniacian-Santonian (89.8-83.6 Ma) around 2706-2751 mbsf, still the boundary cannot be determined with certainty. Following the occurrence of Globotruncanita elevata, 207 208 distinctive of Campanian age, the Santonian-Campanian boundary (83.6 Ma) was placed at 2697 mbsf. The occurrence of *Globotruncanita aegyptiaca* indicative of a late Campanian age 209 210 (ca. 75 Ma) and the last occurrence of Globotruncana ventricosa enabled to place the 211 Campanian-Maastrichtian boundary (72.1 Ma) around 2580 mbfs. The presence of 212 Abathomphalus mayaroensis around 2553 mbsf is indicative of a late Maastrichtian age. No 213 characteristic Danian fauna was identified. The Globanomalina chapmani characteristic of a 214 Selandian-Thanetian age was observed around the interval 1550-1340 mbsf, however it was joint by species like *Globigerinatheka index*, which is characteristic of Eocene (ca. 40 Ma). 215

216	The boundary Maastrichtian-Danian (66 Ma) was then placed at 2553 mbsf, and the subsequent
217	Cenozoic upper interval is tied to an Eocene age (33.9 Ma) around 1340 mbsf.
218	Based on these planktonic foraminifera fossil assemblages, intervals and age boundaries
219	were established following the ICS chronostratigraphic chart v2023/06 (Cohen et al., 2013;
220	Updated 2023). The age model was built using linear regressions between seven tie-points: the
221	Aptian-Albian boundary (113 Ma; 4170 mbsf), the Mid-Cenomanian (97.2 Ma; 3624 mbsf),
222	the Cenomanian-Turonian boundary (93.9 Ma; 3570 mbsf), the Santonian-Campanian
223	boundary (83.6 Ma; 2697 mbsf), the Campanian-Maastrichtian boundary (72.1 Ma; 2580
224	mbsf), The Maastrichtian-Danian (66 Ma; 2553), and the Eocene stage tentatively at 1340 mbsf
225	(ca. 34 Ma).
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- 227 **3.2 Analytical techniques**
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# 3.2.1 X-ray diffraction (XRD) analyses

230 XRD analyses were performed for 77 samples for both bulk rock and clay fractions ( $< 2\mu m$ ) in the Biogéosciences Laboratory (UMR CNRS 6282) at the University of Burgundy, France 231 232 (Supplementary table S1). The bulk rock samples were grinded and set in disks adapters 233 whereas the clay fractions were prepared as oriented glass slides. Both set of measurements 234 were performed on a Bruker D8 Endeavor diffractometer with Cu-Ka radiations, a LynxEye 235 detector coupled to a Ni filter, and under 40-kV voltage with 25-mA intensity. For the clay 236 fraction (< 2µm), each oriented glass slide was measured three times following the protocol set 237 by Moore and Reynolds (1989), including both ethylene glycol and heating (490 °C) 238 preparation stages. The main intensity peaks were identified using the MacDiff software 4.2.5 239 (Petschick, 2001), and the mineral proportions, both for bulk and clay fractions, were 240 determined by calculating the area under the basal reflections peaks  $(d_{001})$ . For the clay

fractions (<  $2\mu$ m) this calculation was done using the resulting diffractograms from the ethylene glycol treatment (Moore and Reynolds, 1989).

243 3.2.2 Transmission Electron Microscopy (TEM) observations

TEM analyses were performed on clay fractions from 6 samples (Supplementary table S1) at the Agroécologie Institut from the Institut National de Recherche pour l'agriculture, l'alimentation et l'environnement (INRAE) in the University of Burgundy, France. The samples were dispersed in distilled water and 2% v/v butylamine solution and then set on a carbon formvar Cu grid. Images were taken using a MET Hitachi 4800.

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250 3.2.3 X-ray fluorescence (XRF) analyses

Forty samples were processed for major and minor oxides abundances (Supplementary table S1) using XRF at the Institut Sciences de la Terre, at the University of Lausanne, Switzerland. Fused disks were prepared from 1.2 g calcined sample powder mixed with Lithium-Tetraborate (1:5 proportions). The measurements were done using a wavelength dispersive PANalytical Axios<sup>mAX</sup> spectrometer fitted with a 4 kW Rh X-ray tube. Rock silicate reference materials were used for calibration.

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# 3.2.4 Chemical processing for trace element and isotopic analyses

Twenty four samples were grinded and processed (1.3 to 1.5 g of powder) through sequential leaching following a procedure based on the protocols described in Bayon et al., 2002 and Gutjahr et al., 2007 (Supplementary table S1). Carbonates were first removed using 10% v/v acetic acid reacting overnight with the samples. After a rinsing stage with ultrapure 18.2 M $\Omega$  water, the Fe-Mn oxyhydroxide phases were removed using a 0.5 M hydroxylamine hydrochloride solution in 20% v/v acetic acid reacting with the samples for 48 hours. After another rinsing stage with ultrapure 18.2 M $\Omega$  water, organic matter was removed using a 5%  $H_2O_2$  solution reacting with the samples for 48 hours, to finally be rinsed again with ultrapure 18.2 M $\Omega$  water. Clay fractions (< 2 $\mu$ m) were separated after this leaching sequence using particle decantation according to the Stokes law.

269 After leaching and sampling of the clay fraction ( $< 2\mu m$ ) the samples were dried and processed by alkaline fusion following the protocol described in Bayon et al., 2009. Three 3 270 271 certified reference standards (BHVO-2, MAG-1, and BCR-2) from the United States 272 Geological Survey (USGS) were also included during the alkaline fusion. Around 50 mg of 273 each sample were placed into carbon crucibles together with 0.6 g of NaOH, 1.2 g of Na<sub>2</sub>O<sub>2</sub>, 274 and 0.5 g of Tm spike. The crucibles underwent a heating stage at 650 °C during 12 minutes in 275 a furnace, to then be cooled down with ultrapure 18.2 M $\Omega$  water, resulting in the precipitation of Fe-hydroxides. The samples were then centrifuged, and finally dissolved in 3,3 ml of 6 M 276 277 HCl. From this solution, 0.3 ml were extracted for trace elements analyses. The Tm spike was added as an internal standard in the quantification of trace elements following the method 278 279 described in Barrat et al., 1996.

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281 3.2.4.1 Trace elements

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Trace elements concentrations were measured by inductively coupled plasma mass spectrometry (ICP-MS) on a Thermo Scientific X-Series II<sup>®</sup>, at the Pole Spectrométrie Océan in Brest, France (Supplementary table S2). We used standard bracketing with the Tm spike to quantify trace elements as described in Barrat et al., 1996 and Freslon et al., 2011. Accuracy control and reproducibility were assessed analysing the 3 USGS reference standards: BHVO-2, MAG-1, and BCR-2 (Supplementary table S3). Deviations of trace elements from the reference BHVO-2 (Jochum et al., 2016) were below 14%. For MAG-1 and BCR-2 standards, some elements presented higher deviations, in particular the light rare earth elements (LREE)
at around 20% (Supplementary table S3).

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293 3.2.4.2 Hf-Nd Isotopic analyses

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Extraction of Hf and Nd for 24 samples was performed by ion chromatography using 295 296 singled customed columns with AG50W-X8 (200-400 mesh) resin for REE + Hf, Ln Spec resin 297 (50-100 µm) for the Nd fraction, and Ln Spec resin (100-150 µm) for the Hf fraction. Column 298 chromatography was performed applying the adapted separation protocol for Hf and Nd 299 described in Gaitan et al. (2023) for the low pressure, automated column chromatography, 300 PrepFAST-MC<sup>®</sup> system device. Isotopic measurements were performed using a MC-ICP-MS 301 Neptune Plus<sup>®</sup> (ThermoFisher-Scientific) at the École Normale Supérieure (ENS) of Lyon, 302 France. Errors are reported as 2 standard deviation (2 sd).

303 Hafnium isotope measurements performed at the ENS consisted of a static acquisition 304 method of 40 cycles, with an average intensity of 4.3 V throughout the session. Ratios were 305 corrected for mass bias through exponential law and using the stable isotope ratio <sup>176</sup>Hf/<sup>177</sup>Hf = 0.7235. Mass bias corrected  ${}^{176}$ Hf/ ${}^{177}$ Hf ratios were normalized to a JMC 475 value of 306 0.282163 (Blichert-Toft et al., 1997). Repeated measurements of bracketed JMC 475 standards 307 308 during the analytical session gave a ratio of  ${}^{176}\text{Hf}/{}^{177}\text{Hf} = 0.282270 \pm 0.000006$  (21 ppm, n = 17). Reference standards BHVO-2 and BCR-2 gave <sup>176</sup>Hf/<sup>177</sup>Hf normalised isotopic values of 309 310  $0.283103 \pm 0.000020$  (70 ppm, n = 3) and  $0.282869 \pm 0.000004$  (14 ppm, n = 1) respectively. Both standards within the range of reference values determined by Weis et al. (2005). Four 311 procedural blanks for Hf were also measured, obtaining values between 27 and 44 pg 312 313 (Supplementary table S3).

314 Neodymium isotope measurements performed at the ENS consisted of a static acquisition method of 40 cycles, with an average intensity of 5 V. Ratios were corrected for mass bias 315 through exponential law and using the stable isotope ratio  ${}^{143}Nd/{}^{144}Nd = 0.7219$ . Mass bias 316 corrected <sup>143</sup>Nd/<sup>144</sup>Nd ratios were normalized to a JNdi-1 value of 0.512115 (Tanaka et al., 317 2000). Repeated measurements of bracketed JNdi-1 standards during the analytical session 318 gave a ratio  ${}^{143}Nd/{}^{144}Nd = 0.511980 \pm 0.000009$  (18 ppm, n = 16). Reference standards BHVO-2 319 and BCR-2 gave  $^{143}Nd/^{144}Nd$  normalised isotopic values of 0.512984 ±0.000013 (25 ppm, n = 320 321 3) and 0.512644  $\pm$ 0.000009 (17 ppm, n = 2) respectively. Both standards are consistent with 322 the reference values determined by Weis et al., (2005). Four procedural blanks for Nd were measured as well, obtaining values between 18 and 313 pg, which are approximately  $1 \times 10^3$ 323 324 times lower than the average concentrations obtained for the samples (Supplementary table 325 S3).

326 Results for hafnium and neodymium isotopic compositions are reported in epsilon notation 327 relative to the deviation from the chondritic reference material (CHUR) reported in Bouvier et 328 al. 2008,  ${}^{176}$ Hf/ ${}^{177}$ Hf<sub>CHUR</sub> = 0.282785 and  ${}^{143}$ Nd/ ${}^{144}$ Nd<sub>CHUR</sub> = 0.512630:

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$$\varepsilon_{Hf} = \left( \left[ \frac{\left( \frac{176Hf}{177Hf} \right)_{Sample}}{\left( \frac{176Hf}{177Hf} \right)_{CHUR}} - 1 \right] * 10^4 \right) \text{ and } \varepsilon_{Nd} = \left( \left[ \frac{\left( \frac{143Nd}{144Nd} \right)_{Sample}}{\left( \frac{143Nd}{144Nd} \right)_{CHUR}} - 1 \right] * 10^4 \right)$$

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331  $ε_{Nd}$  and  $ε_{Hf}$  values are also reported age corrected (t=80 Ma; Table 1). The  $ε_{Hf}$  values were 332 corrected for radioactive decay of <sup>176</sup>Lu to <sup>177</sup>Hf using the measured Lu and Hf concentrations 333 (<sup>176</sup>Lu/<sup>177</sup>Hf = Lu/Hf \* 0.1419), the <sup>176</sup>Lu radioactive decay constant λ (1.867 ± 0.008 x 10<sup>-11</sup> 334 year<sup>-1</sup>; Söderlund et al., 2004), and the present-day <sup>176</sup>Lu/<sup>177</sup>Hf<sub>(CHUR)</sub> value of 0.336 (Bouvier 335 et al., 2008). The  $ε_{Nd}$  values were corrected for radioactive decay of <sup>147</sup>Sm to <sup>144</sup>Nd using the 336 measured Sm and Nd concentrations (<sup>147</sup>Sm/<sup>144</sup>Nd = Sm/Nd \* 0.6049), the <sup>147</sup>Sm radioactive 337 decay constant λ (6.54 x 10<sup>-12</sup> year<sup>-1</sup>; Lugmair and Marti, 1977), and the present-day

338	<sup>147</sup> Sm/ <sup>146</sup> Nd <sub>(CHUR)</sub> value of 0.1960 (Bouvier et al., 2008). Age corrections were done to 80 Ma
339	to interpret the data. This age represents the average age of the analysed time interval.
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341  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$  values from clay fractions display a correlation that is denominated as the "Clay 342 Array" (Bayon et al., 2016). Offsets in  $\varepsilon_{Hf}$  values from this array are expressed as  $\Delta \varepsilon_{Hf}$ 343 (Equation 1) and account for the intensity in chemical weathering (Bayon et al., 2016).  $\Delta \varepsilon_{Hf}$ 344 values were calculated using  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$  values at present time (t = 0) and age-corrected to 80 345 Ma, these values are reported in Table 1.

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$$\Delta \varepsilon_{\text{Hf(t)}} = \varepsilon_{\text{Hf(t)}} - (0.78 * \varepsilon_{\text{Nd(t)}} + 5.23)$$
(Equation 1)

### **4. Results**

# 349 4.1 Bulk rock composition and clay mineral assemblages – XRD analyses 350

351 Bulk rock composition measured shows a predominance of quartz and phyllosilicates (i.e., 352 clays and micas) along the section, with calcite only becoming predominant after the Santonian 353 (~2800 mbsf; ca. 84 Ma) (Supplementary table S4; Fig. 2). At the base, during the Aptian-354 Albian interval (ca. 114-110 Ma), quartz content increases reaching values around 88%. 355 Above, quartz content decreases while phyllosilicates content increases, reaching values up to 356 35% in the Turonian (ca. 90 Ma). An increase in calcite concentrations is observed at the Cenomanian (ca. 96 Ma). Since the Coniacian, calcite content increases in parallel with to a 357 358 decrease in quartz content, while phyllosilicates proportions remain stable at values around 359 35%. The increase in calcite content continues up to the Maastrichtian-Paleogene transition, where it reaches values around 50%. Between the Danian and the Ypresian, quartz and 360 361 phyllosilicates content increase up to values around 35% and 40%, respectively. Calcite content decreases from the Danian onwards reaching to values close to 20%, only showing a positive excursion around the Danian-Selandian boundary with values close to 55%. Plagioclase and K-Feldspar are present throughout the section but in minor proportions. The percentages along the section are around 8%, except for sample O-A1/-91 at the base of the section with 16% and 19% for K-Feldspar and plagioclase, respectively. Other minerals like pyrite, halite, and gypsum, are also present along the section but in proportions lower than 10% (Supplementary table S4).

369

370 The clay assemblages at the base of the section are dominated by smectite with values close 371 to 75% (Supplementary table S4; Fig. 2). However, in the late Aptian (ca. 113 Ma), smectite 372 disappears simultaneously to the appearance of illite-smectite mixed layer, which reaches 373 values around 10%. During the Aptian-Albian interval, illite and chlorite increase up to values 374 around 50% and 30%, respectively. By the early Cenomanian (ca. 100 Ma), illite values decrease to around 15%, following an increase in illite-smectite and kaolinite. From the 375 376 Cenomanian to the Turonian, most of the clay assemblages remain stable, with values around 35% for both chlorite and illite, 15% for kaolinite, and around 10% for illite-smectite. Chlorite 377 378 displays an increase around the Turonian-Coniacian transition and chlorite-smectite mixed layers start to appear in the late Coniacian (ca. 86 Ma), whereas illite decreases to values around 379 380 20%. By late Coniacian (ca. 87 Ma), illite-smectite and kaolinite disappear concomitantly to 381 the appearance of smectite, with illite also showing a decrease down to 10%. Since, smectite 382 increases steadily up to the Danian (ca. 65 Ma) where it peaks at values around 75%. 383 Concomitant to such increase, chlorite-smectite abundance decreases and reaches values 384 around 25% in the Maastrichtian-Danian transition. Illite oscillates during this period but displays a subtle increase reaching values close to 30% in the Maastrichtian-Danian transition. 385 386 During the Danian-Ypresian interval, smectite remains the most abundant assemblage with

387	values close to 70%, while illite and chlorite-smectite decrease significantly to v	alues around
388	8%, however during the Thanetian, illite showed a continuous increase up to v	alues around
389	20%.	

390

- 391 4.2 Hf Nd isotopic compositions
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393 The measured  $\Box_{\rm Hf}$  values range between -4.8 and -2.1  $\Box$ -units throughout the section (Fig.

394 2; Table 1). During the Albian,  $\Box_{Hf}$  values remain stable around -3  $\Box$ -units, only to decrease

395 down to -4.3 🛛 -units in the Cenomanian (ca. 95 Ma). Above, values show a slight increasing

396 trend up to the late Santonian (ca. 85 Ma), where they reach values around -2 - units. During

397 the early Campanian there is a drop in the values down to -3.7 -units, follow by a very minor

increasing trend that continues up to the early Danian (ca. 65 Ma), where they reach back to

399 values around -2.5 -units. Lastly, the values decrease during the Danian-Ypresian to reach a

400 minimum of -4.9 []-units. Measured []<sub>Nd</sub> values range between -9.7 and -8.1 []-units

401 throughout the section (Fig. 2; Table 1). From the Albian up to the early-Coniacian (ca. 89

402 Ma),  $\square_{\rm Nd}$  values display a slight increase from -8.8 to -8.4  $\square$ -units, apart from a negative

- 403 excursion (-9.7  $\Box$ -units) in the late Albian (ca. 102 Ma). Above the Coniacian, the  $\Box_{Nd}$  values
- 404 display a long subtle negative trend up to the Thanetian (ca. 58 Ma), where they reach values
- 405 around -9.7 []-units. However, in this interval there is a minor increasing period during the
- 406 Danian, where values show a show a very minor increase from -9.3 to -9.1 (Fig. 2)
- 407

- Table 1. Hf-Nd isotopic compositions. Depth interval represents the range from where the sample was taken. Reported depth is the depth used to calculate the sample age and plot the data in Fig. 2. Depth and depth interval are reported in meters below the sea floor (mbsf). \*Ages were calculated using an age model built from the biostratigraphy report by Valicenti et al. (1993). Error accounts for the standard error.  $\Delta \varepsilon_{Hf}$  was calculated using the clay array correlation:  $\Delta \varepsilon_{Hf} = \varepsilon_{Hf} - (0.78 * \varepsilon_{Nd} + 5.23)$  published in Bayon et al. (2016).
- 412  $~~2\square$  is the standard deviation calculated for both  $\epsilon_{\rm Nd}$  and  $\epsilon_{\rm Hf}.$

Sample	Depth interval	Depth (mbsf)	Age (Ma) <sup>*</sup>	<sup>143</sup> Nd/ <sup>144</sup> Nd	ŝ <sub>Ni</sub>	20	<sup>176</sup> Hf/ <sup>177</sup> Hf	\$ <sub>Hf</sub>	20	\$ <sub>P4(80)</sub>	\$Hf(80)	∆چ <sub>if(80)</sub>
_	(mbst)											
O-A1/1	2145 - 50	2147	55	0.512149	-9.39	0.12	0.282647	-4.87	0.16	-8.45	-3.60	-2.24
O-A1/5	2235 - 45	2240	57	0.512132	-9.72	0.12	0.28267	-4.08	0.21	-8.76	-2.87	-1.27
O-A1/10	2346 - 55	2349	60	0.512138	-9.59	0.13	0.282667	-4.16	0.18	-8.66	-2.97	-1.45
O-A1/14	2439 - 45	2442	63	0.512158	-9.21	0.14	0.282707	-2.77	0.14	-8.36	-1.59	-0.30
O-A1/16	2481 - 90	2485	64	0.512164	-9.1	0.13	0.282716	-2.44	0.18	-8.17	-1.29	-0.15
O-A1/18	2529 - 35	2532	65	0.512159	-9.19	0.14	0.282707	-2.75	0.15	-8.30	-1.63	-0.39
O-A1/21	2562 - 71	2560	67	0.51215	-9.37	0.06	0.282715	-2.48	0.2	-8.47	-1.36	0.01
O-A1/23	2577 - 83	2580	72	0.512152	-9.33	0.16	0.282711	-2.63	0.18	-8.44	-1.49	-0.13
<b>O-A1/24</b>	2583 - 92	2588	73	0.512142	-9.51	0.14	0.282679	-3.74	1.8	-8.64	-2.62	-1.11
O-A1/27	2604 - 07	2605	74	0.512147	-9.42	0.17	0.282701	-2.96	0.2	-8.52	-1.86	-0.44
O-A1/30	2619 - 25	2622	76	0.51215	-9.37	0.15	0.282704	-2.88	0.24	-8.46	-1.73	-0.36
O-A1/33	2637 - 43	2640	78	0.512145	-9.47	0.11	0.282693	-3.26	0.26	-8.55	-2.16	-0.72
O-A1/37	2661 - 67	2665	80	0.512157	-9.22	0.17	0.282692	-3.3	0.25	-8.23	-2.09	-0.90
O-A1/37 (2)	2661 - 67	2665	80	0.512167	-9.04	0.12	0.282695	-3.18	0.17	-8.23	-2.09	-0.90
O-A1/40	2685 - 94	2690	83	0.512167	-9.04	0.13	0.282725	-2.14	0.29	-8.13	-1.01	0.11
O-A1/43	2790 - 96	2792	85	0.512172	-8.94	0.17	0.282695	-3.17	0.32	-7.98	-2.12	-1.12
O-A1/46	2943 - 46	2944	86	0.5122	-8.4	0.14	0.282696	-3.15	0.28	-7.42	-2.13	-1.57
O-A1/49	3102 - 08	3105	88	0.512214	-8.11	0.17	0.282704	-2.87	0.4	-7.14	-1.77	-1.44
O-A1/51	3201 - 07	3204	89	0.512201	-8.37	0.14	0.282667	-4.17	0.36	-7.45	-3.18	-2.59
O-A1/52	3255 - 61	3258	90	0.5122.08	-8.22	0.13	0.282668	-4.12	0.35	-7.32	-3.20	-2.73
O-A1/57	3492 - 98	3494	93	0.512214	-8.12	0.13	0.282685	-3.53	0.31	-7.14	-2.53	-2.19
O-A1/60	3582 - 91	3587	95	0.512206	-8.27	0.1	0.282661	-4.38	0.36	-7.31	-3.41	-2.94
O-A1/63	3624 - 27	3624	97	0.512183	-8.72	0.14	0.282693	-3.24	0.3	-7.73	-2.32	-1.52
O-A1/66	3693 - 96	3694	99	0.512132	-9.71	0.14	0.282698	-3.09	0.29	-8.63	-2.06	-0.56
O-A1/70	3792 - 98	3795	102	0.512175	-8.87	0.15	0.282694	-3.21	0.26	-7.84	-2.41	-1.53

414

#### 4.3 Major and REE elements

415

416 Major oxides concentrations are reported in the supplementary table S5. The SiO<sub>2</sub> wt% is 417 predominant among the samples with percentages between 24% and 82%. Second in 418 abundance are CaO wt%, Fe<sub>2</sub>O<sub>3</sub> wt%, and Al<sub>2</sub>O<sub>3</sub> wt%, with concentrations reaching up to 28%, 419 19%, and 15%, respectively. Other major oxides like MgO wt%, K<sub>2</sub>O wt%, and Na<sub>2</sub>O wt%, 420 are present in percentages lower than 4 wt%, whereas major oxides like TiO<sub>2</sub> wt%, MnO wt%, 421 P<sub>2</sub>O<sub>5</sub> wt%, Cr<sub>2</sub>O<sub>3</sub> wt%, and NiO wt%, are percentages lower than 1 wt%.

422

423 REE elements for the clay fraction ( $< 2\mu m$ ) are reported together with other trace elements 424 in supplementary table S2. The REE concentrations for samples below sample O-A1/43 (below 425 2793 mbsf – 85 Ma) display a flat pattern with a very subtle depletion in HREE over LREE 426 (Fig. 3). In contrast, samples in the upper part of the section (above 2793 mbsf - 85 Ma) show 427 REE patterns with a slight depletion in LREE and MREE. Additionally, these samples show positive Eu anomalies (Eu/Eu\* > 1; Supplementary table S2). Such anomalies are likely 428 429 introduced due to undercorrected isobaric interferences associated to high Ba contents (>2000 430 ppm; Supplementary table S2)(e.g., Lidman et al., 2019; Zhu and Itoh, 2021), therefore Eu 431 values were removed from the REE plot.

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- 435 **4**

# 4.4 Transmission electron microscopy

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437 Six samples evenly distributed along the section were analysed through TEM imagery (Fig.
438 4; Supplementary table S1). At the base of the section (sample O-A1/88; 4151mbsf – 112 Ma)
439 clay particles show rather regular textures with defined edges, specially illite particles that are

440 present as elongated euhedral prisms (Fig. 4A). By the low-middle part of the section (sample 441 O-A1/61; 3607 mbsf – 96 Ma) particles texture show a rather altered structure, and illite-442 smectite mixed-layers (I-S R<sub>1</sub>) start to appear. The texture shows smaller, thinner, elongated 443 and more fibrous particles (Fig. 4B; 'hairy-like' type). Above (sample O-A1/50; 3155 mbsf -88 Ma), illite-smectite mixed-layers (I-S R<sub>1</sub>) seem less altered, but still rather long and fibrous 444 445 (Fig. 4C). Kaolinite also starts to become more frequent, with a characteristic hexagonal regular shaped texture (Fig. 4C). By the upper-middle part of the section (O-A1/32; above 2635 mbsf 446 447 - 77 Ma), clay particles show predominantly flaky-shape texture, characteristic of smectite 448 particles (Fig. 4D). Thin-elongated particles interpreted as illite also become more frequent 449 (Fig. 4D). 450

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- 456 5. Discussion
- 457 5.1 Clay minerals origin
- 458

459 Detrital clay minerals deposited in sedimentary basins are useful tracers to understand the evolution of surface continental processes (e.g., Bougeault et al., 2017; Chamley, 1989; 460 461 Corentin et al., 2023, 2022; Dera et al., 2009; Gaitan et al., 2023). However, clay minerals can 462 also form within sedimentary basins through processes such as hydrothermal activity, reverse weathering, alteration of volcanic ash falls, or in deep-sea conditions by the interaction between 463

464 sediments constituents and seawater (Chamley, 1989; Velde, 1995; Warr, 2022). Clay mineral 465 assemblages can also be modified after formation or deposition during burial diagenesis in the 466 basin (Chamley, 1989; Ketzer et al., 2003; Lanson et al., 2002). Therefore, to interpret 467 variations in clay mineral assemblages in terms of surface continental processes, clay minerals 468 at sedimentary basins should be identified first as detrital particles inherited from the 469 landmasses where they were formed and without significant alteration.

The clay record of borehole O-A1 shows from top to bottom a clear and progressive 470 471 influence of burial diagenesis, which is common for the depths range observed in borehole O-A1 (i.e., > 2500 mbsf; Fig. 2). At the top of the section, smectite predominates with minor 472 presence of illite and chlorite-smectite mixed layers. Smectite disappears progressively with 473 474 depth, illite-smectite (R<sub>1</sub>) layers start to appear, and chlorite becomes more abundant in 475 chlorite-smectite mixed layers, until only illite and chlorite are remaining at the base of the 476 section (Fig. 2). This transition in clay assemblages is diagnostic of burial diagenesis. Smectites 477 are the most sensitive clay minerals and start to transform into illite-smectite mixed layers 478 around 70°C through illitization process, until a full transformation into illite is completed 479 (Chamley, 1989; Lanson et al., 2009; Nadeau et al., 1985; Surdam and Crossey, 1987; Warr, 480 2022). Illitization and chloritization processes lead to the complete transformation into illite 481 and/or chlorite depending on the chemistry of fluids and enclosed rocks, in particular the 482 availability of K and Mg (Chamley, 1989; Lanson et al., 2009; Warr, 2022). The presence of 483 kaolinite from the middle part of the section (ca. 3000-3600 mbsf) to the top can still represent 484 a remnant of a detrital component, as kaolinite is often more resistant and preserved in diagenetically affected sediment deposits under 120-140 °C (Chamley, 1989; Lanson et al., 485 486 2002; Warr, 2022). The smectite peak at the very bottom of the section (ca. 4180 mbsf) is highly unusual at such depths. In this case, this peak is rather related to interlayered chlorite-487 488 smectite (i.e., corrensite). The presence of corrensite has been reported in metabasalts and related to hydrothermal environments (Bettison-Varga, 1997; Inoue and Utada, 1991), which would be consistent in this case as the sediments at the bottom of the section are in contact with the basaltic basement that due to the depth and temperature could have undergone low-degree metamorphism.

The TEM images across the section support this interpretation (Fig. 4). Clay mineral 493 494 textures in the upper part of the section (Campanian-Ypresian; 2650-2100 mbsf) indicate a rather detrital origin (Fig. 4D). The flake-shaped textures on smectite particles, and the 495 496 polygonal shaped texture of illite and chlorite are indicative of detrital origin (Chamley, 1989; 497 Clauer, 1990; Fagel et al., 2001). Below this upper part (2650 mbsf; Campanian), the progressive appearance of Illite-Smectite  $(R_1)$  particles with fibrous textures is already 498 499 indicative of illitization, consistent with burial diagenetic processes (Bauluz et al., 2012; 500 Lanson et al., 2009; Pearson and Small, 1988). Such textures become more prominent with 501 depth, as interlayered particles become smaller, thinner, and occur as elongated fibres (i.e., 502 'hairy-like' textures — Fig. 4B-C), which is commonly observed in diagenetic affected 503 sequences (Lanson et al., 2009; Nadeau et al., 1985). A semi-quantitative approach to interpret 504 these clay mineral textures is the calculation of the Kübler or 'crystallinity' index on illite-505 smectite interstratified clay particles at the 10-Å X-ray diffraction peak. High crystallinity (i.e., low FWHM — Full Width at Half-Maximum height) is associated to clays recrystallised during 506 507 diagenesis, whereas lower crystallinity (i.e., high FWHM) is associated to detrital non-altered 508 clay particles (Kübler and Jaboyedoff, 2000; Jaboyedoff et al., 2001). This index shows the 509 progressive effect of diagenesis between 2500 and 3800 mbsf, with a progressive increase in 510 crystallinity along this interval of the section (FS2). The section above 2500 mbsf shows a shift 511 to lower FWHM values, these values are associated to the absence of illite-smectite interstratified clay particles, which is also observed in the clay mineralogy with a significant 512 513 increase in the smectite content above 2500 mbsf (Fig. 2). This level appears to represent the

514 threshold between the pristine detrital clay particles and the clay minerals altered by burial 515 diagenesis processes. Similarly, the values around 0.4 observed below 3800 msbf (FS2) seem 516 to be linked to a change in lithology in the borehole, shifting from claystones to more silty-517 sandy sediments. Therefore, illite-smectite interstratified clay particles are rather absent in this 518 interval and their crystallinity cannot be assessed. REE concentrations can also offer insights 519 into the origin of clay minerals (Abbott et al., 2019; Bayon et al., 2023; Cullers et al., 1975). 520 Excluding the Eu/Eu\* anomalies as they are likely related to undercorrected high Ba contents 521 in the samples (Fig. 3; see Major and REE elements 4.3) (e.g., Lidman et al., 2019; Zhu and 522 Itoh, 2021), the REE exhibit predominantly flat patterns, which are commonly observed in detrital clays (Fagel et al., 1992; Zhao et al., 2021). Although there is some variability within 523 524 the REE patterns, with part of the samples showing a slight depletion in LREE, these patterns 525 resemble those observed in clay particles derived from large river basins (Bayon et al., 2015). 526 More importantly, the REE patterns do not display any Ce anomalies accompanied by marked 527 enrichments in HREE characteristic of seawater patterns, thus suggesting that authigenic 528 imprints during the early diagenesis stages are mostly absent (Nozaki, 2001; Zhao et al., 2021). 529 The burial diagenetic imprint observed on clay minerals does not appear either to have 530 impacted the REE concentrations. The clay mineral assemblages affected by burial diagenesis from below the middle part of the section (2700 mbsf; Santonian) display in fact the most 531 532 typical shale-like flat patterns (Fig. 3).

The clay mineralogy, the clay particle textures, and the Kübler index indicate clearly that the clay minerals assemblages, at least between 2500 and 3800 mbsf, have been modified by burial diagenesis, although REE patterns seem to be unaffected. This interval within the borehole corresponds to most of the late Cretaceous section, and therefore the changes in clay mineralogy from borehole O-A1 cannot be interpreted in terms of changes in continental denudation processes along the South African Plateau.

#### 539 **5.2 Sediment provenance**

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Sediment provenance in ancient environments has been largely assessed through neodymium isotopes (e.g., Goldstein et al., 1984; Hindshaw et al., 2018; Peucker-Ehrenbrink et al., 2010). This isotopic system is particularly suited for provenance studies because it is relatively unaffected by weathering processes (Goldstein and Jacobsen, 1988; Hindshaw et al., 2018; McCulloch and Wasserburg, 1978) and as Nd is a quite immobile element, it is resistant to diagenetic effects (e.g., Chakrabarti et al., 2007; Clauer and Chaudhuri, 1995). This is supported by the absence of correlation between  $\Box_{Nd(t)}$  and the Kübler Index (FS3), suggesting

548 that burial diagenesis has not affected the evolution of  $\square_{Nd}$  values at this site.

549 The  $\square_{Nd(80)}$  values from site O-A1 are relatively stable around -8.5  $\square$ -units throughout

550	the studied interval, except for a radiogenic excursion encompassing the Cenomanian to
551	Coniacian interval where values reach about -7. The recorded clay $\epsilon_{Nd(80)}$ values are compatible
552	with a mixed contribution of different sediment sources located on the western-central areas of
553	the South African Plateau (Fig. 5; Gaitan et al., 2023). The increase from -8.6 to -7.1 observed
554	in clay $\Box_{Nd(80)}$ values between the Cenomanian to Coniacian interval (ca. 97-87 Ma) points to
555	an increase in sediment input from rather radiogenic sources (Fig. 6). Such contribution could
556	have come from the Saldania Belt located in the southwestern African margin (Chemale et al.,
557	2011; mean $\Box_{Nd(80)} = -6$ ; Supplementary table S6) or the Karoo basalts (Hawkesworth et al.,

558 1984; mean  $\Box_{Nd(80)} = -4.8$ ; Supplementary table S6). A concomitant increase in P<sub>2</sub>O<sub>5</sub> wt% and

559 clay  $\square_{Nd(80)}$  values during the Cenomanian to Coniacian interval tends to support contributions

560 from the Karoo basalts rather than the Saldania Belt (Fig. 6). P<sub>2</sub>O<sub>5</sub> wt% content is typically

561 high (> 0.4 wt%) in volcanic rocks, including Continental Flood Basalts (CFB) like the Karoo

basalts (e.g., Carlson, 1991; Jourdan et al., 2007). From the Coniacian onwards, the clay  $\Box_{Nd(80)}$ 

values decrease, reaching values around -8.5 in the Campanian (ca. 80 Ma), and remaining stable up to the Ypresian (Fig. 6). These more negative values indicate increased contributions from more unradiogenic units (i.e., lower  $\Box_{Nd}$  values), such as the Namaqua Belt (Reid, 1997;

566 mean  $\Box_{Nd(80)} = -15.7$ ; Supplementary table S6), the Cape Belt (Dia et al., 1990; mean  $\Box_{Nd(80)} =$ 

567 -10; Supplementary table S6) or the Karoo sediments (Dia et al., 1990; mean  $\Box_{Nd(80)} = -8.9$ ;

Supplementary table S6), all located at southwestern – central areas of the South African
Plateau (Fig. 5). Given the nature of these units which include very heterogeneous lithologies,
further analyses including trace elements and other isotopic systems will be required to better
constrain their relative contribution to the sediment sources at site O-A1.

572 The changes in contribution from the sediment sources presented here can be partially 573 associated to the tectonic uplift experienced by the South African Plateau throughout the late 574 Cretaceous. The tectonic uplift of the Plateau involved a margin tilting, with an early uplift 575 phase (ca. 93-80 Ma) affecting predominantly the eastern margin, followed by a second phase 576 (ca. 80-70 Ma) affecting predominantly the western margin (Baby et al., 2019, 2018a; Braun et al., 2014; Stanley et al., 2021). This tilting has been inferred based on changes in 577 578 sedimentation rates from the basins adjacent to the south African margin (Baby et al., 2019, 2018a; Braun et al., 2014; Stanley et al., 2021). Although the onset of the early uplift stage is 579 580 set around 93 Ma (Baby et al., 2019, 2018a; Braun et al., 2014; Stanley et al., 2021), 581 thermochronological data reports erosional cooling as early as 100 Ma (e.g., Flowers and 582 Schoene, 2010; Stanley et al., 2013; Wildman et al., 2015). This eastern margin uplift could 583 have brought sediments from central areas of the Plateau to the Cape Basin, which could be 584 consistent with the increased input of radiogenic sediments derived from the Karoo basalts.

585 Similarly, the switch to more unradiogenic clay  $\Box_{Nd(80)}$  values after the Coniacian (ca. 87 Ma)

could be related to the input of sediments from unradiogenic units (e.g., Namaqua, Karoo sediments or Cape Belt), which could have been progressively incorporated during the later phase of this tectonic uplift stage (ca. 93-80 Ma; Fig. 6). The  $\Box_{Nd(80)}$  values stabilise and remain

quite stable around -8.5 after the early Campanian (ca. 80 Ma), suggesting that the sediment
sources contributions did not change during the later uplift stage of the South African Plateau
(ca. 80-70 Ma).

Along the South African Plateau a main drainage reconfiguration has been identified around the Maastrichtian (ca. 70 Ma), suggesting that the Kalahari and Karoo paleorivers merged into the modern Orange river system due to tectonic activity (De Wit et al., 2009; De Wit, 1999; Goudie, 2005). Our data suggests that previous reconfigurations might have occurred and could explain the shift between the contribution of the Karoo basalts — recorded 597 by the radiogenic  $\Box_{Nd(80)}$  excursion and the associated higher P<sub>2</sub>O<sub>5</sub> wt% content — to a more

598	relevant contribution of sediments derived from the Namaqua belt, Karoo sediments, and Cape
599	Belt, recorded by the decrease in $\Box_{Nd(80)}$ during the Coniacian-Campanian interval. Unlike the
600	early uplift stage, the late uplift stage (ca. 80-70 Ma) of the western margin does not seem to
601	have had a major impact in the sediment sources contribution, as by the Campanian (ca. 80
602	Ma), the $\square_{Nd(80)}$ values have already stabilised and remained quite stable up to the Ypresian
603	(Fig. 5 and 6). This stability in the values also indicate that the suggested drainage
604	reconfiguration around the Maastrichtian (ca. 70 Ma) (De Wit et al., 2009; De Wit, 1999;
605	Goudie, 2005) did not have an impact in sediment sources, only in sediments distribution in
606	the Cape Basin. This is observed in the contrasting changes in sediment fluxes observed
607	between the Orange and Cape Basins (Fig. 6), both located adjacent to the southwestern
608	African margin.
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610	5.3 Evolution of silicate chemical weathering in the southwestern African
611	margin
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613	The evolution of silicate chemical weathering is often investigated using major element
614	ratios and indices (e.g., Baiyegunhi et al., 2017; Dinis et al., 2020). Among these indices the
615	most commonly used is the CIA index (Chemical Index of Alteration; Nesbitt and Young,
616	1982). However, at site O-A, the impact of burial diagenesis across the section could have
617	perturbed the index, especially as Na and K are very mobile elements and could have been
618	impacted by illitization processes. In addition, this index is sensitive to sediment sorting and

619 changes in sediment sources, which vary at this site as indicated by changes in lithology and in

620 clay  $\square_{Nd(80)}$  values. Hence, the use of the CIA is not adequate for site O-A1. To investigate the

621 evolution of silicate chemical weathering, we employed the combined Sm-Nd and Lu-Hf 622 isotopic systems from clay fractions, which have already been used in both modern and ancient

623 environments and have been shown to be independent to changes in sediment sources (e.g.,

Bayon et al., 2012, 2016, Corentin et al., 2023). The clay  $\Box_{Nd}$  and  $\Box_{Hf}$  values are not correlated

to the Kübler index (FS3), suggesting that the burial diagenetic processes identified in this site did not affect the clay isotopic compositions. During magmatic processes, the Sm-Nd and Lu-Hf isotope systems behave alike, with mantellic material enriched in Sm and Lu, leading with time to more radiogenic  $\Box_{Nd}$  and  $\Box_{Hf}$  values in mantellic material when compared to crustal

629 material. This is highlighted by the correlation of crustal and mantellic  $\Box_{Nd}$  and  $\Box_{Hf}$  material,

630 called the "Terrestrial Array" (Vervoort et al., 2011). Conversely, during surface continental

631 processes these isotope systems behave differently, leading to a decoupling between  $\Box_{Nd}$  and

 $G_{Hf}$  in weathering products. Autogenic processes within sediment transport such as mineral

sorting are partly responsible for this decoupling, mainly due to the control exerted by zircon over  $\varepsilon_{\text{Hf}}$  values, as zircon is characterized by high concentrations and highly unradiogenic Hf values (Bayon et al., 2006; Carpentier et al., 2014; Patchett et al., 1984). In consequence, the 636 clay fraction — usually devoid of zircons in comparison to coarser fractions — displays higher 637  $\varepsilon_{\rm Hf}$  values (Carpentier et al., 2014; Chauvel et al., 2014; Garçon et al., 2013). The clay fraction 638 of sediments display a distinct correlation between  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$  values denominated the "Clay 639 Array" (Bayon et al., 2016). Within the clay fraction, it has been shown that the offset in  $\varepsilon_{\rm Hf}$ 640 values from the clay array, expressed as  $\Delta \varepsilon_{\text{Hf}}$ , is dependent on silicate weathering intensity (Bayon et al. 2016). Enhanced weathering would tend to augment the release of radiogenic Hf 641 642 from easy weatherable and Lu-enriched mineral phases (Bayon et al., 2006) into the solutes from which clay minerals precipitate, therefore they exhibit an increase in their  $\varepsilon_{Hf}$  values. 643

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While zircons are typically present in trace amounts within clay fractions, if they are 645 646 present in substantial quantities they can have an impact on the evolution of  $\varepsilon_{Hf}$  on clay 647 fractions. In our study, Zr concentrations of the analyzed samples oscillate between 120 and 648 180 ppm (Supplementary table S2), which falls within the range of the zircon-poor clay fraction and thus discards any significative influence of zircons in our  $\varepsilon_{Hf}$  values (Bayon et al., 2016; 649 650 Corentin et al., 2022). Most importantly, there is no correlation between Zr concentrations and 651 the  $\Delta \varepsilon_{Hf}$  and  $\varepsilon_{Hf}$  values (FS4), indicating that changes in  $\Delta \varepsilon_{Hf}$  are not controlled by zircon 652 abundance, but can be attributed mainly to changes in silicate chemical weathering intensity.

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Following Equation 1,  $\Delta \varepsilon_{Hf(80)}$  values were calculated to assess changes in silicate chemical weathering intensity. As for  $\varepsilon_{Nd}$  values, an age correction of t = 80 Ma has been uniformly applied to the  $\varepsilon_{Hf}$  values to calculate the  $\Delta \varepsilon_{Hf(80)}$  (Table 1). This correction corresponds to the average age of the analysed time interval. Based on Bayon et al. (2016), the increase in  $\Delta \varepsilon_{Hf(80)}$  values observed from the Cenomanian to the Maastrichtian (ca. 95-68 Ma; Fig. 6) points to a long-term enhancement of silicate chemical weathering intensity in the South African Plateau, that only decreases after the Cretaceous-Paleogene (K-Pg) boundary. The 661 initial increase in  $\Delta \epsilon_{Hf(80)}$  values between the Cenomanian and the Coniacian is concomitant to 662 the suggested weathering of the Karoo basalts, as previously shown by  $\epsilon_{Nd(80)}$  values and P<sub>2</sub>O<sub>5</sub> 663 wt%. However, while there is a shift to more unradiogenic  $\epsilon_{Nd(80)}$  values around the Coniacian 664 (ca. 87 Ma), indicating the input of a different sediment source, the  $\Delta \epsilon_{Hf(80)}$  still shows a slight 665 increase in the intensity of silicate chemical weathering that extends up to the Maastrichtian.

666 The new combined  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$  data from site O-A1 support the enhancement in silicate 667 chemical weathering processes in the South African Plateau during the late Cretaceous, which has been depicted at the more distal site DSDP 361 (Gaitan et al., 2024). Both sites display 668 mostly concordant sediment sources as suggested by the  $\varepsilon_{Nd(80)}$  values (Fig. 6). However, the 669 670 geochemical signals recorded at both sites display differences in their evolution. While records 671 from borehole O-A1 show an enhancement in silicate chemical weathering through an increase 672 in  $\Delta \varepsilon_{\text{Hf}(80)}$  values between the Cenomanian to the Maastrichtian (Fig. 6), records from DSDP 673 core 361 show an increase from the Campanian to the Danian. Part of this discrepancy can be 674 related to the uncertainties in the age model from DSDP core 361, which are more pronounced 675 for the Turonian-Santonian interval (Gaitan et al., 2023). However, during the Maastrichtian-676 Danian interval where the age model is better calibrated for both sites their  $\Delta \varepsilon_{\text{Hf}(80)}$  records clearly differ. At borehole O-A1, the decrease in chemical weathering starts at the K-Pg 677 678 boundary, whereas at DSDP core 361  $\Delta \varepsilon_{\text{Hf}(80)}$  continues to increase and only begins to decrease 679 at the end of the Danian (ca. 58 Ma; Fig. 6). This time lag is also observed in the  $\varepsilon_{Nd(80)}$  values. 680 Clay  $\varepsilon_{Nd(80)}$  values from borehole O-A1 remain rather stable from the Campanian upwards 681 whereas clay  $\varepsilon_{Nd(80)}$  values at DSDP core 361 display a pronounced decrease of about 2  $\varepsilon$ -units 682 in the Maastrichtian-Danian interval, increasing again during the Danian-Thanetian interval 683 (Fig. 6). These differences in the isotopic record during the K-Pg transition could be related to 684 the input of different source material into the abyssal plain where the DSDP core 361 is located, 685 which would explain differences in  $\varepsilon_{Nd(80)}$  values between the two sites at that time. As the DSDP core 361 is located in a distal area, it has already been suggested that deep-bottom currents could have brought allogenic sediments from afar to this site (Murphy and Thomas, 2013; Natland, 1978). Likewise, aeolian transport of fine particles could incorporate allogenic material to distal areas. Overall, all these factors could have contributed to a divergence between borehole O-A1 and DSDP core 361 sources.

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# 5.4 Insights on climate, erosion, and weathering.

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Clay mineralogy is commonly used to assess the relative evolution of physical erosion (e.g., 694 695 Corentin et al., 2023, 2022; John et al., 2012), however, the clay mineralogy from site O-A1 696 cannot be used. The clay assemblages have been clearly affected by burial diagenetic processes 697 at least below 2500 mbsf (ca. 66 Ma), thus their content is not indicative of surface continental 698 processes below this depth. Published clay mineral assemblages reported for the abyssal plain 699 of the Cape Basin (i.e., DSDP 361) show an increase in illite and chlorite over smectite during the Campanian-Danian interval (Gaitan et al., 2023). This increase in primary clays over 700 701 secondary clays has been interpreted as a relative increase in physical erosion processes, and it 702 is coherent with existing thermochronological analyses across the South African Plateau that 703 have constrained rock cooling episodes interpreted as enhanced denudation during the mid-late 704 Cretaceous (Brown et al., 2002; Gallagher and Brown, 1999; Green et al., 2017; Kounov et al., 705 2009, 2013; Stanley et al., 2013, 2015; Stanley and Flowers, 2020; Tinker et al., 2008; 706 Wildman et al., 2015, 2016). The onset of these cooling episodes around the margins -707 interpreted primarily as erosional cooling — are dated around 100 Ma and extending up to 70 708 Ma (Brown et al., 2002; Gallagher and Brown, 1999; Green et al., 2017; Kounov et al., 2009, 709 2013; Tinker et al., 2008; Wildman et al., 2015, 2016). Additional AHe data (Apatite-Helium)

710 from inner areas of the Plateau have also shown an eastern erosional wave between ca. 120 and

711 60 Ma (Stanley et al., 2015, 2013; Stanley and Flowers, 2020).

712 These rock cooling episodes have been used to link tectonic uplift with enhanced physical 713 erosion and increases in sedimentation rates along the south African Margin during the late 714 Cretaceous (Baby et al., 2019, 2018a; Braun et al., 2014; Stanley et al., 2021). This period also 715 corresponds to the main increase in  $\Delta \varepsilon_{\text{Hf}(80)}$  associated to enhanced weathering of the margin, 716 suggesting that the tectonic uplift of the South African Plateau during the late Cretaceous has 717 resulted in an enhancement of both physical erosion and chemical weathering. The coupled 718 response of both processes suggests that denudation in this area likely operated under transport-719 limited conditions (Gabet and Mudd, 2009; West et al., 2005), in which erosion and weathering 720 are positively correlated. This transport-limited regime has been recently suggested to be 721 optimal for CO<sub>2</sub> withdrawal through silicate reactions (Bufe et al., 2024), particularly because 722 in these settings chemical weathering reactions are not kinetically-limited and therefore the 723 intensity of silicate chemical weathering reactions is not limited or reduced (Caves Rugenstein 724 et al., 2019; Gabet and Mudd, 2009; Maher and Chamberlain, 2014).

725 The increase in silicate chemical weathering recorded by  $\Delta \varepsilon_{\text{Hf}(80)}$  data starting by the 726 Cenomanian-Turonian is concomitant to the onset of the global climatic cooling recorded during the late Cretaceous (O'Brien et al. 2017). This increase during cooling climate 727 728 conditions hints to tectonics over climate as the main driver of weathering processes, as cooler 729 conditions tend to hamper the kinetics of silicate chemical weathering reactions (Maher, 2010; 730 Riebe et al., 2004). More importantly, this enhanced weathering episode could in contrast have 731 contributed to the onset of this cooling trend via silicate reactions. In fact, other sites on the 732 West and South African margin (Corentin et al., 2024; Gaitan 2023), and on the eastern South American margin (Corentin et al., 2023, 2022), have also recorded increases in silicate 733

chemical weathering, which in spite starting around the Santonian, have been suggested to have

- 735 contributed to the long-term global climate trend recorded during the late Cretaceous.
- 736

737 The increase in  $\Delta \varepsilon_{\text{Hf}(80)}$  in site O-A1 also precedes an increase in Ba (ppm) and CaO wt% 738 during the Santonian-Maastrichtian interval (ca. 85-65 Ma) suggesting also an increase in 739 marine productivity (Fig. 7A). The increase in CaO wt% could be initially linked to carbonate 740 productivity, especially due to the presence of the planktonic foraminifera used to establish the 741 age model of the site. However, to determine whether this increase is associated to marine 742 productivity or to a detrital input, the fraction of Ba can be normalised to Al to determine the 743 biogenic component of Ba (Ba<sub>xs</sub>), which can be used as a proxy to determine export production 744 (i.e., carbon flux to the deep ocean) (Bridgestock et al., 2019; Dymond et al., 1992; Eagle et 745 al., 2003; Liu et al., 2018; Paytan and Griffith, 2007). The evolution of Ba<sub>xs</sub> shows indeed an 746 increase during the Santonian-Maastrichtian interval, but more importantly appears to show a positive correlation with  $\Delta \varepsilon_{\text{Hf}(80)}$  (Fig. 7B), suggesting that enhanced silicate chemical 747 748 weathering could represent a source of nutrients that could have promoted and sustained high 749 productivity in this setting (e.g., Pogge von Strandmann et al., 2013; Yobo et al., 2021). This correlation between increases in silicate chemical weathering based on  $\Delta \varepsilon_{Hfclays}$  and export 750 751 production, has already been proposed as a mechanism in which silicate chemical weathering 752 can impact atmospheric CO<sub>2</sub> levels through enhancement of organic carbon burial by increases in nutrient release, in addition to the CO2 drawdown from silicate weathering reactions (Chen 753 754 et al., 2023). This hypothesis however, needs to be further confirmed by additional analyses such as TOC (total organic content),  $\delta^{13}C$ , or organic phosphorus, that could give independent 755 756 insights on the potential evolution of export production (Schoepfer et al., 2015).

757 Our new data from site O-A1 reveals an onset of enhanced silicate chemical weathering as 758 early as the Cenomanian (ca. 95 Ma; Fig. 6), also potentially contributing to promote and 759 sustain primary productivity between the Santonian and the Maastrichtian (Fig. 7). This 760 suggests, for the first time, that the increase in silicate chemical weathering across the South 761 African Plateau could have contributed as one of the triggering factors in the atmospheric CO<sub>2</sub> 762 withdrawal and subsequent onset of the global cooling trend during the late Cretaceous. First, through atmospheric CO<sub>2</sub> withdrawal by silicate weathering reactions, and later in the 763 764 Santonian - Maastrichtian by potentially enhancing primary productivity and organic carbon storage sustained by enhanced nutrient supply from the South African continent. Still, coupled 765 766 geochemical-climate models are needed to constrain quantitatively the impact of these 767 processes in atmospheric CO<sub>2</sub> withdrawal and on the global climate.

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#### 770 **6. Conclusion**

Our data from borehole O-A1 suggest, for the first time, that tectonically driven enhancement of silicate chemical weathering along the South African Plateau might have played a crucial role in the triggering of the late Cretaceous global cooling trend that started around the Cenomanian-Turonian transition. Additionally, the impact of enhanced silicate chemical weathering in  $CO_2$  seems to extend beyond the silicate mineral reactions, and also could be potentially related to the enhancement of primary productivity and organic carbon storage, which is a process that also captures carbon.

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Our new  $\Delta \varepsilon_{\rm Hf(80)}$  data indicate an increase in silicate chemical weathering from the 779 780 Cenomanian up to the Maastrichtian. The early phase (i.e., Cenomanian-Coniacian) of 781 enhanced erosion and silicate chemical weathering is linked to the denudation of the Karoo 782 basalts. This is supported by  $\varepsilon_{Nd(80)}$  and  $P_2O_5$  wt% data, which show patterns relatable to the incorporation of sediments derived from basaltic volcanic rocks. The  $\varepsilon_{Nd(80)}$  data suggest that 783 the main sediment source of the Cape Basin are units from the western-central areas of the 784 South African Plateau, including the Karoo basalts, Karoo sediments, and potentially the Cape, 785 786 Saldania, and Namaqua belts. Yet, there are some discrepancies between the DSDP core 361 787 and the borehole O-A1 record. Discrepancies in the  $\Delta \varepsilon_{\text{Hf}(80)}$  evolution during the Cenomanian-788 Santonian interval (ca. 97-83 Ma) are likely linked to age model uncertainties from the DSDP 789 core 361. However, discrepancies during the Maastrichtian-Thanetian interval appear to 790 indicate that drainage reconfiguration and changes in sediment distribution and source, altered 791 the denudation record between both locations during this period. Overall, our study highlights 792 the pertinence of extended basin analysis using multiple records across the basin to understand

and interpret denudation records, and to link them with changes in response to climatic andtectonic forcings.

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# 797 Figure Captions

Figure 1. Map of the source-to-sink system composed by the South African plateau and the
Cape basin. O-A1 borehole (33.161539°S; 16.822406°E) is located at the continental slope of
the Cape Basin. DSDP core 361 is located in the abyssal plain of the basin (35.066167°S;
15.448500°E). Topography and bathymetry data compiled from NASA JPL, (2013).

802 Figure 2. Biostratigraphy (planktonic foraminifera), sedimentary log, bulk mineralogy, clay 803 assemblages, and present-day  $\square_{Nd}$ - $\square_{Hf}$  isotope compositions of the analysed samples from 804 borehole O-A1. Units refer to the classification established by Baby et al., (2019). Planktonic 805 foraminifera abbreviations: *R*. appenninica=Rotalipora appenninica; Н. 806 helvetica=Helvetoglobotruncana helvetica; M. coronata=Marginotruncana coronata; G. 807 elevata = Globotruncanita elevata; G. aegyptiaca = Globotruncanita aegyptiaca; A. 808 mayaroensis = Abathomphalus mayaroensis. Error bars represent the standard deviation (2 $\square$ ) 809 calculated for both  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$ . Both  $\varepsilon_{Nd}$  and  $\varepsilon_{Hf}$  isotopic values are reported for t = 0.

Figure 3. REE concentrations of the clay fraction (< 2□m) from borehole O-A1. REE contents</li>
were normalized to the Post-Archean average Australian Shale (PAAS; Taylor and McLennan,
1985).

Figure 4. A) TEM image from sample O-A1/88 exhibiting long elongated regular-shaped illite
particles. B) TEM image from sample O-A1/61 showing interlayered illite-smectite highly
altered, thin, and fibrous. C) TEM image from sample O-A1/50 showing interlayered illite-

816 smectite, showing an elongated fibrous texture. At the bottom of the image a hexagonal 817 kaolinite particle can be observed as well. **D**) TEM image from sample O-A1/32 showing 818 smectite particles and regular polygonal illite and chlorite particles.

819 Figure 5. Lithological units from the South African Plateau. The  $\Box_{Nd}$  data compilation are as

follow: 1) Karoo Basalts (Hawkesworth et al., 1984) 2) Cape Belt and Karoo Sediments (Dia
et al., 1990) 3) Namaqua Belt (Reid, 1997) 4) Saldania Belt (Chemale et al., 2011). Modified
after Gaitan et al. (2023).

823 **Figure 6.** P<sub>2</sub>O<sub>5</sub> wt%,  $\Box_{Nd(80)}$  and  $\Box \Box_{Hf(80)}$  values for borehole O-A1 (continental slope),  $\Box_{Nd(80)}$ 

and  $\Box_{Hf(80)}$  values are also reported for the DSDP 361 core (abyssal plain; Gaitan et al., 2023). Values were calculated at t = 80 Ma. Palaeoprecipitations of the South African plateau are based on flora data set (Rayner et al., 1997; Sandersen, 2006; Smith, 1986). Sedimentation rates of the Cape and Orange Basins are estimations done by Baby et al. (2019).  $\Box^{18}$ O records from global benthic foraminifera are compiled from Friedrich et al. (2012); including data from Zachos et al. (2008) for the Cenozoic. Sample O-A1/37 was the only sample with a replicate for both the  $\Box_{Nd}$  and  $\Box_{Hf}$  values, the values were averaged, and the mean was used for the

plots. A smoothed curve and its 95% confidence interval were calculated for the  $P_2O_5$  wt% using Kernel regressions 1) Rock cooling interval based on multiple AFT and AHe thermochronology data across the South African Plateau (Brown et al., 2002; Gallagher and Brown, 1999; Green et al., 2017; Kounov et al., 2009, 2013; Raab et al., 2002; Stanley et al., 2013, 2015, 2015; Stanley and Flowers, 2020; Tinker et al., 2008; Wildman et al., 2015, 2016) 2) Timing of the South African Plateau margins uplift is based on the erosion-deposition model
proposed by Baby et al. (2019).

Figure 7. A) Evolution of Ba (ppm) and CaO wt% along the borehole O-A1. B) Linear correlation and its 95% confidence interval between  $Ba_{xs}$  and  $\Box \varepsilon_{Hf(80)}$ .  $Ba_{xs}$  was calculated as:  $Ba_{xs} = Ba_{total} - \left(\frac{Ba}{Al}\right)_{ter} \times Al_{total}$ . (Ba/Al)<sub>ter</sub> ratio value used for the calculations was 0.0075, which reflects the average upper crust (Dymond et al., 1992).

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FS1. Sedimentary log of borehole O-A1 and the age model used to determine the ages along
the section. Planktonic foraminifera were used to build the age model. Abbreviations: M.
coronata = Marginotruncana coronata. Species used to define the anchor points (white circles)
of the age model are shown on the graph.

FS2. Kûbler index of crystallinity along the borehole O-A1 with the evolution on lithology andclay assemblages.

**FS3.** Kûbler index compared to the evolution of  $\Box_{Nd}$ ,  $\Box_{Hf}$ , and  $\Box\Box_{Hf}$ . No correlation is found between the index and the isotope measurements.

**FS4.**  $\square_{\text{Hf}}$  vs Zr (in ppm) and  $\square_{\text{Hf}}$  vs Zr (in ppm). Zr content is assumed as an indicator of the presence of zircon. The R<sup>2</sup> coefficient for both graphs shows low correlation between the values.

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# 1303 Data availability

1304 All the data are provided in the supplementary materials attached to this paper.

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